PH20104 Planets and Exoplanets

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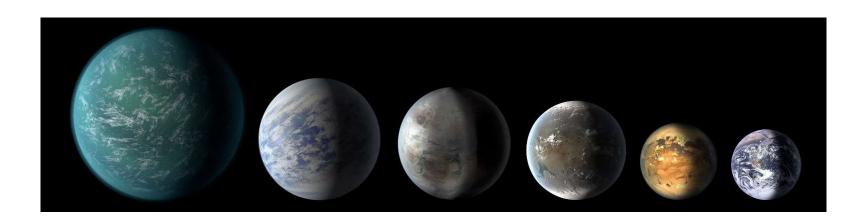
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Week #	Tuesday 10:15-11:05	Thursday 11:15-12:05	Friday 10:15-11:05
	4E3.10	8W3.22	5W2.3
19	8th February	15 th February	11 th February
	Lecture	Lecture	Lecture
20	15th February	17th February	18th February
	Lecture	Lecture	Lecture
21	22nd February	24 th February	25th February
	Lecture	Lecture	PC#1, group A
22	1 st March	3 rd March	4 th March
	Lecture	Lecture	PC#1, group B
23	8 th March	10 th March	11 th March
	Lecture	Lecture	PC #2, group A
24	15th March	17 th March	18 th March
	Lecture	Lecture	PC#2, group B
25	22 nd March	24 th March	25 th March
	Revision Lecture	"Planets" exam workshop	Office hour
26	29 th March	31 st March	1 st April
	Lecture	Lecture	Lecture
27	5th April	7th April	8th April
	Lecture	Lecture	Lecture
Easter Vacation	12 th April	14 th April	15 th April
Easter Vacation	19 th April	21st April	22 nd April
30	26th April	28th April	29 th April
	PC#3, group A	PC#3, group B	Office hour
31	3 rd May	5 th May	6 th May
	Revision Lecture	"Exoplanets" exam workshop	Office hour

Sample exam on Moodle – Past exams on Library webpage
Revision "checklist" + workshop at the end of each part of the unit.

PE

GM

What this unit is about

Present the latest discoveries in Planetary Physics and explain the science
Show how all fields of Physics combine to make these discoveries
Introduce the key tools used in the study of planets and exoplanets
Summarise current scientific challenges

"If you're teaching today what you were teaching 5 years ago, either the field is dead or you are"

(Noam Chomsky, interview to New Scientist, 17 March 2012)

What this unit is <u>not</u> about

Encyclopaedic description of each and every technique

Last-minute developments, especially if not relevant or not validated

Long demonstrations of key equations: they can be found in textbooks

Long presentations of historical developments: we will focus on current science

Unit outline

- 1. Taking Stock A grand tour of the Solar System
- 2. Measuring the Solar System Traditional and new techniques
- **3. Planetary Missions** Design, from orbital mechanics to science
- 4. Radar Mapping—Basics of radar scattering, applications to planetary imaging
- **5. Atmospheric processes**
- 6. Planetary magnetospheres

Revisions + exam workshop

Exoplanets – Introduction

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1. Taking Stock

Topic 1a: Rocky planets

Mercury

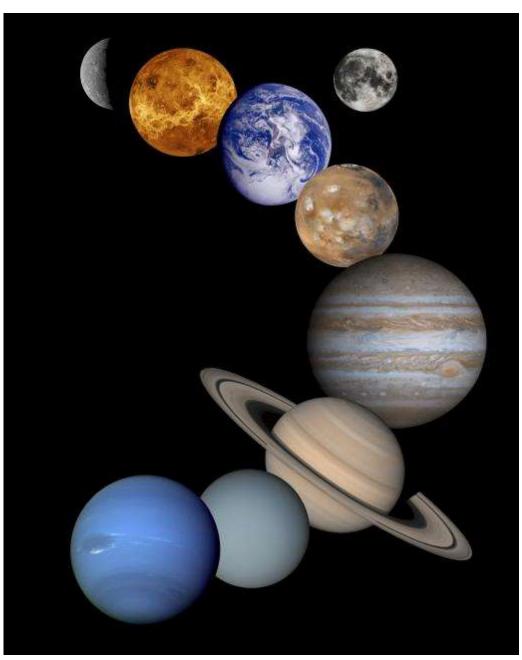
Venus

Earth

Moon

Mars

The Solar System



Large – Mostly rocks and gas

Sun 1,390,000 km diameter

Mercury 4,879 km

Venus 12,104 km

Earth 12,756 km + Moon 3,475 km 146 million km from Sun (1 AU)

