

# OPTIMIZING 3D MODELS FROM 2D IMAGES

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## Abstract

Here comes the abstract

## 1 Introduction

### 1.1 Related work

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### 1.3 Organization of this thesis

## 2 Projective Geometry

## 3 Skyline detection

### 3.1 Introduction

The sky and the earth are separated by a skyline in images. The detection of this skyline has proven to be a very succesful computer vision application in a wide range of domains. In this domain it is used to provide a countour of a building. This contour is in a next step used to refine a 3D model of this building.

The organisation of this chapter is as follows. First related work on skyline detection is discussed, then a new algorithm of the skyline algorithm is described and finally some results are presented.

### 3.2 Related work

A lot of related work on skyline detection is done and it is used in a wide range of domains. [1] yields a good introduction of different skyline detection techniques, these are listed below.

### 3.2.1 Cloud detection for Mars Exploration Rovers (MER)

Mars Exploration Rovers (MER) are used to detect clouds and dust devils on Mars. In [1] their approach is to first identify the sky (equivalentl, the skyline) and then determine if there are clouds in the region segmented as sky.

### 3.2.2 Horizon detection for Unmanned Air Vehicles (UAV)

In this domain, the horizon detector for UAVs can take advantage of the high altitude of the vehicle and therefor the horizon can be approximated to be a straight line. This turns the detection problem into a line-fitting problem. Ofcourse this work is not applicable for detecting a building countour as the straight line assumption doesn't work. But it needs to be mentioned that from this idea some inspiration on line fitting is done because the building countour has straight line segments.

### 3.2.3 Planetary Rover localisation

In [9] they use the skyline detection in planetary rovers, their approach is to combine the detected skyline with a given map of the landscape (hills, roads) to detect its current location. The advantage of their technique is the simplicity and effectiveness of the algorithm which makes it suitable for this project. A big drawback is that it is geared toward speed over extremely high accuracy because it is interactive system where an operator refines the skyline.

As mentioned in the introduction, in this project we use the skyline to extract the building contour to eventually update a 3D model which is a brand new purpose of skyline detection. There is no user interaction present, and the accuracy is a matter of high importance. This makes it different from existing skyline techniques and caution should be taken by using existing algorithms. From the related work the Planetary Rover localisation [9] seemed to fit most on this project. Therefor method [9] is used, but as a basis, and a custom algorithm with higher accuracy is developed. This is explained in the next section.

## 3.3 The Algorithm

### 3.3.1 The original algorithm

The skyline detection algorithm as described in [9] works as follows: The frames are first preprocessed by converting them to Gaussian smoothed images. The skyline of a frame is then detected by analysing the the image columns seperately. The smoothed intensity gradient is calculated from top to bottom. This is done by taking the derivative of the gaussian smoothed image. The system takes the first pixel with gradient higher then a threshold to be classified as a skyline element. This is done for every column in the image. The result is a set of coordinates of length  $W$ , where  $W$  is the width of the image, that represent the skyline.

Taking the smoothed intensity gradient is the most basic method of edge detection and has the disadvantage that it is not robust to more vague edges. This is not surprising as its purpose was an interactive system where the user refines the result. It is clear that an optimization is needed.

### 3.3.2 The optimization

The column based approach seems to be very useful and is therefore unchanged. The effectiveness of the algorithm is totally dependent of the method of edge detection and the preprocessing of the images. The original algorithm uses the smoothed intensity gradient as a way of detecting edges. This is a very basic method and more sophisticated edge detection algorithms are present.

To select a proper edge detector, a practical study is done on the different Matlab built in edge detection techniques. The output of the different edge detection techniques was studied and the Sobel edge detector came with the most promising results. The Sobel edge detector outputs a binary image, therefore the column inlier threshold method is replaced by finding the first white pixel. This is as the original algorithm done from top to bottom for every column in the image.

To make the algorithm more precise, two preprocessing steps are introduced. First the contrast of the image is increased, this makes sharp edges stand out more. Secondly the image undertakes a Gaussian blur, this removes a large part of the noise.

The system now has several parameters which has to be set manually by the user:

- contrast,
- intensity (window size) of Gaussian blur,
- Sobel edge detector threshold,

*Should I write down what parameter values I used or is this of too much detail*

If the user introduces a new dataset these parameters need to be changed as the image quality and lightning condition are probably different.

