NSCpp: Non-Standard Cosmologies in C++

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Abstract

Program summary:

Program title: NSCpp.

Developer's respository link: https://github.com/dkaramit/NSCpp.

Programming language: C++ and python.

Licensing provisions: MIT license.

Nature of problem:

Solution method: Embedded Runge-Kutta for the numerical integration of the equation of motion. The user may choose between explicit and Rosenborck methods, or implement their own Butcher tableau. For the various interpolations, the library uses cubic splines.

Restrictions:

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1 Introduction

2 Physics background

2.1 Notation

3 NSCpp usage

The latest stable version of NSCpp is available at github.com/dkaramit/NSCpp/tree/stable, which can also be obtained by running: 1

```
git clone -b stable https://github.com/dkaramit/NSCpp.git
```

It is important to note that NSCpp relies on NaBBODES [1] and SimpleSplines [2]. These are two libraries developed independently, and in order to get NSCpp with the latest version of these libraries, one needs to run following commands

```
git clone https://github.com/dkaramit/NSCpp.git
cd NSCpp
git submodule init
git submodule update --remote
```

This downloads the master branch of NSCpp; and NaBBODES [1] and SimpleSplines [2] as submodules. This guaranties that NSCpp uses the most updated version of these libraries, although it may not be stable.

Once everything is downloaded successfully, we can go inside the NSCpp directory, and run "bash configure.sh" and "make". The bash script configure.sh, just writes some paths to some files, formats the data files provided in an acceptable format (in section 4.3 the format is explained), and makes some directories. The makefile is responsible for compiling some examples and checks, as well as the shared libraries that needed for the python interface. If everything runs successfully, there should be two new directories exec and lib. Inside exec, there are several executables that ready to run, in order to ensure that the code runs (e.g. no segmentation fault occurs).

Although there are various options available at compile-time, we first discuss how NSCpp can be used, in order for the role of these options to be clear.

3.1 First steps

```
#include"src/NSC/NSCSolve.hpp"
```

Notice that if the this .cpp file is not in the root directory of NSCpp, we need to compile it using the flag -Ipath-to-root, "path-to-root" the relative path to the root directory of NSCpp; e.g. if the .cpp is in the MiMeS/UserSpace/Cpp/NSC directory, this flag should be -I../../.

```
Setting up the solver

nsc::NSC<LD,SOLVER,METHOD<LD>> BE(TEND,c,Ti,ratio,umax,TSTOP,

initial_step_size,minimum_step_size, maximum_step_size,

absolute_tolerance, relative_tolerance, beta,

fac_max, fac_min, maximum_No_steps);
```

Here, LD should be the numeric type to be used; it is recommended to use long double, but other choices are also available as we discuss later. Moreover Solver and Method depend on the type of Runge-Kutta (RK) the user chooses. The available choices are shown in table3.

The various parameters are as follows:

¹Instructions on how git can be installed can be found in https://github.com/git-guides/install-git.

1.

2.

- 3. umax: If u > umax the integration stops (remember that $u = \log(a/a_i)$). Typically, this should be a large number (~ 1000), in order to avoid stopping the integration before the axion begins to evolve adiabatically.
- 4. TSTOP: If the temperature drops below this, integration stops.
- 5. initial_stepsize (optional): Initial step the solver takes.
- 6. maximum_stepsize (optional): This limits the step-size to an upper limit.
- 7. minimum_stepsize (optional): This limits the step-size to a lower limit.
- 8. absolute_tolerance (optional): Absolute tolerance of the RK solver
- 9. relative_tolerance (optional): Relative tolerance of the RK solver.
- 10. beta (optional): Controls how agreesive the adaptation is. Generally, it should be around but less than 1.
- 11. fac_max, fac_min (optional): The stepsize does not increase more than fac_max, and less than fac_min. This ensures a better stability. Ideally, fac_max= ∞ and fac_min= 0, but in reality one must tweak them in order to avoid instabilities.
- 12. maximum_No_steps (optional): Maximum steps the solver can take. Quits if this number is reached even if integration is not finished.

```
BE.solveNSC();
```

3.1.1 Using nsc in python

The modules for the python interface are located in NSCpp/src/interfacePy. Although the usage of the classes is similar to the C++ case, it is worth showing explicitly how the python interface works. One should keep in mind that the various template arguments discussed in the C++ case have to be chosen at compile-time. That is, for the python interface, one needs to choose the numeric type, and RK method to be used when the shared libraries are compiled. This is done by assigning the relevant variable in NSCpp/Definitions.mk before running "make". The various options are discussed in section 4.2, and outlined in table 4.

