

Welcome to the MPAS tutorial practice guide

This web page is intended to serve as a guide through the practice exercises of this tutorial. Exercises are split into seven sections, corresponding to the seven practical sessions of the tutorial. The details of the exercises in each section can be viewed by clicking on the headers.

In case you would like to refer to the lecture slides, you can find links to PDF versions of each lecture, below.

Monday, 9 September 2019

MPAS Overview

Obtaining and building MPAS-Atmosphere

Running global MPAS, part 1: Initialization for real-data applications

<<< Practice session 1>>>

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Running global MPAS, part 1: Initialization for idealized test cases

Mesh structure

<u>Visualization / analysis tools</u>

Runnning MPAS, part 2: Rotating meshes, streams and I/O, etc.

<<< Practice session 3>>>

Tuesday, 10 September 2019

Dynamics: overview and configuration

Regional MPAS: overview

Regional MPAS: creating a mesh and generating ICs and LBCs

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MPAS software: Registry, pools, and logging

How to add a passive tracer with sources and sinks

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Physics in MPAS

MPAS development: using git and GitHub

Diagnostics framework, and an example of adding a new diagnostic

<<< Practice session 7>>>

MPAS mesh generation

New MPAS capabilities under development

In order to work through these exercises on your own computer, you will need to <u>download the tar file with all input files</u>. Note that the download is about 3.2 GB, and unpacking the .tar.gz file will require about 21.5 GB of space.

After unpacking the input files, you will need to set the environment variable \$MPAS_TUTORIAL to the absolute path of the resulting mpas_tutorial directory.

While going through this practice guide, you may find it helpful to have a copy of the MPAS-Atmosphere Users' Guide available in another window.

Click here to open the Users' Guide in a new window.

Throughout the tutorial, we'll be working directly in our home directories for simplicity. We'll also need to ensure that the following modules are loaded before beginning:

\$ module list

Currently Loaded Modules:

Generally, we would need to install an MPI implementation (e.g., MPICH, OpenMPI, or MVAPICH) as well as parallel-netCDF, netCDF4, and PIO libraries ourselves — if these didn't already exist on our system — before proceeding with these exercises.

1. Compiling MPAS, and creating static files and idealized ICs

In this session, our goal is to obtain a copy of the MPAS source code directly from the MPAS GitHub repository. We'll then compile both the init_atmosphere_model and atmosphere_model programs before using the former to begin the processing of time-invariant, terrestrial fields for use in a real-data simulation. While this processing is taking place, in another terminal window, we'll use any extra time to practice preparing idealized initial conditions.

1.1 Obtaining the MPAS-Model source code

Since the default shell on the classroom machines is *tcsh*, all shell commands throughout this tutorial will be written for tcsh. However, if you prefer a different shell and don't mind mentally translating, e.g., setenv commands to export commands, then feel free to switch shells.

As described in the lectures, the MPAS code is distributed directly from the GitHub repository where it is developed. While it's possible to navigate to https://github.com/MPAS-Dev/MPAS-Model and obtain the code by clicking on the "Releases" tab, it's much faster to *clone* the repository directly on the command-line. From within our \${HOME} directory, we can:

```
$ git clone https://github.com/MPAS-Dev/MPAS-Model.git
```

Cloning the MPAS-Model repository may take a few seconds, and at the end of the process, the output to the terminal should look something like the following:

```
Cloning into 'MPAS-Model'...
remote: Enumerating objects: 71, done.
remote: Counting objects: 100% (71/71), done.
remote: Compressing objects: 100% (47/47), done.
remote: Total 46608 (delta 38), reused 42 (delta 23), pack-reused 46537
Receiving objects: 100% (46608/46608), 19.65 MiB | 23.01 MiB/s, done.
Resolving deltas: 100% (35848/35848), done.
Checking out files: 100% (1540/1540), done.
```

We should now have an MPAS-Model directory:

That's it! We now have a copy of the latest release of the MPAS source code.

1.2 Compiling the MPAS init_atmosphere and atmosphere cores

As mentioned in the lectures, there are two "cores" that need to be compiled from the same MPAS source code: the *init_atmosphere* core and the *atmosphere* core.

Before beginning the compilation process, it's worth verifying that we have MPI compiler wrappers in our path, and that environment variables pointing to the netCDF, parallel-netCDF, and PIO libraries are set:

```
$ which mpif90
/usr/local/cmpwrappers/mpif90

$ which mpicc
/usr/local/cmpwrappers/mpicc

$ echo $NETCDF
/usr/local/netcdf-4.7.0-gcc

$ echo $PNETCDF
/usr/local/parallel-netcdf-1.11.2-gfortran

$ echo $PIO
/usr/local/pio-1.9.23-gcc
```

If all of the above commands were successful, your shell environment should be sufficient to allow the MPAS cores to be compiled.

We'll begin by compiling the *init_atmosphere* core, producing an executable named init_atmosphere_model.

After changing to the MPAS-Model directory

```
$ cd MPAS-Model
```

we can issue the following command to build the init_atmosphere core with the gfortran compiler:

```
$ make -j4 gfortran CORE=init_atmosphere PRECISION=single
```

Note the inclusion of PRECISION=single on the build command; the default is to build MPAS cores as double-precision executables, but single-precision executables require less memory, produce smaller output files, run faster, and — at least for MPAS-Atmosphere — seem to produce results that are no worse than double-precision executables.

In the build command, we have also added the -j4 flag to tell make to use four tasks to build the code; this should reduce the time to compile the init atmosphere model executable significantly!

After issuing the make command, above, compilation should take just a couple of minutes, and if the compilation was successful, the end of the build process should have produced messages like the

```
following:
```

If compilation of the <code>init_atmosphere</code> core was successful, we should also have an executable file named <code>init_atmosphere</code> model:

Note, also, that default namelist and streams files for the <code>init_atmosphere</code> core have also been generated as part of the compilation process: these are the files named <code>namelist.init_atmosphere</code> and <code>streams.init_atmosphere</code>.

Now, we're ready to compile the *atmosphere* core. If we try this without cleaning any of the common infrastructure code, first, e.g.,

\$ make gfortran CORE=atmosphere PRECISION=single

we would get an error like the following:

To remove the init_atmosphere_model_model executable and clean the MPAS infrastructure, run: make clean CORE=init_atmosphere_model

To preserve all executables except atmosphere_model and clean the MPAS infrastructure, run: make clean CORE=atmosphere

Alternatively, AUTOCLEAN=true can be appended to the make command to force a clean, build a new atmosphere_model executable, and preserve all other executables.

After compiling one MPAS core, we need to clean up the shared infrastructure before compiling a different core. We can clean all parts of the infrastructure that are needed by the *atmosphere* core by running the following command:

```
$ make clean CORE=atmosphere
```

Then, we can proceed to compile the atmosphere core in single-precision with:

```
$ make -j4 gfortran CORE=atmosphere PRECISION=single
```

The compilation of the atmosphere_model executable can take several minutes or longer to complete (depending on which compiler is used). Similar to the compilation of the *init_atmosphere* core, a successful compilation of the *atmosphere* core should give the following message:

Now, our MPAS-Model directory should contain an executable file named atmosphere_model, as well as default namelist.atmosphere and streams.atmosphere files:

```
$ 1s -1L
-rwxr-xr-x 1 class114 cbet 11572304 Sep 7 15:09 atmosphere model
```

```
        -rwxr-xr-x
        1
        class114
        cbet
        213424
        Sep
        7
        15:08
        build_tables

        -rw-r-r--
        1
        class114
        cbet
        20580056
        Sep
        7
        15:07
        CAM_ABS_DATA.DBL

        -rw-r-r---
        1
        class114
        cbet
        18208
        Sep
        7
        15:07
        CAM_AEROPT_DATA.DBL

        -rw-r-r---
        1
        class114
        cbet
        261
        Sep
        7
        15:09
        default_inputs

        -rw-r-r---
        1
        class114
        cbet
        261
        Sep
        7
        15:07
        GENPARM.TBL

        -rw-r-r---
        1
        class114
        cbet
        3131
        Sep
        7
        14:55
        INSTALL

        -rw-r-----
        1
        class114
        cbet
        27607
        Sep
        7
        14:55
        INSTALL

        -rw-r-----
        1
        class114
        cbet
        27607
        Sep
        7
        14:55
        Makefile

        -rw-r-----
        1
        class114
        cbet
        174
        Sep
        7
        15:07
        OZONE_DA
```

Note, also, the presence of files named stream_list.atmosphere.*. These are lists of output fields that are referenced by the streams.atmosphere file.

The new files like CAM_ABS_DATA.DBL, LANDUSE.TBL, RRTMG_LW_DATA, etc. are look-up tables and other data files used by physics schemes in MPAS-Atmosphere. These files should all be symbolic links to files in the src/core atmosphere/physics/physics wrf/files/ directory.

If we have init_atmosphere_model and atmosphere_model executables, then compilation of both MPAS-Atmosphere cores was successful, and we're ready to proceed to the next sub-section.

1.3 Static, terrestrial field processing

For our first MPAS simulation in this tutorial, we'll use a quasi-uniform 240-km mesh. Since we'll later be working with a variable-resolution mesh, it will be helpful to maintain separate sub-directories for these two simulations.

To begin processing of the static, terrestrial fields for the 240-km quasi-uniform mesh, let's change to our \${HOME} directory and make a new sub-directory named 240km uniform:

```
$ cd ${HOME}
$ mkdir 240km_uniform
$ cd 240km_uniform
```

We can find a copy of the 240-km quasi-uniform mesh at \${MPAS_TUTORIAL}/meshes /x1.10242.grid.nc. To save disk space, we'll symbolically link this file into our directory:

```
$ ln -s ${MPAS_TUTORIAL}/meshes/x1.10242.grid.nc .
```

We will also need the init_atmosphere_model executable, as well as copies of the namelist.init_atmosphere and streams.init_atmosphere files. We'll make a symbolic link to the executable, and copies of the other files (since we will be changing them):

```
$ 1n -s ${HOME}/MPAS-Model/init_atmosphere_model .
$ cp ${HOME}/MPAS-Model/namelist.init_atmosphere .
$ cp ${HOME}/MPAS-Model/streams.init_atmosphere .
```

Before running the init_atmosphere_model program, we'll need to set up the namelist.init_atmosphere file as described in the lectures:

```
&nhyd_model
    config_init_case = 7
/

&data_sources
    config_geog_data_path = '/classroom/wrfhelp/mpas/geog/'
    config_landuse_data = 'MODIFIED_IGBP_MODIS_NOAH'
    config_topo_data = 'GMTED2010'
    config_vegfrac_data = 'MODIS'
    config_albedo_data = 'MODIS'
    config_maxsnowalbedo_data = 'MODIS'
    config_supersample_factor = 3
/
&preproc_stages
    config_static_interp = true
    config_native_gwd_static = true
    config_vertical_grid = false
```

```
config_met_interp = false
config_input_sst = false
config_frac_seaice = false
```

Note, in particular, the setting of the path to the geographical data sets, and the settings to enable only the processing of static data and the "GWD" static fields.

After editing the namelist.init_atmosphere file, it will also be necessary to tell the init_atmosphere_model program the name of our input grid file, as well as the name of the "static" output file that we would like to create. We do this by editing the streams.init_atmosphere file, where we first set the name of the grid file for the "input" stream:

and then set the name of the static file to be created for the "output" stream:

When interpolating geographical fields to a mesh, the "surface" and "lbc" streams can be ignored (but, you should not delete the "surface" or "lbc" streams, or the init_atmosphere_model program will complain).

