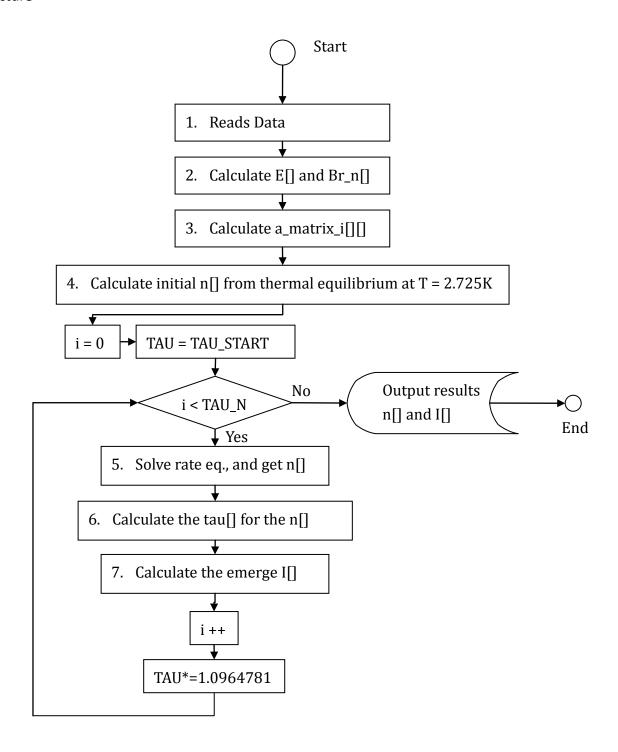
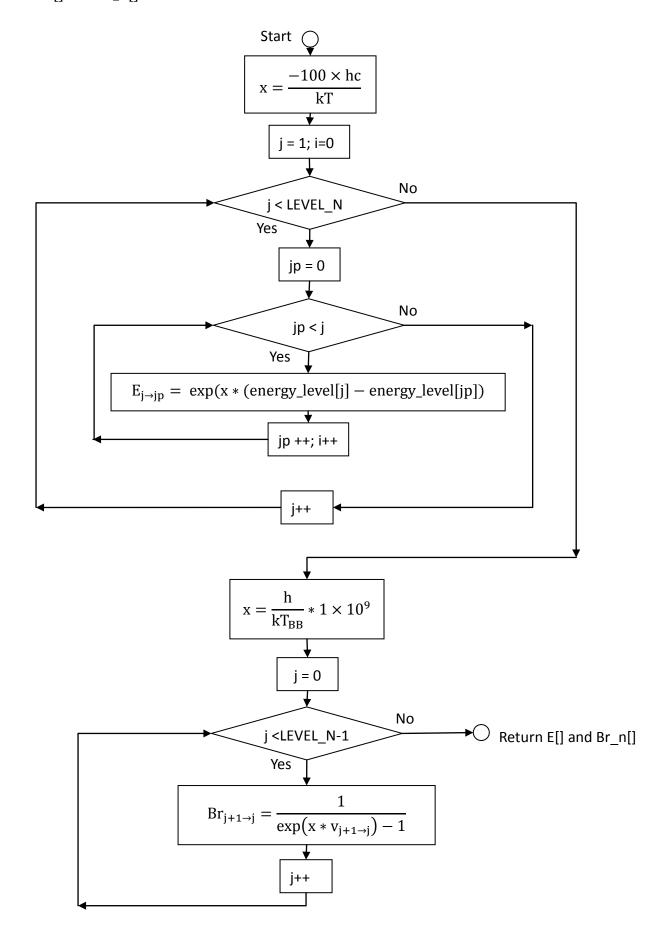
Flow Chart for c3.4 (2010.05.04)

