

Mapping The Milky Way at 21 cm Hydrogen Line

Mir Sakhawat Hossain,¹★

¹*International Astrostatistics Association, Milano, Italy*

¹*Department of Mathematics, Kabi Nazrul Government College, Laxmi Bazar, Dhaka-1100, Bangladesh*

Accepted XXX. Received YYY; in original form ZZZ

ABSTRACT

After discovery of radiation from Galactic Hydrogen gas clouds in 1951 the 21 cm wavelength hyperfine line of atomic Hydrogen has become the best method to study spectral features in radio astronomy. It has been used as an important tracer for the distribution and velocity of gas clouds in the Inter stellar that has helped enormously to understand the galactic structure. For undergraduate level laboratory experiment it can be assembled a radio frequency receiving system at a low cost to study galactic structure. This paper presents an observation to determine the spiral structure of Milky Way galaxy which was carried out within galactic coordinate longitude 6° to 225° and latitude 0° to 35° with a low cost Haystack model type radio frequency receiving system 2.3 metre SALSA radio telescope at Onsala Space Observatory in Sweden maintained by Chalmers University of Technology. Velocity components of Hydrogen gas clouds were calculated in different galactic longitudes and latitudes as function of distance from galactic centre to plot spiral galactic arms. This observation was made in frequency switching mode and the telescope was remotely operated by internet.

Key words: Galaxy: general – Galaxy: structure – radio lines: galaxies – surveys

1 INTRODUCTION

Neutral Hydrogen(H I) at ground state level is abundant and uniformly distributed element throughout the interstellar medium(ISM). It is the most ubiquitous element in low density regions of ISM but detectable in $\lambda \approx 21$ cm or ≈ 1420 MHz where H₂ is symmetric but not detectable at the radio frequencies. In 1933 Karl Guthe Jansky detected first extraterrestrial radio frequency(Jansky 1933). After that in 45 Then Van de Hulst predicted 21 cm wavelength emission(C J Bakker 1945). The same frequency line also detected by Muller and Oort(Muller & Oort 1951) in the same year. A preliminary survey was taken by Christiansen and Hindman(Christiansen & Hindman 1952) in Australia. They made this survey with a 7.5-m paraboloid and movable radio antenna with beam width between half power 1.9° in horizontal and 2.7° in vertical direction and it covered galactic longitude -10° to $+10^\circ$ at galactic plane. In Netherlands Muller and Westerhout(Muller et al. 1957) took an extended neutral HI line profile survey and made a catalogue approximately in galactic latitude $\pm 20^\circ$ and longitude 318° to 220° . Within these periods angular resolution has been developed from 30 deg to 30- μ s(Kellermann & Moran 2001; Middelberg & Bach 2008). Recently all sky mapping in HI line based on EBHIS and GASS has been completed(Bekhti et al. 2016) with angular resolution 16.2'' and sensitivity $\sigma_{rms} = 43$ mK. Santo and Ashraf(Santo & Uddin 2013) carried out a galactic survey to map Milky Way galaxy in galactic longitude 0° to 225° at galactic

plane using SALSA radio telescope which was built for EU HOU project(Doran et al. 2007). Considering this observation we have completed our observation using the SALSA radio telescope in extended galactic coordinates i.e., galactic longitude 6° to 225° and latitude 0° to 35°. We have discussed here galactic geometry for observable parameters, observation details, data analysis and results with plotting and importance of this project.

2 THEORY

2.1 Hyperfine Splitting of Hydrogen

Neutral Hydrogen consists of a motionless proton(positive charged $+e$) and moving electron(negative charged $-e$) e.g. Fig 1. Electron orbits around proton for the mutual attraction of opposite charges. The derivation is as follows Griffiths(Griffiths 2016). We can imagine electron is orbiting around nucleus(proton). From the view of electron proton is orbiting electron. This circling creates a magnetic field \mathbf{B} in the frame of electron which causes torque on spinning electron. It tends to align its magnetic moment($\boldsymbol{\mu}$) along the direction of magnetic field. So Hamiltonian

$$H = -\boldsymbol{\mu} \cdot \mathbf{B} \quad (1)$$

★ E-mail: s.hossain18@gmail.com

We can calculate magnetic field of proton (\mathbf{B}) and dipole moment of electron ($\boldsymbol{\mu}$) using Biot Savart law which is

$$\mathbf{B} = \frac{\mu_0 I}{2r} \quad (2)$$

Moreover magnetic field \mathbf{B} and angular momentum \mathbf{L} is in the same direction, so

$$\mathbf{B} = \frac{1}{4\pi\epsilon_0} \frac{e}{mc^2 r^3} \quad (3)$$

The direction of magnetic moment $\boldsymbol{\mu}$ and spin \mathbf{S} are the same, so

