Bayesian model selection and averaging with mombf

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The mombf package implements Bayesian model selection (BMS) and model averaging (BMA) for several regression (linear, asymmetric linear, median and quantile regression) and mixture models. This is the main package implementing the family of non-local prior (NLP) distributions (briefly reviewed here, see Johnson and Rossell (2010, 2012) for a more detailed treatment), although other priors are also implemented (e.g. Zellner's prior in regression, Normal-IWishart priors for mixtures). The main features are:

- Density, cumulative density, quantiles and random numbers for NLPs
- BMS in linear regression (Section 1, Johnson and Rossell (2010, 2012)).
- BMA in linear regression (Section 4, Rossell and Telesca (2016)).
- Exact BMS and BMA under orthogonal and block-diagonal regression (Section 5, Papaspiliopoulos and Rossell (2016)).
- BMS and BMA for certain generalized linear models (Section 6, Johnson and Rossell (2012); Rossell et al. (2013))
- BMS in linear regression with non-normal residuals (Rossell and Rubio, 2016). Particular cases are Bayesian versions of asymmetric least squares, median and quantile regression.
- BMS for mixture models (Section 7, currently only Normal mixtures), see Fúquene et al. (2018).

This manual introduces some basic notions underlying NLPs and illustrates the use of R functions implementing the main operations required for model selection and averaging. Most of these are internally implemented in C++ so, while they are not optimal in any sense they are designed to be minimally scalable to large sample sizes n and high dimensions (large p).

1 Basics on non-local priors

The basic motivation for NLPs is what one may denominate the model separation principle. The idea is quite simple, suppose we are considering a (possibly infinite) set of probability models M_1, M_2, \ldots for an observed dataset y, if these models overlap then it becomes hard to tell which of them generated y. The notion is important because the vast majority of applications consider nested models: if say M_1 is nested within M_2 then these two models are not well-separated. Intuitively, if y are truly generated from M_1 then M_1 will receive high integrated likelihood however that for M_2 will also be relatively large given that M_1 is contained in M_2 . We remark that the notion remains valid when none of the posed models are true, in that case M_1 is the model of smallest dimension minimizing Kullback-Leibler divergence to the data-generating distribution of y. A usual mantra is that performing Bayesian model selection via posterior model probabilities (equivalently, Bayes factors) automatically incorporates Occam's razor, e.g. M_1 will eventually be favoured over M_2 as the sample size $n \to \infty$. This statement is correct but can be misleading: there is no guarantee that parsimony is enforced to an adequate extent, indeed it turns out to be insufficient in many practical situations even for small p. This issue is exacerbated for large p to the extent that one may even loose consistency of posterior model probabilities (Johnson and Rossell, 2012) unless sufficiently strong sparsity penalities are introduced into the model space prior.

Intuitively, NLPs induce a probabilistic separation between the considered models which, aside from being philosophically appealing (to us), one can show mathematically leads to stronger parsimony. When we compare two nested models and the smaller one is true the resulting BFs converge faster to 0 than when using conventional priors and, when the larger model is true, they present the usual exponential convergence rates in standard procedures. That is, the extra parsimony induced by NLPs is data-dependent, as opposed to inducing sparsity by formulating sparse model prior probabilities or increasingly vague prior distributions on model-specific parameters.

To fix ideas we first give the general definition of NLPs and then proceed to show some examples. Let $y \in \mathcal{Y}$ be the observed data with density $p(y \mid \theta)$ where $\theta \in \Theta$ is the (possibly infinite-dimensional) parameter. Suppose we are interested in comparing a series of models M_1, M_2, \ldots with corresponding parameter spaces $\Theta_k \subseteq \Theta$ such that $\Theta_j \cap \Theta_k$ have zero Lebesgue measure for $j \neq k$ and, for precision, there exists an l such that $\Theta_l = \Theta_j \cap \Theta_k$ so that the whole parameter space is covered.

Definition 1 A prior density $p(\theta \mid M_k)$ is a non-local prior under M_k iff

 $\lim p(\theta \mid M_k) = 0$ as $\theta \to \theta_0$ for any $\theta_0 \in \Theta_i \subset \Theta_k$.

In words, $p(\theta \mid M_k)$ vanishes as θ approaches any value that would be consistent with a submodel M_j . Any prior not satisfying Definition 1 is a local prior (LP). As a canonical example, suppose that $y = (y_1, \ldots, y_n)$ with independent $y_i \sim N(\theta, \phi)$ and known ϕ , and that we entertain the two following models:

$$M_1: \theta = 0$$
$$M_2: \theta \neq 0$$

Under M_1 all parameter values are fully specified, the question is thus reduced to setting $p(\theta \mid M_2)$. Ideally this prior should reflect one's knowledge or beliefs about likely values of θ , conditionally on the fact that $\theta \neq 0$. The left panel in Figure 1 shows two LPs, specifically the unit information prior $\theta \sim N(0,1)$ and a heavy-tailed alternative $\theta \sim \text{Cachy}(0,1)$ as recommended by Jeffreys. These assign $\theta = 0$ as their most likely value a priori, even though $\theta = 0$ is not even a possible value under M_2 , which we view as philosophically unappealing. The right panel shows three NLPs (called MOM, eMOM and iMOM, introduced below). Their common defining feature is their vanishing as $\theta \to 0$, thus probabilistically separating M_2 from M_1 or, to borrow terminology from the stochastic processes literature, inducing a repulsive force between M_1 and M_2 . As illustrated in the figure beyond this defining feature the user is free to choose any other desired property, e.g. the speed at which $p(\theta \mid M_2)$ vanishes at the origin, prior dispersion, tail thickness or in multivariate cases the prior dependence structure.

Once the NLP has been specified inference proceeds as usual, e.g. posterior model probabilities are

$$p(M_k \mid y) = \frac{p(y \mid M_k)p(M_k)}{\sum_{j} p(y \mid M_j)p(M_j)}$$
(1)

where $p(y \mid M_k) = \int p(y \mid \theta) dP(\theta \mid M_k)$ is the integrated likelihood under M_k and $p(M_k)$ the prior model probability. Similarly, inference on parameters can be carried out conditional on any chosen model via $p(\theta \mid M_k, y) \propto p(y \mid \theta) p(\theta \mid M_k)$ or via Bayesian model averaging $p(\theta \mid y) = \sum_k p(\theta \mid M_k, y) p(M_k \mid y)$. A useful construction (Rossell and Telesca, 2016) is that any NLP can be expressed as

$$p(\theta \mid M_k) = \frac{p(\theta \mid M_k)}{p^L(\theta \mid M_k)} p^L(\theta \mid M_k) = d_k(\theta) p^L(\theta \mid M_k), \tag{2}$$

where $p^L(\theta \mid M_k)$ is a LP and $d_k(\theta) = p(\theta \mid M_k)/p^L(\theta \mid M_k)$ a penalty term. Simple algebra shows that

$$p(y \mid M_k) = p^L(y \mid M_k)E^L(d_k(\theta) \mid M_k, y), \tag{3}$$

where $E^L(d_k(\theta) \mid M_k, y) = \int d_k(\theta) dP^L(\theta \mid M_k, y)$ is the posterior mean of the penalty term under the underlying LP. That is, the integrated likelihood under a NLP is equal to that under a LP times the posterior expected penalty under that LP. The construction allows to use NLPs in any situation where LPs are already implemented, all one needs is $p^L(y \mid M_k)$ or an estimate thereof and posterior samples under $p^L(\theta \mid y, M_k)$. We remark that most functions in mombf do not rely on construction (2) but instead work directly with NLPs, as this is typically more efficient computationally. For instance, there are closed-form expressions and Laplace approximations to $p(y \mid M_k)$ (Johnson and Rossell, 2012), and one may sample from $p(\theta \mid M_k, y)$ via simple latent truncation representations (Rossell and Telesca, 2016).