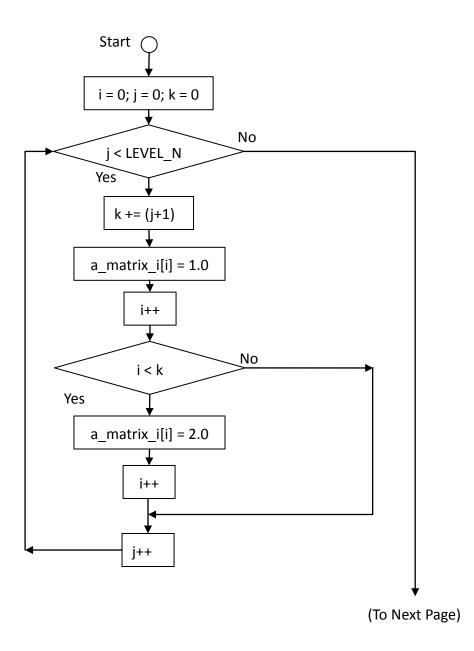
Main Structure

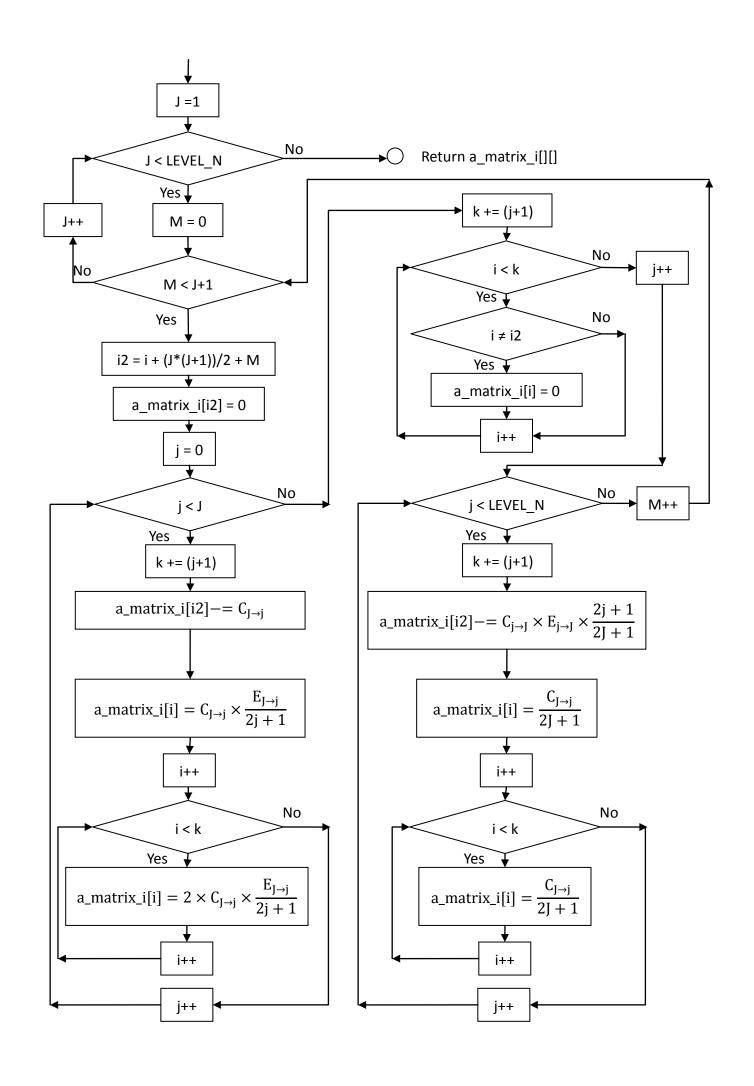


2. Calculate E[] and Br_n[]

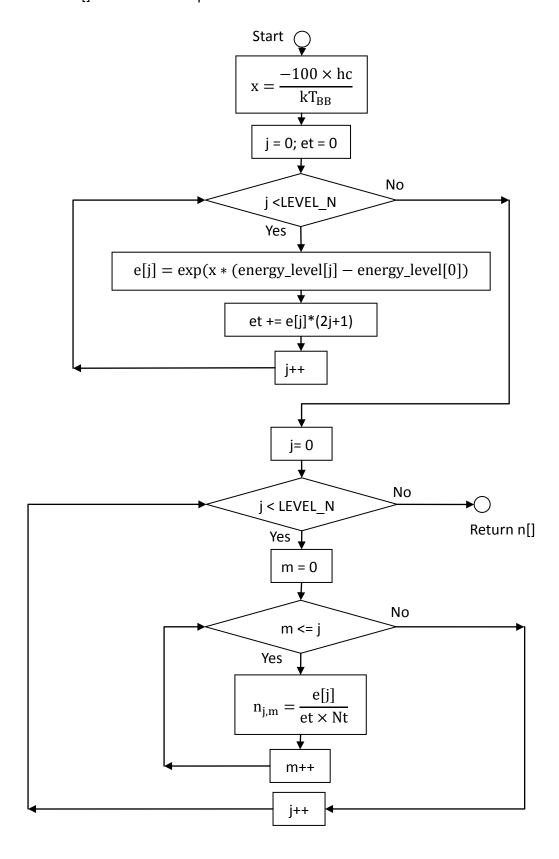


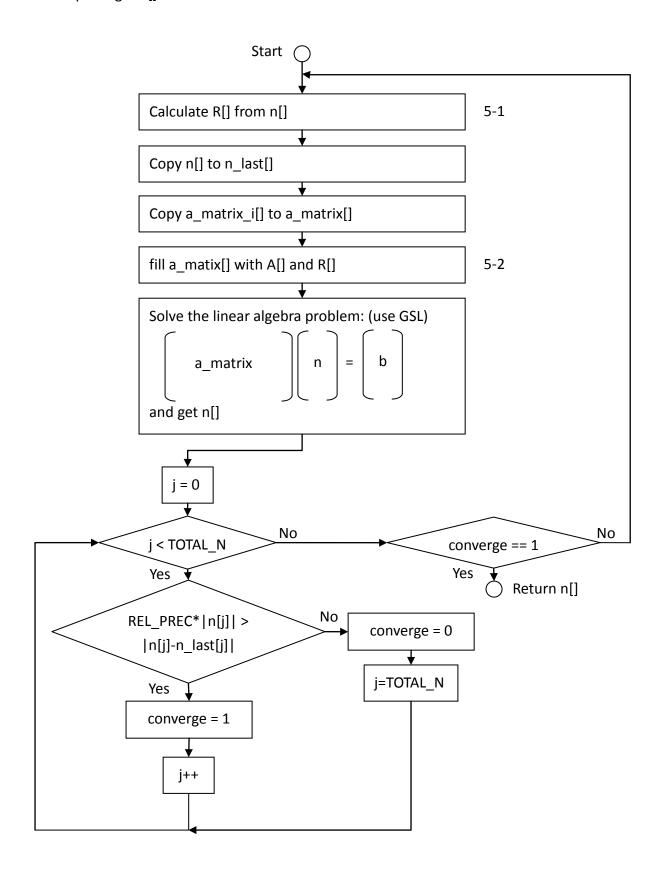