## 3.4 Results

The system assumes that the first sharp edge (seen from top to bottom) is always the skyline/building edge. This gives rise to some outliers, for example a streetlight or a tree. These outliers are removed as described in the next section. The Skyline detector without outlier removal has an accuracy of 80 %

Some results on the Floriande dataset can be seen in Figure 3.4. *TODO quote Isaac here*

dit is een retest a 3.4



(a) caption2



(b) 1

Figure 1: caption1

## 4 Skyline projection

### 4.1 Introduction

The final product of this research is an accurate 3D model of an urban landscape. This is accomplished in several steps, so far the first step: skyline detection is explained. The next step is to create a basic 3D model of the urban landscape. Then the retrieved skyline of the previous section is used to update the basic 3D model of the building. This is illustrated in Figure 2

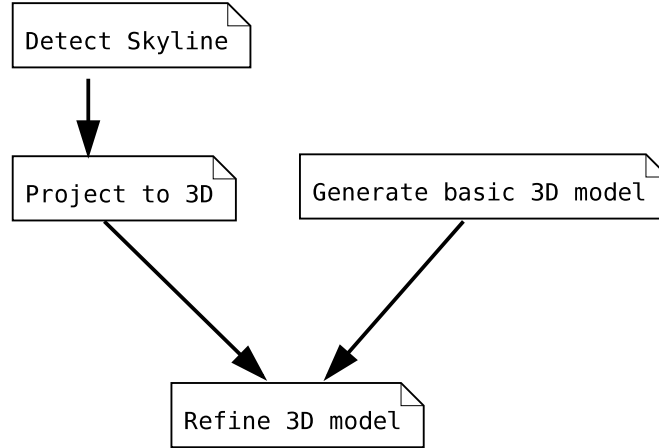


Figure 2: Situation scheme

## 4.2 3D modelling

First a rough 3D model of the urban landscape has to be generated. This is done by taking a top-view Google maps image of the scene and extract the contour of the building, this is done manually. The height of the building is estimated, also manually, and together with the contour a rough 3D model is generated.

*TODO convert to algorithm or method*

## 4.3 Project to 3D space

When the skylines of the 2D images are projected to 3D it can give accurate information about the contour of the building(s). This is used to refine the basic 3D model. Next is explained how the skylines are projected to 3D.

A skyline consist of different skylinepixels. Every 2D skyline pixel presents a 3D point in space. No information is known about the distance from the 3D point to the camera that took the picture. What is known is de 2D location of the pixel which reduces the possible points in 3D space to an infinite line. This line is known and spanned by two coordinates:

- The camera center
- $K'p$ , where  $K$  is the Calibration matrix of the camera and  $p$  is the homogeneous pixel coordinate.

*I don't remember the theory behind it and can't find it in Isaac's paper. Would you explain this Isaac?*

For every skyline pixel a line spanned by the above two coordinates is derived.

## 4.4 Intersect with building

The lines derived as described by the previous section are not enough to refine the 3D model because it is still unknown which skyline part belongs to which part of the 3D model.

Therefor these lines of possibel pixel locations need to be reduced to the actual 3D location of the pixel. This is done by intersecting them with the walls of the rough 3D model. This is done as follows.

The building is first divided into different walls. Every wall of the building spans a plane. Intersections are calculated between the lines and the planes of the building walls.

*Isaac, should I put a intersection formula down here or is this trivial?*

The algorithm returns  $w$  intersections for every skylinepixel (where  $w$  is the number of walls). This is because both the lines and planes infinite and have a very low change of being parallel.

Next challenge is to reduce the number of intersection for every skylinepixel to one. In other words, to determine the wall that is responsible for that pixel. This is used to update the 3D model at the right place.

The details of this process is explained next:

## 4.5 Find most likely wall

### 4.5.1 When is a wall responsible?

Lets define the intersection between the projected skylinepixel line and the plane of the wall as intersection point  $isp$ . And the wall sides as  $w1, w2, w3, w4 \in W$ . And  $d$  as a distance measure which is explained later.

If a certain wall is responsible for a pixel, the intersection (i.e. projected pixel) must lie either

(1) Somewhere on the wall

or

(2) On a small distance  $d$  from that wall

(1) is calculated by testing if the pixel lies inside the polygonal representation of the wall. This is done using the Matlabs in-polygon algorithm. If the in-polygon test succeeds  $d$  is considered to be 0.