Up to this point we kept the discussion as generic as possible, Section 2 illustrates NLPs for variable selection and mixture models. For extensions to other settings see Consonni and La Rocca (2010) for directed acyclic graphs under an objective Bayes framework, Chekouo et al. (2015) for gene regulatory networks, Collazo et al. (2016) for chain event graphs, or ? for finite mixture models. We also remark that this manual focuses mainly on practical aspects. Readers interested in theoretical NLP properties should see Johnson and Rossell (2010) and Rossell and Telesca (2016) for an asymptotic characterization under asymptotically Normal models with fixed $\dim(\Theta)$, essentially showing that $E^{L}(d_{k}(\theta) \mid M_{k}, y)$ leads to stronger parsimony, ? for similar results in mixture models, and Rossell and Rubio (work in progress) for robust linear regression with non-normal residuals where data may be generated by a model other than those under consideration (M-complete). Regarding high-dimensional results Johnson and Rossell (2012) prove that under certain linear regression models $p(M_t \mid y) \stackrel{P}{\longrightarrow} 1$ as $n \to \infty$ where M_t is the data-generating truth when using NLPs and $p = O(n^{\alpha})$ with $\alpha < 1$. The authors also proved the conceptually stronger result that $p(M_t \mid y) \stackrel{P}{\longrightarrow} 0$ under LPs, which implies that NLPs are a necessary condition for strong consistency in high dimensions (unless extra parsimony is induced via $p(M_k)$ or increasingly diffuse $p(\theta \mid M_k)$ as n grows, but this may come at a loss of signal detection power). Shin et al. (2015) extend the consistency result to ultra-high linear regression with $p = O(e^{n^{\alpha}})$ with $\alpha < 1$ under certain specific NLPs.

2 Some default non-local priors

2.1 Regression models

Definition 1 in principle allows one to define NLPs in any manner that is convenient, as long as $p(\theta \mid M_k)$ vanishes for any value θ_0 that would be consistent with a submodel of M_k . mombf implements some simple priors that lead to convenient implementation and interpretation while offering a reasonable modelling flexibility, but naturally we encourage everyone to come up with more sophisticated alternatives as warranted by their specific problem at hand. It is important to distinguish between two main strategies to define NLPs, namely imposing additive vs. product penalties. Additive NLPs were historically the first to be introduced (Johnson and Rossell, 2010) and primarily aimed to compare only two models, whereas product NLPs were introduced later on (Johnson and Rossell, 2012) for the more general setting where considers multiple models. Throughout let $\theta = (\theta_1, \ldots, \theta_p) \in \mathbb{R}^p$ be a vector of regression coefficients and ϕ a dispersion parameter such as the residual variance in linear regression.

Suppose first that we wish to test $M_1: \theta = (0, ..., 0)$ versus $M_2: \theta \neq (0, ..., 0)$. An additive NLP takes the form $p(\theta \mid M_k) = d(q(\theta))p^L(\theta \mid M_k)$, where $q(\theta) = \theta'V\theta$ for some positive-definite $p \times p$ matrix V, the penalty $d(q(\theta)) = 0$ if and only if $q(\theta) = 0$ and $p^L(\theta \mid M_k)$ is an arbitrary LP with the only restriction that $p(\theta \mid M_k)$ is proper. For instance,

$$p_{M}(\theta \mid \phi, M_{k}) = \frac{\theta' V \theta}{p \tau \phi} N(\theta; 0, \tau \phi V^{-1})$$

$$p_{E}(\theta \mid \phi, M_{k}) = c_{E} e^{-\frac{\tau \phi}{\theta' V \theta}} N(\theta; 0, \tau \phi V^{-1})$$

$$p_{I}(\theta \mid \phi, M_{k}) = \frac{\Gamma(p/2)}{|V|^{\frac{1}{2}} (\tau \phi)^{\frac{p}{2}} \Gamma(\nu/2) \pi^{p/2}} (\theta' V \theta)^{-\frac{\nu+p}{2}} e^{-\frac{\tau \phi}{\theta' V \theta}}$$

$$(4)$$

are the so-called moment (MOM), exponential moment (eMOM) and inverse moment (iMOM) priors, respectively, and c_E is the moment generating function of an inverse chi-square random variable evaluated at -1. By default V = I, but naturally other choices are possible.

Suppose now that we wish to consider all 2^p models arising from setting individual elements in θ to 0. Product NLPs are akin to (4) but now the

penalty term $d(\theta)$ is a product of univariate penalties.

$$p_{M}(\theta \mid \phi, M_{k}) = \prod_{i \in M_{k}} \frac{\theta_{i}^{2}}{\tau \phi_{k}} N(\theta_{i}; 0, \tau \phi_{k})$$

$$p_{E}(\theta \mid \phi, M_{k}) = \prod_{i \in M_{k}} \exp\left\{\sqrt{2} - \frac{\tau \phi_{k}}{\theta_{i}^{2}}\right\} N(\theta_{i}; 0, \tau \phi_{k}),$$

$$p_{I}(\theta \mid \phi, M_{k}) = \prod_{i \in M_{k}} \frac{(\tau \phi_{k})^{\frac{1}{2}}}{\sqrt{\pi} \theta_{i}^{2}} \exp\left\{-\frac{\tau \phi_{k}}{\theta_{i}^{2}}\right\}. \tag{5}$$

This implies that $d(\theta) \to 0$ whenever any individual $\theta_i \to 0$, in contrast with (4) which requires the whole vector $\theta = 0$. More generally, one can envision settings requiring a combination of additive and product penalties. For instance in regression models for continuous predictors product penalties are generally appropriate, but for categorical predictors one would like to either include or exclude all the corresponding coefficients simultaneously, in this sense additive NLPs resemble group-lasso type penalties and have the nice property of being invariant to the chosen reference category. At this moment mombf primarily implements product NLPs and in some cases additive NLPs, we plan to incorporate combined product and addivite penalties in the future.

Figure 1 displays the prior densities in the univariate case, where (4) and (5) are equivalent. p_M induces a quadratic penalty as $\theta \to 0$, and has the computational advantage that for posterior inference the penalty can often be integrated in closed-form, as it simply requires second order moments. p_E and p_I vanish exponentially fast as $\theta \to 0$, inducing stronger parsimony in the Bayes factors than p_M . This exponential term converges to 1 as $q(\theta)$ increases, thus the eMOM has Normal tails and the iMOM can be easily checked to have tails proportional to those of a Cauchy. Thick tails can be interesting to address the so-called *information paradox*, namely that the posterior probability of the alternative model converges to 1 as the residual sum of squares from regressing y on a series of covariates converges to 0 (Liang et al., 2008; Johnson and Rossell, 2010), although in our experience this is often not an issue unless n is extremely small. The priors above can be extended to include nuisance regression parameters that are common to all models, and also to consider higher powers $q(\theta)^r$ for some r>1, but for simplicity we restrict attention to (4).