After editing the streams.init_atmosphere file, we can begin the processing of the static, time-invariant fields for the 240-km quasi-uniform mesh:

```
$ ./init_atmosphere_model
```

The processing of the static, time-invariant fields will take some time — perhaps up to an hour. In order to verify that the processing is proceeding as expected, we can open another terminal window and change directories to the 240km_uniform directory to check on the progress of the init_atmosphere_model program:

```
$ cd ${HOME}/240km_uniform
$ tail -f log.init_atmosphere.0000.out
```

We should expect to see lines of output similar to the following being printed every few seconds:

```
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/28801-30000.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/30001-31200.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/31201-32400.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/31201-32400.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/33601-34800.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/33601-34800.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/34801-36000.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/36001-37200.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/37201-38400.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/38401-39600.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/39601-40800.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/40801-42000.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/40801-42000.03601-04800
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/00001-01200.04801-06000
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/01201-02400.04801-06000
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/02401-03600.04801-06000
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/02401-03600.04801-06000
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/02401-03600.04801-06000
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/02401-03600.04801-06000
${MPAS_TUTORIAL}/geog/topo_gmted2010_30s/02401-03600.04801-06000
```

If lines similar to the above are being periodically written, we can kill the tail process with CTRL-C before proceeding to the next sub-section.

Important note: If lines of output are not being periodically written to the terminal when running the tail command as described above, look for a file named log.init_atmosphere.0000.err and detemine what went wrong before going on to the next sub-section! Exercises in the second session of this tutorial will require the successful processing of static, time-invariant fields!

1.4 Preparing idealized initial conditions

While the init_atmosphere_model program is processing static, time-invariant fields for a real-data simulation in the 240km_uniform directory, we can try initializing and running an idealized simulation in a separate directory.

The init_atmosphere_model program supports the creation of idealized initial conditions for the following cases:

- 3-d baroclinic wave on the sphere
- 3-d supercell thunderstorm on a doubly-periodic Cartesian plane
- 2-d mountain wave in the xz-plane

For each of these idealized cases, prepared input files are available on the MPAS-Atmosphere download

page. We'll go through the process of creating initial conditions for the idealized supercell case; the general process is similar for the other two cases.

The prepared input files for MPAS-Atmosphere idealized cases may be found by going to the MPAS-Homepage and clicking on the "MPAS-Atmosphere download" link at the left. Then, clicking on the "Configurations for idealized test cases" link will take you to the download links for idealized cases.

Although the input files for these idealized cases can be downloaded through the web page, we can also download, e.g., the supercell input files with wget once we know the URL. Beginning from our \${HOME} directory, we can download the supercell input file archive and unpack it with:

```
$ cd ${HOME}
$ wget http://www2.mmm.ucar.edu/projects/mpas/test_cases/v7.0/supercell.tar.gz
$ tar xzvf supercell.tar.gz
```

After changing to the resulting supercell directory, the README file will give an overview of how to initialize and run the test case.

Once the initial conditions have been created as described in the README file, it's helpful to verify that there were no errors; the end of the log.init_atmosphere.0000.out file should look something like the following:

```
Total log messages printed:

Output messages = 304

Warning messages = 13

Error messages = 0

Critical error messages = 0
```

Assuming there were no errors, the simulation can be run using up to 4 MPI tasks with the mpiexec command:

```
$ mpiexec -n 4 ./atmosphere_model
```

The simulation may take about an hour to run. During one of the later practical sessions, if you would like to revisit the simulation to see the results, running the supercell.ncl NCL script should produce plots of the time evolution of several model fields:

```
$ ncl supercell.ncl
```

2. Creating real-data ICs and running a simulation

Having compiled both the init_atmosphere and atmosphere cores, and also having begun the interpolation of time-invariant, terrestrial fields to create a "static" file for real-data simulations, we'll interpolate atmospheric and land-surface fields to create complete initial conditions for an MPAS simulation in this section. We'll also process 10 days' worth of SST and sea-ice data to update these fields periodically as the model runs. Then, we'll start a five-day simulation, which will take around an hour to complete. Once the model has started running, we can use any extra time to compile the convert mpas utility program.

2.1 Interpolating real-data initial conditions

In the previous session, we started the process of interpolating static geographical fields to the 240-km, quasi-uniform mesh in the 240km_uniform directory. If this interpolation was successful, this directory should contain two new files: log.init_atmosphere.0000.out and x1.10242.static.nc.

```
$ cd ${HOME}/240km_uniform
$ ls -l log.init_atmosphere.0000.out x1.10242.static.nc
```

The end of the log.init_atmosphere.0000.out file should show that there were no errors:

```
Total log messages printed:

Output messages = 2875

Warning messages = 10

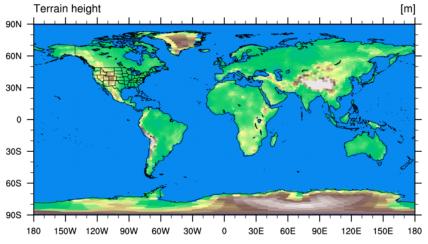
Error messages = 0

Critical error messages = 0
```

We can also make a plot of the terrain elevation field in the x1.10242.static.nc file with an NCL script named plot_terrain.ncl. We'll discuss the use of NCL scripts for visualization in a later session, so for now we can simply execute the following commands to plot the terrain field:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/plot_terrain.ncl .
$ setenv FNAME x1.10242.static.nc
$ ncl plot terrain.ncl
```

Running the script may take up to a minute, but the result should be a figure that looks like the following:



Now that we have convinced ourselves that the processing of static, geographical fields was successful, we can proceed to interpolate atmosphere and land-surface initial conditions for our 240-km simulation. Generally, it is necessary to use the *ungrib* component of the WRF Pre-Processing System to prepare atmospheric datasets in the *intermediate format* used by MPAS. For the purposes of this tutorial, we will simply assume that these data have already been processed with the *ungrib* program in \${MPAS TUTORIAL}/met data/.

To interpolate the NCEP GFS data valid at 0000 UTC on 10 September 2014 to our 240-km mesh, we will symbolically link the GFS:2014-09-10_00 to our working directory:

```
$ ln -s ${MPAS_TUTORIAL}/met_data/GFS:2014-09-10_00 .
```

Then, we will need to edit the namelist.init_atmosphere as described in the lectures to instruct the init_atmosphere_model program to interpolate the GFS data to our mesh. The critical items to set in the namelist.init_atmosphere file are:

```
&nhyd model
   config_init_case = 7
    config_start_time = '2014-09-10_00:00:00'
&dimensions
    config_nvertlevels = 41
    config_nsoillevels = 4
    config_nfglevels = 38
    config_nfgsoillevels = 4
&data sources
    config_met_prefix = 'GFS'
    config_use_spechumd = false
&vertical_grid
   config_ztop = 30000.0
    config_nsmterrain = 1
    config_smooth_surfaces = true
    config dzmin = 0.3
   config_nsm = 30
    config_tc_vertical_grid = true
    config_blend_bdy_terrain = false
&interpolation_control
    config_extrap_airtemp = 'linear'
&preproc stages
    config static interp = false
   config_native_gwd_static = false
    config_vertical_grid = true
    config_met_interp = true
    config_input_sst = false
    config_frac_seaice = true
```

You may find it easier to copy the default namelist.init_atmosphere file from \${HOME}/MPAS-Model/namelist.init_atmosphere before making the edits highlighted above.

When editing the namelist.init_atmosphere file, the key changes are to the starting time for our real-data simulation, the prefix of the *intermediate file* that contains the GFS atmospheric and land-surface fields, and the pre-processing stages that will be run. For this simulation, we'll also use just 41 vertical layers in the atmosphere, rather than the default of 55, so that the simulation will run faster.

After editing the namelist.init_atmosphere file, we will also need to set the name of the input file in the streams.init atmosphere file to the name of the "static" file that we just produced:

```
filename_template="x1.10242.static.nc" input interval="initial only"/>
```

and we will also need to set the name of the output file, which will be the MPAS real-data initial conditions file:

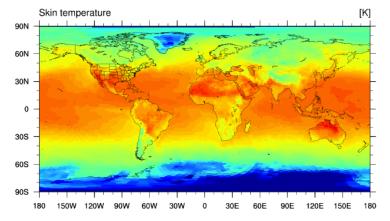
Once we've made the above changes to the namelist.init_atmosphere and streams.init_atmosphere files, we can run the init_atmosphere_model program:

```
$ ./init_atmosphere_model
```

The program should take just a minute or two to run, and the result should be an x1.10242.init.nc file and a new log.init_atmosphere.0000.out file. Assuming no errors were reported at the end of the log.init_atmosphere.0000.out file, we can plot, e.g., the skin temperature field in the x1.10242.init.nc file using the plot_tsk.ncl NCL script:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/plot_tsk.ncl .
$ setenv FNAME x1.10242.init.nc
$ ncl plot tsk.ncl
```

Again, we'll examine the use of NCL scripts for visualization later, so for now we only need to verify that a plot like the following is produced:



If a plot like the above is produced, then we have completed the generation of an initial conditions file for our 240-km real-data simulation!

2.2 Generating SST and sea-ice update files

Because MPAS-Atmosphere — at least, when run as a stand-alone model — does not contain prognostic equations for the SST and sea-ice fraction, these fields would remain constant if not updated from an external source; this is, of course, not realistic, and it will generally impact the quality of longer model simulations. Consequently, for MPAS-Atmosphere simulations longer than roughly a week, it is typically necessary to periodically update the sea-surface temperature (SST) field in the model. For real-time simulations, this is generally not an option, but it is feasible for *retrospective* simulations, where we have observed SST analyses available to us.

In order to create an SST update file, we will make use of a sequence of intermediate files containing SST and sea-ice analyses that are available in \${MPAS_TUTORIAL}/met_data. Before proceeding, we'll link all of these SST intermediate files into our working directory:

```
$ ln -sf ${MPAS_TUTORIAL}/met_data/SST* .
```

Following Section 8.1 of the MPAS-Atmosphere Users' Guide, we must edit the namelist.init_atmosphere file to specify the range of dates for which we have SST data, as well as the frequency at which the SST intermediate files are available. The key namelist options that must be set are shown below; other options can be ignored.

```
&nhyd_model
    config_init_case = 8
    config_start_time = '2014-09-10_00:00:00'
    config_stop_time = '2014-09-20_00:00:00'
/

&data_sources
    config_sfc_prefix = 'SST'
    config_fg_interval = 86400
/
```

```
&preproc_stages
    config_static_interp = false
    config_static_interp = false
    config_vertical_grid = false
    config_met_interp = false
    config_input_sst = true
    config_frac_seaice = true
/
```

Note in particular that we have set the *config_init_case* variable to **8**! This is the initialization case used to create surface update files, instead of real-data initial condition files.

As before, we also need to edit the streams.init_atmosphere file, this time setting the name of the SST update file to be created by the "surface" stream, as well as the frequency at which this update file should contain records:

Note that we have set both the $filename_template$ and $output_interval$ attributes of the "surface" stream. The output_interval should match the interval specified in the namelist for the $config_fg_interval$ variable. The other streams ("input", "output", and "lbc") can remain unchanged — the input file should still be set to the name of the static file.

After setting up the namelist.init_atmosphere and streams.atmosphere files, we're ready to run the init atmosphere model program:

```
$ ./init atmosphere model
```

After the job completes, as before, the log.init_atmosphere.0000.out file should report that there were no errors. Additionally, the file x1.10242.sfc_udpate.nc file should have been created.

We can confirm that the x1.10242.sfc_udpate.nc contains the requested valid times for the SST and sea-ice fields by printing the "xtime" variable from the file; in MPAS, the "xtime" variable in a netCDF file is always used to store the times at which records in the file are valid.

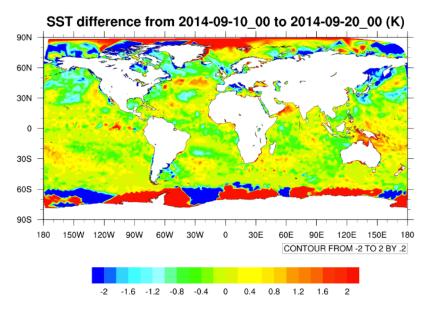
```
$ ncdump -v xtime x1.10242.sfc_update.nc
```

This ncdump command should print the following times:

If the times shown above were printed, then we have successfully created an SST and sea-ice update file for MPAS-Atmosphere. As a final check, we can use the NCL script from \${MPAS_TUTORIAL}/ncl_scripts/plot_delta_sst.ncl to plot the difference in the SST field between the last time in the surface update file and the first time in the file.