(2) Note that in this case the pixel is treated as an inlier because the 3D model is sparse and the height of the building is estimated. It is calculated as follows: First the distances from  $isp$  to the four wall sides are calculated. For every wall the minimum distance is stored.

$$\min_{w \in W} d(isp, w)$$

This is done for every wall. The wall with the smallest distance is the one that

most likely presents the pixel:

$$\arg \min_{W \in Walls} (\min_{w \in W} d(isp, w))$$

If there are two (or even more) walls that are classified equally well to present the pixel (that is if they both succeed the in-polygon or have exactly the same  $d$  value) then the nearest wall is selected. The nearest wall is calculated by taking the wall with the smallest Euclidean distance from the  $isp$  to the camera center. The main algorithm is now explained. In the next section the intersection point - wall distance  $d$  is explained.

#### 4.5.2 Calculate the intersection point - wall distance

Every wall is rectangular and consist four wallsides, a wallside spans an infinite line.

The intersection point ( $isp$ ) is projected orthogonally on these four lines. The projected  $isp$  is denoted as  $isp_{proj}$ . *TODO image? TODO projection formula?*

If  $isp$  is close to a wallside,  $e(isp, isp_{proj})$  (where  $e$  is defined as the Euclidean distance) is small. Note that this doesn't mean that if  $e(isp, isp_{proj})$  is small  $isp$  is always close to the wall.

In Figure 3 the distance between  $p$  and  $p_{proj}$  is very small but  $p$  lies far away from the wall. This is ofcourse because  $p$  is projected to the infinite line spanned by the wallside  $w$  and in this case it got unfortunately projected far next to the wallside. Because of this artefact, it is not robust to calculate the perpendicular projection distance. To make the distant measure more robust  $d$  is calculated differently if  $isp_{proj}$  lies outside the wallside segment of the line. In this case the Euclidean distance between  $isp_{proj}$  and the closest corner-point of the wallside is returned. This is illustrated in Figure 3.

#### 4.5.3 Determine whether $isp_{proj}$ lies on, to the left or to the right of a wallsegment

To determine whether the  $isp_{proj}$  lies on, to the left or to the right of a wallsegment a efficient algorithm is designed. As will be clear in a few moments the projection calculation can be skipped,  $isp$  is used instead together with a computational cheap trick.

Take a look at Figure ?? again, consider the angles between the wallsegment and the the wallsegments endpoints to  $p, q$  and  $r$ . For  $q$  this is  $w, w0q$ , and  $w, w1q$ , note that both angles are acute. This is not true for  $p$  and  $r$  where one of the angles is always obtuse. What can be concluded is that if both angles are equal to or less than  $90^\circ$  then the  $isp_{proj}$  will be located on the wallsegment  $w$ . If not, the  $isp_{proj}$  lies to the left or to the right of the wallsegment  $w$  according to whether one of the angles is acute or obtuse. The angles are acute or obtuse

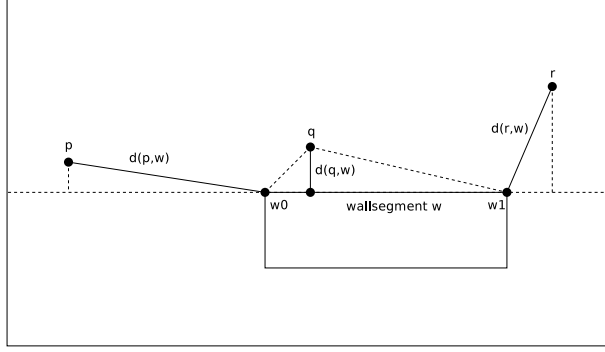


Figure 3: The behavior of the distance measure  $d$ . For  $p$  and  $q$  the euclidean distance to the closest corner point is taken. For  $q$  the perpendicular projection distance is taken

if the dot product of the vectors involved are respectively positive or negative. To summarize: determining in which region the  $isp_{proj}$  lies is boiled down to two dot product calculations with the advantage that the actual projection calculation can be skipped.

*Todo come back on outlier removal*

#### 4.6 Results

### 5 Update 3D model

This module is not finished yet.

### 6 References

- [1] Castano, Automatic detection of dust devils and clouds on Mars.
- [9] Cozman, Outdoor visual position estimation for planetary rovers.