We now briefly discuss how to set the prior dispersion τ and propose default values, though we encourage everyone to think which values are more appropriate for the current problem at hand. By default we set τ so that $P(|\theta/\sqrt{\phi}| > 0.2) = 0.99$, that is the signal-to-noise ratio $|\theta_j|/\sqrt{\phi}$ is a priori

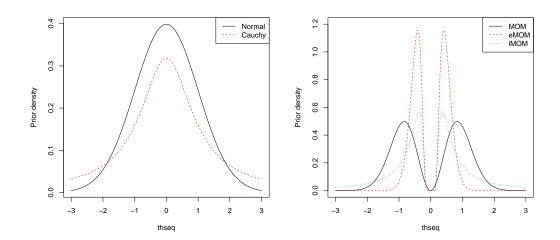


Figure 1: Priors for θ under a model $M_2: \theta \neq 0$. Left: local priors. Right: non-local priors

expected to be > 0.2. The reason for choosing the 0.2 threshold is that in many applications smaller signals are not practically relevant (e.g. the implied contribution to the R^2 coefficient is small). The function priorp2g facilitates finding τ for any given threshold. Other useful functions are pmom, pemom and pimom for distribution functions, and qmom, qemom and qimom for quantiles. Prior densities can be evaluated and plotted as shown below.

```
> library(mombf)
> thseq <- seq(-3,3,length=1000)
> plot(thseq,dnorm(thseq),type='l',ylab='Prior density')
> lines(thseq,dt(thseq,df=1),lty=2,col=2)
> legend('topright',c('Normal','Cauchy'),lty=1:2,col=1:2)
> thseq <- seq(-3,3,length=1000)
> plot(thseq,dmom(thseq,tau=.348),type='l',ylab='Prior density',ylim=c(0,1.2))
> lines(thseq,demom(thseq,tau=.119),lty=2,col=2)
> lines(thseq,dimom(thseq,tau=.133),lty=3,col=4)
> legend('topright',c('MOM','eMOM','iMOM'),lty=1:3,col=c(1,2,4))
```

Another way to elicit τ is to mimic the Unit Information Prior and set the prior variance to 1, for the MOM prior this leads to the same τ as the earlier rule $P(|\theta/\sqrt{\phi}| > 0.2) = 0.99$.

Another helpful function for prior elicitation of τ is mode2g, which returns the τ value associated to a given prior mode $m = \arg\max_{|\theta|} p(|\theta|)$, for instance m = 0.4 in our example below. The intuition is that NLPs place

most of the prior mass on $|\theta| > m$, which implicitly means we have little interest in detecting $|\theta| < m$.

```
> prior.mode <- .4^2
> taumom <- mode2g(prior.mode,prior='normalMom')
> tauimom <- mode2g(prior.mode,prior='iMom')
> taumom
[1] 0.08
> tauimom
[1] 0.16
```

2.2 Mixture models

Let $y = (y_1, \ldots, y_n)$ be the observed data, where n is the sample size. Denote by M_k a mixture model with k components and density

$$p(y_i \mid \vartheta_k, M_k) = \sum_{j=1}^k \eta_j p(y_i \mid \theta_j), \tag{6}$$

where $\eta = (\eta_1, \dots, \eta_k)$ are the mixture weights, $\theta = (\theta_1, \dots, \theta_k)$ componentspecific parameters, and $\vartheta_k = (\eta, \theta)$ denotes the whole parameter vector. For instance, for Normal mixtures $\theta_j = (\mu_j, \Sigma_j)$, where μ_j is the mean and Σ_j the covariance matrix. A standard prior choice for such location-scale mixtures is the Normal-IW-Dir prior

$$\tilde{p}(\vartheta_k \mid M_k) = \left[\prod_{j=1}^k N(\mu_j; \mu_0, g\Sigma_j) \mathrm{IW}(\Sigma_j; \nu_0, S_0) \right] \mathrm{Dir}(\eta; \tilde{q}) \tag{7}$$

for some given prior parameters $(\mu_0, g, \nu_0, S_0, \tilde{q})$. This choice defines a local prior, which we view as unappealing in this setting: it assigns positive prior density to two components having the same parameters and, if $\tilde{q} \leq 1$, also to having some mixture weights $\eta_i = 0$.

A non-local prior on ϑ_k must penalize parameter values that would make the k-component mixture equivalent to a mixture with < k components. As formalized in Fúquene et al. (2018) for any generically identifiable mixture this is achieved by penalizing zero weights $(\eta_j = 0)$ and any two components having the same parameter values (i.e. $\theta_j = \theta_l$ for $j \neq l$). Specifically, for location scale mixtures where $\theta_j = (\mu_j, \Sigma_j)$ we consider the MOM-IW-Dir prior $p(\vartheta_k \mid M_k) =$

$$C_k^{-1} \left[\prod_{j < l} (\mu_j - \mu_l)' A^{-1} (\mu_j - \mu_l) \right] \left[\prod_{j=1}^k N(\mu_j; \mu_0, gA) \text{IW}(\Sigma_j; \nu_0, S_0) \right] \text{Dir}(\eta; q)$$
(8)

where C_k is the prior normalization constant (implemented in mombf), $A^{-1} = \frac{1}{k} \sum_{j=1}^{k} \Sigma_{j}^{-1}$ the average precision matrix, (μ_0, g, ν_0, S_0) are given prior parameters (mombf implements defaults for all of them) and one must have q > 1 for (8) to define a non-local prior.

Computing Bayes factors and posterior model probabilities turns out to be surprisingly easy due to two results by Fúquene et al. (2018). First, the integrated likelihood under a MOM-IW-Dir prior

$$p(y \mid M_k) = \tilde{p}(y \mid M_k) \int d_k(\vartheta_k) \tilde{p}(\vartheta_k \mid y, M_k),$$

where $\tilde{p}(y \mid M_k)$ is the integrated likelihood associated to the Normal-IW-Dir prior, $\tilde{p}(\vartheta_k \mid y, M_k) \propto p(y \mid \vartheta_k, M_k) \tilde{p}(\vartheta_k \mid M_k)$ the corresponding posterior and $d_k(\vartheta_k) =$

$$C_k^{-1} \left[\prod_{j < l} (\mu_j - \mu_l)' A^{-1} (\mu_j - \mu_l) \right] \left[\prod_{j=1}^k N(\mu_j; \mu_0, gA) \operatorname{IW}(\Sigma_j; \nu_0, S_0) \right] \frac{\operatorname{Dir}(\eta; q)}{\operatorname{Dir}(\eta; \tilde{q})}$$
(9)

is a penalty term straightforward to estimate provided one has access to posterior samples from $\tilde{p}(\vartheta_k \mid y, M_k)$.

