VECTOR CALCULUS

Many physical quantities are **vectors**, which change with time or with position. How do we describe this mathematically?

In this part of the unit, we will deal with differentiation & integration first in Cartesian coordinates (x, y, z), then in other "coordinate systems".

Contents:

- 1. Differentiation of vectors & space curves
- 2. Differentiation of scalar and vector fields
- 3. Integration of scalar and vector fields
- 4. non-Cartesian coordinates
- 5. Integration in non-Cartesian coordinates
- 6. Differentiation in non-Cartesian coordinates
- 7. Vector integral theorems

Sections are not equal in length.

Each has a "Problem Sheet" and an "Extra Sheet".

1. Differentiation of vectors & space curves

Consider a vector quantity **a** which changes according to a scalar *t*.

i.e. \mathbf{a} is really $\mathbf{a}(t)$

[eg: If the position of a particle \mathbf{r} moves with time $\Rightarrow \mathbf{r} = \mathbf{r}(t)$].

The **components** of **a** also change with *t* so write

$$\mathbf{a}(t) = a_x(t)\mathbf{i} + a_y(t)\mathbf{j} + a_z(t)\mathbf{k}$$

Note that the Cartesian basis vectors **i**, **j** and **k** are FIXED.

The derivative of \mathbf{a} with respect to t is defined by

$$\frac{d\mathbf{a}}{dt} = \lim_{\delta t \to 0} \left[\frac{\mathbf{a}(t + \delta t) - \mathbf{a}(t)}{\delta t} \right] = \lim_{\delta t \to 0} \left[\frac{\delta \mathbf{a}}{\delta t} \right],$$

or
$$\left| \frac{d\mathbf{a}}{dt} = \frac{da_x}{dt} \mathbf{i} + \frac{da_y}{dt} \mathbf{j} + \frac{da_z}{dt} \mathbf{k} \right|$$
 (1)

A good example to keep in mind is

position, velocity, acceleration.

Position: $\mathbf{r}(t)$

NB: r is ALWAYS position vector on this course.

Velocity:
$$\mathbf{v}(t) = \frac{d\mathbf{r}}{dt}$$

Acceleration:
$$\mathbf{a}(t) = \frac{d\mathbf{v}}{dt} = \frac{d^2\mathbf{r}}{dt^2}$$

If
$$\mathbf{r}(t) = x(t)\mathbf{i} + y(t)\mathbf{j} + z(t)\mathbf{k}$$
 then

$$\mathbf{v}(t) = v_x(t)\mathbf{i} + v_y(t)\mathbf{j} + v_z(t)\mathbf{k},$$

with

$$v_x(t) = \frac{dx(t)}{dt}$$
 $v_y(t) = \frac{dy(t)}{dt}$ $v_z(t) = \frac{dz(t)}{dt}$

etc

Example

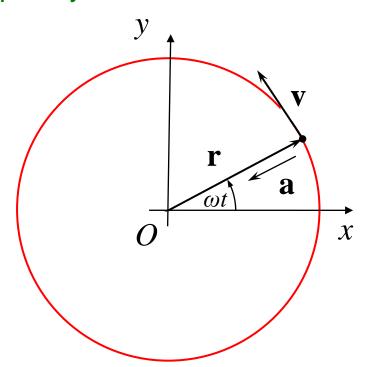
$$\mathbf{r} = \underbrace{\cos(\omega t)}_{x(t)} \mathbf{i} + \underbrace{\sin(\omega t)}_{y(t)} \mathbf{j}.$$

$$\mathbf{v} = \frac{d\mathbf{r}}{dt} = \underbrace{-\omega \sin(\omega t)}_{v_x(t)} \mathbf{i} + \underbrace{\omega \cos(\omega t)}_{v_y(t)} \mathbf{j}$$

$$\mathbf{a} = \frac{d\mathbf{v}}{dt} = \underbrace{-\omega^2 \cos(\omega t)}_{a_x(t)} \mathbf{i} \underbrace{-\omega^2 \sin(\omega t)}_{a_y(t)} \mathbf{j}.$$

So
$$\mathbf{a}(t) = -\omega^2 \mathbf{r}(t)$$

This is just a description of **circular motion** with angular frequency ω .



1.1 Rules for differentiation of vectors

(i)
$$\frac{d}{dt}(c\mathbf{a}) = c\frac{d\mathbf{a}}{dt}$$
 (c is a constant)

(ii)
$$\frac{d}{dt}(\mathbf{a} + \mathbf{b}) = \frac{d\mathbf{a}}{dt} + \frac{d\mathbf{b}}{dt}$$

(iii)
$$\frac{d}{dt}(\phi(t)\mathbf{a}(t)) = \phi \frac{d\mathbf{a}}{dt} + \frac{d\phi}{dt}\mathbf{a}$$
 (ϕ is a scalar fn of t)

(iv)
$$\frac{d}{dt}(\mathbf{a} \cdot \mathbf{b}) = \mathbf{a} \cdot \frac{d\mathbf{b}}{dt} + \frac{d\mathbf{a}}{dt} \cdot \mathbf{b}$$