3. Calculate a_matrix_i[][]

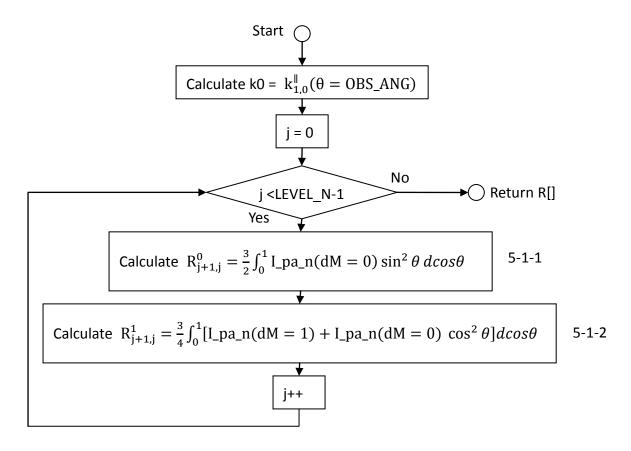




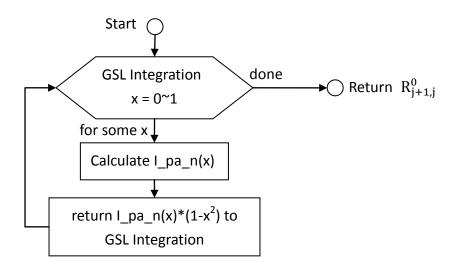
4. Calculate initial n[] from thermal equilibrium at T = 2.725K