The second result is the following simple connection between Bayes factors and empty-component posterior probabilities:

$$\frac{\tilde{p}(y \mid M_{k-1})}{\tilde{p}(y \mid M_k)} = \frac{1}{ka_k} \sum_{j=1}^k P(n_j = 0 \mid y, M_k)$$

where $a_k = \Gamma(kq)\Gamma(n+kq-q)/(\Gamma(kq-q)\Gamma(n+kq))$ and n_j is the number of individuals allocated to cluster j. The right-hand side requires the average marginal posterior probability of one cluster being empty $P(n_j = 0 \mid y, M_k)$. Given T posterior samples $\vartheta_k^{(1)}, \ldots, \vartheta_k^{(T)}$ this can be estimated in a Rao-Blackwellized fashion as $\hat{P}(n_j = 0 \mid y, M_k) =$

$$\frac{1}{T} \sum_{t=1}^{T} P(n_j = 0 \mid y, \vartheta_k^{(t)}, M_k) = \frac{1}{T} \sum_{t=1}^{T} \prod_{i=1}^{n} P(z_i \neq j \mid y, \vartheta_k^{(t)}, M_k),$$

where $z_i \in \{1, ..., k\}$ for i = 1, ..., n are latent cluster indicators. mombf implements this computational strategy, see Section 7 for examples.

We remark that $\hat{P}(n_j = 0 \mid y, M_k)$ only requires cluster allocation probabilities $P(z_i \neq j \mid y, \vartheta_k^{(t)}, M_k)$. The latter are readily available as a byproduct of any standard MCMC algorithm and for any parametric mixture model. That is relative to running a standard MCMC one can obtain

 $\hat{P}(n_j = 0 \mid y, M_k)$ at very little cost, further the estimator is readily applicable to non-conjugate models (in constrast to other estimators, see Marin (2008)).

3 Variable selection for linear models

The main function for model selection is modelSelection, which returns model posterior probabilities under linear regression models allowing for Normal, asymmetric Normal, Laplace and asymmetric Laplace residuals. A second interesting function is nlpMarginal, which computes the integrated likelihood for a given model under the same settings. We illustrate their use with a simple simulated dataset. Let us generate 100 observations for the response variable and 3 covariates, where the true regression coefficient for the third covariate is 0.

```
> set.seed(2011*01*18)
> x <- matrix(rnorm(100*3),nrow=100,ncol=3)
> theta <- matrix(c(1,1,0),ncol=1)
> y <- x %*% theta + rnorm(100)</pre>
```

To start with we assume Normal residuals (the default). We need to specify the prior distribution for the regression coefficients, the model space and the residual variance. We specify a product iMOM prior on the coefficients with prior dispersion tau=.131, which targets the detection of standardized effect sizes above 0.2. Regarding the model space we use a Beta-binomial(1,1) prior (Scott and Berger, 2010). Finally, for the residual variance we set a minimally informative inverse gamma prior. For defining other prior distributions see the help for msPriorSpec (e.g. momprior, emomprior and zellnerprior can be used to define MOM, eMOM and Zellner priors, respectively).

```
> priorCoef <- imomprior(tau=.131)
> priorDelta <- modelbbprior(alpha.p=1,beta.p=1)
> priorVar <- igprior(.01,.01)</pre>
```

modelSelection enumerates all models when its argument enumerate is set to TRUE, otherwise it uses a Gibbs sampling scheme to explore the model space (saved in the slot postSample). It returns the visited model with highest posterior probability and the marginal posterior inclusion probabilities for each covariate (when using Gibbs sampling these are estimated via Rao-Blackwellization to improve accuracy).

```
> fit1 <- modelSelection(y=y, x=x, center=FALSE, scale=FALSE,
+ priorCoef=priorCoef, priorDelta=priorDelta, priorVar=priorVar)</pre>
```

```
Computing posterior probabilities..... Done.
> fit1$postMode
x1 x2 x3
 1 1 0
> fit1$margpp
        x1
                   x2
1.00000000 1.00000000 0.04138239
> postProb(fit1)
  modelid family
      1,2 normal 9.586176e-01
   1,2,3 normal 4.138239e-02
2
       1 normal 2.765878e-13
       2 normal 9.613353e-14
3
6
      1,3 normal 1.221807e-15
7
      2,3 normal 6.836394e-16
         normal 8.690008e-20
1
5
        3 normal 2.871131e-22
> fit2 <- modelSelection(y=y, x=x, center=FALSE, scale=FALSE,
+ priorCoef=priorCoef, priorDelta=priorDelta, priorVar=priorVar,
+ enumerate=FALSE, niter=1000)
Greedy searching posterior mode... Done.
Running Gibbs sampler..... Done.
> fit2$postMode
x1 x2 x3
 1 1 0
> fit2$margpp
                              x3
        x1
                   x2
1.00000000 1.00000000 0.04138239
> postProb(fit2,method='norm')
  modelid family
     1,2 normal 0.95861761
    1,2,3 normal 0.04138239
> postProb(fit2,method='exact')
  modelid family
     1,2 normal 0.97
    1,2,3 normal 0.03
```

Enumerating models...

The highest posterior probability model is the simulation truth, indicating that covariates 1 and 2 should be included and covariate 3 should be excluded. fit1 was obtained by enumerating the $2^3 = 8$ possible models, whereas fit2 ran 1,000 Gibbs iterations, delivering very similar results. postProb estimates posterior probabilities by renormalizing the probability of each model conditional to the set of visited models when method='norm' (the default), otherwise it uses the proportion of Gibbs iterations spent on each model.

Below we run model Selection again but now using Zellner's prior, with prior dispersion set to obtain the so-called Unit Information Prior. The posterior mode is still the data-generating truth, albeit its posterior probability has decreased substantially. This illustrates the core issue with NLPs: they tend to concentrate more posterior probability around the true model (or that closest in the Kullback-Leibler sense). This difference in behaviour relative to LPs becomes amplified as the number of considered models becomes larger, which may result in the latter giving a posterior probability that converges to 0 for the true model (Johnson and Rossell, 2012).

```
> priorCoef <- zellnerprior(tau=nrow(x))</pre>
> fit3 <- modelSelection(y=y, x=x, center=FALSE, scale=FALSE, niter=10^2,
+ priorCoef=priorCoef, priorDelta=priorDelta, priorVar=priorVar,
+ method='Laplace')
Enumerating models...
Computing posterior probabilities..... Done.
> postProb(fit3)
  modelid family
      1,2 normal 7.214937e-01
    1,2,3 normal 2.785063e-01
8
        1 normal 1.079508e-13
2
3
        2 normal 3.565310e-14
6
      1.3 normal 1.096444e-14
7
      2,3 normal 3.827255e-15
1
          normal 3.640151e-20
5
        3 normal 1.394484e-21
```

Finally, we illustrate how to relax the assumption that residuals are Normally distributed. We may set the argument family to 'twopiecenormal', 'laplace' or 'twopiecelaplace' to allow for asymmetry (for two-piece Normal and two-piece Laplace) or thicker-than-normal tails (for Laplace and asymmetric Laplace). For instance, the maximum likelihood estimator under Laplace residuals is equivalent to median regression and under asymmetric Laplace residuals to quantile regression, thus these options can be interpreted as robust alternatives to Normal residuals. A nice feature is that regression

coefficients can still be interpreted in the usual manner. These families add flexibility while maintaining analytical and computational tractability, e.g. they lead to convex optimization and efficient approximations to marginal likelihoods, and additionally to robustness we have found they can also lead to increased sensitivity to detect non-zero coefficients. Alas, computations under Normal residuals are inevitably faster, hence whenever this extra flexibility is not needed it is nice to be able to fall back onto the Normal family, particularly when p is large. modelSelection and nlpMarginal incorporate this option by setting family=='auto', which indicates that the residual distribution should be inferred from the data. When p is small a full model enumeration is conducted, but when p is large the Gibbs scheme spends most time on models with high posterior probability, thus automatically focusing on the Normal family when it provides a good enough approximation and resorting to one of the alternatives when warranted by the data.