$$\boldsymbol{\mu} = \frac{q}{2m} \mathbf{S} \quad (4)$$

If charge of electron is e , mass of electron is m_e and spin of electron is \mathbf{S}_e then magnetic moment of electron is,

$$\boldsymbol{\mu}_e = -\frac{e}{m_e} \mathbf{S}_e \quad (5)$$

If mass of proton is m_p , spin of proton is \mathbf{S}_p and g-factor is g_p (measured value is 5.59) then magnetic moment of proton is,

$$\boldsymbol{\mu}_p = \frac{g_p e}{2m_p} \mathbf{S}_p \quad (6)$$

In accordance with classical electrodynamics magnetic field produced by dipole $\boldsymbol{\mu}$ of proton sets up magnetic field

$$\mathbf{B} = \frac{\mu_0}{4\pi r^3} [3(\boldsymbol{\mu} \cdot \hat{\mathbf{r}})\hat{\mathbf{r}} - \boldsymbol{\mu}] + \frac{2\mu}{3}\delta^3(\mathbf{r}) \quad (7)$$

So Hamiltonian of the electron in the magnetic field due to magnetic dipole moment of proton is

$$\begin{aligned} H'_{hf} = & \frac{\mu_0 g_p e^2}{8\pi m_p m_e} \frac{[(3\mathbf{S}_e \cdot \hat{\mathbf{r}})(\mathbf{S}_e \cdot \hat{\mathbf{r}})] - \mathbf{S}_p \cdot \mathbf{S}_e}{r^3} + \\ & \frac{\mu_0 g_p e^2}{3m_p m_e} \mathbf{S}_p \cdot \mathbf{S}_e \delta^3(\mathbf{r}) \end{aligned} \quad (8)$$

In accordance with perturbation the first order correction is the expectation value of the perturbing Hamiltonian

$$\begin{aligned} E'_{hf} = & \frac{\mu_0 g_p e^2}{8\pi m_p m_e} \left\langle \frac{[(3\mathbf{S}_e \cdot \hat{\mathbf{r}})(\mathbf{S}_e \cdot \hat{\mathbf{r}})] - \mathbf{S}_p \cdot \mathbf{S}_e}{r^3} \right\rangle + \\ & \frac{\mu_0 g_p e^2}{3m_p m_e} \langle \mathbf{S}_p \cdot \mathbf{S}_e \rangle |\psi(0)|^2 \end{aligned} \quad (9)$$

In the ground state level the wave function is spherically symmetrical and the first expectation value vanishes. So we get

$$E'_{hf} = \frac{\mu_0 g_p e^2}{3\pi m_p m_e a^3} \langle \mathbf{S}_p \cdot \mathbf{S}_e \rangle \quad (10)$$

This is called spin-spin coupling because of the dot product of two spin \mathbf{S}_p and \mathbf{S}_e . For this coupling the individual spin angular momenta are not conserved. So the good states are eigen vectors of total spin

$$\mathbf{S} \equiv \mathbf{S}_e + \mathbf{S}_p \quad (11)$$

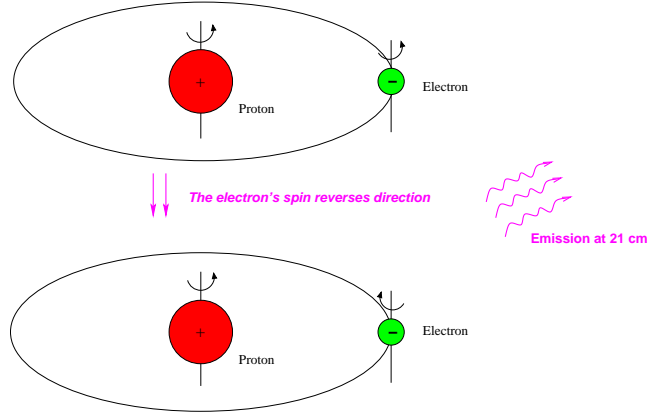


Figure 1. Illustration of the 21 cm transition of the hydrogen atom, caused by the energy change when the electron's spin changes from parallel with the proton's spin to antiparallel.