The two relevant classes are defined in the modules interfacePy.AxionMass and interfacePy.Axion, and can be loaded in a python script as

```
from sys import path as sysPath
sysPath.append('path_to_src')
from interfacePy.NSC import NSC
```

It is important that 'path_to_src' provides the relative path to the NSCpp/src directory. For example, if the script is located in NSCpp/UserSpace/Python, 'path_to_src' should be '../../src'.

The solver We can define an NSC instance as follows

```
BE=NSC(TEND,c,Ti,ratio,umax,TSTOP,
initial_step_size,minimum_step_size, maximum_step_size, absolute_tolerance,
relative_tolerance, beta, fac_max, fac_min, maximum_No_steps)
```

Here the input parameters are the same as in the C++ case, and outlined in table 1. Moreover, the usage of the class can be found by running ?NSC after loading the module.

Using the defined variable (BE in this example), we can simply run

BE.solveNSC()

in order to solve the BEs. In contrast to the C++ implementation, this only gives us access to the points where the behaviour changes; the corresponding variables are BE.TE1,BE.TE2,BE.TD1,BE.TD2. In order to get the evolution of the densities, we need to run

BE.getPoints()

This will make numpy [?] arrays that contain the ... Moreover, we can run the following

BE.getErrors()

4 Assumptions and user input

4.1 Restrictions

4.2 Options at Compile-time

The user has a number of options regarding different aspects for the code. If NSCpp is used without using the available makefiles, then they must use the correct values for the various template arguments, explained in Appendix A. The various choices we for the shared libraries used by the python interface are given in NSCpp/Definitions.mk while the corresponding options for the C++ examples are in the Definitions.mk files inside the subdirectories of NSCpp/UserSpace/Cpp. The options correspond to different variables, which are

- 1. rootDir: The relative path of root directory of NSCpp.
- 2. LONG: This sets the numeric types for the C++ examples. It should be either long or omitted. If omitted, the type of the numeric values is double (double precision). On the other hand, if LONG=long, the type is long double. Generally, using double should be enough. For the sake of numerical stability, however, it is advised to always use LONG=long, as it a safer option. The reason is that the axion angle redshifts, and can become very small, which introduces "rounding errors". Moreover, if the parameters absolute_tolerance or absolute_tolerance are chosen to be below ~ 10⁻⁸, then double precision numbers may not be enough, and LONG=long is preferable. This choice comes at the cost of speed; double precision operations are usually preformed much faster. It is important to note that LONG defines a macro with the same name (in the C++ examples), which then defines the macro (again in the C++ examples) as #define LD LONG double. The macro LD, then is used as the corresponding template argument in the various classes. We point out again that if one chooses not to use the makefile files, the template arguments need to be known at compile-time. So the user has to define them in the code.
- 3. LONGpy: the same as LONG, but for the python interface. One should keep in mind that this cannot be changed inside python scripts. It just instructs ctypes what numeric type to use. Since the preferred way to compile the shared libraries is via running "make" in the root directory of NSCpp, this variable needs to be defined inside NSCpp/Definitions.mk. By default, this variable is set to long, since this is the most stable choice in general.
- 4. SOLVER: NSCpp uses the ordinary differential equation (ODE) integrators of ref. [1]. Currently, there are two available choices; Rosenbrock and RKF. The former is a general embedded Rosenbrock implementation and it is used if SOLVER=1, while the latter is a general explicit embedded Runge-Kutta implementation and can be chosen by using SOLVER=2 (a brief description of how these algorithms are implemented can be found in Appendix ??). By default inside the Definitions.mk files SOLVER=1, because the axion EOM tends to oscillate rapidly. However, in

some cases, a high order explicit method may also work. Note that this variable defines a macro that is then used as the second template argument of the mimes::Axion<LD,Solver,Method> class. The preferred way to do it in the shared libraries is via the NSCpp/Definitions.mk file, however, the user if free to compile everything in a different way. In this case, the various Definitions.mk files, are not being used, and the user must define the relevant arguments in the code where NSCpp is used.

- 5. METHOD: Depending on the type of solver, there are some available methods. ²
 - For SOLVER=1, the available methods are METHOD=RODASPR2 and METHOD=ROS34PW2. The RODASPR2 choice is a fourth order Rosenbrock-Wanner method (more information can be found in ref. [?]). The ROS34PW2 choice corresponds to a third order Rosenbrock-Wanner method [?].
 - For SOLVER=2, the only reliable method available in NaBBODES is the Dormand-Prince [?] chosen if METHOD=DormandPrince, which is an explicit Runge-Kutta method of seventh order.