For instance, in the example below there's roughly 0.95 posterior probability that residuals are Normal, hence the Gibbs algorithm would spend most time on the (faster) Normal model. The two-piece Normal and two-piece Laplace (also known as asymmetric Laplace) incorporate an asymmetry parameter $\alpha \in [-1,1]$, where $\alpha = 0$ corresponds to the symmetric case (i.e. Normal and Laplace residuals). We set a NLP on $\operatorname{atanh}(\alpha) \in (-\infty, infty)$ so that under the asymmetric model we push prior mass away from $\alpha = 0$, which intuitively means we are interested in finding significant departures from asymmetry and otherwise we fall back onto the simpler symmetric case.

```
> priorCoef <- imomprior(tau=.131)</pre>
> fit4 <- modelSelection(y=y, x=x, center=FALSE, scale=FALSE,
+ priorCoef=priorCoef, priorDelta=priorDelta, priorVar=priorVar,
+ priorSkew=imomprior(tau=.131),family='auto')
Enumerating models...
Computing posterior probabilities..... Done.
> head(postProb(fit4))
  modelid
                    family
4
      1,2
                    normal 9.472921e-01
8
     1,2,3
                    normal 4.089348e-02
20
       1,2
                   laplace 8.904186e-03
12
      1,2 twopiecenormal 2.693917e-03
24
     1,2,3
                   laplace 1.024381e-04
28
       1,2 twopiecelaplace 6.088958e-05
```

All examples above use modelSelection, which is based on product NLPs (5). mombf also provides some (limited) functionality for additive NLPs (4). The code below contains an example based on the Hald data, which has

```
We load the data and fit a linear regression model.
> data(hald)
> dim(hald)
[1] 13 5
> lm1 <- lm(hald[,1] ~ hald[,2] + hald[,3] + hald[,4] + hald[,5])
> summary(lm1)
Call:
lm(formula = hald[, 1] ~ hald[, 2] + hald[, 3] + hald[, 4] +
Residuals:
    Min
             10 Median
                             3Q
                                    Max
-3.1750 -1.6709 0.2508 1.3783 3.9254
Coefficients:
            Estimate Std. Error t value Pr(>|t|)
(Intercept) 62.4054
                        70.0710
                                 0.891
                                          0.3991
hald[, 2]
                         0.7448
                                  2.083
                                          0.0708 .
              1.5511
hald[, 3]
                                          0.5009
              0.5102
                         0.7238
                                  0.705
hald[, 4]
              0.1019
                         0.7547
                                  0.135
                                          0.8959
             -0.1441
                         0.7091 -0.203
hald[, 5]
                                          0.8441
Signif. codes: 0 âĂŸ***âĂŹ 0.001 âĂŸ**âĂŹ 0.01 âĂŸ*âĂŹ 0.05 âĂŸ.âĂŹ 0.1 âĂŸ âĂŹ 1
Residual standard error: 2.446 on 8 degrees of freedom
Multiple R-squared: 0.9824,
                                    Adjusted R-squared:
                                                         0.9736
F-statistic: 111.5 on 4 and 8 DF, p-value: 4.756e-07
> V <- summary(lm1)$cov.unscaled
> diag(V)
                             hald[, 3]
                                          hald[, 4]
                                                        hald[, 5]
 (Intercept)
                hald[, 2]
820.65457471
               0.09271040
                            0.08756026
                                         0.09520141
                                                       0.08403119
```

n=13 observations, a continuous response variable and p=4 predictors.

Bayes factors between a fitted model and a submodel where some of the variables are dropped can be easily computed from the lm output using functions mombf and imombf. As an example here we drop the second coefficient from the model. Parameter g corresponds to the prior dispersion τ in our notation. There are several options to estimate numerically iMOM Bayes factors (for MOM they have closed form), here we compare adaptive quadrature with a Monte Carlo estimate.

4 Parameter estimation

A natural question after performing model selection is obtaining estimates for the parameters. Rossell and Telesca (2016) developed a general posterior sampling framework for NLPs based on Gibbs sampling an augmented probability model that expresses NLPs as mixtures of truncated distributions. The methodology is implemented in function rnlp. Its basic use is simple, by setting parameter msfit to the output of modelSelection the function produces posterior samples under each model γ visited by modelSelection. The number of samples is proportional to its posterior probability $p(\gamma \mid y)$, thus averaging the output gives Bayesian model averaging estimates $E(\theta \mid y) = \sum_{\gamma} E(\theta \mid \gamma, y) p(\gamma \mid y)$ and likewise for the residual variance $E(\phi \mid y)$.

```
> priorCoef <- momprior(tau=.348)</pre>
> priorDelta <- modelbbprior(alpha.p=1,beta.p=1)</pre>
> fit1 <- modelSelection(y=y, x=x, center=FALSE, scale=FALSE,
+ priorCoef=priorCoef, priorDelta=priorDelta)
Enumerating models...
Computing posterior probabilities..... Done.
> th <- rnlp(y=as.numeric(y),x=x,msfit=fit1,priorCoef=priorCoef,niter=10000)</pre>
> colMeans(th)
                                             phi
      beta1
                  beta2
                               beta3
> head(th)
        beta1
                beta2 beta3
                                   phi
[1,] 1.0699785 0.8691159 0 0.8907935
[2,] 0.9656519 0.8487383
                           0 1.2229253
[3,] 0.9649019 0.8758050
                           0 1.0516350
[4,] 0.9050806 0.8189167
                           0 0.8862611
[5,] 0.9087373 0.9623837
                           0 1.0739552
[6,] 1.1247194 0.8844062
                           0 0.8097169
```

From this output we can obtain model-specific posterior means.

```
> model <- apply(th!=0,1,function(z) paste(which(z),collapse=','))</pre>
> table(model)
model
1,2,3,4
         1,2,4
   250
          9750
> colMeans(th[model=='1,2,4',])
    beta1
             beta2
                      beta3
                                phi
1.0061922 0.9433184 0.0000000 1.0016814
> colMeans(th[model=='1,2,3,4',])
     beta1
               beta2
 1.0323933 0.9419133 -0.1601307 0.9851272
```

Another interesting use of rnlp is to obtain posterior samples under a generic non-local posterior

$$p(\theta \mid y) \propto d(\theta)N(\theta; m, V),$$

where $d(\theta)$ is a non-local prior penalty and $N(\theta; m, V)$ is the normal posterior that one would obtain under the underlying local prior. For instance, suppose our prior is proportional to Zellner's prior times a product MOM penalty $p(\theta) \propto \prod_j \theta_j^2 N(\theta; 0, n\tau\phi(X'X)^{-1})$ where ϕ is the residual variance, then the posterior is proportional to $p(\theta \mid y) \propto \prod_j \theta_j^2 N(\theta; m, V)$ where $m = s_\tau(X'X)^{-1}X'y$, $V = \phi(X'X)^{-1}s_\tau^2$ where $s_\tau = n\tau/(1+n\tau)$ is the usual ridge regression shrinkage factor. We may obtain posterior samples as follows. Note that the posterior mean is close to that obtained above.

```
> tau= 0.348
> shrinkage= nrow(x)*tau/(1+nrow(x)*tau)
> V= shrinkage * solve(t(x) %*% x)
> m= as.vector(shrinkage * V %*% t(x) %*% y)
> phi= mean((y - x%*%m)^2)
> th= rnlp(m=m,V=phi * V,priorCoef=momprior(tau=tau))
> colMeans(th)
    beta1    beta2    beta3
1.0038342    0.9289234 -0.1617318
```