(v)
$$\frac{d}{dt}(\mathbf{a} \times \mathbf{b}) = \mathbf{a} \times \frac{d\mathbf{b}}{dt} + \frac{d\mathbf{a}}{dt} \times \mathbf{b}$$
 (care with order)

(vi)
$$\frac{d}{dt}\mathbf{a}(s) = \frac{d\mathbf{a}}{ds}\frac{ds}{dt}$$
 (a is a function of s, which is a scalar function of t)

All these can be proved by looking at components

Example of the use of these rules:

Consider a vector $\mathbf{a}(t)$ with constant **magnitude** a. ie $\mathbf{a}(t) \cdot \mathbf{a}(t) = a^2$.

Though the magnitude is constant, the direction need not be, so the derivative of $\mathbf{a}(t)$ need not be zero. Differentiating,

LHS:
$$\frac{d}{dt} [\mathbf{a}(t) \cdot \mathbf{a}(t)] = \mathbf{a} \cdot \frac{d\mathbf{a}}{dt} + \frac{d\mathbf{a}}{dt} \cdot \mathbf{a} = 2\mathbf{a} \cdot \frac{d\mathbf{a}}{dt} \quad \text{(iv)}$$

RHS: $\frac{d}{dt}(a^2) = 0$, as a^2 is a **constant.**

$$\therefore \mathbf{a} \cdot \frac{d\mathbf{a}}{dt} = 0$$

So either $\frac{d\mathbf{a}}{dt} = 0$ [so **a** is a constant vector],

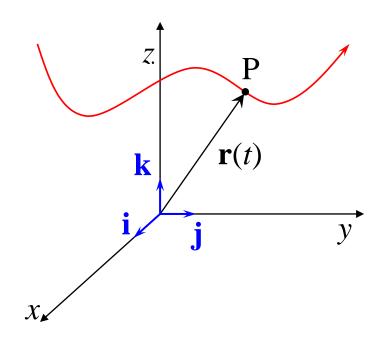
or $\frac{d\mathbf{a}}{dt}$ is perpendicular to \mathbf{a}

The latter case is what we have for circular motion, where the position vector \mathbf{r} is a vector of constant magnitude; we found that the velocity vector \mathbf{v} is always perpendicular to \mathbf{r} .

1.2 Space Curves

How do we describe *trajectories* of particles moving through space?

Think about point P whose coordinates x(t), y(t), and z(t) are continuous functions of time t. As t increases, P traces out a **curve in space**.



eg
$$P = \mathbf{r}(t) = \underbrace{\cos(\omega t)}_{x(t)} \mathbf{i} + \underbrace{\sin(\omega t)}_{y(t)} \mathbf{j} + \alpha t \mathbf{k}$$

This is a space curve with circular motion in 2d (the xy plane in this case), with angular frequency ω , and constant motion, with speed α , in z. This is a HELIX.

Q: How do we write a known trajectory or curve in the form $\mathbf{r}(t) = x(t)\mathbf{i} + y(t)\mathbf{j} + z(t)\mathbf{k}$?

1.3 Parameterisation of curves

This is a very useful technique.

Key point is that all lines (curves) are 1-dimensional.

We can : follow any line by increasing or decreasing a parameter that takes us along the line.

Let such a parameter be u

We parameterise a curve by finding expressions for x(u), y(u), and z(u) for that curve.

Eg (2d): take the straight line y = 3x + 4.

Let u = x. Then the line is followed as u is varied, provided we set y = 3u + 4 and z = 0.

The position vector varies with parameter u as

$$\mathbf{r}(u) = x(u)\mathbf{i} + y(u)\mathbf{j} + z(u)\mathbf{k}$$
$$= u\mathbf{i} + (3u + 4)\mathbf{j}$$

Note that any curve can be parameterised an infinite number of ways – pick the one that suits you best!

Eg let v = 3x. Then the same line y = 3x + 4 can be written

$$\mathbf{r}(v) = \frac{v}{3}\mathbf{i} + (v+4)\mathbf{j}.$$

For space curves, two useful parameters are time *t* and distance along the curve *s*.

More 2d examples of parameterisation

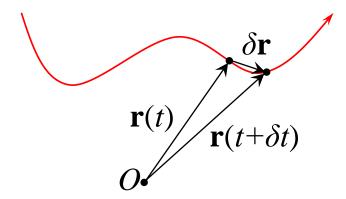
- Express $y = 3x^2$ as $\mathbf{r}(t)$.

- Express $x^2 + y^2 = 9$ as $\mathbf{r}(u)$.

- Express $\mathbf{r}(u) = e^{u}\mathbf{i} + e^{-u}\mathbf{j}$ as y(x).

1.4 Tangent to a space curve

Consider 2 adjacent points $\mathbf{r}(t)$ and $\mathbf{r}(t+\delta t)$ on a space curve.



