## GAMMA-RAY SPECTRAL ANALYSIS OF THREE ENERGETIC MILLISECOND PULSARS

Pulsar Name	f	$g_1$	$h_1/h_2$	$E_{\mathrm{cut}}$	Energy Flux $(10^{-11} \text{ erg cm}^{-2} \text{ s}^{-1})$
J0633+1746	0.76	0.85	0.933	$2.2 \pm 0.1$	$423.3 \pm 1.2$
J0835-4510	0.16	0.92	0.927	$3.0 \pm 0.1$	$906 \pm 2$
J0007+7303	0.65	0.94	0.967	$4.7\pm0.2$	$40 \pm 0.4$
J0030+0451	0.60	0.88	0.947	$1.8 \pm 0.2$	$6.14 \pm 0.18$
J0437-4715	0.38	0.95	0.927	$1.1\pm0.3$	$1.67 \pm 0.11$

Table 1: The best-fit parameters of the two-layer model and observational properties for a few pulsars. Both young pulsars and MSPs are included.

## Bibliography