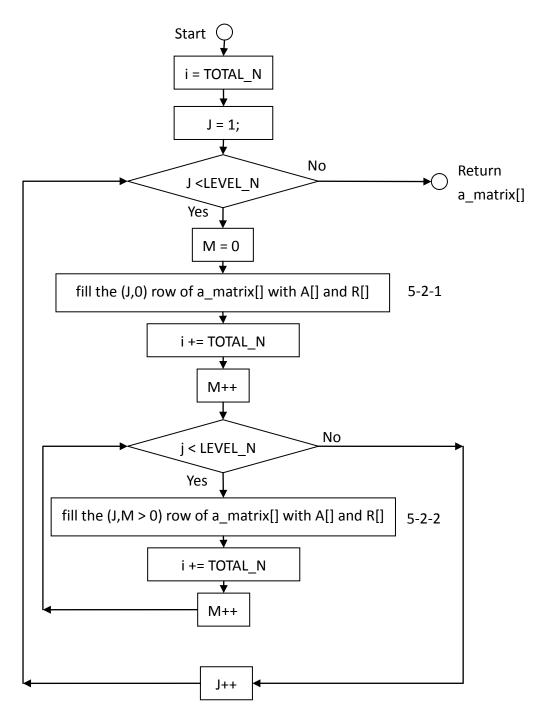




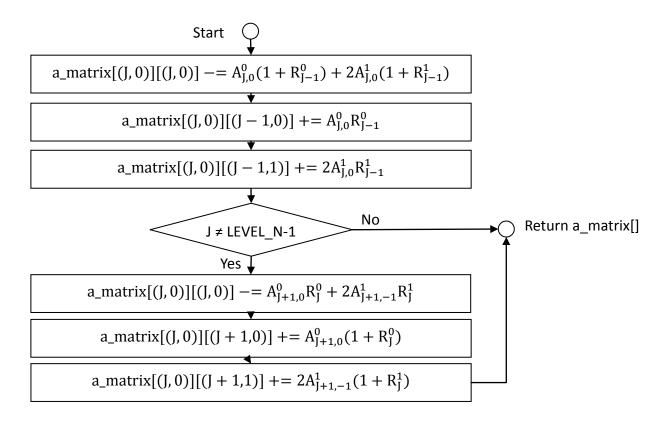
5-1-1
$$R_{j+1,j}^0 = \frac{3}{2} \int_0^1 I_p a_n (dM = 0) \sin^2 \theta \, d\cos \theta$$



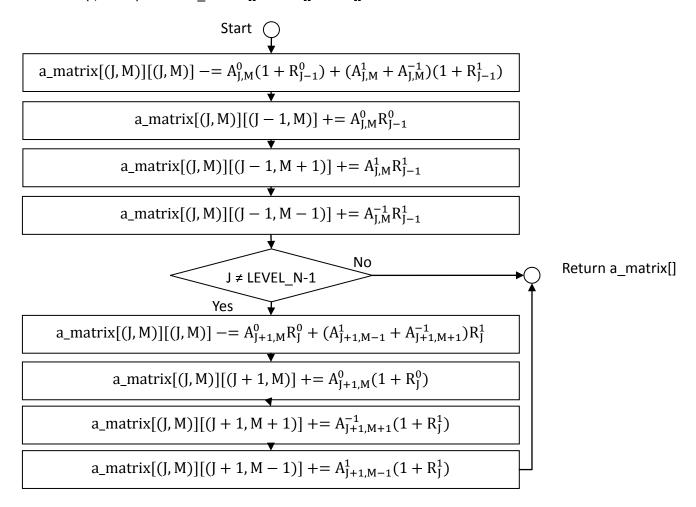
5-1-2
$$R_{j+1,j}^1 = \frac{3}{4} \int_0^1 [I_pa_n(dM = 1) + I_pa_n(dM = 0) \cos^2 \theta] d\cos \theta$$
 use the similar way as 5-1-1