5 Exact inference for block-diagonal regression

Papaspiliopoulos and Rossell (2016) proposed a fast computational framework to compute exact posterior model probabilities, variable inclusion probabilities and parameter estimates for Normal linear regression when the X'X

matrix is block-diagonal. Naturally this includes the important particular case of orthogonal regression where X'X is diagonal. The framework performs a fast model search that finds the best model of each size (i.e. with $1,2,\ldots,p$ variables) and a fast deterministic integration to account for the fact that the residual variance is uknown (the residual variance acts as a "cooling" parameter that affects how many variables are included, hence the associated uncertainty must be dealt with appropriately). The function postModeOrtho tackles the diagonal X'X case and postModeBlockDiag the block-diagonal case.

The example below simulates n=210 observations with p=200 variables where all regression coefficients are 0 except for the last three $(0.5,\,0.75,\,1)$ and the residual variance is one. We then perform variable selection under Zellner's and the MOM prior.

```
> set.seed(1)
> p <- 200; n <- 210
> x <- scale(matrix(rnorm(n*p),nrow=n,ncol=p),center=TRUE,scale=TRUE)
> S \leftarrow cov(x)
> e <- eigen(cov(x))</pre>
> x <- t(t(x %*% e$vectors)/sqrt(e$values))</pre>
> th <- c(rep(0,p-3),c(.5,.75,1)); phi <- 1
> y <- x %*% matrix(th,ncol=1) + rnorm(n,sd=sqrt(phi))</pre>
> priorDelta=modelbinomprior(p=1/p)
> priorVar=igprior(0.01,0.01)
> priorCoef=zellnerprior(tau=n)
> pm.zell <-
+ postModeOrtho(y,x=x,priorCoef=priorCoef,priorDelta=priorDelta,
+ priorVar=priorVar,bma=TRUE)
> head(pm.zell$models)
              modelid
          200,199,198 0.828818155
5
       200,199,198,54 0.037323461
107
       200,199,198,11 0.006070065
      200,199,198,186 0.004133534
108
       200,199,198,36 0.003679338
110
    200,199,198,54,11 0.000316667
> priorCoef=momprior(tau=0.348)
> pm.mom <- postModeOrtho(y,x=x,priorCoef=priorCoef,priorDelta=priorDelta,
+ priorVar=priorVar,bma=TRUE)
> head(pm.mom$models)
              modelid
4
          200,199,198 9.779392e-01
5
       200,199,198,54 1.144910e-02
107
       200,199,198,11 1.209685e-03
      200,199,198,186 7.314291e-04
```

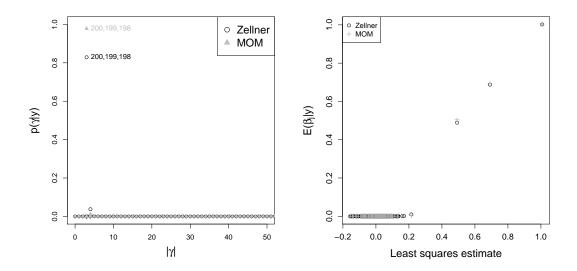


Figure 2: Posterior probability under simulated orthogonal data

```
110 200,199,198,36 6.262828e-04 6 200,199,198,54,11 1.717659e-05
```

postModelBlockDiag returns a list with the best model of each size and its corresponding (exact) posterior probability, displayed in Figure 2 (left panel). It also returns marginal inclusion probabilities and BMA estimates, shown in the right panel. The code required to produce these figures is below.

```
> par(mar=c(5,5,1,1))
> nvars <- sapply(strsplit(as.character(pm.zell$models$modelid),split=','),length)</pre>
> plot(nvars,pm.zell$models$pp,ylab=expression(paste("p(",gamma,"|y)")),
+ xlab=expression(paste("|",gamma,"|")),cex.lab=1.5,ylim=0:1,xlim=c(0,50))
> sel <- pm.zell$models$pp>.05
> text(nvars[sel],pm.zell$models$pp[sel],pm.zell$models$modelid[sel],pos=4)
> nvars <- sapply(strsplit(as.character(pm.mom$models$modelid),split=','),length)
> points(nvars,pm.mom$models$pp,col='gray',pch=17)
> sel <- pm.mom$models$pp>.05
> text(nvars[sel],pm.mom$models$pp[sel],pm.mom$models$modelid[sel],pos=4,col='gray')
> legend('topright',c('Zellner','MOM'),pch=c(1,17),col=c('black','gray'),cex=1.5)
> par(mar=c(5,5,1,1))
> ols \leftarrow (t(x) %*% y) / colSums(x^2)
> plot(ols,pm.zell$bma$coef,xlab='Least squares estimate',
+ ylab=expression(paste('E(',beta[j],'|y)')),cex.lab=1.5,cex.axis=1.2,col=1)
> points(ols,pm.mom$bma$coef,pch=3,col='darkgray')
> legend('topleft',c('Zellner','MOM'),pch=c(1,3),col=c('black','darkgray'))
```

We now illustrate similar functionality under block-diagonal X'X. To this end we consider a total p = 100 variables split into 10 blocks of 10

variables each, generated in such a way that they all have unit variance and within-blocks pairwise correlation of 0.5. The first block has three non-zero coefficients, the second block two and the remaining blocks contain no active variables.

```
> set.seed(1)
> p <- 100; n <- 110
> blocksize <- 10
> blocks <- rep(1:(p/blocksize),each=blocksize)</pre>
> x <- scale(matrix(rnorm(n*p),nrow=n,ncol=p),center=TRUE,scale=TRUE)
> S \leftarrow cov(x)
> e <- eigen(cov(x))
> x \leftarrow t(t(x %*% e$vectors)/sqrt(e$values))
> Sblock <- diag(blocksize)
> Sblock[upper.tri(Sblock)] <- Sblock[lower.tri(Sblock)] <- 0.5
> vv <- eigen(Sblock)$vectors
> sqSblock <- vv %*% diag(sqrt(eigen(Sblock)$values)) %*% t(vv)
> for (i in 1:(p/blocksize)) x[,blocks==i] <- x[,blocks==i] %*% sqSblock</pre>
> th <- rep(0,ncol(x))
> th[blocks==1] <- c(rep(0,blocksize-3),c(.5,.75,1))
> th[blocks==2] < c(rep(0,blocksize-2),c(.75,-1))
> phi <- 1
> y <- x %*% matrix(th,ncol=1) + rnorm(n,sd=sqrt(phi))</pre>
```