This variable defines a macro (with the same name) that is passed as the third template parameter of mimes::Axion<LD,Solver,Method> (i.e. METHOD<LD> in the place of Method). If the compilation is not done via the makefile files, the user must define the relevant template arguments in the code.

- 6. CC: The C++ compiler that one chooses to use. The default option is CC=g++, which is the GNU C++ compiler, and is available for all systems. Another option is to use the clang compiler, which is chosen by CC=clang -lstdc++. NSCpp is mostly tested using g++, but clang also seems to work (and the resulting executables are sometimes faster), but the user has to make sure that their version of the compiler of choice supports the C++ 17 standard, otherwise NSCpp probably will not work.
- 7. OPT: Optimization level of the compiler. By default, this is OPT=03, which produces executables that are marginally faster than OPT=01 and OPT=02, but significantly faster than OPT=00. There is another choice, OPT=Ofast, but it can cause various numerical instabilities, and is generally considered dangerous although we have not observed any problems when running NSCpp.

It is important to note, once again, that the variables that correspond to template arguments must be known at compile time. Thus, if the compilation is done without the help of the various makefile files, the template arguments must be given, otherwise compilation will fail. ³ For example, the choice LONG=long, SOLVER=1, and METHOD=RODASPR2 will be used to compile the shared libraries (and C++ example in NSCpp/UserSpace/Cpp/NSC) with nsc::NSC<long double,1,RODASPR2<long double>>.

4.3 User input

4.3.1 Compile-time input

Files NSCpp requires files that provide data for the relativistic degrees of freedom (RDOF) of the plasma, and the anharmonic factor. Although NSCpp is shipped with the standard model RDOF found in [?], and a few points for $f(\theta_{\text{peak}})$ introduced in eq. (??), the user is free change them via the corresponding variables in NSCpp/Paths.mk. Moreover, there is a set of data for the QCD axion mass as calculated in ref.[?]. The variables pointing to these data files are cosmoDat, axMDat, and anFDat, for the RDOF, axion mass, and the anharmonic factor; respectively.

The format of the files has to be the following:

²It is worth mentioning that NaBBODES is built in order to be a template for all possible Rosenbrock and explicit Runge-Kutta embedded methods, and one can provide their own Butcher tableau if they want to use another method, as shown in Appendix ??.

³In C++ the template arguments are part of the definition of a class; if the template arguments are not known, the class is not even constructed.

- The RDOF data must be given in three columns; T (in GeV), h_{eff} , and g_{eff} .
- The axion mass data must be given in two columns; T (in GeV), χ (in GeV⁴). Here, χ is defined as in eq. (??). The user can provide a function instead of data for the axion mass, by leaving the axMDat variable empty.
- The data for the anharmonic factor must be give in two columns θ_{peak} $f(\theta_{\text{peak}})$; with increasing θ_{peak} .

The paths to these files should be given at compile time. That is, once Paths.mk changes, we must run "bash configure.sh" and then "make" in order to make sure that they will be used. The user can change the content of the data files (without changing their paths), in order to use them without compiling NSCpp again. However, the user has to make sure that all the files are sorted so that the values of first column increase (with no duplicates or empty lines). In order to ensure this, it is advised to run "bash FormatFile.sh path-to-file" (in Appendix ?? there are some details on MiMeS/src/FormatFile.sh), in order to format the file (that should exist in "path-to-file") so that it complies with the requirements of NSCpp.

These paths are stored as strings in NSCpp/src/misc_dir/path.hpp at compile-time (they are defined as constexpr), and can be accessed once this header file is included. The corresponding variables are cosmo_PATH, chi_PATH, and anharmonic_PATH, for the path to data file of the plasma quantities, $\chi(T)$, and $f(\theta_{\text{peak}})$; respectively. Although, the axion mass data file may be omitted – since the axion mass is defined by the user, the variable chi_PATH is still useful if the axion mass is defined via a data file, as it is automatically converted to an absolute path. ⁴

5 Acknowledgements

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6 Summary

Appendix

A C++ classes

B NSCpp python interface

C Quick guide to the user input

We present tables..., with the various available run-time inputs, required files, and template arguments. In table 4 we show the available compile-time options, that can be used when compiling using the various makefile files.

⁴Absolute paths have the advantage to be accessible from everywhere else in the system. Thus, executables that seek the corresponding files can be called and copied easily.