By applying total spin states the expected perturbation value can be written as terms of eigen operators as follows

$$\mathbf{S}_p \cdot \mathbf{S}_e = \frac{1}{2} (S^2 - S_e^2 - S_p^2) \quad (12)$$

But both of electron and proton have spin 1/2, so $S_e^2 = S_p^2 = (3/4)\hbar^2$. In the triplet state (parallel spins) total spin is 1 and so $S^2 = 2\hbar^2$. In the singlet state total spin is 0 and $S^2 = 0$. Thus

$$E'_{hf} = \frac{4g_p \hbar^4}{3m_p m_e^2 c^2 a^4} \begin{cases} +\frac{1}{4} & \text{(triplet)} \\ -\frac{3}{4} & \text{(singlet)} \end{cases} \quad (13)$$

The spin-spin coupling breaks the spin degeneracy of the ground state lifting the triplet configuration and depressing the singlet. The energy gap is evidently

$$\Delta E = \frac{4g_p \hbar^4}{3m_p m_e^2 c^2 a^4} = 5.88 \times 10^{-6} \text{ eV} \quad (14)$$

The frequency of the photon emitted in a transition from the triplet to the singlet state is,

$$\nu = \frac{\Delta E}{h} = 1420 \text{ MHz} \quad (15)$$

The corresponding wavelength is $c/\nu = 21 \text{ cm}$ which is part of micro wave region. In a single Hydrogen atom this transition occurs once per $\approx 10^7$ years but enormous amount of Hydrogen in spiral arms of Milky Way galaxy causes pervasive and ubiquitous forms of radiation which is observable by radio telescope.

3 METHODS, OBSERVATIONS, SIMULATIONS ETC.

Normally the next section describes the techniques the authors used. It is frequently split into subsections, such as Section 3.1 below.

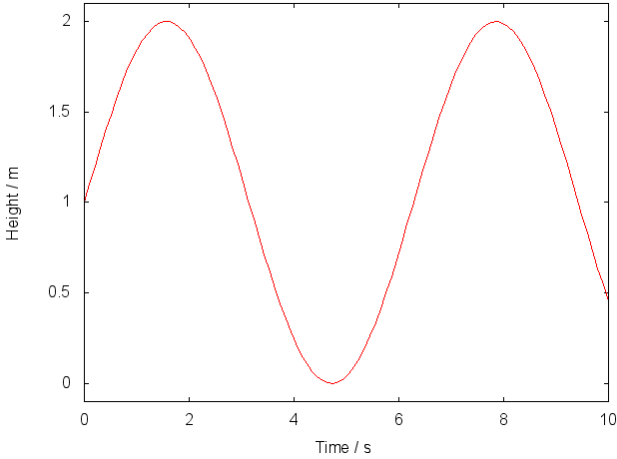


Figure 2. This is an example figure. Captions appear below each figure. Give enough detail for the reader to understand what they're looking at, but leave detailed discussion to the main body of the text.

Table 1. This is an example table. Captions appear above each table. Remember to define the quantities, symbols and units used.

A	B	C	D
1	2	3	4
2	4	6	8
3	5	7	9

3.1 Maths

Simple mathematics can be inserted into the flow of the text e.g. $2 \times 3 = 6$ or $v = 220 \text{ km s}^{-1}$, but more complicated expressions should be entered as a numbered equation:

$$x = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}. \quad (16)$$

Refer back to them as e.g. equation (16).

3.2 Figures and tables

Figures and tables should be placed at logical positions in the text. Don't worry about the exact layout, which will be handled by the publishers.

Figures are referred to as e.g. Fig. 2, and tables as e.g. Table 1.

4 CONCLUSIONS

The last numbered section should briefly summarise what has been done, and describe the final conclusions which the authors draw from their work.

ACKNOWLEDGEMENTS

The Acknowledgements section is not numbered. Here you can thank helpful colleagues, acknowledge funding agencies, telescopes and facilities used etc. Try to keep it short.

REFERENCES

- Bekhti N. B., et al., 2016, *Astronomy & Astrophysics*, 594, A116
 C J Bakker H. C. v. d. H., 1945, *Radiogolven uit het wereldruim.* s-Gravenhage : Martinus Nijhoff
 Christiansen W. N., Hindman J. V., 1952, *Australian Journal of Chemistry*, 5, 437
 Doran R., et al., 2007, in *European Planetary Science Congress 2007*. p. 940, <http://adsabs.harvard.edu/abs/2007epsc.conf..940D>
 Griffiths D. J., 2016, *Introduction to quantum mechanics*. Cambridge University Press
 Jansky K. G., 1933, *Nature*, 132, 66
 Kellermann K. I., Moran J. M., 2001, *Annual Review of Astronomy and Astrophysics*, 39, 457
 Middelberg E., Bach U., 2008, *Reports on Progress in Physics*, 71, 066901
 Muller C. A., Oort J. H., 1951, *Nature*, 168, 357
 Muller C. A., Westerhout G., et al., 1957, *Bulletin of the Astronomical Institutes of the Netherlands*, 13, 151
 Santo T. R., Uddin S. A., 2013, *International Journal of Astronomy*, 2, 37

APPENDIX A: SOME EXTRA MATERIAL

If you want to present additional material which would interrupt the flow of the main paper, it can be placed in an Appendix which appears after the list of references.

This paper has been typeset from a \LaTeX file prepared by the author.