SCHWARZCHILD BLACK HOLE SOLUTION.

Vijay Chawan, 13th February.

The line element for spherically symmetric object(after some processing)* is written as;

$$ds^{2} = A(r, t) dt^{2} - B(r, t) dr^{2} - r^{2} (d\theta^{2} + \sin^{2}\theta d\phi^{2}).$$

For static case, A and B will be independent of t. Which is the case with Schwarzschild Black Hole.

In[•]:=

Print["Covarient metric"]

inversemetric // MatrixForm (*contravarient metric*)

Covarient metric in $r-\theta-\phi$ coordinates

$$\begin{pmatrix} A[r] & 0 & 0 & 0 \\ 0 & -B[r] & 0 & 0 \\ 0 & 0 & -r^2 & 0 \\ 0 & 0 & 0 & -r^2 \sin[\theta]^2 \end{pmatrix}$$

Covarient metric

$$\begin{pmatrix} \frac{1}{A[r]} & 0 & 0 & 0 \\ 0 & -\frac{1}{B[r]} & 0 & 0 \\ 0 & 0 & -\frac{1}{r^2} & 0 \\ 0 & 0 & 0 & -\frac{Csc[\theta]^2}{r^2} \end{pmatrix}$$

Calculating Christoffel Connection, Riemann tensor, Ricci tensor using the above metric.

$$\Gamma_{\mu\nu}{}^{\lambda} = 1/2 g^{\lambda\sigma} (\partial_{\mu} g_{\sigma\nu} + \partial_{\nu} g_{\mu\sigma} - \partial_{\sigma} g_{\mu\nu})$$

$$R_{\sigma\mu\nu}{}^{\lambda} = \partial_{\mu}\Gamma^{\lambda}{}_{\nu\sigma} - \partial_{\nu}\Gamma^{\lambda}{}_{\mu\sigma} + \Gamma^{\lambda}{}_{\mu\rho}\Gamma^{\rho}{}_{\nu\sigma} - \Gamma^{\lambda}{}_{\nu\rho}\Gamma^{\rho}{}_{\mu\sigma}$$

$$R_{\sigma v} = R_{\sigma \lambda v}^{\lambda}$$

Note: metric[[1,1]] will give the first column and first row element that is tt element of metric.

Christoffel Connection

```
In[ • ]:=
     crishtoffel[\lambda_{-}, \mu_{-}, v_{-}] :=
       crishtoffel[\lambda, \mu, v] = Simplify \left[\frac{1}{2} * \left(\sum_{\sigma=1}^{4} inversemetric[[\lambda, \sigma]] * \right)\right]
               (D[metric[[\sigma, v]], coords[[\mu]]] + D[metric[[\mu, \sigma]],
                   coords[[v]]] - D[metric[[\mu, v]], coords[[\sigma]]]);
```

Riemann tensor

Simplify D[crishtoffel[
$$\lambda$$
, σ , μ] := riemanntensor[λ , σ , μ , ν] = D[crishtoffel[λ , σ , μ], coords[[ν]]] +
$$\sum_{\rho=1}^{4} (\text{crishtoffel}[\lambda, \mu, \rho] * \text{crishtoffel}[\rho, \sigma, \nu]) - \sum_{\rho=1}^{4} (\text{crishtoffel}[\lambda, \nu, \gamma] * \text{crishtoffel}[\gamma, \sigma, \mu])];$$

Ricci tensor and Ricci scalar

```
In[ • ]:=
     riccitensor[\mu_{-}, v_{-}] :=
       riccitensor[\mu, \nu] = Simplify[\sum_{i=1}^{4} riemanntensor[\lambda, \mu, \lambda, \nu];
     ricciscalar =
       Simplify \left[\sum_{i=1}^{4}\sum_{j=1}^{4}\text{inversemetric}[[\mu, v]]*\text{riccitensor}[\mu, v]\right];
```