postModeBlockDiag performs the model search using an algorithm nick-named "Coolblock" (as it is motivated by treating the residual variance as a cooling parameter). Briefly, Coolblock visits a models of sizes ranging from 1 to p and returns the best model for that given size, thus also helping identify the best model overall.

```
> priorCoef=zellnerprior(tau=n)
> priorDelta=modelbinomprior(p=1/p)
> priorVar=igprior(0.01,0.01)
> pm <- postModeBlockDiag(y=y,x=x,blocks=blocks,priorCoef=priorCoef,
+ priorDelta=priorDelta,priorVar=priorVar,bma=TRUE)
> head(pm$models)
   modelid nvars
                               pp.upper
                         pp
1
              0 1.764684e-24 1.764684e-24
2
              1 5.764046e-22 5.764046e-22
3
       6,7
              2 4.600662e-22 4.600662e-22
4
     2,6,7
              3 5.250774e-23 5.250774e-23
5
   2,5,6,7
              4 1.139264e-24 1.139264e-24
              5 1.085583e-26 1.085583e-26
6 2,4,5,6,7
> head(pm$postmean.model)
   modelid X1
                   X2 X3
                               Х4
                                        Х5
                                                  X6
                                                           X7 X8 X9 X10
```

```
0.0000000
                                0 0.0000000 0.0000000 0.0000000 1.1069287
3
                0 0.0000000
                                0 0.0000000 0.0000000 0.8493494 0.6822540
                                                                                                0
                                0 0.0000000 0.0000000 0.6091185 0.4420231
                                                                                                0
                0 0.7206928
    2,5,6,7
                0 0.5756777
                                0 0.0000000 0.5800601 0.4641034 0.2970081
                                                                                                0
                                0 0.4956023 0.4809396 0.3649830 0.1978876
6 2,4,5,6,7
                0 0.4765573
                                                                                       0
                                                                                                0
  X11 X12 X13 X14 X15 X16 X17 X18 X19 X20 X21 X22 X23 X24 X25 X26 X27
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```

Figure 3 shows a LASSO-type plot with the posterior means under the best model of each size visited by Coolblock. We appreciate how the truly active variables 8, 9, 10, 19 and 20 are picked up first.

```
> maxvars=50
> ylim=range(pm$postmean.model[,-1])
> plot(NA,NA,xlab='Model size',
+ ylab='Posterior mean given model',
```

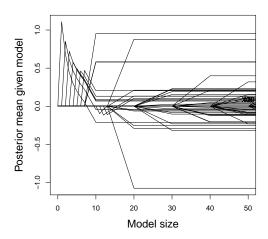


Figure 3: Coolblock algorithm: posterior mean of regression coefficients under best model of each size

```
+ xlim=c(0,maxvars),ylim=ylim,cex.lab=1.5)
> visited <- which(!is.na(pm$models$pp))
> for (i in 2:ncol(pm$postmean.model)) {
+ lines(pm$models$nvars[visited],pm$postmean.model[visited,i])
+ }
> text(maxvars, pm$postmean.model[maxvars,which(th!=0)+1],
+ paste('X',which(th!=0),sep=''), pos=3)
```

6 Bayes factors for generalized linear models

At the moment mombf provides some limited functionality for generalized linear models. pmomLM implements probit models following the framework in Rossell et al. (2013), see also Nikooienejad et al. (2016) for an alternative framework that bypasses the need to sample θ to improve the mixing of the Markov Chain.

As an illustration we simulate n=50 observations from a probit regression model with p=2 correlated predictors and $\theta=(\log(2),0)$. The predictors are stored in the matrix \mathbf{x} , the success probabilities in the vector \mathbf{p} and the observed responses in the vector \mathbf{y} . For reproducibility purposes we set the random number generator seed.

```
> set.seed(4*2*2008)
> n <- 50; theta <- c(log(2),0)
```

```
> x <- matrix(NA,nrow=n,ncol=2)</pre>
> x[,1] \leftarrow rnorm(n,0,1); x[,2] \leftarrow rnorm(n,.5*x[,1],1)
> p <- pnorm(x %*% matrix(theta,ncol=1))</pre>
> y \leftarrow rbinom(n,1,p)
> th <- pmomLM(y=y,x=x,xadj=rep(1,n),niter=10000)
Running MCMC.....Done.
> head(th$postModel)
      [,1] [,2]
[1,]
        1
[2,]
        1
              0
[3.]
        1
[4,]
        1
              0
[5,]
              1
        1
[6,]
> table(apply(th$postModel,1,paste,collapse=','))
 0,0 0,1 1,0 1,1
 237
       19 7482 1262
```

An alternative to pmomLM, which considers all possible models under product NLPs, is to compute Bayes factors between pairs of models under additive NLPs using functions momknown and imomknown.

Before computing Bayes factors, we fit a probit regression model with the function glm. The maximum likelihood estimates are stored in thetahat and the asymptotic covariance matrix in V. These functions take as primary arguments a vector of regression coefficients and their covariance matrix, and hence they can be used in any setting where one has a statistic that is asymptotically sufficient and normally distributed. The resulting Bayes factors are approximate. The functions also allow for the presence of a dispersion parameter sigma, i.e. the covariance of the regression coefficients is sigma*V, but they assume that sigma is known. The probit regression model that we simulated has no over-dispersion and hence it corresponds to sigma=1. We first compare the full model with the model resulting from excluding the second covariate (the first term corresponds to the intercept), setting g=0.5 for illustration.

```
> glm1 <- glm(y~x[,1]+x[,2],family=binomial(link = "probit"))
> thetahat <- coef(glm1)
> V <- summary(glm1)$cov.scaled
> g <- .5
> bfmom.1 <- momknown(thetahat[2],V[2,2],n=n,g=g,sigma=1)
> bfimom.1 <- imomknown(thetahat[2],V[2,2],n=n,nuisance.theta=2,g=g,sigma=1)
> bfmom.1
```

```
[,1]
[1,] 4.262401
> bfimom.1

[,1]
[1,] 3.336888
```

Both priors result in evidence for including the first covariate. We now check whether the second covariate can be dropped.

```
> bfmom.2 <- momknown(thetahat[3],V[3,3],n=n,g=g,sigma=1)
> bfimom.2 <- imomknown(thetahat[3],V[3,3],n=n,nuisance.theta=2,g=g,sigma=1)
> bfmom.2

[,1]
[1,] 0.02784354
> bfimom.2

[,1]
[1,] 0.008250121
```

Both Mom and iMom BF provide strong evidence in favor of the simpler model, i.e. excluding x[,2]. To compare the full model with the model that has no covariates (i.e. only the constant term remains) we use the same routines, passing a vector as the first argument and a matrix as the second argument.

Based on the resulting BF being close to 1, it is not clear whether the full model is preferable to the model with no covariates.

The BF can be used to easily compute posterior probabilities for each of the four considered models: no covariates, only x[,1], only x[,2] and both x[,1] and x[,2]. We assume equal probabilities a priori.

```
> prior.prob <- rep(1/4,4)
> bf <- c(bfmom.0,bfmom.1,bfmom.2,1)
> pos.prob <- prior.prob*bf/sum(prior.prob*bf)
> pos.prob
```

The model with the highest posterior probability is the one including only x[,1], *i.e.* the correct model, and the model with the lowest posterior probability is that including only x[,2].

7 Model selection for mixtures

bfnormmix is the main function to obtain posterior probabilities on the number of mixture components, as well as posterior samples for any given number of components. We start by simulating univariate Normal data truly arising from a single component.