Mars 6,794 km

Jupiter 142,984 km

Saturn 120,536 km

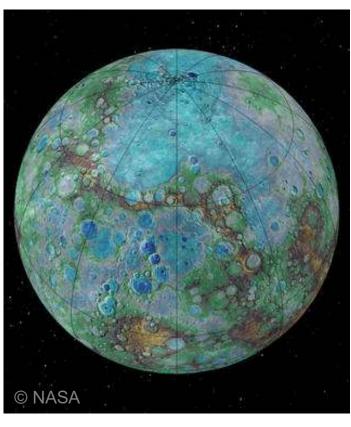
Uranus 51,118 km

Neptune 49,528 km 4.5 billion km from Sun (30.07 AU)

... and Pluto ...

Mercury

Whe deeper Seen higher



Smallest planet (a bit larger than the Moon)

58 million km (0.39 AU) from Sun

Orbits the Sun in 88 (Earth) days

1 Mercury day = 59 Earth days

Thin atmosphere: O₂, Na, H₂, He, K ejected from surface by solar wind

Surface: 430°C (day) to -180°C (night)

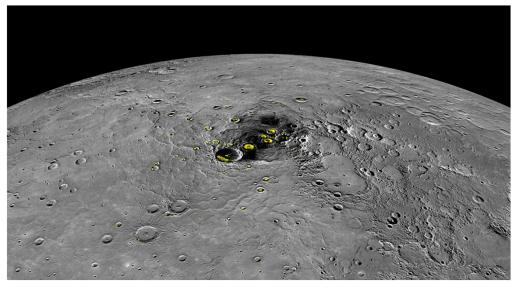
Only 2 missions: Mariner 10, 1974-75

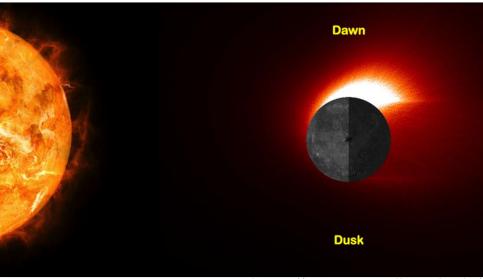
MESSENGER, 2011

Rocky surface – Traces of volcanism and tectonics – Impact craters



Mercury





https://messenger.jhuapl.edu/

© NASA

Presence of water ice at the poles (detected with Earth radiotelescope and orbiting spectrometers)

Calcium exosphere?Single source region – Why? How?

Magnetic field (1% Earth's) - Interacts with solar wind's magnetism

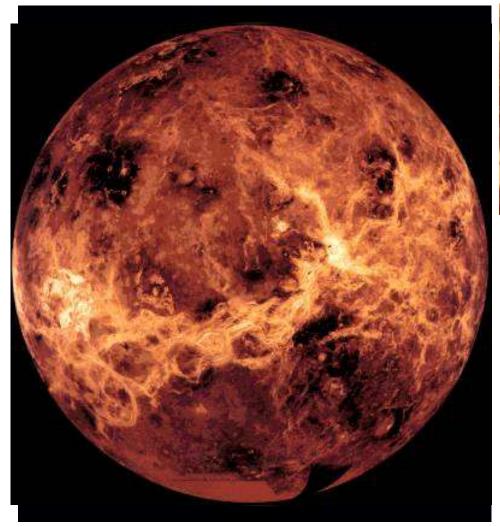
Bepi-Colombo (ESA)

2-spacecraft mission: launched 2018 (arrival 2025)

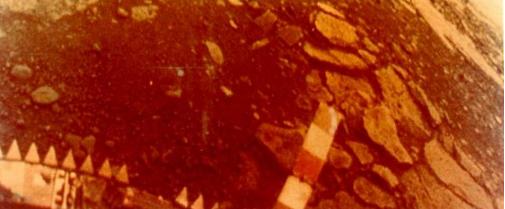
Some of its aims:

Is core liquid or solid? Still active tectonically?
No iron at surface?
Testing general relativity (proximity to Sun)

Venus – Earth's evil twin?