Printing Ricci tensor

```
In[ • ]:=
    riccitensormetric =
       {{riccitensor[1, 1], riccitensor[1, 2], riccitensor[1, 3], riccitensor[1, 4]},
        {riccitensor[2, 1], riccitensor[2, 2], riccitensor[2, 3], riccitensor[2, 4]},
        {riccitensor[3, 1], riccitensor[3, 2], riccitensor[3, 3], riccitensor[3, 4]},
        {riccitensor[4, 1], riccitensor[4, 2], riccitensor[4, 3], riccitensor[4, 4]}};
```

riccitensormetric // MatrixForm

Out[]//MatrixForm=

MatrixForm=
$$\begin{pmatrix}
-\frac{A'[r]B'[r]}{4B[r]^2} + \frac{\frac{A'[r]^2}{r} - \frac{A''[r]^2}{4A[r]} + \frac{A''[r]}{2}}{B[r]} & 0 & 0 \\
0 & \frac{A[r](4A[r] + rA'[r])B'[r] + rB[r](A'[r]^2 - 2A[r]A''[r])}{4 r A[r]^2 B[r]} & 0 \\
0 & 0 & \frac{1}{2}\left(2 - \frac{2 + \frac{rA'[r]}{A[r]}}{B[r]} + \frac{rB'[r]}{B[r]^2}\right) \\
0 & 0 & 0 & \frac{\sin[\theta]^2(-rB[r]A'[r])}{2} + \frac{\sin[\theta]^2(-rB[r]A'[r])}{2}$$

The Ricci tensor components,

$$R_{11} = -\frac{A'[r]B'[r]}{4B[r]^2} + \frac{\frac{A'[r]}{r} - \frac{A'[r]^2}{4A[r]} + \frac{A''[r]}{2}}{B[r]}$$

$$R_{22} = \frac{A[r] (4 A[r] + r A'[r]) B'[r] + r B[r] (A'[r]^2 - 2 A[r] A''[r])}{4 r A[r]^2 B[r]}$$

$$R_{33} = \frac{1}{2} \left(2 - \frac{2 + \frac{rA'[r]}{A[r]}}{B[r]} + \frac{rB'[r]}{B[r]^2} \right)$$

$$R_{44} = \frac{\sin[\theta]^2 \left(-r B[r] A'[r] + A[r] \left(-2 B[r] + 2 B[r]^2 + r B'[r]\right)\right)}{2 A[r] B[r]^2}$$

$$R_{11}$$
 can be written as $\frac{A''}{2B} + \frac{A'}{rB} - \frac{A'}{4B} \left(\frac{A'}{A} + \frac{B'}{B} \right)$

$$R_{22}$$
 can be written as $-\frac{A''}{2A} + \frac{B'}{rB} + \frac{A'}{4A} \left(\frac{A'}{A} + \frac{B'}{B}\right)$

$$R_{33}$$
 can be written as $1 - \frac{1}{B} - \frac{r}{2B} \left(\frac{A'}{A} - \frac{B'}{B} \right)$

1). Multiply R_{11} with B/A and add to R_{22} :

The result should be $\frac{B'}{A} = -\frac{A'}{B}$, thus, A*B = constant. The space is asymptotically flat, thus, for r tends to infinity B and A tends to 1, this gives the constant to be 1.

2). Use
$$\frac{B'}{A} = -\frac{A'}{B}$$
 is R_{33} :

The result is $A + r^*A' = 1$ which implies $d(rA)/dr = d(r + r^*A')$ constant), that is rA = r + constant.

Thus, A = 1 + constant/r and B = 1/(1 + constant/r).

To find the constant now.

In the weak field limit, $g_{00} = (1 - 2GM/r)$, where r >> 2GM.

In our case, A = 1 + constant/r. have the same form as that of g_{00} weak field limit case, thus constant = 2GM.

The line element is then,

$$ds^{2} = (1 - 2 GM/r) dt^{2} - (1 - 2 GM/r)^{-1} dr^{2} - r^{2} (d\theta^{2} + Sin[\theta]^{2} d\phi^{2})$$

Hence, Schwarzschild Solution is achieved.