	User run-time input for solving the BEs.
	· · · · · · · · · · · · · · · · · · ·
TSTOP	Once $T < TSTOP$, integration stops. Typical value: 10^{-4} GeV.
initial_stepsize	Initial step-size of the solver. Default value: 10^{-2} .
minimum_stepsize	Lower limit of the step-size. Default value: 10^{-8} .
maximum_stepsize	Upper limit of the step-size. Default value: 10^{-2} .
absolute_tolerance	Absolute tolerance of the RK solver. Default value: 10^{-8} .
relative_tolerance	Relative tolerance of the RK solver. Default value: 10^{-8} .
beta	Aggressiveness of the adaptation strategy. Default value: 0.9.
fac_max, fac_min	The step-size does not change more than fac_max and less than
	fac_min within a trial step. Default values: 1.2 and 0.8, respectively.
maximum_No_steps	If integration needs more than maximum_No_steps integration stops.
	Default value: 10^7 .

Table 1: Table of the constructor arguments of the nsc::NSC<LD,Solver,Method> class.

Required data files, with corresponding variables in NSCpp/Paths.mk.		
cosmoDat	Relative path to data file with T (in GeV), $h_{\rm eff}$, $g_{\rm eff}$. If the path	
	changes one must run bash configure.sh and make.	

Table 2: Paths to the required data files. Variables defined in the NSCpp/Paths.mk files, and used when running bash configure.sh in the root directory of NSCpp.

Template arguments.				
LD	This template argument appears in all classes of NSCpp. The preferred choice is long double. However, in many cases double can be used. The user should be careful, as the later can lead to an inaccurate result; especially for low tolerances, and small values of θ .			
Solver	This is the second template argument of the nsc::NSC <ld,solver,method> class. The available choices are Solver=1 for Rosenbrock method, and Solver=2 for explicit RK method.</ld,solver,method>			
Method	The third template argument of the nsc::NSC <ld,solver,method> class. Its value depends on the choice of Solver; For Solver=1, Method can be either RODASPR2<ld> (fourth order) or ROS34PW2<ld> (third order). For Solver=2, Method can only be DormandPrince<ld> (seventh order). Notice that the definitions of the various method classes, also need a template argument,LD, that must be the same as the first template argument of the nsc::NSC<ld,solver,method> class. If one defines their own Butcher table, then they would have to follow their definitions and assumptions.</ld,solver,method></ld></ld></ld></ld,solver,method>			

Table 3: Template arguments of the various NSCpp classes.

References

[1] D. Karamitros, NaBBODES: Not a Black Box Ordinary Differential Equation Solver in C++, 2019.

User compile-time options. Variables in the various Definitions.mk files.				
rootDir	The relative path of root directory of NSCpp. Relevant only when compiling using make. Available in all Definitions.mk.			
LONG	long for long double or empty for double. This is defines a macro in the source files of the various C++ examples. Available in Definitions.mk inside the various subdirectories of NSCpp/UserSpace/Cpp.			
LONGpy	long or empty. Same as LONG, applies in the python modules. Available in NSCpp/Definitions.mk.			
SOLVER	In order to use a Rosenbrock method SOLVER=1. For explicit RK method, SOLVER=2. This defines a macro that is passes as the second template argument of mimes::Axion <ld,solver,method>. The corresponding variable in NSCpp/Definitions.mk applies to the python modules. The variable in NSCpp/UserSpace/Cpp/NSC/Definitions.mk applies to the example in the same directory.</ld,solver,method>			
METHOD	Depending on the solver, this variable should name one of its available methods. For SOLVER=1, METHOD=RODASPR2(fourth order) or ROS34PW2(third order). For SOLVER=2, METHOD=DormandPrince (seventh order). There is a macro (METHOD) used by the shared library NSCpp/lib/libNSC.so. The corresponding variable in NSCpp/Definitions.mk applies to the python modules. The variable in NSCpp/UserSpace/Cpp/NSC/Definitions.mk applies to the example in that directory.			
Compiler options				
CC	The preferred C++ compiler (g++ by default). Corresponding variable in all Definitions.mk files.			
OPT	Available options are OPT=01, 02, 03 (be default). This variable defines the optimization level of the compiler. The variable can be changed in all Definitions.mk files. In the root directory of NSCpp, the optimization level applies to the python modules (i.e. the shared libraries), while in the subdirectories of NSCpp/UserSpace/Cpp it only applies to example inside them.			

Table 4: User compile-time input and options. These are available in the various Definitions.mk files, which are used when compiling using make.

^[2] D. Karamitros, SimpleSplines: A header-only library for linear and cubic spline interpolation in C++, 2021.