© NASA



USSR Academy of Sciences

Earth size – 0.7 AU from Sun

40+ missions (recently: Magellan, Venus Express and Akatsuki)

Orbits the Sun in 225 Earth days
1 Venus day = 243 Earth days
Retrograde spin

Thick atmosphere: CO₂, clouds of sulfuric acid

Greenhouse effect

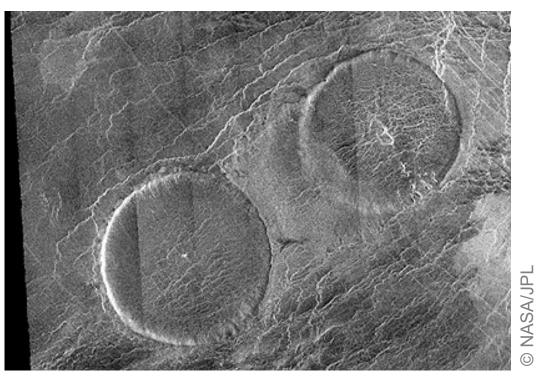
Surface: 490 °C – Pressure: 92 bars lvery high)

Rocky surface – Volcanism and tectonics – Impact craters when to the contract of the contract

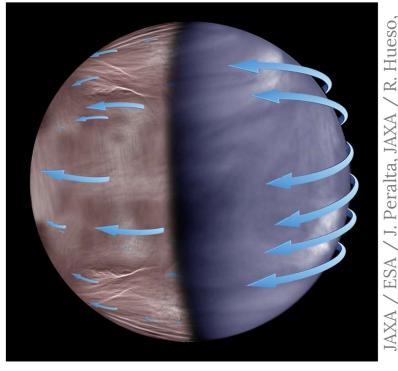
volca o every where

active!

Venus



Magellan radar image



Venus Express / Akatsuki 3.8 μm (IR) / 360 nm (UVI)

Open scientific questions:

Very thick and dynamic atmosphere: wind speeds > 400 km/h at surface

Equatorial jets and high-altitude winds?

Origin and significance of persistent atmospheric structures near mountains?

Is it still volcanically active? What about phosphine (PH₃) in the atmosphere?

Age of last resurfacing: 300 – 500 My? Dates from impact craters and samples?

Earth and the Moon

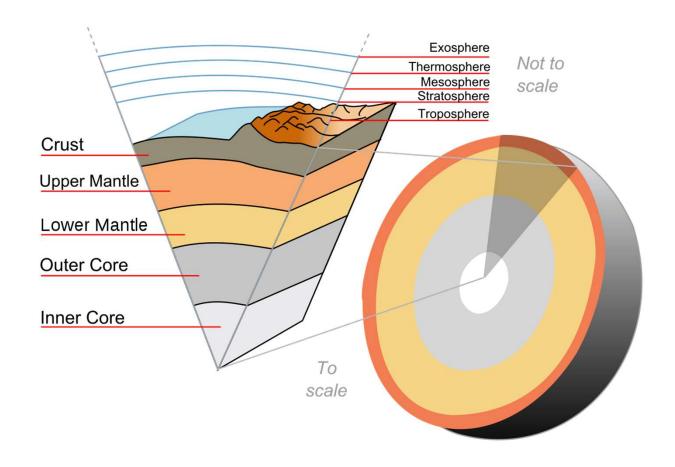




Close companionship (384,402 km average distance) – Same age? (4.53 billion years)

The **only** planetary bodies <u>directly</u> sampled with *human* **and** *robotic* presence.