The separation of the 2 points is

$$\delta \mathbf{r} = \mathbf{r}(t + \delta t) - \mathbf{r}(t) \approx \frac{d\mathbf{r}}{dt} \delta t$$
 [from defin of $\frac{d\mathbf{r}}{dt}$]

In the limit that $\delta t \to 0$ this approximation becomes exact – then $\delta \mathbf{r}$ becomes **parallel to the local** direction of the curve, which is the tangent direction at that point.

- \therefore tangent is in direction of $d\mathbf{r} = \lim_{\delta t \to 0} \delta \mathbf{r}$
- \therefore tangent is in direction of $\frac{d\mathbf{r}}{dt}$ [from above]

So UNIT TANGENT VECTOR
$$\hat{\mathbf{T}} = \frac{\frac{d\mathbf{r}}{dt}}{\left|\frac{d\mathbf{r}}{dt}\right|}$$
.

Also,
$$\frac{d\mathbf{r}}{dt} = \mathbf{v}$$
 (velocity), and $\left| \frac{d\mathbf{r}}{dt} \right| = |\mathbf{v}|$ (speed).

tangent is in direction of velocity, with

$$\hat{\mathbf{T}} = \frac{\mathbf{v}}{|\mathbf{v}|} = \hat{\mathbf{v}}.$$

Space curves can also be expressed in terms of a position vector as a function of **distance travelled along the curve**, *s*.

ie consider $\mathbf{r}(s)$ instead of $\mathbf{r}(t)$

From above, tangent vector is in direction of $d\mathbf{r}$, and the magnitude of $d\mathbf{r}$ must be ds, the distance travelled in dt.

$$\therefore \quad \hat{\mathbf{T}} = \frac{d\mathbf{r}}{ds}$$

But
$$\frac{d\mathbf{r}(s)}{dt} = \frac{d\mathbf{r}(s)}{ds} \frac{ds}{dt} = \hat{\mathbf{T}} \frac{ds}{dt}$$
 which means we

have 3 useful ways of writing the speed of a particle:

Speed =
$$|\mathbf{v}| = \left| \frac{d\mathbf{r}}{dt} \right| = \frac{ds}{dt}$$

NOTE: if we know \mathbf{r} as a function of t, then use

$$\left| \frac{d\mathbf{r}}{dt} \right| = \frac{ds}{dt}$$
 to find the distance travelled as a function

of t, s(t). Simply differentiate \mathbf{r} with respect to t, find the magnitude of the resulting vector, then integrate with respect to t.

Example

Consider circular motion again...

If
$$\mathbf{r}(t) = a\cos(\omega t)\mathbf{i} + a\sin(\omega t)\mathbf{j}$$
 [a is the radius]

then
$$\mathbf{v}(t) = \frac{d\mathbf{r}}{dt} = -a\omega\sin(\omega t)\mathbf{i} + a\omega\cos(\omega t)\mathbf{j}$$
,

$$\left| \frac{d\mathbf{r}}{dt} \right| = \left[a^2 \omega^2 \sin^2(\omega t) + a^2 \omega^2 \cos^2(\omega t) \right]^{\frac{1}{2}} = a\omega$$

so
$$\hat{\mathbf{T}} = \frac{\frac{d\mathbf{r}}{dt}}{\left|\frac{d\mathbf{r}}{dt}\right|} = -\sin(\omega t)\mathbf{i} + \cos(\omega t)\mathbf{j}$$
 for any t .

For example, to find $\hat{\mathbf{T}}$ at the point (0,a) on the curve,

- (i) find t. Here $x(t) = 0 = a\cos(\omega t)$ and $y(t) = a = a\sin(\omega t)$. This is true for $\omega t = \frac{\pi}{2}$ (and other values...)
- (ii) find $\hat{\mathbf{T}}$. Here $\hat{\mathbf{T}} = -\sin(\frac{\pi}{2})\mathbf{i} + \cos(\frac{\pi}{2})\mathbf{j} = -\mathbf{i}$.

Also,
$$\left| \frac{d\mathbf{r}}{dt} \right| = \frac{ds}{dt} = a\omega$$
, so $s = a\omega t + c$ [c is integration constant]

1.5 Newton's Law in vector form

Newton's Law "Force = mass x acceleration" is really a vector differential equation

$$\mathbf{F} = m \frac{d^2 \mathbf{r}}{dt^2}.$$

So, to find the trajectory $\mathbf{r}(t)$ of an object subjected to a force \mathbf{F} we solve equations like this one.

For example, the electromagnetic force on a particle of charge q in electric and magnetic fields \mathbf{E} and \mathbf{B} is

$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

so the full "equations of motion" for the particle are

$$m\frac{d^2\mathbf{r}}{dt^2} = q\mathbf{E} + q\frac{d\mathbf{r}}{dt} \times \mathbf{B}.$$

In 3d, we solve 3 ODEs, one for each component:

$$F_{x} = m\frac{d^{2}x}{dt^{2}} \qquad F_{y} = m\frac{d^{2}y}{dt^{2}} \qquad F_{z} = m\frac{d^{2}z}{dt^{2}}$$