1 Analytic Solution

Typically, in fluid dynamics, the scalar function known as the velocity potential denoted by φ, is what is desired. Once the velocity potential is known, the system is solved because $\vec{\nabla}\varphi = \mathbf{u}$, and in the case of a free surface, the deformation of the surface can readily be found from the velocity potential as well. The velocity potential will be found in the coming sub-chapters, and then the energy deposited into the neutron star will be calculated.

1.1 Eigenfunctions of the Laplacian

The first step in analytically solving for the velocity potential will be to find the eigenfunctions of Laplace's equation, $\nabla^2\varphi = 0$, since the velocity potential is conservative. Since this is a partial differential equation we will assume the solution is the product of univariate functions, $\varphi = f(r)g(\theta)h(z)T(t)$. Expanding the Laplacian for cylindrical systems gives

$$\frac{1}{r}\frac{\partial}{\partial r}\left(r\frac{\partial\varphi}{\partial r}\right) + \frac{1}{r^2}\left(\frac{\partial^2\varphi}{\partial\theta^2}\right) + \frac{\partial^2\varphi}{\partial z^2} = 0.$$

The temporal component can be divided out since the Laplacian does not involve time, and will be found later on. Substituting in our assumed form, and dividing by φ, yields

$$\frac{1}{fr}\frac{\partial}{\partial r}(rf') + \frac{1}{r^2}\frac{g''}{g} + \frac{h''}{h} = 0. \tag{1}$$

Where primes denote the derivative with respect to the function's only variable. Notice that the function $h(z)$ has been separated from both $f(r)$, and $g(\theta)$, and since each function is univariate, it must be the case that the last term is constant,

$$\frac{h''}{h} = k^2.$$

We cleverly force this constant to be positive to satisfy the boundary conditions, namely, that $h(-\infty) = 0$. The most general solution to this is of course exponentials,

$$h(z) = Ae^{kz} + Be^{-kz}.$$

Since the velocity potential has to vanish at negative infinity, this is simplified to

$$h(z) = e^{kz}, \tag{2}$$

without loss of generality we can assume the constant is one, since it can be absorbed into $T(t)$. By substituting in k^2 and multiplying through by r^2, (1) becomes

$$\frac{r}{f}\frac{\partial}{\partial r}(rf') + k^2r^2 + \frac{g''}{g} = 0, \tag{3}$$

and once again, $g(\theta)$ has been separated from $f(r)$, and so the last term must be constant –

$$\frac{g''}{g} = -\mu^2.$$

The constant is forced to be negative since $g(\theta)$ is expected to be periodic, and not exponential. Of course, the general solution to this differential equation is

$$g(\theta) = A\sin(\mu\theta) + B\cos(\mu\theta).$$

However, our system is cylindrically symmetric; there is no way to differentiate one value of θ to another, thus, it is expected that φ has no θ dependence, and consequentially, $g(\theta)$

must be constant, the only way this is satisfied is if $\mu = 0$, so that

$$g(\theta) = 1. \tag{4}$$

By substituting $\mu = 0$ into (3), and multiplying through by f we obtain

$$r\frac{\partial}{\partial r}(rf') + (k^2r^2 - 0^2)f = 0,$$

which indeed is Bessel's differential equation of order 0. Therefore,

$$f(r) = AJ_0(kr) + BY_0(kr),$$

but, once again, B must equal zero since $Y_0(0)$ is not finite which is unphysical. The radial part thus simplifies to

$$f(r) = J_0(kr). \tag{5}$$

By combining (2), (4), and (5) we find that the velocity potential is

$$\varphi \propto e^{kz}T(t)J_0(kr),$$

where k is the eigenvalues. Clearly though, the only restriction on k is that it must be a non-negative real number, all of which are a valid solution to Laplace's equation. Therefore, the total solution must be the sum of all of these, or since k is valid within an interval, we have the integral form

$$\varphi = \int_0^\infty e^{kz}T(t)J_0(kr)dk. \tag{6}$$

The temporal component can be solved for using equations of fluid dynamics.