Earth

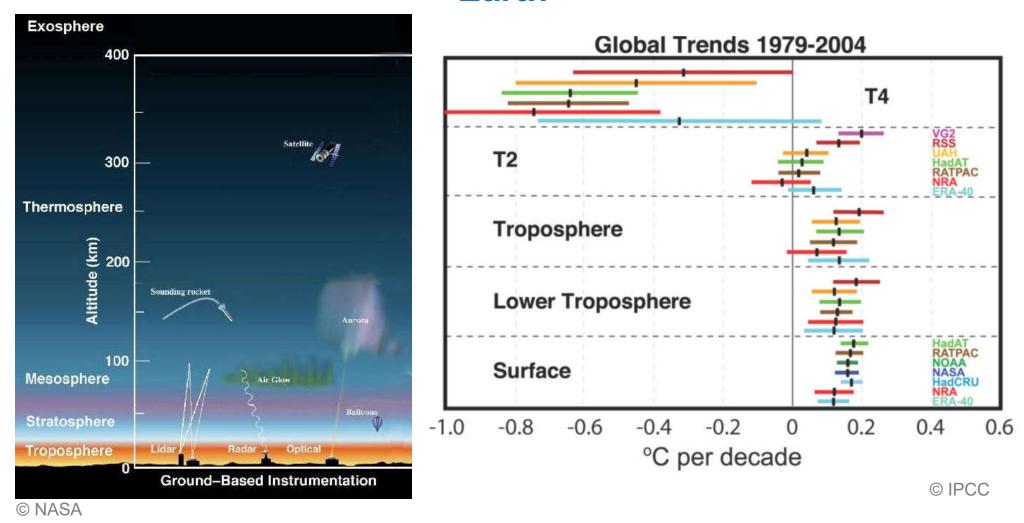


1 A.U. from Sun – Differentiated planet with rocky surface and >71% water cover

Complex surface and interior processes

By Surachit - Self-made, based on the public domain image File:Earth-crust-cutaway-english.png by Jeremy Kemp This W3C-unspecified vector image was created with Inkscape., CC BY-SA 3.0, https://commons.wikimedia.org/w/index.php?curid=2584710

Earth

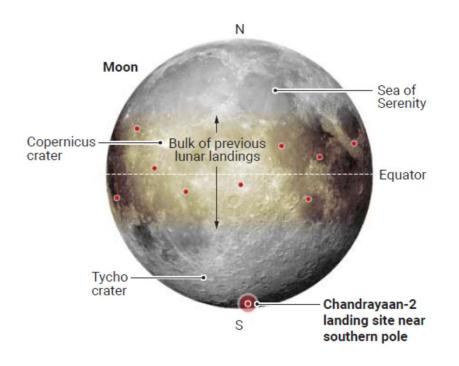


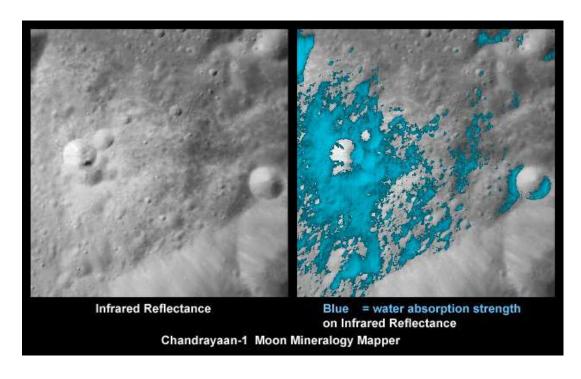
Atmosphere: 78% N₂, 21% O₂, 1% others

Complex and variable atmosphere – <u>Undergoing significant changes</u> (including in the oceans)

Variable magnetosphere

Moon





1,737 km diameter – Differentiated and same age as Earth?

12 "Moon-walkers", 100+ robotic missions

Tenuous atmosphere: He, Ar, Ne, CH₃, CH₄, CO₂ + sodium, potassium

Complex and variable atmosphere – Variable magnetosphere

From चन्द्रयान to 玉兔

Chandrayaan 3 (India):

launch 2021, repeat of Chandrayaan 2 (2019)

Orbiter on far side Water mapper

Lander and rover (600 km from south pole)
Langmuir probe: plasma and
electric charge in regolith
Seismometer for moonquakes

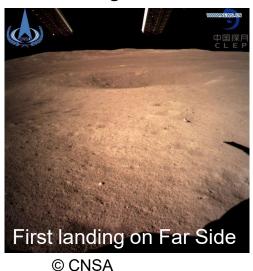
More missions in next decade – Google Lunar X Prize (March 2018)

Far-side observatory? South Pole Base? Sample return to Earth (Chang'e 5): December 2020

Jade Rabbit missions (China) 2014



Chang'e 4 2019



Still many questions about:

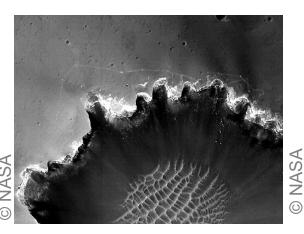
the formation of the Moon, its mineral deposits, and the presence of water ...

Artemis Gateway

Mars – Next stop?







½ size of Earth − 1.5 AU from Sun Many robotic expeditions, some on-going

Orbits the Sun in 687 Earth days
1 Mars day = 24.6 hours

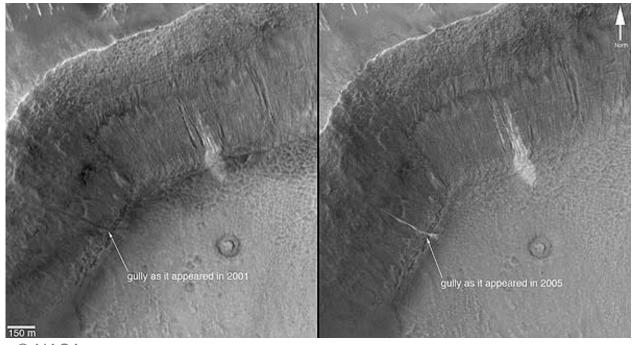
Thin atmosphere: mostly CO₂, N₂, Ar Complex atmosphere (winds, dust storms)

Surface: -89 °C to -31 °C – 9 mb pressure

Rocky planet – Strong weather (planet-scale storms) – Polar caps Large volcanoes – Flat plains

No (current) magnetosphere

Mars



Open questions:

Free-flowing water?

Deposits and habitability

Atmosphere circulation Methane anomaly

Structure of the core

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NASA Mars Reconnaissance Orbiter (MRO) – 50,000 orbits in Oct. 2017

ExoMars 2016 Trace Gas Orbiter (since Oct. 2016) – Lander Schiaparelli lost on descent

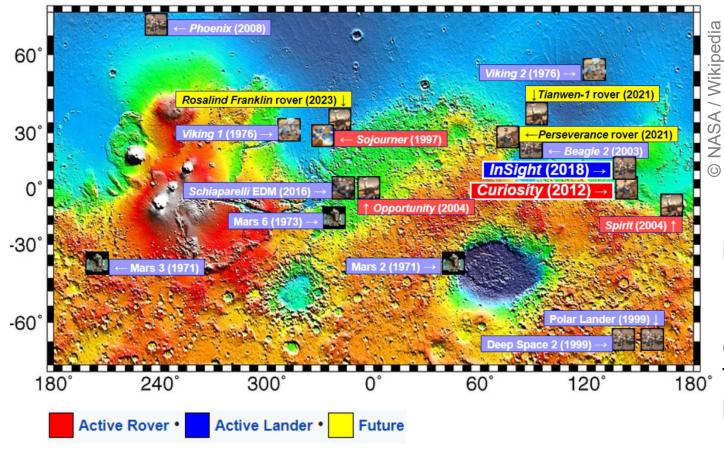
Search for methane and other trace atmospheric gases (signatures of active biological or geological processes)

NASA InSight (Nov. 2018) – Sub-surface mapping (heat flows and marsquakes)

ESA ExoMars Rover (launch Summer 2022) – and more missions planned ...

Tianwen-1 (China) – Orbit insertion scheduled 10 Feb. 2021 – Rover landing in July

Mars missions in 2021



Perseverance (NASA) Landing 18 Feb. 2021

Rosalind Franklin/ExoMars (ESA/Roscosmos) To be launched 2022 Landing June 2023



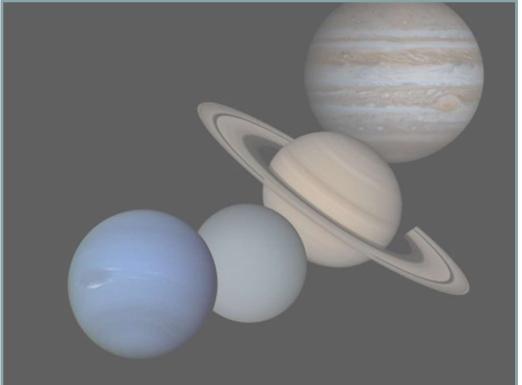
Mars



Hope (مسبار الأمل) (UAE)
Orbit insertion 9 Feb. 2021

Rocky planets - Summary





Planetary bodies within the frost line

(point between the orbits of Mars and Jupiter where the material is cool enough for volatile icy compounds to remain solid)

All bodies have been mapped and explored: Only manned missions on Earth and Moon. Only samples from Earth and Moon.

Great variety: s

sizes / surfaces

atmospheres

magnetospheres

Water present on:

Mercury

Earth + Moon

Mars

New discoveries still coming in ... Solar system formation: still debated