```
> set.seed(1)
> n=200; k=1:3
> x= rnorm(n)
```

We run bfnormmix based on B=10000 MCMC iterations and default prior parameters. The prior dispersion g in (8) is an important parameter that determines how separated one expects components to be a priori. Roughly speaking its default value assigns 0.95 prior probability to any two components giving a multimodal density, this default helps focus the analysis on detecting well-separated components and enforcing parsimony whenever components are poorly separated. The Dirichlet prior parameter q is also important, bfnormmix allows one to specify different values for the MOM-IW-Dir prior in (8) (argument q) and the Normal-IW-Dir prior in (7) (argument q.niw). The default q=3 helps discard components with small weights, since these could be due to over-fitting or outliers.

```
> fit= bfnormmix(x=x,k=k,q=3,q.niw=1,B=10^4)
> postProb(fit)
                  pp.niw logprobempty logbf.momiw
                                                     logpen logbf.niw
     pp.momiw
1 1 0.95924296 0.75292130
                                        0.000000 0.000000 0.000000
                                 -Inf
                                        -3.488004 -2.073395 -1.414609
2 2 0.02931620 0.18297543
                            -3.888696
3 3 0.01144084 0.06410328
                            -3.566264
                                       -4.428955 -1.965490 -2.463465
       model
1 Normal, VVV
2 Normal, VVV
3 Normal, VVV
```

Under a MOM-Dir-IW prior one assigns high posterior probability $P(M_1 | y) \approx 0.96$ to the data-generating truth, for the Normal-Dir-IW this probability is ≈ 0.753 .

We can use postSamples to retrieve samples from the Normal-Dir-IW posterior, and coef to obtain the coresponding posterior means.

```
> mcmcout= postSamples(fit)
> names(mcmcout)
[1] "k=1" "k=2" "k=3"
> names(mcmcout[[2]])
                    "mu"
[1] "eta"
                                   "cholSigmainv" "momiw.weight"
> colMeans(mcmcout[[2]]$eta)
[1] 0.5345456 0.4654544
> parest= coef(fit)
> names(parest)
[1] "k=1" "k=2" "k=3"
> parest[[2]]
$eta
[1] 0.5345456 0.4654544
$mu
[1] -0.11035687 0.06271486
$Sigmainv
$Sigmainv[[1]]
         [,1]
[1,] 10.55558
$Sigmainv[[2]]
         [,1]
[1,] 11.36628
```

We remark that one can reweight these samples to obtain posterior samples from the non-local MOM-Dir-IW posterior, these weights are given by the term $d_k(\vartheta_k)$ in (9) and are stored in momiw.weight. As illustrated below in our example under the 2-component model the location parameters under the MOM-Dir-IW posterior are more separated than under the Normal-Dir-IW, as one would expect from the repulsive force between components.

```
> w= postSamples(fit)[[2]]$momiw.weight
> apply(mcmcout[[2]]$eta, 2, weighted.mean, w=w)
[1] 0.5618042 0.4381958
> apply(mcmcout[[2]]$mu, 2, weighted.mean, w=w)
[1] -0.1818446 0.3205265
```

To further illustrate we simulate another dataset, this time where the data-generating truth are k=2 components. Interestingly the MOM-IW-Dir favours k=2 more than the Normal-IW-Dir and both clearly discard k=1, as expected given that the two components are well-separated (the means are 3 standard deviations away from each other).

```
> set.seed(1)
> n=200; k=1:3; probs= c(1/2,1/2)
> mu1= -1.5; mu2= 1.5
> x= rnorm(n) + ifelse(runif(n) probs[1],mu1,mu2)
> fit= bfnormmix(x=x,k=k,q=3,q.niw=1,B=10^4)
> postProb(fit)
       pp.momiw
                       pp.niw logprobempty logbf.momiw
                                                           logpen logbf.niw
1 1 2.257427e-47 3.771374e-47
                                      -Inf
                                                0.0000 0.0000000
                                                                     0.0000
2 2 8.083230e-01 6.103454e-01 -111.703635
                                              107.1945 0.7941503 106.4003
3 3 1.916770e-01 3.896546e-01
                                 -4.166356
                                              105.7553 -0.1962352 105.9516
       model
1 Normal, VVV
2 Normal, VVV
3 Normal, VVV
```

References

- T. Chekouo, F.C. Stingo, J.D. Doecke, and K.-A. Do. mirna—target gene regulatory networks: A bayesian integrative approach to biomarker selection with application to kidney cancer. *Biometrics*, 71(2):428–438, 2015.
- Rodrigo A Collazo, Jim Q Smith, et al. A new family of non-local priors for chain event graph model selection. *Bayesian Analysis*, (in press), 2016.
- G. Consonni and L. La Rocca. On moment priors for Bayesian model choice with applications to directed acyclic graphs. In J.M. Bernardo, M.J. Bayarri, J.O. Berger, A.P. Dawid, D. Heckerman, A.F.M Smith, and M. West, editors, *Bayesian Statistics 9 - Proceedings of the ninth Valencia interna*tional meeting, pages 119–144. Oxford University Press, 2010.
- J. Fúquene, M.F.J. Steel, and D. Rossell. On choosing mixture components via non-local priors. arXiv 1604.00314v3, pages 1–43, 2018.
- V.E. Johnson and D. Rossell. Prior densities for default bayesian hypothesis tests. *Journal of the Royal Statistical Society B*, 72:143–170, 2010.
- V.E. Johnson and D. Rossell. Bayesian model selection in high-dimensional settings. *Journal of the American Statistical Association*, 24(498):649–660, 2012.

- F. Liang, R. Paulo, G. Molina, M.A. Clyde, and J.O. Berger. Mixtures of g-priors for Bayesian variable selection. *Journal of the American Statistical Association*, 103:410–423, 2008.
- C.P. Marin, J. M. Robert. Approximating the marginal likelihood in mixture models. *Bulleting of the Indian Chapter of ISBA*, 1:2–7, 2008.
- Amir Nikooienejad, Wenyi Wang, and Valen E Johnson. Bayesian variable selection for binary outcomes in high dimensional genomic studies using non-local priors. *Bioinformatics*, btv764:1–8, 2016.
- O. Papaspiliopoulos and D. Rossell. Scalable variable selection and model averaging under block-orthogonal design. *arXiv*, pages 1–19, 2016.
- D. Rossell and F.J. Rubio. Tractable bayesian variable selection: beyond normality. *arXiV*, 1609.01708:1–59, 2016.
- D. Rossell and D. Telesca. Non-local priors for high-dimensional estimation. Journal of the American Statistical Association, (in press):1–33, 2016.
- D. Rossell, D. Telesca, and V.E. Johnson. High-dimensional Bayesian classifiers using non-local priors. In *Statistical Models for Data Analysis XV*, pages 305–314. Springer, 2013.
- J.G. Scott and J.O Berger. Bayes and empirical Bayes multiplicity adjustment in the variable selection problem. *The Annals of Statistics*, 38(5): 2587–2619, 2010.
- M. Shin, A. Bhattacharya, and V.E. Johnson. Scalable Bayesian variable selection using nonlocal prior densities in ultrahigh-dimensional settings. *Texas A&M University (technical report)*, pages 1–33, 2015.