Because the surface update file contains no information about the latitude and longitude of grid cells, we need to specify two environment variables when running this script: one to give the name of the surface update file, and the second to give the name of the grid file, which contains cell latitude and longitude fields:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/plot_delta_sst.ncl .
$ setenv FNAME x1.10242.sfc_update.nc
$ setenv GNAME x1.10242.static.nc
$ ncl plot_delta_sst.ncl
```



In this plot, we have masked out the SST differences over land, since the values of the field over land are not representative of actual SST differences, but may represent differences in, e.g., skin temperature or 2-m air temperature, depending on the source of the SST analyses.

2.3 Model integration

Assuming that an initial condition file and a surface update file have been created, we're ready to run the MPAS-Atmosphere model itself!

Working from our 240km_uniform directory, we'll begin by creating a symbolic link to the atmosphere_model executable that we compiled in \${HOME}/MPAS-Model. We'll also make copies of the namelist.atmosphere and streams.atmosphere files as well. Recall from when we compiled the model that there are stream_list.atmosphere.* files that accompany the streams.atmosphere file; we'll copy these as well.

```
$ ln -s ${HOME}/MPAS-Model/atmosphere_model .
$ cp ${HOME}/MPAS-Model/namelist.atmosphere .
$ cp ${HOME}/MPAS-Model/streams.atmosphere .
$ cp ${HOME}/MPAS-Model/stream_list.atmosphere.* .
```

You'll also recall that there were many look-up tables and other data files that are needed by various physics parameterizations in the model. Before we can run MPAS-Atmosphere, we'll need to create symbolic links to these files as well in our working directory:

```
$ ln -s ${HOME}/MPAS-Model/src/core_atmosphere/physics/physics_wrf/files/* .
```

Before running the model, we will need to edit the namelist.atmosphere file, which is the namelist for the MPAS-Atmosphere model itself. The default namelist.atmosphere is set up for a 120-km mesh; in order to run the model on a different mesh, there are several key parameters that depend on the model resolution:

- \bullet config_dt the model timestep (delta-t) in seconds.
- \bullet config_len_disp the length-scale for explicit horizontal diffusion, in meters.

To tell the model the date and time at which integration will begin (important, e.g., for computing solar radiation parameters), and to specify the length of the integration to be run, there are two other parameters that must be set for each simulation:

- config_start_time the start time of the integration.
- \bullet config_run_duration the length of the integration.

Lastly, when running the MPAS-Atmosphere model in parallel, we must tell the model the prefix of the filenames that contain mesh partitioning information for different MPI task counts:

 config_block_decomp_file_prefix — the file prefix (to be suffixed with processor count) of the file containing mesh partitioning information.

Accordingly, we must edit at least these parameters in the namelist.atmosphere file; other parameters are described in Appendix B of the MPAS-Atmosphere Users' Guide, and do not necessarily need to be changed:

```
&nhyd_model
    config_dt = 1200.0
    config_start_time = '2014-09-10_00:00:00'
    config_run_duration = '5_00:00:00'
    config_len_disp = 240000.0
/
```

```
&decomposition
    config_block_decomp_file_prefix = 'x1.10242.graph.info.part.'
/
```

For a relatively coarse mesh like the 240-km mesh, we can also call the radiation schemes less frequently to allow the model to run a little more quickly. Let's set the radiation calling interval to 1 hour:

```
&physics
    config_radtlw_interval = '01:00:00'
    config_radtsw_interval = '01:00:00'
/
```

After changing the parameters shown above in the namelist.atmosphere file, we must also set the name of the initial condition file, the name of the surface update file, and the interval at which we would like to read the surface update file in the streams.atmosphere file:

Having set up the namelist.atmosphere and streams.atmosphere files, we're almost ready to run the model in parallel. You'll recall from the lectures that we need to supply a *graph partition file* to tell MPAS how the horizontal domain is partitioned among MPI tasks. Accordingly, we'll create a symbolic link to the partition file for 4 MPI tasks for the 240-km mesh in our working directory:

```
$ ln -s ${MPAS_TUTORIAL}/meshes/x1.10242.graph.info.part.4 .
```

As a general rule, a mesh with N grid columns should be run on at most N/160 processors in MPAS-Atmosphere to make efficient use of the processors. So, for the mesh in this exercise, which has 10242 grid columns, we could in principle use up to about 64 MPI tasks while still making relatively efficient use of the computing resources. For this exercise, however, we only have 4 cores available to us.

Now, we should be able to run MPAS using 4 MPI tasks:

```
$ mpiexec -n 4 ./atmosphere_model
```

Once the model begins to run, it will write information about its progress to the log.atmosphere.0000.out file. From another terminal window, we can use tail -f to follow the updates to the log file as the model runs:

```
$ tail -f log.atmosphere.0000.out
```

If the model has started up successfully and begun to run, we should see messages like this

```
Begin timestep 2014-09-10_06:00:00
--- time to run the LW radiation scheme L_RADLW =T
--- time to run the SW radiation scheme L_RADSW =T
--- time to run the convection scheme L_CONV =T
--- time to apply limit to accumulated rainc and rainnc L_ACRAIN =F
--- time to apply limit to accumulated radiation diags. L_ACRADT =F
--- time to calculate additional physics_diagnostics =F
split dynamics-transport integration 3

global min, max w -0.177687 0.512867
global min, max u -112.127 112.008
Timing for integration step: 5.78077 s
```

written for each timestep that the model takes. If not, a log.atmosphere.0000.err file may have been created, and if so, it may have messages to indicate what might have gone wrong.

The model simulation will take about an hour to complete. As it runs, you may notice that several netCDF output files are being created:

- diag.2014.09.*.nc These files contain mostly 2-d diagnostic fields that were listed in the stream_list.atmosphere.diagnostics file.
- history.2014.09.*.nc These files contain mostly 3-d prognostic and diagnostic files from the model.
- restart.2014.09.*.nc These are model restart files that are essentially checkpoints of the model state; a simulation can be re-started from any of these checkpoint/restart files.

If the model has successfully begun running, we can move to the next sub-section to obtain and compile the convert mpas utility for interpolating model output files to a lat-lon grid for quick visualization.

2.4 Obtaining and compiling the convert_mpas program

We saw in earlier sections that NCL scripts may be used to visualize fields in MPAS netCDF files. Although these scripts can produce publication-quality figures, many times all that is needed is a way to quickly inspect the fields in a file. The <code>convert_mpas</code> program is designed to interpolate fields from the unstructured MPAS mesh to a regular lat-lon grid for easy checking, e.g., with noview.

As with the MPAS-Model code, the convert_mpas code may be obtained most easily by cloning a GitHub repository. Working from our \${HOME} directory, we'll download the code in this way:

```
$ cd ${HOME}
$ git clone https://github.com/mgduda/convert_mpas.git
$ cd convert_mpas
```

The convert_mpas utility is compiled via a simple *Makefile*. In general, one can edit the Makefile in the main convert mpas directory.

We can build the convert mpas utility by simply running make with no command-line arguments:

```
$ make
```

If compilation was successful, there should now be an executable file named <code>convert_mpas</code>. So that we can run this program from any of our working directories, we'll add our current directory to our shell <code>\${PATH}</code> environment variable:

```
$ setenv PATH ${HOME}/convert mpas:${PATH}
```

You may like to take a few minutes to read through the **README**.md file to familiarize yourself with the usage of the **convert_mpas** program. We'll have a chance to exercise the **convert_mpas** utility in the next practical session.

3. Running MPAS with variable resolution, and basic visualization

In the previous sessions, we were able to make a five-day, real-data simulation on a 240-km quasi-uniform mesh. In this session, we'll practice preparing a variable-resolution, 240km – 48km mesh with the <code>grid_rotate</code> tool, after which we'll proceed as we did for the quasi-uniform case to process static, terrestrial fields on this variable-resolution mesh.

As we saw in the first session, processing the static fields can take some time, so while this processing is taking place, we'll practice using the <code>convert_mpas</code> tool to interpolate output from our 240-km quasi-uniform simulation to a lat-lon grid for quick visualization.

Once the static field processing has completed for the 240km – 48km variable-resolution mesh, we'll interpolate the initial conditions fields and start a simulation on this mesh. This simulation will take several hours to complete, so we'll want to get this running so that it will have completed by the next session.

3.1 Obtaining, building, and running the grid_rotate program

A common theme with MPAS is the need to obtain source code from *git* repositories hosted on GitHub. The <code>grid_rotate</code> utility is no exception in this regard. One difference, however, is that the <code>grid_rotate</code> code is contained in a repository that houses many different tools for MPAS cores, including the MPAS-Albany Land Ice core, the MPAS-Ocean core, and the MPAS-Seaice core.

The <code>grid_rotate</code> utility is found in the MPAS-Tools repository, which we can clone into our home directory with the following commands:

```
$ cd ${HOME}
$ git clone https://github.com/MPAS-Dev/MPAS-Tools.git
```

Within the MPAS-Tools repository, the grid_rotate code resides in the mesh_tools/grid_rotate sub-directory. We can change to this directory and build the grid_rotate utility with the make command:

```
$ cd MPAS-Tools/mesh_tools/grid_rotate
$ make
```

The netCDF library on these machines is quite old, so there may be some build warnings about missing ne-config commands; in this particular case, these can be ignored, but in general, one should be concerned with build warnings.

So that we can more easily run the <code>grid_rotate</code> program from any working directory, we'll add our current working directory to our shell <code>\${PATH}</code>.

```
$ setenv PATH ${HOME}/MPAS-Tools/mesh tools/grid rotate:${PATH}
```

Now that we have a means to rotate the location of refinement in a variable-resolution MPAS mesh, we can begin the process of preparing such a mesh. For this tutorial, we'll be using a mesh that refines from 240-km grid spacing down to about 48-km grid spacing over an elliptic region.

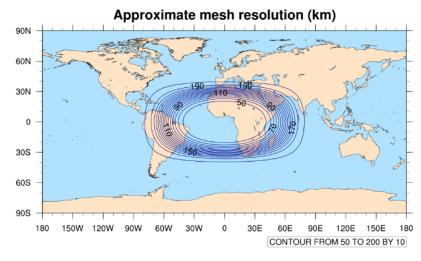
When creating initial conditions and running the 240-km quasi-uniform simulation, we worked in a sub-directory named 240km_uniform. For our variable-resolution simulation, we'll work in a sub-directory named 240-48km_variable to keep our work separated. Let's create the 240-48km_variable directory and link in the variable-resolution mesh from \${MPAS_TUTORIAL}/meshes into that directory:

```
$ cd ${HOME}
$ mkdir 240-48km_variable
$ cd 240-48km_variable
$ ln -s ${MPAS TUTORIAL}/meshes/x5.30210.grid.nc .
```

The x5.30210.grid.nc mesh contains 30210 grid cells and refines by a factor of five (from 240-km grid spacing to 48-km grid spacing). The refinened region is centered over (0.0 lat, 0.0 lon), and we can get a quick idea of the size of the refinement by plotting contours of the approximate mesh resolution with the mesh_resolution.ncl NCL script:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/mesh_resolution.ncl .
$ setenv FNAME x5.30210.grid.nc
$ ncl mesh resolution.ncl
```

When running the mesh_resolution.ncl script, a plot like that below should be produced.



Using the <code>grid_rotate</code> tool, we can change the location and orientation of this refinement to an area of our choosing. Before running the <code>grid_rotate</code> program, we need a copy of the input namelist that is read by this program:

```
$ cp ${HOME}/MPAS-Tools/mesh_tools/grid_rotate/namelist.input .
```

As described in the lectures, we'll set the parameters in the namelist.input file to relocate the refined region in the mesh. For example, to move the refinement over South America, we might use:

```
config_original_latitude_degrees = 0
  config_original_longitude_degrees = 0

config_new_latitude_degrees = -19.5
  config_new_longitude_degrees = -62
  config_birdseye_rotation_counter_clockwise_degrees = 90
//
```

Feel free to place the refinement wherever you would like!

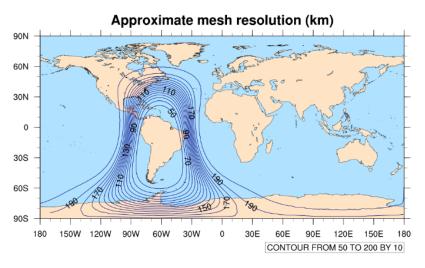
After making changes to the namelist.input file, we can rotate the x5.30210.grid.nc file to create a new "grid" file with any name we choose. For example:

```
$ grid_rotate x5.30210.grid.nc SouthAmerica.grid.nc
```

As with the original x5.30210.grid.nc file, we can plot the approximate resolution of our rotated grid file with commands like the following:

```
$ setenv FNAME SouthAmerica.grid.nc
$ ncl mesh_resolution.ncl
```

Of course, the name <code>southAmerica.grid.nc</code> should be replaced with the name that you have chosen for your rotated grid when setting the <code>FNAME</code> environment variable. In our example, the resulting plot looks like the one below.



It may take some iteration to get the refinement just where you would like it. You can go back and edit namelist.input file, then re-run the grid_rotate program, and finally, re-run the ncl mesh resolution.ncl command until you are satisfied!

3.2 Static, terrestrial field processing

Interpolating static, geographical fields to a variable-resolution MPAS mesh works in essentially the same way as it does for a quasi-uniform mesh. So, the steps taken in this section should seem familiar — they mirror what we did in Section 1.3!

To begin the process, we'll symbolically link the init_atmosphere_model executable into our 240-48km_variable directory, and we'll copy the default namelist.init_atmosphere and streams.init atmosphere files as well:

```
$ In -s ${HOME}/MPAS-Model/init_atmosphere_model .
$ cp ${HOME}/MPAS-Model/namelist.init_atmosphere .
$ cp ${HOME}/MPAS-Model/streams.init_atmosphere .
```

As before, we'll set the following variables in the namelist.init_atmosphere file:

```
&nhyd model
    config_init_case = 7
&data sources
    __config_geog_data_path = '/classroom/wrfhelp/mpas/geog/'
    config landuse data = 'MODIFIED IGBP MODIS NOAH
    config_topo_data = 'GMTED2010'
    config_vegfrac_data = 'MODIS'
    config_albedo_data = 'MODIS'
    config_maxsnowalbedo_data = 'MODIS'
    config_supersample_factor = 3
&preproc_stages
   config static interp = true
   config native gwd static = true
   config_vertical_grid = false
   config_met_interp = false
   config_input_sst = false
    config_frac_seaice = false
```

In editing the streams.init_atmosphere file, we'll choose the filename for the "input" stream to match whatever name we chose for our rotated grid file:

and we'll set the filename for the "output" stream to follow the same naming convention, but using "static" instead of "grid":

With these changes, we can run the ${\tt init_atmosphere_model}$ program:

```
$ ./init atmosphere model
```

As before, processing of the static, geographical fields may take up to an hour.

3.3 Using the convert_mpas utility

While the static, geographical fields are being processed for our 240km – 48km variable-resolution mesh, we can experiment with the use of the convert_mpas utility to interpolate output from our 240-km, quasi-uniform simulation.

If you didn't have a chance to download and compile the <code>convert_mpas</code> program back in Section 2.4, you will need to go back to that section before proceeding with the rest of this section.

In another terminal window, we'll change to the 240km_uniform directory, and we may also need to add the directory containing the convert_mpas utility to our shell \${PATH} variable again:

```
$ cd ${HOME}/240km_uniform
$ setenv PATH ${HOME}/convert_mpas:${PATH}
```

In the 240km_uniform directory, we should see output files from MPAS that are named history*nc and diag*nc:

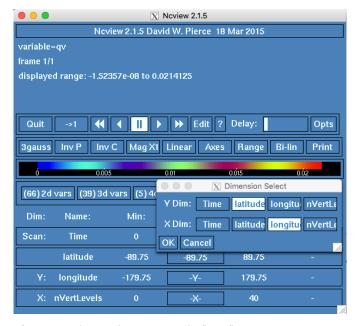
As a first try, we can interpolate the fields from the final model "history" file, history.2014-09-15_00.00.00.nc, to a 0.5-degree lat-lon grid with the command:

```
$ convert_mpas history.2014-09-15_00.00.00.nc
```

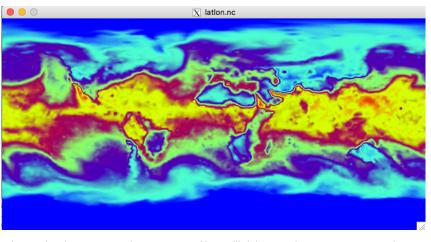
The interpolation should take less than a minute, and the result should be a file named latlon.nc. Let's take a look at this file with the neview utility:

\$ ncview latlon.nc

Depending on which field you've chosen to view first, you may need to change the coordinate axes that are displayed in ncview, e.g.,



After setting the coordinate axes in the "Axes" menu, we can, e.g., view the qv field, which may look something like the plot below at the lowest model level.



When we're done viewing the latlon.nc file, we'll delete it. The convert_mpas utility attempts to append to existing latlon.nc files when converting a new MPAS netCDF file, and this often doesn't work as expected, so to avoid confusion, it's best to remove the latlon.nc file we just created with

\$ rm latlon.nc

before converting another MPAS output file.

As we just saw, it can take upwards of a minute to interpolate a single MPAS "history" file (at least, on these classroom machines). If we are only interested in a small subset of the fields in an MPAS file, we can restrict the interpolation to just those fields by providing a file named <code>include_fields</code> in our working directory.

Using an editor of your choice, create a new text file named include_fields with the following contents:

u10 v10 precipw

After saving the file, we can re-run the <code>convert_mpas</code> program as before to produce a <code>latlon.nc</code> file with just the 10-meter surface winds and precipitable water:

- \$ convert_mpas history.2014-09-15_00.00.00.nc
- \$ ncview latlon.nc

The convert_mpas program should have taken just a few seconds to run; if not, make sure the latlon.nc file was deleted before re-running convert_mpas. The only fields in the new latlon.nc file should be "precipw", "u10", and "v10".

Rather than converting just a single file at a time, we can convert multiple MPAS output files and write the results to the same latlon.nc file. The key to doing this is to use the first command-line argument to convert_mpas to specify a file that contains horizontal (SCVT) mesh information, and to let subsequent command-line arguments specify MPAS netCDF files that contain fields to be remapped.

To interpolate the "precipw", "u10", and "v10" fields for our entire 5-day simulation, we can use the following commands:

- \$ rm latlon.nc
- \$ convert_mpas x1.10242.init.nc history.*.nc
- \$ ncview latlon.nc

Note, again, that the first command-line argument to <code>convert_mpas</code> is only used to obtain SCVT mesh information by the <code>convert_mpas</code> program when multiple command-line arguments are given, and all remaining command-line arguments specifie netCDF files with fields to be interpolated!

When viewing a netCDF file with multiple time records, you can step through the time periods in neview with the "forward" and "backward" buttons to the right and left of the "pause" button:



As a final exercise in viewing MPAS output via the convert_mpas and noview programs, we'll zoom in on a part of the globe, looking at the 250 hPa vertical vorticity field over East Asia on a 1/10-degree lat-lon grid.

To do this, we will first need to edit the include_fields file so that it contains just the following line:

vorticity_250hPa

Then, we'll create a new file named target_domain with the editor of our choice. This file should contain the following lines:

```
startlat=30
endlat=60
startlon=90
endlon=150
nlat=300
nlon=600
```

After editing the include_fields file and creating the target_domain file, we can remove the old latlon.nc file and use the convert_mpas program to interpolate the 250 hPa vorticity field from our diag.*.nc files:

```
$ rm latlon.nc
$ convert_mpas x1.10242.init.nc diag*nc
$ ncview latlon.nc
```

If the processing of the static, geographical fields for the 240km – 48km variable-resolution mesh has not finished, yet, you can feel free to experiment more with the convert mpas program!

3.4 Interpolating real-data initial conditions

Once the processing of the static, geographical fields has completed in the 240-48km_variable directory, we should have a new "static" file as well as a log.init_atmosphere.0000.out file. In our example, the static file is named southAmerica.static.nc, but your file should have whatever name was specified in the definition of the "output" stream in the streams.init atmosphere file.

Before proceeding, it's a good idea to verify that there were no errors reported in the log.init_atmosphere.0000.out file. We can also make a plot of the terrain height field as we did for the 240-km quasi-uniform simulation:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/plot_terrain.ncl .
$ setenv FNAME SouthAmerica.static.nc
$ ncl plot terrain.ncl
```

Once we have convinced ourselves that the static, geographical processing was successful, we can proceed as we did for the 240-km quasi-uniform simulation. We'll create a symbolic link to the NCEP GFS intermediate file valid at 0000 UTC on 10 September 2014:

```
$ ln -s ${MPAS_TUTORIAL}/met_data/GFS:2014-09-10_00 .
```

As before, we'll set up the namelist.init_atmosphere file to instruct the init_atmosphere_model program to interpolate meteorological and land-surface initial conditions:

```
&nhyd model
   config_init_case = 7
    config_start_time = '2014-09-10_00:00:00'
   config_nvertlevels = 41
   config_nsoillevels = 4
   config nfglevels = 38
   config_nfgsoillevels = 4
&data_sources
   config_met_prefix = 'GFS'
   config_use_spechumd = false
&vertical_grid
   config_ztop = 30000.0
   config_nsmterrain = 1
   config_smooth_surfaces = true
   config_dzmin = 0.3
   config_nsm = 30
    config_tc_vertical_grid = true
    config_blend_bdy_terrain = false
&interpolation control
   config_extrap_airtemp = 'linear'
&preproc stages
   config_static_interp = false
   config_native_gwd_static = false
   config_vertical_grid = true
   config_met_interp = true
   config_input_sst = false
   config frac seaice = true
```

Besides the namelist.init_atmosphere file, we must also edit the XML I/O configuration file, streams.init_atmosphere, to specify the name of the static file that will serve as input to the init_atmosphere_model program, as well as the name of the MPAS-Atmosphere initial condition file to be created. Specifically, we must set the filename_template for the "input" stream to the name of our static file:

```
filename_template="SouthAmerica.static.nc"
input interval="initial only"/>
```

and we must set the name of the initial condition file to be created in the "output" stream:

After editing the namelist.init_atmosphere and streams.atmosphere files, and linking the intermediate file to our run directory, we can run the init_atmosphere_model program.

```
$ ./init_atmosphere_model
```

Once the init_atmosphere_model program finishes, as always, it is best to verify that there were no error messages reported in the log.init_atmosphere.0000.out file. We should also have a file whose name matches the filename we specified for the "output" stream in the streams.init_atmosphere file. In our example, we should have a file named SouthAmerica.init.nc.

If all was successful, we are ready to run a variable-resolution MPAS simulation. For this simulation, we won't prepare an SST and sea-ice surface update file for the model; however, if you would like to try doing so following what was done in Section 2.2, feel free to do so!

3.5 Model integration

Assuming that an initial condition file, and, optionally, a surface update file have been created, we're ready to run a variable-resolution, 240km – 48km simulation with refinement over a part of the globe that interests us.

As a first step, you may recall that we need to create symbolic links to the atmosphere_model executable, as well as to the physics look-up tables, into our working directory:

```
$ ln -s ${HOME}/MPAS-Model/atmosphere_model .
$ ln -s ${HOME}/MPAS-Model/src/core_atmosphere/physics/physics_wrf/files/* .
```

We will also need *copies* of the namelist.atmosphere, streams.atmosphere, and stream_list.atmosphere.* files:

```
$ cp ${HOME}/MPAS-Model/namelist.atmosphere .
$ cp ${HOME}/MPAS-Model/streams.atmosphere .
$ cp ${HOME}/MPAS-Model/stream list.atmosphere.* .
```

Next, we will need to edit the namelist.atmosphere file. Recall from before that the default namelist.atmosphere is set up for a 120-km mesh; in order to run the model on a different mesh, there are several key parameters that depend on the model resolution:

- config_dt the model timestep (delta-t) in seconds.
- config_len_disp the length-scale for explicit horizontal diffusion, in meters.

The model timestep, <code>config_dt</code> must be set to a value that is appropriate for the smallest grid distance in the mesh. In our case, we need to choose a time step that is suitable for a 48-km grid distance; a value of 240 seconds is reasonable. In the lecture on MPAS dynamics we will say more about how to choose the model timestep.

Similarly, we need to set the mixing length-scale according to the smallest nominal grid distance in a variable-resolution mesh, so we can set config len disp to 48000 meters.

To tell the model the date and time at which integration will begin, and to specify the length of the integration to be run, there are two other parameters that must be set for each simulation:

- config_start_time the start time of the integration.
- \bullet config_run_duration the length of the integration.

In order to make maximal use of the time between this practical session and the next, we'll try to run a 5-day simulation starting on 2014-09-10_00.

Lastly, when running the MPAS-Atmosphere model in parallel, we must tell the model the prefix of the filenames that contain mesh partitioning information for different MPI task counts:

 config_block_decomp_file_prefix — the file prefix (to be suffixed with processor count) of the file containing mesh partitioning information.

For the 240km - 48km variable-resolution mesh, we can find a mesh partition file with 4 partitions at \${MPAS_TUTORIAL}/meshes/x5.30210.graph.info.part.4. Let's symbolically link that file to our run directory now before specifying a value of 'x5.30210.graph.info.part.' for the config_block_decomp_prefix in our namelist.atmosphere file.

```
$ ln -s ${MPAS_TUTORIAL}/meshes/x5.30210.graph.info.part.4 .
```

Before reading further, you may wish to try making the changes to the namelist.atmosphere yourself.

Once you've done so, you can verify that the options highlighted below match the changes you have

made.

```
&nhyd_model
    config_dt = 240.0
    config_start_time = '2014-09-10_00:00:00'
    config_run_duration = '5_00:00:00'
    config_len_disp = 48000.0
//
&decomposition
    config_block_decomp_file_prefix = 'x5.30210.graph.info.part.'
//
```

After changing the parameters shown above in the namelist.atmosphere file, we must also set the name of the initial condition file in the streams.atmosphere file:

Optionally, if you decided to create an SST update file for this variable-resolution simulation in the previous section, you can also edit the "surface" stream in the streams.atmosphere file to specify the name of that surface update file, as well as the interval at which the model should read updates from it:

Having set up the namelist.atmosphere and streams.atmosphere files, we're ready to run the model in parallel:

```
$ mpiexec -n 4 ./atmosphere_model
```

Once the model begins to run, you may like to open another terminal window, change to the 240-48km_variable directory, and to "tail" the log.atmosphere.0000.out file with the "-f" option to follow the progress of the MPAS simulation:

```
$ tail -f log.atmosphere.0000.out
```

If the model has started up successfully and begun to run, we should see messages like this:

```
Begin timestep 2014-09-10_00:12:00
--- time to run the LW radiation scheme L_RADLW =F
--- time to run the SW radiation scheme L_RADSW =F
--- time to run the convection scheme L_CONV =T
--- time to apply limit to accumulated rainc and rainnc L_ACRAIN =F
--- time to apply limit to accumulated radiation diags. L_ACRADT =F
--- time to calculate additional physics_diagnostics =F
split dynamics-transport integration 3

global min, max w -1.06611 0.722390
global min, max u -117.781 118.251
Timing for integration step: 4.18681 s
```

This simulation will take several hours to run. Now may be a convenient time to ask any questions that you may have!

4. Running limited-area simulations

Having gained experience in running global simulations, we are now well positioned to run a limited-area (i.e., a regional) simulation with MPAS-Atmosphere. We will begin by defining the simulation domain using the MPAS-Limited-Area tool to cover any part of the globe that interests us. Then, we will subset the "static" fields from an existing 60-km quasi-uniform static file.

Once we are satisfied with the geographic coverage of our regional static fields, we will interpolate initial and lateral boundary conditions from the 0.5-degree CFSv2 datset. Given ICs and LBCs for our regional domain, we can begin running a 3-day reigonal simulation from 2019-09-01 0000 UTC through 2019-09-04 0000 UTC.

After our simulation has run (or, even after the first few model output files are available), we will practice using NCL and the convert_mpas tool to visualize the regional output.

4.1 Defining a regional domain and subsetting static fields

To begin the work of defining a regional simulation domain for MPAS-Atmosphere, we will first obtain a copy of the MPAS-Limited-Area tool. As with the MPAS model code and MPAS-Tools code, we will clone the MPAS-Limited-Area repository from GitHub. From within our \${HOME} directory, we can do this with the following commands:

```
$ cd ${HOME}
$ git clone https://github.com/MiCurry/MPAS-Limited-Area.git
```

After changing to the MPAS-Limited-Area directory that should have been created by the git clone command, we can prepare a file describing our desired regional domain. Depending on the type of region that we intend to define (circular, elliptical, channel, or polygon), we can use any of the example region definition files in the docs/points-examples/ sub-directory as a starting point.

For the purpose of illustration, we'll use the japan.ellipse.pts file from docs/points-examples/ as a starting point to set up an elliptical domain over the Mediterranean Sea; however, you may set up your regional domain over any other part of the globe using any of the available region types. After copying this file to a file named mediterranean.pts in our current working directory with

```
$ cp docs/points-examples/japan.ellipse.pts mediterranean.pts
```

we can edit the file so that its contents look like the following:

Name: Mediterranean Type: ellipse Point: 37.9, 18.0 Semi-major-axis: 3200000 Semi-minor-axis: 1700000 Orientation-angle: 100

Again, you can set up any alternative regional domain as you prefer — the Mediterranean domain is only used here as an example!

To aid in locating your regional domain, it may be helpful to use OpenStreetMap, where you can rightclick anywhere on a map and choose "Show address" to find the latitude and longitude of a point.

Having prepared a region definition file, we can use the <code>create_region</code> tool to create a subset of an existing global "static" file for our specified region. We can symbolically link a global, quasi-uniform, 60-km static file to our working directory with:

```
$ ln -s ${MPAS_TUTORIAL}/meshes/x1.163842.static.nc .
```

Then, we can create a regional static file with a command like:

```
$ ./create_region mediterranean.pts x1.163842.static.nc
```

The create_region tool should take just a few seconds to run, and the result should be a new static file and a new graph.info file, named according to the "Name" we specified in our region defintion file, e.g,:

```
Mediterranean.static.nc
Mediterranean.graph.info
```

At this point, we may like to plot one of the time-invariant, terrestrial fields from our new regional static file. We can plot the terrain height field using an NCL script, regional_terrain.ncl, which we will first copy to our working directory:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/regional_terrain.ncl .
```

This NCL script uses a custom color map. In order for NCL to find this color map, we'll need to set the NCARG COLORMAPS environment variable as follows:

```
$ setenv NCARG_COLORMAPS ${MPAS_TUTORIAL}/ncl_colormaps/:$NCARG_ROOT/lib/ncarg/colormaps/
```

Before running the script, we will also need to set the latitude and longitude view point for the orthographic projection used by the script. In the regional_terrain.ncl script, locate the section that looks like the following, and modify the cenlat and cenlon values so that they coincide roughly with the center of your regional domain.

Additionally, we will need to set the **GNAME** environment variable to the name of our regional static file; e.g., for our Mediterranean example:

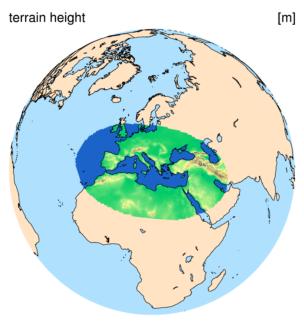
```
$ setenv GNAME Mediterranean.static.nc
```

We can then run the script with

```
$ ncl regional_terrain.ncl
```

to produce a PNG image that should look something like the following:

Mediterranean.static.nc



Before proceeding to the next section, you may need to iterate between (1) modifying your region definition file, (2) running the create_region tool to create a new netCDF file, and (3) plotting the new regional terrain field with the regional_terrain.ncl script.

4.2 Interpolating regional initial conditions (ICs)

Now that we have a regional static file for our domain of interest, we can prepare initial conditions for this domain in essentially the same way as we prepared initial conditions for global simulations. The key difference is that we will blend the terrain field along the regional domain boundaries with the terrain from our initial conditions data.

Let's create a new directory in \${HOME} to run our regional simulation:

```
$ mkdir ${HOME}/regional
$ cd ${HOME}/regional
```

We can then move the regional static file and graph.info files to this directory:

```
$ mv ${HOME}/MPAS-Limited-Area/Mediterranean.static.nc .
$ mv ${HOME}/MPAS-Limited-Area/Mediterranean.graph.info .
```

As with the global simulations from earlier sessions, we will need to link the <code>init_atmosphere_model</code> executable to our run directory, and we will also need to <code>copy</code> the <code>namelist.init_atmosphere</code> and <code>streams.atmosphere</code> files, which we will modify for our regional domain.

```
$ In -s ${HOME}/MPAS-Model/init_atmosphere_model .
$ cp ${HOME}/MPAS-Model/namelist.init_atmosphere .
$ cp ${HOME}/MPAS-Model/streams.init_atmosphere .
```

For our regional simulation, we will use 0.5-degree CFSv2 data valid on 2019-09-01 0000 UTC for initial conditions. We can symbolically link the CFSV2 intermediate file to our run directory as follows:

```
$ ln -s ${MPAS_TUTORIAL}/met_data/CFSV2:2019-09-01_00 .
```

Unlike in past exercises, we can also try interpolating the initial conditions using multiple MPI tasks. Recall that we use the Metis tool to create mesh partition files for a specified number of MPI tasks. Using Metis's <code>gpmetis</code> tool, we can create a partition file for 4 MPI tasks from the <code>Mediterranean.graph.info</code> file that was generated by MPAS-Limited-Area. The following command should be sufficient:

```
$ gpmetis -minconn -contig -niter=200 Mediterranean.graph.info 4
```

Even if we choose not to interpolate the initial conditions in parallel, we will ultimately need the mesh partition file to run the model in parallel.

Now, we are ready to edit our namelist.init_atmosphere and streams.atmosphere files. To test your understanding, you may like to first try preparing these files yourself for the creation of initial conditions for your regional domain. Once you've finished editing these files, you can compare against the changes highlighted below (click on the light blue header bars). Remember that your static file and graph.info file may have different names, depending on where your regional domain is located.

The edited namelist.init_atmosphere file

The edited streams.init_atmosphere file

For namelist and streams settings that were not modified, either the default value of the setting is acceptable, or the setting is not used when interpolating time-varying initial conditions. Now is also a good time to check that you understand why we needed to modify the settings that we did in the namelist.init_atmosphere and streams.atmosphere files.

After checking your modifications to the namelist.init_atmosphere and streams.atmosphere files, above, we can run the init_atmosphere_model program with 4 MPI tasks to create regional initial conditions:

```
$ mpiexec -n 4 ./init_atmosphere_model
```

As always, it's important to verify that there were no errors during the execution of the init_atmosphere_model program; the end of the log.init_atmosphere.0000.out file should report zero errors:

```
Total log messages printed:

Output messages = 464

Warning messages = 6

Error messages = 0

Critical error messages = 0
```

If there were no errors, we should now have a regional initial conditions file; for our example regional domain the file should be named Mediterranean.init.nc. We're now ready to generate lateral boundary conditions for our domain in the next section! We're now ready to generate lateral boundary conditions for our domain in the next section!

4.3 Interpolating regional lateral boundary conditions (LBCs)

The interpolation of lateral boundary conditions for a regional simulation domain is conceptually similar to the interpolation of initial conditions: the key differences are: (1) fields are interpolated at several time periods, and (2) the set of fields to be interpolated are a subset of those required for initial conditions (they include: theta, rho, u, w, and moisture species).

We will use the CFSv2 reanalyses for lateral boundary conditions. Three days' worth of six-hourly CFSv2 fields are available in the \${MPAS_TUTORIAL}/met_data directory; we can symbolically link all of these CFSv2 files to our run directory with the following command:

```
$ ln -sf ${MPAS_TUTORIAL}/met_data/CFSV2* .
```

All that is needed now is to once again edit the namelist.init_atmosphere and streams.init_atmosphere files to set up the interpolation of LBCs using "init case" 9. The relevant settings are highlighted in the sections, below (click the light blue headers to expand each edited file).

The edited namelist.init_atmosphere file

```
&nhyd_model
    config_init_case = 9
    config_start_time = '2019-09-01_00:00:00'
config_stop_time = '2019-09-04_00:00:00'
    config_theta_adv_order = 3
    config_coef_3rd_order = 0.25
&dimensions
    config_nvertlevels = 55
    config_nsoillevels = 4
    config_nfglevels = 38
    config_nfgsoillevels = 4
&data_sources
   config geog data path = '/glade/work/wrfhelp/WPS GEOG/'
    config met prefix = 'CFSV2'
    config_sfc_prefix = 'SST'
    config_fg_interval = 21600
    config_landuse_data = 'MODIFIED_IGBP_MODIS_NOAH'
    config topo data = 'GMTED2010'
    config_vegfrac_data = 'MODIS'
    config_albedo_data = 'MODIS'
    config maxsnowalbedo data = 'MODIS'
    config_supersample_factor = 3
    config_use_spechumd = false
&vertical_grid
    config_ztop = 30000.0
    config_nsmterrain = 1
    config_smooth_surfaces = true
    config_dzmin = 0.3
    config nsm = 30
    config_tc_vertical_grid = true
    config_blend_bdy_terrain = false
```

```
%interpolation_control
    config_extrap_airtemp = 'linear'
/
&preproc_stages
    config_static_interp = false
    config_native_gwd_static = false
    config_vertical_grid = true
    config_met_interp = true
    config_input_sst = false
    config_input_sst = false
    config_frac_seaice = true
/
&io
    config_pio_num_iotasks = 0
    config_pio_stride = 1
/
&decomposition
    config_block_decomp_file_prefix = 'Mediterranean.graph.info.part.'
/
```

The edited streams.init_atmosphere file

```
<streams>
<immutable_stream name="input"</pre>
                   type="input"
                   filename_template="Mediterranean.init.nc"
                   input_interval="initial_only" />
<immutable_stream name="output"</pre>
                   type="output"
                   filename_template="foo.nc"
                   packages="initial_conds"
                   output interval="initial_only" />
<immutable stream name="surface"</pre>
                   type="output"
                   filename_template="x1.40962.sfc_update.nc"
                   filename_interval="none"
                   packages="sfc_update"
                   output interval="86400" />
<immutable_stream name="lbc"</pre>
                   type="output"
                   filename_template="lbc.$Y-$M-$D_$h.$m.$s.nc"
                   filename_interval="output_interval"
                   packages="lbcs"
                   output_interval="6:00:00" />
</streams>
```

There are a few points worth making about the changes in the streams.init_atmosphere file.

- Because the interpolation of LBCs is not only a horizontal interpolation to the regional mesh, but also
 a vertical interpolation to the vertical coordinate surfaces, the "input" stream must read from a file
 that contains a valid "zgrid" field. The initial conditions file contains horizontal and vertical mesh
 information, and so we use it as the file from which the "input" stream reads.
- The MPAS stream manager infrastructure will complain if an input stream and an output stream share the same filename_template, so the filename_template for the "output" stream must be set to some other filename; since "init case" 9 only writes the "lbc" stream, we an make up any filename we like for the "output" stream.
- The output_interval for the "lbc" stream must match the value of the config_fg_interval namelist option. In this regional example, using "21600" or "6:00:00" are equivalent (we could even specify "360:00" if we would like!).

After having edit our namelist.init_atmosphere and streams.init_atmosphere files, we are ready to run the init_atmosphere_model program again using 4 MPI tasks:

```
$ mpiexec -n 4 init atmosphere model
```

After the init_atmosphere_model program completes, it is a good idea (as always!) to verify that there were no errors reported at the end of the log.init_atmosphere.0000.out file:

```
Total log messages printed:

Output messages = 2742
Warning messages = 0
Error messages = 0
Critical error messages = 0
```

If there were no errors, we should now have "lbc" files in our run directory:

```
$ ls -l lbc*nc

-rw-r--r- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-01_00.00.00.nc

-rw-r--r- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-01_06.00.00.nc
```

```
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-01_12.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-01_18.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-02_00.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-02_00.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-02_12.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-02_18.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-02_18.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-03_00.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-03_00.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-03_18.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-03_18.00.00.nc
-rw-r-r-- 1 class112 cbet 15469296 Sep 5 16:37 lbc.2019-09-03_18.00.00.nc
```

We're now ready to run a regional simulation in the next section!

4.4 Running a regional simulation

Given regional initial conditions and lateral boundary conditions, we are almost ready to run a regional simulation. The only differences from running a global simulation are: (1) we need to set a namelist option to enable the application of LBCs in each model timestep, and (2) we need to set up an LBC input stream for the model.

As in previous exercises, we will need to symbolically link the atmosphere_model executable, as well as the look-up tables used by physics schemes, into our run directory. We can do this with the following commands:

```
$ ln -s ${HOME}/MPAS-Model/atmosphere_model .
$ ln -s ${HOME}/MPAS-Model/src/core_atmosphere/physics/physics_wrf/files/* .
```

We also need copies of the default namelist.atmosphere, streams.atmosphere, and stream list.atmosphere.* files, which we will edit for our simulation.

```
$ cp ${HOME}/MPAS-Model/namelist.atmosphere .
$ cp ${HOME}/MPAS-Model/streams.atmosphere .
$ cp ${HOME}/MPAS-Model/stream list.atmosphere.* .
```

As with global simulations, the key namelist parameters that must be modified are:

- config_dt the model integration timestep; roughly six times the minimum grid distance in km is a reasonable first guess
- config_start_time the starting time for the simulation
- config run duration the duration of the simulation
- config_len_disp the explicit horizontal mixing length scale; the minimum grid distance in meters is an appropriate setting
- config_block_decomp_file_prefix the prefix (excluding the final number) of the file providing
 mesh partition information

Different from global simulations, we must also set the namelist option config_apply_lbcs = true!

You may like to try editing the namelist.atmosphere file yourself before checking your changes against those highlighted, below (click on the light blue header).

The edited namelist.atmosphere file

```
&nhvd model
   config time integration order = 2
   config_dt = 360.0
   config_start_time = '2019-09-01_00:00:00'
   config_run_duration = '3_00:00:00'
   config_split_dynamics_transport = true
   config number of sub steps = 2
   config_dynamics_split_steps = 3
   config h mom eddy visc2 = 0.0
   config_h_mom_eddy_visc4 = 0.0
   config_v_mom_eddy_visc2 = 0.0
   config_h_theta_eddy_visc2 = 0.0
   config_h_theta_eddy_visc4 = 0.0
   config_v_theta_eddy_visc2 = 0.0
   config_horiz_mixing = '2d_smagorinsky'
   config_len_disp = 60000.0
   config_visc4_2dsmag = 0.05
   config_w_adv_order = 3
   config theta adv order = 3
   config_scalar_adv_order = 3
   config_u_vadv_order = 3
   config_w_vadv_order = 3
   config_theta_vadv_order = 3
   config_scalar_vadv_order = 3
   config scalar advection = true
   config_positive_definite = false
   config monotonic = true
   config coef 3rd order = 0.25
   config_epssm = 0.1
   config_smdiv = 0.1
```

```
&damping
    config_zd = 22000.0
    config_xnutr = 0.2
&limited area
    config_apply_lbcs = true
&io
    config_pio_num_iotasks = 0
    config_pio_stride = 1
&decomposition
    config_block_decomp_file_prefix = 'Mediterranean.graph.info.part.'
    config_do_restart = false
&printout
   config print global minmax vel = true
    config_print_detailed_minmax_vel = false
&IAU
    config_IAU_option = 'off'
    config IAU window length s = 21600.
&physics
    config_sst_update = false
   config_sstdiurn_update = false
    config_deepsoiltemp_update = false
   config_radtlw_interval = '00:30:00'
    config_radtsw_interval = '00:30:00'
    config_bucket_update = 'none'
    config_physics_suite = 'mesoscale_reference'
&soundings
    config sounding interval = 'none'
```

In the streams.atmosphere file, we will need to set the filename_template for the "input" stream to the name of our regional initial conditions file. We will also need to set the input_interval for the "lbc_in" stream to the interval at which we have prepared lateral boundary conditions files.

After editing the streams.atmosphere file, you can compare your changes against those highlighted, below (click on the light blue header).

The edited streams.atmosphere file

Once we have finished editing the namelist.atmosphere and streams.atmosphere files, we can launch our regional simulation using 4 MPI tasks:

```
$ mpiexec -n 4 ./atmosphere model
```

This simulation will take a while to run. However, once the first few netCDF output files have been produced, we can proceed to the next section to practice visualizing regional model output.

4.5 Visualizing regional output

If the regional simulation ran successfully — even if it has just begun running and has produced several time periods of output — we can visualize the output using two different methods.

In order to take a quick first look at the output, we can use the convert_mpas tool to interpolate our output to a regular latitude-longitude grid, after which we can view the interpolated fields using noview.

Let's first check whether the path to the convert_mpas program is in our \$PATH environment variable with:

```
$ which convert mpas
```

If the "which" command prints something like

```
convert_mpas: Command not found.
```

we can set our \$PATH environment variable by typing:

```
$ setenv PATH ${HOME}/convert_mpas:$PATH
```

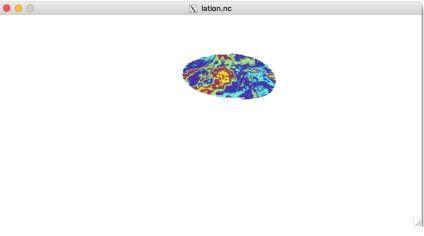
To start, we can simply interpolate the full set of fields from all diagnostics output files to a global 0.5-degree lat-lon grid with the following command:

```
$ convert_mpas Mediterranean.static.nc diag*nc
```

We can then view the resulting latlon.nc file with ncview:

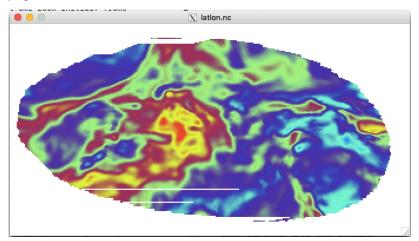
```
$ ncview latlon.nc
```

Because the default target grid for the convert_mpas program is a global, 0.5-degree grid, viewing, e.g., the "precipw" field might look like the image, below.



As we did in Section 3.3, we can try to interpolate to a lat-lon grid that is windowed around our regional domain. Following Section 3.3 to prepare a target_domain file, can you produce a latlon.ne file that looks something like the following?

Remember: it's generally best to remove any existing lation.nc file before re-running the convert_mpas program!



As mentioned in the lectures, the <code>convert_mpas</code> program may produce some interpolation artifacts as of the time of this tutorial. We expect to have this fixed in the near future, but for now, we can ignore any interpolation artifacts and animate through all time slices of several model diagnostic fields in <code>ncview</code> to gain confidence that our regional simulation is producing plausible results.

After we've taken some time to browse the interpolated diagnostic fields with noview, we can make use of NCL to produce figures more suitable for use in publications or presentations.

In \${MPAS_TUTORIAL}/ncl_scripts is a template script, regional_cells.ncl, for plotting the individual cell values from a regional MPAS netCDF file. We can copy this script to our run directory with the following command:

```
$ cp ${MPAS_TUTORIAL}/ncl_scripts/regional_cells.ncl .
```

By default, this script plots the vegetation type field, <code>ivgtyp</code>, with the view-point set to 20.0 N latitude, 80.0 E longitude. With the editor of your choice, open the <code>regional_cells.ncl</code> file and locate the section of the script with the following (it should be around line 68):

```
;
; Center latitude and longitude;
cenlat = 20.0
cenlon = 80.0
```

Let's change the center latitude and longitude values for the view-point to the approximate center of our regional domain. For our example Mediterranean domain, the following settings might be reasonable, but if your regional domain is located over a different part of the globe, you should modify these values as necessary.

```
;
; Center latitude and longitude
;
cenlat = 38.0
```

```
cenlon = 20.0
```

In general, the regional_cells.ncl script reads three environment variables:

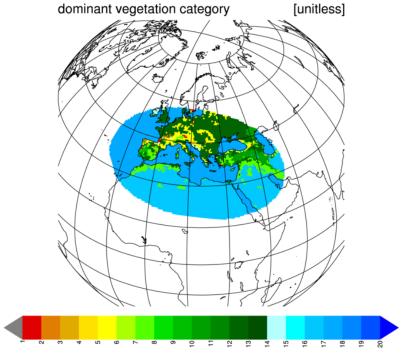
- FNAME the name of the file containing the field to be plotted
- GNAME the name of the file containing the mesh information
- T for time-varying fields, the time period to plot (0 is the first time in a file)

In order to plot the vegetation type field from our initial conditions files, the following settings of these variables will work:

```
• FNAME = Mediterranean.init.nc
• GNAME = Mediterranean.init.nc
```

After setting these environment variables and running the NCL script, we should have a plot equivalent to the one, below.

2019-09-01_00:00:00



The part of the regional_cells.ncl script that handles the actual plotting is general enough to handle any field — we just need to set an appropriate plotting range and color table in the script when changing fields. Next, we can try to modify the script to plot the precipitable water ("precipw") field using a different color map.

Before proceeding, you may like to open $\underline{\text{the NCL Color Table Gallery}}$ in a new tab to see the available NCL color tables.

With the editor of your choice, open the regional_cells.ncl file and locate the section of the script (beginning around line 84) where the field to be plotted, the color map, and the contour levels are set. Let's modify this section of the script so that it looks like the following:

```
; The field to be plotted
;
h = f->precipw(iTime,:)

;
; The color map to use
; (see https://www.ncl.ucar.edu/Document/Graphics/color_table_gallery.shtml)
;
colormap = "GMT_ocean"
;
; The number of colors in the color map
;
nColors = 80
;
; The contour intervals
; E.g., fspan(1, 10, 10) yields (/1, 2, 3, 4, 5, 6, 7, 9, 10/)
; Typically, we have nColors-1 levels
```

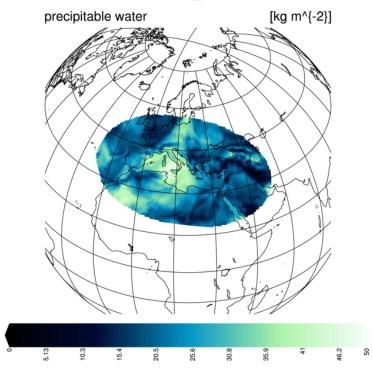
```
;
cnLevels = fspan(0.0, 50.0, nColors-1)
```

We located the "GMT_ocean" color table from the NCL Color Table Gallery link, above. Before re-running the script, we will need to set the **fname** environment variable to the name of one of our diagnostics model output files. If our regional simulation has progressed through at least forecast hour 24, we could use the following:

\$ setenv FNAME diag.2019-09-02_00.00.00.nc

We can then re-run the script to produce a plot like the one below.

2019-09-02_00:00:00



With the remaining time in this session, feel free to experiment with plotting other fields from any of the netCDF output files. Browsing the model output that has been converted to a latitude-longitude grid with the convert_mpas before deciding on the field to plot and the range of values to set in the regional_cells.ncl Script may be helpful.

5. Defining new output streams, and visualization

In this session, we'll restart a simulation from the end of the five-day, global 240-km simulation that we started in the second session. Before restarting the simulation, though, we'll define new output streams.

Optionally, before restarting the simulation, we can also try adding a list of sounding locations for the simulation.

After launching the restart simulation with new output streams (and perhaps some sounding output), we'll practice using some NCL scripts to visualize the first five days of output from the variable-resolution 240-km – 48-km simulation

5.1 Setting up the restart simulation

As we saw in the lectures, setting up a simulation to restart from a previously run simulation is relatively simple. We'll begin by making sure that we're in the directory where we ran the original, 5-day 240-km global simulation:

\$ cd \${HOME}/240km_uniform

Before making any changes to the namelist.atmosphere file, we will first check to see which restart files are available; if our original simulation completed successfully, we should have one restart file per day, valid at the end of each of the five days of the simulation:

```
$ ls -1 restart*nc
-rw-r--r- 1 class101 cbet 196953464 Jul 29 12:26 restart.2014-09-11_00.00.00.nc
-rw-r--r- 1 class101 cbet 196953464 Jul 29 12:31 restart.2014-09-12_00.00.00.nc
-rw-r--r- 1 class101 cbet 196953464 Jul 29 12:35 restart.2014-09-13_00.00.00.nc
-rw-r--r- 1 class101 cbet 196953464 Jul 29 12:40 restart.2014-09-14_00.00.00.nc
-rw-r--r- 1 class101 cbet 196953464 Jul 29 12:44 restart.2014-09-15_00.00.00.nc
```

In principle, we could restart our simulation from any of these files, but we'll begin from the last available file, valid at 0000 UTC on 15 September.

There are just two changes that need to be made to the namelist.atmosphere file: we need to set the starting time of the simulation to 2014-09-10_00:00:00 and we need to set the config_do_restart option to true:

```
&nhyd_model
    config_start_time = '2014-09-15_00:00:00'
/
&restart
    config_do_restart = true
/
```

We can keep all of the other namelist options as they were, including the <code>config_run_duration</code> option — restarting the simulation and running for another five days should be fine, since we had a total of 10 days of SST and sea-ice information in our surface update file.

Before running the model, we'll make a few other changes as described in the next sections!

5.2 Defining new output streams

We may also like to write out some additional fields from our restart simulation at a higher temporal frequency. You may recall from the lectures that we can define new output streams in the streams.atmosphere file.

We will provide an example here of writing out the 10-meter AGL horizontal winds every 60 minutes, but you can feel free to look through Appendix D of the MPAS-Atmosphere Users' Guide to see if there are any other fields that you may like to write out.

Let's define a new output stream named "sfc_winds" in the streams.atmosphere file:

This new stream must be defined somewhere between the opening <streams> and </streams> tags in the streams.atmosphere file.

Note in the stream definition above that the files that will be written will have names like surface_winds.2014-09-15_00.nc, surface_winds.2014-09-15_01.nc, etc. by our choice of the
"filename_template". Also, MPAS will create new files at each output time, rather than packing many
time records into the same file; this is controlled by the "filename_interval" option. Finally, note that the
stream will be written every 60 minutes. Since our simulation starts at 00:00:00 UTC, the minutes and
seconds part of the output files is unnecessary. Of course, we may want to add \$m and \$s to our filename
template if we were writing this stream more frequently than once per hour.

In the next section, we'll describe the steps to write out vertical profiles at specific (lat,lon) locations before we run our restart simulation.

5.3 Adding soundings to the simulation

It may at times also be helpful to write out vertical profiles of basic model fields at specified (lat,lon) locations from a model simulation. In MPAS terminology, this is described as writing "soundings" from the simulation.

In principle, we can write out soundings for as many locations as we would like. Here, we'll describe how to write out soundings from three locations: Boulder, Colorado; McMurdo Station, Antarctica; and New Delhi, India.

We will need to specify the locations for the soundings in a text file named <code>sounding_locations.txt</code> in our model run directory:

```
40.0 -105.25 Boulder
28.7 77.2 NewDelhi
-77.85 166.67 McMurdo
```

The three columns of the file contain the latitude, longitude, and name of the sounding. The latitude and longitude are given as decimal degrees, and the name cannot contain whitespace (since it will be used to name output files).

For this exercise, if you would prefer to write out soundings for different locations, feel free to change the list of locations as you would like!

Lastly, we need to specify how frequently these soundings will be written during the model simulation. This is done through the config sounding interval option in the namelist.atmosphere file. By

default, this option can be found near the bottom of the file. Let's choose to write soundings every three hours:

```
&soundings
config_sounding_interval = '3:00:00'
/
```

Having set the output interval for soundings, we're ready to make a restart simulation.

5.4 Running the restart simulation

Now that we've edited our namelist.atmosphere file for our restart run, we've added a new output stream with surface winds to the streams.atmosphere file, and we've specified a list of locations for which profiles will be written every three hours, we can run the model with 4 MPI tasks:

```
$ mpiexec -n 4 ./atmosphere_model
```

In another terminal window, it's a good idea to change to the 240km_uniform directory and check several points:

- Does running tail -f log.atmosphere.0000.out show timesteps being taken by the model?
- Do you see new output files named surface_winds.2014-09-15_00.nc, etc.?
- Do you see sounding files for Boulder, New Delhi, and McMurdo Station as files with the .snd suffix?

If all of the above check out, our restart simulation is successfully running! In the next section, we'll look at a couple of new NCL scripts for plotting model output.

5.5 Using NCL to visualize output

Now that our restart simulation is running, we can examine in more detail an example NCL script for plotting horizontal fields as individual cells. To try out this script, we can work with output from our 240km – 48km variable-resolution simulation.

In order to plot horizontal fields, we can begin by ensuring that we are in the 240-48km_variable directory before copying over the latlon_cells.ncl script:

```
$ cd ${HOME}/240-48km_variable
$ cp ${MPAS_TUTORIAL}/ncl_scripts/latlon_cells.ncl .
```

This script reads the name of the MPAS netCDF file with SCVT mesh information from and environment variable named GNAME (roughly, "G" means "grid"). The MPAS netCDF file that contains the actual field to be plotted is identified by the environment variable FNAME (roughly, "F" means "field").

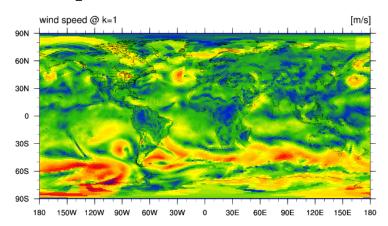
We'll plot a field from the diag.2014-09-11_00.00.00.nc file. The mesh information can be found in the MPAS initial conditions file, which may be named SouthAmerica.init.nc, though you may have chosen a different name corresponding to a different location of refinement. We'll set the GNAME and FNAME environment variables before proceeding:

```
$ setenv GNAME SouthAmerica.init.nc
$ setenv FNAME diag.2014-09-11_00.00.00.nc
```

Remember, your initial conditions file may have a different name!

By default, the latlon_cells.ncl script plots the lowest-model-level wind speed, and the output is saved to a PNG image in a file named cell_plot.png. Running the script should produce this plot, which you can view with the eog command.