1.2 Temporal Component of the Velocity Potential

$$\left(\frac{\partial \varphi}{\partial t} + (g\eta + \Phi)\right)\bigg|_{z=0} = 0$$

by taking the time derivative and using the substitution that

$$\frac{\partial \eta}{\partial t} = \frac{\partial \varphi}{\partial z}\bigg|_{z=0}, \tag{7}$$

by ***, we obtain

$$\left(\frac{\partial^2 \varphi}{\partial t^2} + g\frac{\partial \varphi}{\partial z} + \frac{\partial \Phi}{\partial t}\right)\bigg|_{z=0} = 0. \tag{8}$$

In order to solve this partial differential equation, we must write the gravitational potential as an infinite sum of Bessel functions to match the form of φ. To do so, we take the Hankel transform using the fact that it is self-reciprocal, then

$$\begin{aligned}
\frac{\partial \Phi}{\partial t}\bigg|_{z=0} &= \int_0^\infty \left(\mathscr{H}\,\frac{\partial \Phi}{\partial t}\bigg|_{z=0}\right)(k) J_0(kr)k\,dk, \\
&= Gmv^2t\int_0^\infty \left(\mathscr{H}\,\frac{1}{(r^2+v^2t^2)^{3/2}}\right)(k) J_0(kr)k\,dk, ^1 \\
&= Gmv^2t\int_0^\infty \frac{1}{|vt|}e^{-k|vt|}J_0(kr)k\,dk, *** \\
&= Gmv\,\mathrm{sgn}(t)\int_0^\infty e^{-kv|t|}J_0(kr)k\,dk.
\end{aligned}$$

[1] Note that $\int_0^\infty \sqrt{r}(r^2+v^2t^2)^{-3/2}dr = \Gamma^2(3/4)(\pi v^3t^3)^{-1/2} < \infty$ for $t \neq 0$ which is not concerning since this is true almost everywhere, and so the condition in Theorem *** is satisfied.

Now we can substitute this into (8) to obtain,

$$\int_0^\infty J_0(kr)\ddot{T}(t)dk + g\int_0^\infty kJ_0(kr)T(t)dk + Gmv\int_0^\infty \operatorname{sgn}(t)e^{-kv|t|}J_0(kr)k\,dk = 0,$$
$$\text{or,}\quad \int_0^\infty \left[\frac{\ddot{T}(t)}{k} + gT(t) + Gmv\operatorname{sgn}(t)e^{-kv|t|}\right] J_0(kr)k\,dk = 0.$$

But, this is nothing more than the Hankel transform of the differential equation for T. By taking the Hankel transform of both sides we can remove the integral, giving an ordinary differential equation for the time component:

$$\frac{\ddot{T}(t)}{k} + gT(t) + Gmv\operatorname{sgn}(t)e^{-kv|t|} = 0.$$

Clearly, the homogeneous solution is $T(t) = A\cos(\omega_k t) + B\sin(\omega_k t)$ with $\omega_k^2 = gk$. The form of this differential equations suggests the form $T(t) = Ce^{-kv|t|}$ for the particular solution. Substituting this in yields

$$C\left(k^2v^2\operatorname{sgn}^2(t)e^{-kv|t|} + gke^{-kv|t|}\right) + Gmvk\operatorname{sgn}(t)e^{-kv|t|} = 0.$$

Isolating for C

$$\begin{aligned} C &= \frac{-Gmvk\operatorname{sgn}(t)}{k^2v^2 + gk} \\ &= \frac{Gmv}{g}\frac{-\operatorname{sgn}(t)}{1 + kv^2/g} \end{aligned}$$

as the coefficient, and,

$$T(t) = \frac{Gmv}{g}\frac{1}{1 + kv^2/g}\left(-\operatorname{sgn}(t)e^{-kv|t|}\right) + A\cos(\omega_k t) + B\sin(\omega_k t)$$

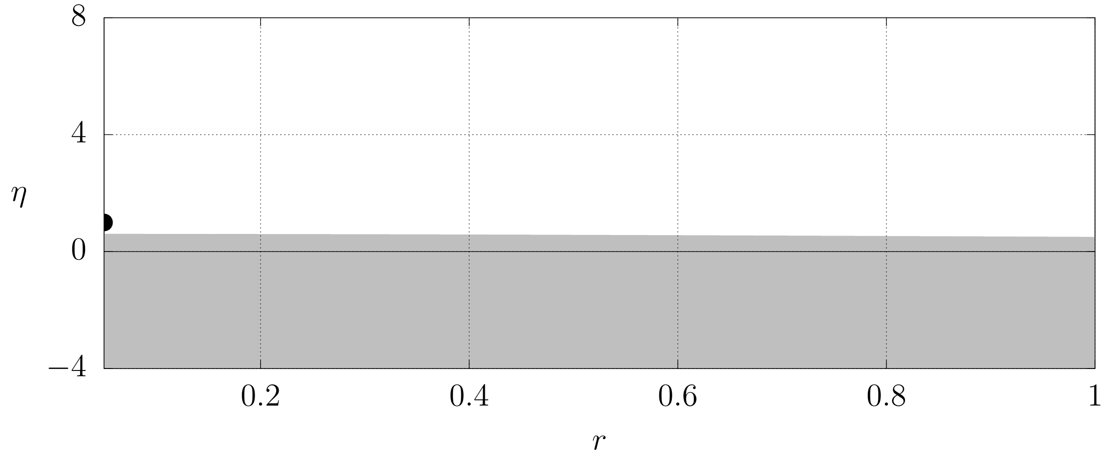
as the full time component of the velocity potential. We can now apply the boundary conditions to find A, and B. Physically, we expect $T(t) \in C^1(-\infty, \infty)$, furthermore, we only expect the sinusoidal terms to contribute at times greater than zero, thus finally, we procure

$$\begin{aligned}
T(t) &= \frac{Gmv}{g} \frac{1}{1 + kv^2/g} \left(-\operatorname{sgn}(t) e^{-kv|t|} + 2H(t) \cos(\omega_k t) \right) dk, \text{ and,} \\
\varphi &= \frac{Gmv}{g} \int_0^\infty \frac{J_0(kr) e^{kz}}{1 + kv^2/g} \left(-\operatorname{sgn}(t) e^{-kv|t|} + 2H(t) \cos(\omega_k t) \right) dk.
\end{aligned}$$

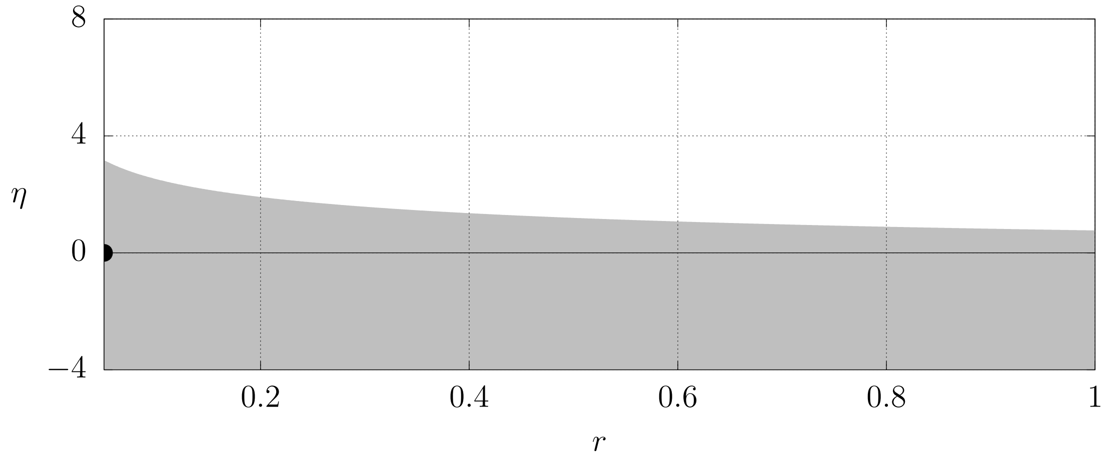
1.3 Deformation of the Surface

From (7) we can isolate for η:

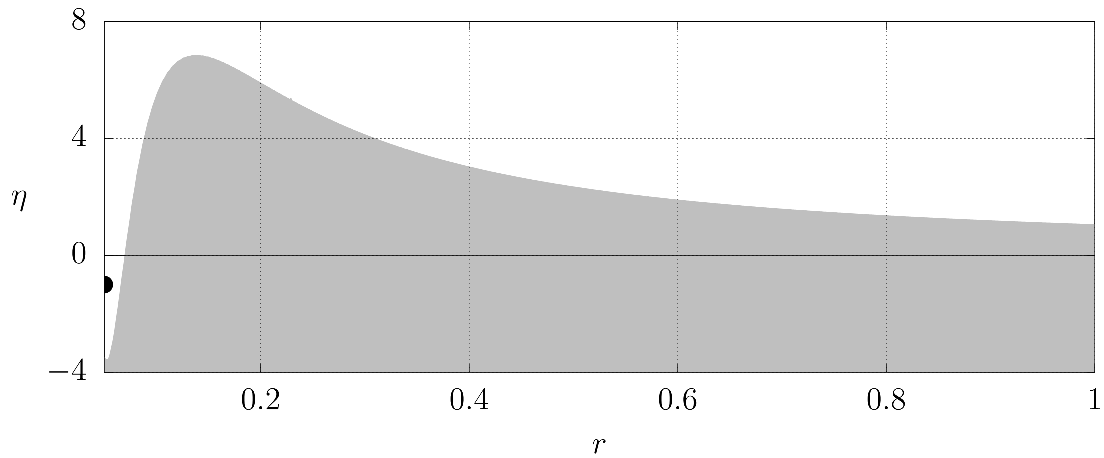
$$\begin{aligned}
\eta &= \int_0^t \left. \frac{\partial \varphi}{\partial z} \right|_{z=0} dt \\
&= \int_0^t \frac{Gmv}{g} \int_0^\infty \frac{k J_0(kr)}{1 + kv^2/g} \left(-\operatorname{sgn}(t) e^{-kv|t|} + 2H(t) \cos(\omega_k t) \right) dk \, dt \\
&= \frac{Gmv}{g} \int_0^\infty \frac{k J_0(kr)}{1 + kv^2/g} \left(\int_0^t -\operatorname{sgn}(t) e^{-kv|t|} + 2H(t) \cos(\omega_k t) \, dt \right) dk \\
&= \frac{Gmv}{g} \int_0^\infty \frac{k J_0(kr)}{1 + kv^2/g} \left(-\operatorname{sgn}(t) \frac{e^{-kv|t|}}{-kv \operatorname{sgn}(t)} + 2H(t) \frac{1}{\omega_k} \sin(\omega_k t) \right) dk \\
&= \frac{Gm}{g} \int_0^\infty \frac{J_0(kr)}{1 + kv^2/g} \left(e^{-kv|t|} + 2H(t) v \sqrt{\frac{k}{g}} \sin(\omega_k t) \right) dk
\end{aligned}$$



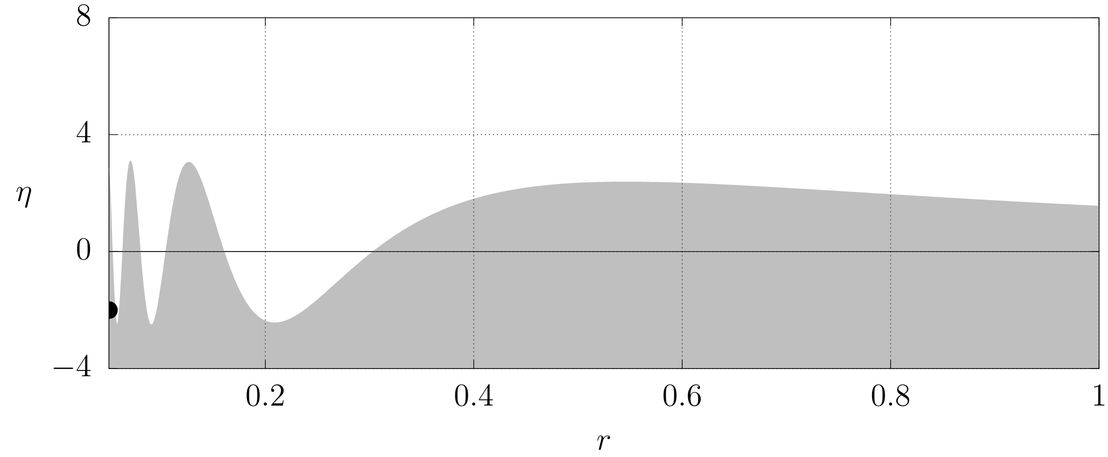
(a) $t = -1$ s.



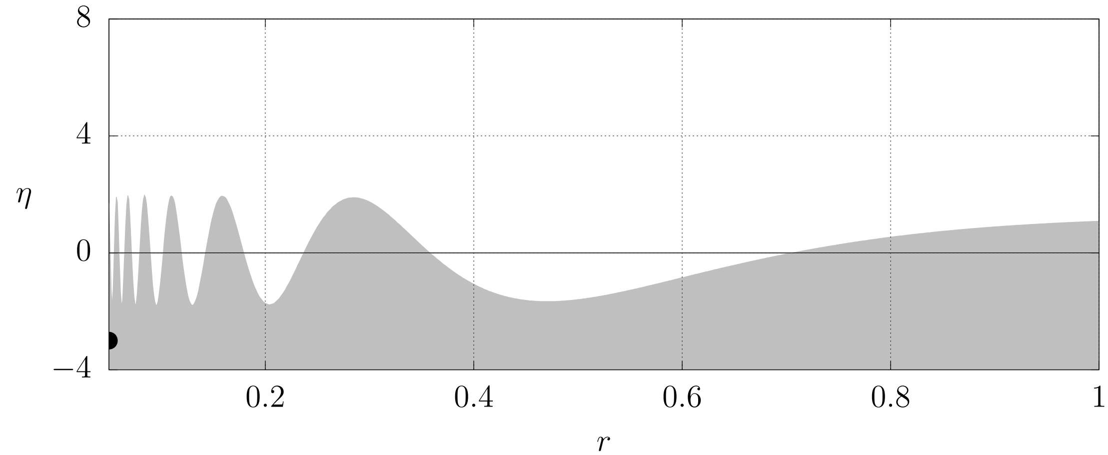
(b) $t = 0$ s.



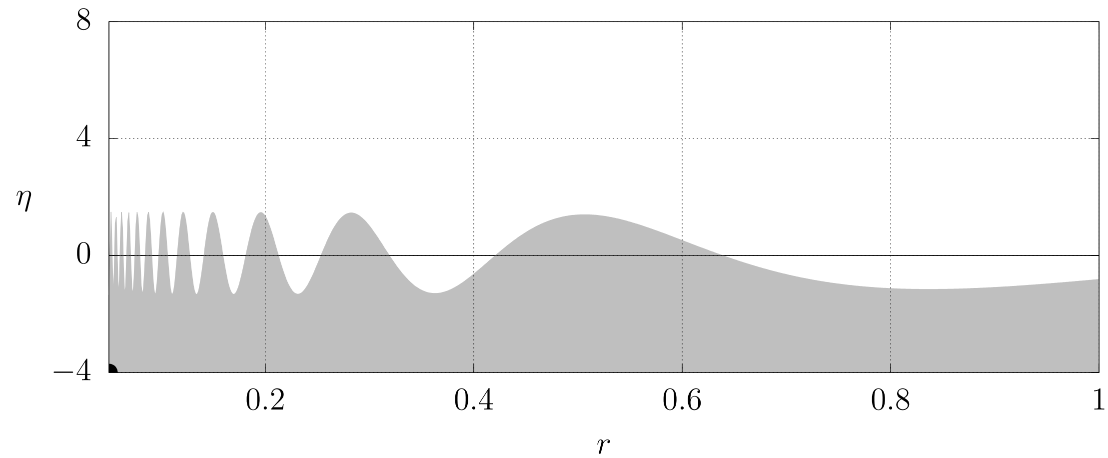
(c) $t = 1$ s.