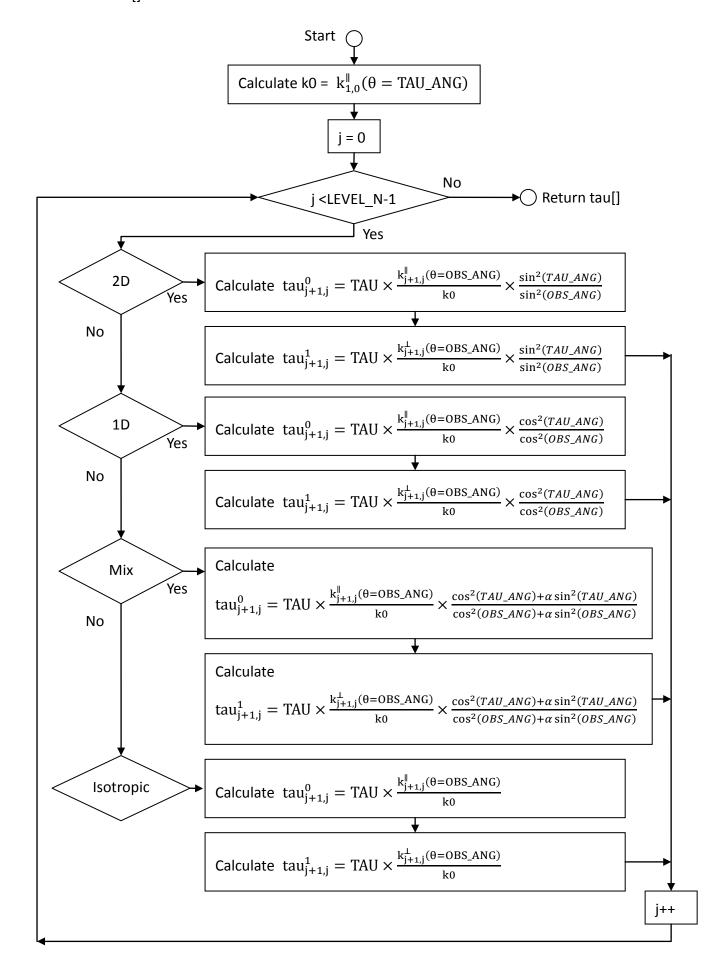
5-2-1 fill the (J,0) row of a_matrix[] with A[] and R[]



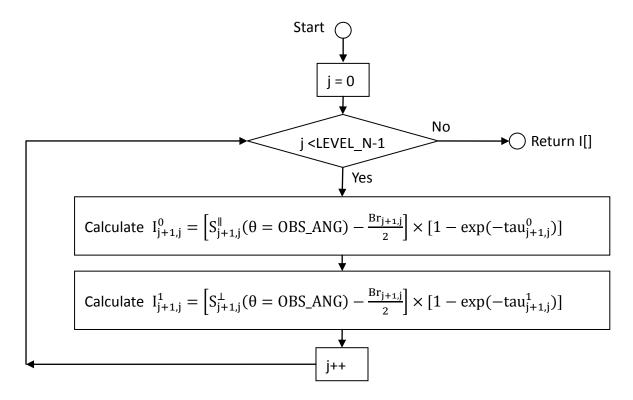
5-2-2 fill the (J,M > 0) row of a_matrix[] with A[] and R[]



6. Calculate the tau[] for each levels



7. Calculate the emerge I[]

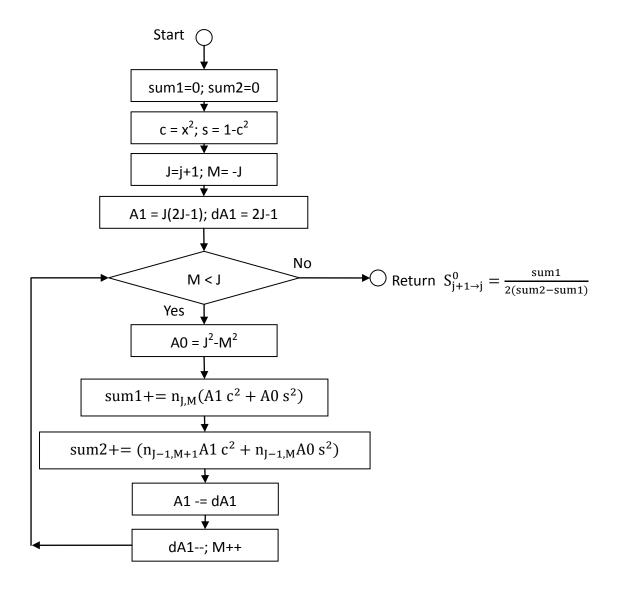


Other function:

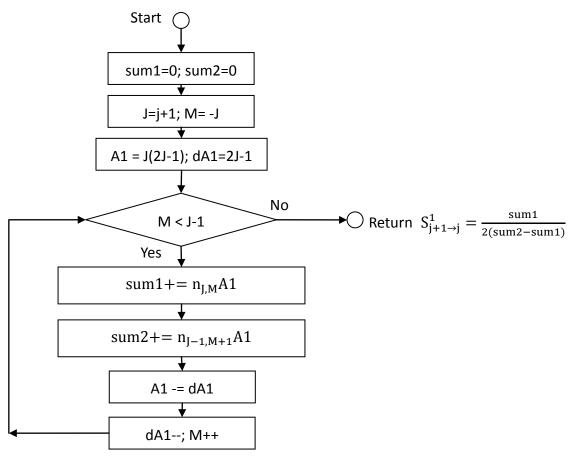
(1) beta_f(
$$\tau$$
 = tau) $\equiv \beta(\tau) = \frac{1-e^{-\tau}}{\tau}$

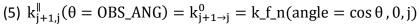
(2) I_pa_n(angle = x = cos
$$\theta$$
, q = dM, tau_d, j) = $S_{j+1\rightarrow j}^q(x) \times (1 - \beta(tau_d)) + \frac{Br_{j+1\rightarrow j}}{2} \times \beta(tau_d)$

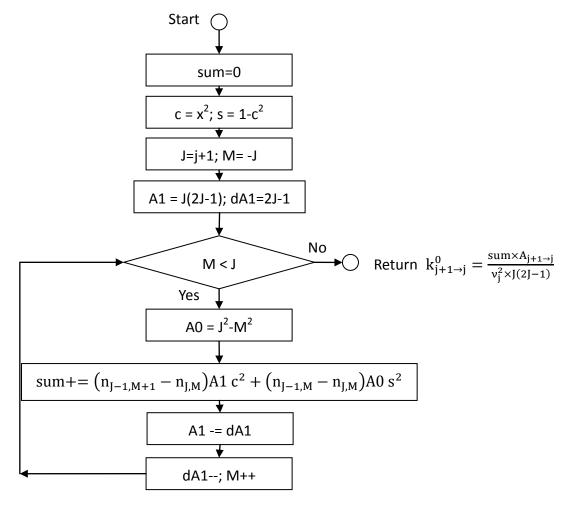
(3)
$$S_{j+1\rightarrow j}^0(x=\cos\theta)=S_{j+1,j}^\parallel(x=\cos\theta)=source_f_n(angle=x=\cos\theta,0,J=j+1)$$



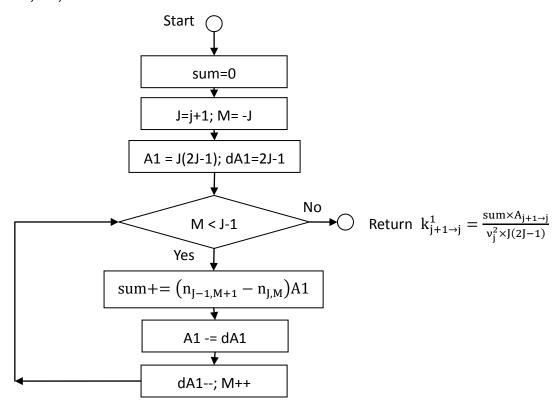
(4) $S_{j+1\rightarrow j}^1=S_{j+1,j}^\perp$ (no angle dependence) = source_f_n(angle = x = cos θ , 1, J = j + 1)







(6) $k_{j+1,j}^{\perp} = k_{j+1\rightarrow j}^{1}$ (no angle dependence) = k_f n(angle = $\cos \theta$, 1, j)



- (7) $Br_{j+1\rightarrow j} = Br_n[j]$: normalized Cosmic Blackbody Radiation intensity for transition j+1 -> j
- (8) $A_{j+1\rightarrow j}=A[\ j\]$: Einstein A coefficients for transition j+1 -> j Note:

$$A_{J,M\to(J-1),M+dM} = A_{J\to(J-1)} \times \begin{cases} \frac{J^2 - M^2}{(2J-1)J} & , dM = 0\\ \frac{(J+M)(J-1+M)}{2(2J-1)J} & , dM = -1\\ \frac{(J-M)(J-1-M)}{2(2J-1)J} & , dM = +1 \end{cases}$$

 $A_{J,M->J-1,M-1}$ is equivalent to the following process:

