

# FIREDRAKE NOTES

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## Key Notes to NEVER FORGET

### "Integration by Parts"

This comes up constantly in FEM stuff, when expressing problems in variational form. The usual spiel is to say "multiply by a test function  $v$  and then integrate by parts", to obtain the desired form. This hides a number of key subtle steps that otherwise look like magic.

First to note is that "integrate by parts" really means "apply (a corollary of) the Divergence Theorem":

**Theorem 1.1** (Divergence Theorem). If  $\Omega$  is a compact subset of  $\mathbb{R}^N$  with a piecewise smooth boundary  $\partial\Omega = \Gamma$ , and if  $\mathbf{F}$  is a continuously differentiable vector field defined on a neighbourhood of  $\Omega$  then we have:

$$\int_{\Omega} (\nabla \cdot \mathbf{F}) \, d\Omega = \oint_{\Gamma} (\mathbf{F} \cdot \mathbf{n}) \, d\Gamma$$

where  $\mathbf{n}$  is the outward pointing unit normal field of the boundary  $\Gamma$ .

**Corollary 1.1.** Replacing  $\mathbf{F}$  with  $\mathbf{F}g$  in the theorem, where  $g$  is a scalar function, we get:

$$\int_{\Omega} \mathbf{F} \cdot (\nabla g) \, d\Omega + \int_{\Omega} g(\nabla \cdot \mathbf{F}) \, d\Omega = \oint_{\Gamma} g \mathbf{F} \cdot \mathbf{n} \, d\Gamma$$

We can apply this corollary to the LHS (i.e. to the terms involving  $u$ ) to rewrite it as the sum of a different volume integral and a surface integral, which can often be made to vanish by applying boundary conditions.

### Example: Linear Poisson Equation

Let us take an initial easy example of the basic linear Poisson problem:

$$\begin{aligned} (-\Delta u) &= f, \text{ on } \Omega \\ u &= 0, \text{ on } \partial\Omega = \Gamma \end{aligned}$$

We multiply both sides by the test function  $v$  and integrate to obtain:

$$\int_{\Omega} (-\Delta u)v \, d\Omega = \int_{\Omega} f v \, d\Omega$$

Now we apply the Corollary to the LHS (replacing  $\mathbf{F}$  with  $\nabla u$  and  $g$  with  $v$ ) to get:

$$\int_{\Omega} \nabla u \cdot \nabla v \, d\Omega + \oint_{\Gamma} v \nabla u \cdot \mathbf{n} \, d\Gamma = \int_{\Omega} (-\Delta u)v \, d\Omega = \int_{\Omega} f v \, d\Omega$$

The second term in the new LHS is a closed line integral of a grad function and thus equal to the difference of its endpoints, which are the same, hence the term is zero, leaving us with the desired variational form  $a(u, v) = L(v)$ :

$$\int_{\Omega} \nabla u \cdot \nabla v \, d\Omega = \int_{\Omega} f v \, d\Omega$$

**Example: Nonlinear Poisson Equation**

Let's now look at the following nonlinear Poisson problem:

$$\begin{aligned} -\nabla \cdot ((1+u)\nabla u) &= f, \text{ in } \Omega \\ u &= 0, \text{ on } \partial\Omega = \Gamma \end{aligned}$$

We multiply by the test function  $v$  and integrate both sides:

$$\int_{\Omega} \left( -\nabla \cdot ((1+u)\nabla u) \right) v \, d\Omega = \int_{\Omega} f v \, d\Omega$$

Again we apply the Corollary to the LHS (replacing  $\mathbf{F}$  with  $((1+u)\nabla u)$  and  $g$  with  $v$ ) to get:

$$\int_{\Omega} ((1+u)\nabla u) \cdot \nabla v \, d\Omega + \oint_{\Gamma} v \left( ((1+u)\nabla u) \cdot \mathbf{n} \right) d\Gamma = \int_{\Omega} \left( -\nabla \cdot ((1+u)\nabla u) \right) v \, d\Omega = \int_{\Omega} f v \, d\Omega$$

Looking again at the surface integral term on the new LHS, we recall the initial condition  $u = 0$  on  $\Gamma$  and thus this term simplifies to:

$$\oint_{\Gamma} v (\nabla u \cdot \mathbf{n}) \, d\Gamma = 0 \quad (\text{closed line integral of a grad function})$$

Leaving us with the desired variational form  $F(u; v) = 0$ :

$$\int_{\Omega} ((1+u)\nabla u) \cdot \nabla v \, d\Omega = \int_{\Omega} f v \, d\Omega$$

## Solving Navier Stokes in Firedrake

This section is based on material from the Navier Stokes tutorials from [1].

The following is the time-dependent Navier Stokes equation:

$$\rho \left( \frac{\partial u}{\partial t} + u \cdot \nabla u \right) = \nabla \cdot \sigma(u, p) + f \quad (1.1)$$

$$\nabla \cdot u = 0 \quad (1.2)$$

where

$$\sigma(u, p) = 2\mu\epsilon(u) - pI,$$

$$\epsilon(u) = \frac{\nabla u + (\nabla u)^T}{2},$$

$f$  is a given force per unit volume.

We could choose to discretize this in time by replacing the time derivative with a difference quotient. There are complications to such a method though, since it has a "saddle-point" structure, and requires particular preconditioners and iterative methods to solve properly. Instead, we shall use one example of what is known as a *splitting method*, which essentially ‘splits’ the above nonlinear problem into numerous easier problems to be solved in sequence.

### Time-Dependent Navier Stokes via "Chorin's Method (IPCS)"

We shall use a variant of what is known as *Chorin's Method*. This variant, called IPCS, solves three linear problems per timestep:

- P1: We compute a ‘tentative velocity’  $u^*$  by advancing the momentum equation by a midpoint finite difference scheme in time, but using the previous pressure  $p^n$  and linearising the nonlinear convective term using  $u^n$ .
- P2: We use the tentative velocity  $u^*$  to compute the new pressure  $p^{n+1}$ .
- P3: We use the new pressure  $p^{n+1}$  to compute the new velocity  $u^{n+1}$ .

Each of these steps has some individual complexity, so we will discuss them in detail:

#### P1 - Computing the Tentative Velocity

We can express this as a linear variational problem, using the following variational form:

$$\underbrace{\left\langle \frac{\rho(u^* - u^n)}{\delta t}, v \right\rangle + \left\langle \rho u^n \cdot \nabla u^n, v \right\rangle}_{1 \ \& \ 2} + \underbrace{\left\langle \sigma(u^{n+\frac{1}{2}}, p^n), \epsilon(v) \right\rangle + \left\langle p^n n, v \right\rangle_{\partial\Omega} - \left\langle \mu \nabla u^{n+\frac{1}{2}} \cdot n, v \right\rangle_{\partial\Omega}}_{3, \ 4 \ \& \ 5} = \left\langle f^{n+1}, v \right\rangle \quad (1.3)$$

where

$$u^{n+\frac{1}{2}} = \frac{u^n + u^{n+1}}{2} \text{ (arithmetic mean)}$$

Let's break down this variational form a little to see where each term comes from:

Terms 1 & 2 come from the LHS of the first equation in the initial problem. We write  $\frac{u^* - u^n}{\delta t}$  for  $\frac{\partial u}{\partial t}$  and then just integrate after multiplying by a test function  $v$ :

$$\begin{aligned} \rho \left( \frac{\partial u}{\partial t} + u \cdot \nabla u \right) &\rightsquigarrow \int_{\Omega} \frac{\rho(u^* - u^n)}{\delta t} v \, dx + \int_{\Omega} \rho(u \cdot \nabla u) v \, dx \\ &= \underbrace{\left\langle \frac{\rho(u^* - u^n)}{\delta t}, v \right\rangle}_1 + \underbrace{\left\langle \rho u^n \cdot \nabla u^n, v \right\rangle}_2 \end{aligned}$$

Terms 3, 4 & 5 come from the term  $\nabla \cdot \sigma(v, p)$  from the RHS of the initial expression. We multiply by  $v$  and integrate by parts to obtain:

$$- \int_{\Omega} (\nabla \cdot \sigma) v \, dx = \int_{\Omega} \sigma \cdot \nabla v \, dx - \int_{\partial\Omega} v(\sigma \cdot n) \, ds$$

Now expressing in  $\langle \cdot, \cdot \rangle$  notation we get:

$$\begin{aligned} - \langle \nabla \cdot \sigma, v \rangle &= \overbrace{\langle \sigma, \nabla v \rangle}^{= \langle \sigma, \epsilon(v) \rangle \because \sigma \text{ is sym \& } \epsilon(v) \text{ is sym part of } \nabla v} - \langle \sigma \cdot n, v \rangle_{\partial\Omega} \\ (\text{expanding } \sigma) &= \langle \sigma, \epsilon(v) \rangle + \langle pn - \mu \nabla u \cdot n - \underbrace{\mu (\nabla u)^T \cdot n}_{\star}, v \rangle_{\partial\Omega} \\ \left( \begin{array}{l} \text{assuming } \star = 0 \text{ (see below),} \\ \text{and using } \nabla u = \left( \frac{\partial u_j}{\partial x_i} \right)_{ij} \text{ convention} \end{array} \right) &= \underbrace{\langle \sigma, \epsilon(v) \rangle}_3 + \underbrace{\langle pn, v \rangle_{\partial\Omega}}_4 - \underbrace{\langle \mu \nabla u \cdot n, v \rangle_{\partial\Omega}}_5 \end{aligned}$$

We should avoid passing over  $\star$  without remarking on what assumption we are making in allowing it to be zero. If we had been solving a problem with a free boundary we could simply set  $\sigma \cdot n = 0$  on the boundary. However we are computing flow through a channel and hence our flow continues at the end into some 'imaginary' further channel or region, so we need to be a little more careful. We therefore make the assertion that "the derivative of the velocity in the direction of the channel is zero at the outflow"<sup>1</sup>, which is what (assuming the  $\nabla u = \left( \frac{\partial u_j}{\partial x_i} \right)_{ij}$  convention) yields  $\star = 0$ .

<sup>1</sup>[https://fenicsproject.org/pub/tutorial/html/.\\_ftut1009.html](https://fenicsproject.org/pub/tutorial/html/._ftut1009.html)[1]

## P2 - Computing $p^{n+1}$ with the tentative velocity $u^*$

To compute the new pressure  $p^{n+1}$ , we solve the following linear variational problem, wherein  $q$  is a scalar-valued test function from the pressure function space:

$$\langle \nabla p^{n+1}, \nabla q \rangle = \langle \nabla p^n, \nabla q \rangle - \Delta t^{-1} \langle \nabla \cdot u^*, q \rangle \quad (1.4)$$

Again let's provide some intuition for where this comes from. Let us take the Navier Stokes momentum equation in terms of  $u^*$  and  $p^{n+1}$  and subtract it from the same equation expressed in terms of  $u^{n+1}$  and  $p^{n+1}$ :

$$\rho \left( \frac{u^{n+1} - u^*}{\Delta t} + u^n \cdot \nabla u^n \right) - \rho \left( \frac{u^* - u^n}{\Delta t} + u^n \cdot \nabla u^n \right) = \nabla \cdot \sigma(u^n, p^{n+1}) - \nabla \cdot \sigma(u^n, p^n)$$

which simplifies to:

$$\begin{aligned} \rho \frac{u^{n+1} - u^*}{\Delta t} &= \nabla \cdot \sigma(u^n, p^{n+1}) - \nabla \cdot \sigma(u^n, p^n) \\ (\text{by defn of } \sigma) &= (\nabla \cdot p^n I) - (\nabla \cdot p^{n+1} I) \\ &= \nabla p^n - \nabla p^{n+1} \end{aligned}$$

and hence yields:

$$\frac{u^{n+1} - u^*}{\Delta t} + \nabla p^{n+1} - \nabla p^n = 0 \quad (1.5)$$

(the  $\rho$  vanishes inexplicably because hey, physics!)

If we then take the divergence and require the  $\nabla \cdot u^{n+1} = 0$  condition from the second Navier Stokes equation 1.2, we get:

$$\frac{-\nabla \cdot u^*}{\Delta t} + \nabla^2 p^{n+1} - \nabla^2 p^n = 0$$

which is a Poisson problem for  $p^{n+1}$ , yielding the variational problem 1.4 we were hoping to arrive at!

## P3 - Computing the new velocity $u^{n+1}$ with the new pressure $p^{n+1}$

Once again we seek a linear variational problem, this time to solve for  $u^{n+1}$  using the updated pressure  $p^{n+1}$  that we just solved in P2. If we look back at equation 1.5 we see that this provides exactly what we need! Multiplying by a test function  $v$  and integrating yields precisely:

$$\langle u^{n+1}, v \rangle = \langle u^*, v \rangle - \Delta t \langle \nabla(p^{n+1} - p^n), v \rangle \quad (1.6)$$

## References

### MIGRATE TO BIBLIO

- [https://en.wikipedia.org/wiki/Divergence\\_theorem](https://en.wikipedia.org/wiki/Divergence_theorem)
- [https://en.wikipedia.org/wiki/Surface\\_integral](https://en.wikipedia.org/wiki/Surface_integral)
- [http://mathinsight.org/gradient\\_theorem\\_line\\_integrals](http://mathinsight.org/gradient_theorem_line_integrals)



## References

- [1] H. P. Langtangen. Solving PDEs in Python: the FEniCS tutorial I. SpringerOpen, 2016.