(d) $t = 2$ s.



(e) $t = 3$ s.



(f) $t = 4$ s.

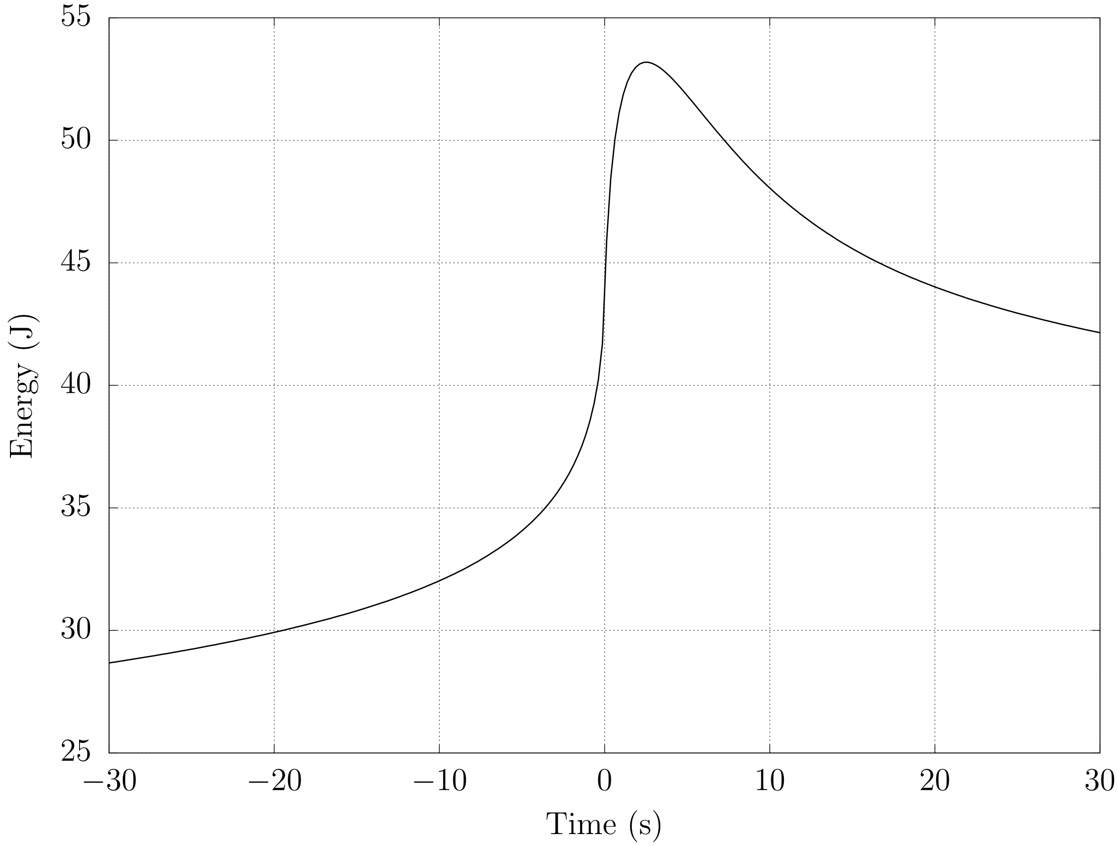


Figure 2: Energy! All variables set to 1.