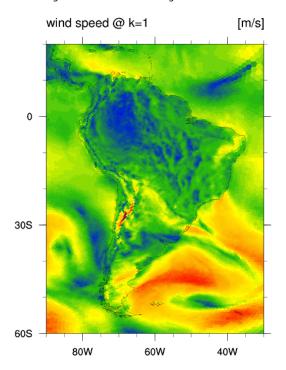
```
$ ncl latlon_cells.ncl
$ eog cell_plot.png
```



Since we are working with a variable-resolution mesh, we may like to set a plotting window to cover only part of the refined region in our simulation. We can set a plotting window in the latlon_cells.ncl script by editing the section of the script that looks like this:

```
; The bounding box for the plot mapLeft = -180.0 mapRight = 180.0 mapBottom = -90.0 mapTop = 90.0
```

Feel free to set the plotting window however you would like for your variable-resolution simulation. For example, if we were to set the plotting window to cover South America before re-running the script, the result might look like the following:



Finally, we may like to plot other fields. We can see which fields were written to the diag.*.nc files using ncdump:

```
$ ncdump -h diag.2014-09-11_00.00.00.nc
```

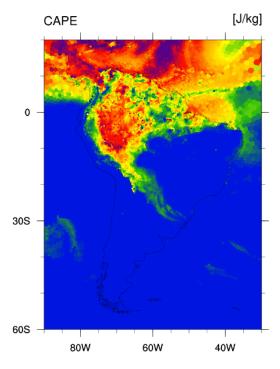
If, for example, we would like to plot Convective Available Potential Energy (CAPE), we could edit the latlon_cells.ncl script near the comment "The field to be plotted" so that it looked like the following:

```
; The field to be plotted
h = f->cape(0,:)

; The field name
field_name = "CAPE"
field_units = "[J/kg]"

; The range of field values to plot (start, stop, number)
cnLevels = fspan(0,3000,50)
```

Generally, we will need to assign the variable to be plotted, specify the plotting name and units for the field, and specify a range of value for the field. Making these changes and re-running the NCL script might produce a plot like the following:



Now that you have a feel for how to edit the latlon_cells.ncl script, feel free to try plotting other fields in any remaining time in this session!

6. Adding new passive tracers

In this session, we'll get some experience in making source-code changes to add a passive tracer to the model. Although the exercises in this session are independent of those in other session, it may be best to finish the exercises from the previous sessions, first, since we will be changing the model source code and re-compiling the model executable.

6.1 Creating initial conditions for the tracer

We'll begin by defining our new passive tracer in the Registry.xml file for the init_atmosphere core. Let's change directories to the copy of the MPAS-Model source code in our \${HOME} directory:

```
$ cd ${HOME}/MPAS-Model
```

Opening up the src/core_init_atmosphere/Registry.xml file with our editor of choice, we will look for the var array section named "scalars"; it should look something like this:

As we saw in the lectures, we will need to add our new passive tracers here. Let's add a new passive tracer named "mytracer1". The array_group that we assign to this tracer can be anything other than "moist". By convention, we'll call the array group for this new tracer "passive". After adding our new tracer, the "scalars" var array should look like this:

After making this change to the Registry.xml file, we'll clean and re-compile the init_atmosphere core.

Recall that, after we make changes to the Registry.xml file, we need to "clean" the build first so that the changes take effect. Later, when we modify only Fortran source code, there will be no need to "clean" before re-compiling. The following two commands should suffice:

```
$ make clean CORE=init_atmosphere
$ make -j4 gfortran CORE=init_atmosphere PRECISION=single
```

If compilation was successful, we should now have a new init_atmosphere_model executable. At this point, if we were to create new initial conditions, the "mytracer1" field would be present in the initial conditions, but it would be zero everywhere. We next need to add code to define the initial conditions for "mytracer1".

To define the initial value of our new tracer, use your preferred editor to open the src/core_init_atmosphere/mpas_init_atm_cases.F file. Search for these two lines of code that call the routine to initialize real-data test cases:

An appropriate place to add a call to our tracer initialization routine would be directly after the call to init_atm_case_gfs. Let's add a new line of code below this to call a new subroutine, which we'll name init_tracers.

Notice that this new routine (which we have yet to define!) takes as input two arguments that are MPAS "pools".

As a start, we will define the init_tracers function so that it simply writes a message to the log file. At the bottom of the src/core_init_atmosphere/mpas_init_atm_cases.F, but before the line "end module init_atm_cases", let's add the following Fortran code:

```
subroutine init_tracers(mesh, state)
  implicit none
  type (mpas_pool_type), intent(in) :: mesh
  type (mpas_pool_type), intent(inout) :: state
  call mpas_log_write('====== Handling tracer initialization =======')
end subroutine init_tracers
```

After saving our code changes, we can re-compile the init_atmosphere core with:

```
$ make -j4 gfortran CORE=init_atmosphere PRECISION=single
```

Again, because we have not changed the Registry.xml file since we last re-compiled, there's no need to "make clean".

Compilation should take just a minute, and if it completes successfully, we're ready to try out our modified code, even though it won't actually create initial conditions for our passive tracer, yet. So that we don't overwrite files from any of our previous sessions, we'll work in a new directory, passive_tracer, inside our home directory:

```
$ cd ${HOME}
$ mkdir passive_tracer
$ cd passive_tracer
```

We'll test out our passive tracer on the 240-km uniform mesh, for which we have already created a "static" file. Now that we have experience in creating initial conditions, the following procedure should seem familiar:

```
$ ln -s ${HOME}/240km_uniform/x1.10242.static.nc .
$ ln -s /classroom/wrfhelp/mpas/met_data/GFS:2014-09-10_00 .
$ ln -s ${HOME}/MPAS-Model/init_atmosphere_model .
$ cp ${HOME}/MPAS-Model/namelist.init_atmosphere .
$ cp ${HOME}/MPAS-Model/streams.init_atmosphere .
```

As in Section 2.1, we'll edit the namelist.init_atmosphere so that it looks like the following:

```
&nhyd_model
    config_init_case = 7
    config_start_time = '2014-09-10_00:00:00'
/
&dimensions
    config_nvertlevels = 41
    config_nsoillevels = 4
    config_nfglevels = 38
    config_nfgsoillevels = 4
/
&data_sources
    config_met_prefix = 'GFS'
    config_use spechumd = false
```

```
/
&vertical_grid
    config_ztop = 30000.0
    config_nsmterrain = 1
    config_smooth_surfaces = true
    config_dzmin = 0.3
    config_nsm = 30
    config_tc_vertical_grid = true
    config_blend_bdy_terrain = false
/
&preproc_stages
    config_static_interp = false
    config_native_gwd_static = false
    config_vertical_grid = true
    config_met_interp = true
    config_input_sst = false
    config_frac_seaice = true
/
```

We'll then edit the filename that will be read by the "input" stream in the streams.init_atmosphere file so that it matches the name of our "static" file:

Also as in Section 2.1, we'll change the name of the initial conditions file to be created:

We should now be able to run the init_atmosphere_model executable to produce an x1.10242.init.no file:

\$./init_atmosphere_model

If there were no errors in the log.init_atmosphere.0000.out file, we can check for the presence of our passive tracer in the x1.10242.init.nc file:

```
$ ncdump -h x1.10242.init.nc | grep mytracer1
```

Besides the creation of the x1.10242.init.nc file, we should also look for this message in the $log.init_atmosphere.0000.out$ file:

```
===== Handling tracer initialization ======
```

If all was successful, we can go back to the MPAS-Model directory and add code to provide initial conditions for our passive tracer. Let's change back to the MPAS source code directory:

\$ cd \${HOME}/MPAS-Model

Using our favorite editor, let's open the src/core_init_atmosphere/mpas_init_atm_cases.F file and change the code for our init_tracers subroutine so that it looks like the following:

subroutine init_tracers(mesh, state)

```
implicit none
type (mpas_pool_type), intent(in) :: mesh
type (mpas_pool_type), intent(inout) :: state
integer, pointer :: index_mytracer1
integer, pointer :: nCells
integer, pointer :: nVertLevels
integer :: k, iCell
real (kind=RKIND), dimension(:), pointer :: latCell
real (kind=RKIND), dimension(:), pointer :: lonCell
real (kind=RKIND), dimension(:,:,:), pointer :: scalars
call mpas log write('===== Handling tracer initialization ======')
call mpas pool get dimension(mesh, 'nCells', nCells)
call mpas_pool_get_dimension(mesh, 'nVertLevels', nVertLevels)
call mpas pool get dimension(state, 'index mytracer1', index mytracer1)
call mpas_pool_get_array(state, 'scalars', scalars)
call mpas_pool_get_array(mesh, 'latCell', latCell)
call mpas_pool_get_array(mesh, 'lonCell', lonCell)
do iCell = 1, nCells
```

After saving our changes, we can re-compile the init_atmosphere core once again:

```
$ make -j4 gfortran CORE=init_atmosphere PRECISION=single
```

We can now try creating new initial conditions with our updated init_tracers routine. Let's change back to the passive_tracer directory, remove the old x1.10242.init.nc file, and create a new initial conditions file:

```
$ cd ${HOME}/passive_tracer
$ rm x1.10242.init.nc
$ ./init_atmosphere_model
```

If the creation of the new x1.10242.init.no file was successful, you may like to try using the convert_mpas utility to see what our new "mytracer1" field looks like!

6.2 Adding the tracer to the model

Now that we have a new passive tracer in our initial conditions, we'll need to edit the Registry.xml file for the atmosphere core so that the model itself will know about the new tracer.

Let's begin by changing back to the model source code directory:

```
$ cd ${HOME}/MPAS-Model
```

We can edit the src/core_atmosphere/Registry.xml file, recalling from the lectures that we not only need to define the tracer, but we also need to define an array that will hold the tendencies for the tracer. Even if we don't supply any sources or sinks, the model needs memory to be allocated for the tracer tendency, e.g., for use in the scalar transport routines.

In the src/core_atmosphere/Registry.xml file, search for the word "var_array" several times until you come to the definition of the scalars var array tag, which should begin with lines that look like this:

Let's add a new XML var tag inside the var_array tag to define our new "mytracer1" tracer; this new tag may look something like the following:

```
<var name="mytracer1" array_group="passive" units="kg kg^{-1}"
    description="Our first passive tracer"/>
```

Remember: it's important that the passive tracer belongs to an array_group other than "moist"; here, we've just used the string "passive", though in principle you could choose another name.

Similarly, we need to define a tendency array for our new passive tracer. Searching for the word "scalars_tend", we should find a section of XML for the scalars_tend var_array that begins with the following lines:

We'll define a new var entry in this var_array for our passive tracer tendency; this new entry might look something like the following:

```
<var name="tend_mytracer1" name_in_code="mytracer1" array_group="passive" units="kg m^{-3} s^{-1}"
    description="Tendency of our first tracer mass per unit volume divided by d(zeta)/dz"/>
```

Don't forget to set the value for array_group to "passive"!

With these changes to the Registry.xml file, we can re-compile the atmosphere core. But, we'll first need to "clean" the code, since the last core that we compiled was the init atmosphere core.

```
$ make clean CORE=atmosphere
$ make -j4 gfortran CORE=atmosphere PRECISION=single
```

If compilation was successful, we can change directories to our passive_tracer directory and prepare to run MPAS:

```
$ cd ${HOME}/passive_tracer
```

You may recall from Section 2.3 that we need the MPAS executable, atmosphere_mode1, as well as a number other files in order to run a simulation. Let's link or copy these files as we did in Section 2.3:

```
$ ln -s ${HOME}/MPAS-Model/atmosphere_model .
$ cp ${HOME}/MPAS-Model/namelist.atmosphere .
$ cp ${HOME}/MPAS-Model/streams.atmosphere .
$ cp ${HOME}/MPAS-Model/stream_list.atmosphere.* .
$ ln -s ${HOME}/MPAS-Model/src/core_atmosphere/physics/physics_wrf/files/* .
```

Now, you can proceed to set up the namelist.atmosphere and streams.atmosphere files as you did in Section 2.3 before starting the model simulation using 4 MPI tasks.

The entire set of scalars is written to the default "history" files, so as soon as the first several output files are available, you can try making plots of our new "mytracer1" passive tracer. Then, we can stop the model simulation and proceed to the next sub-section, where we will add sources and sinks for our new tracer before running a longer simulation.

6.3 Adding sources and sinks

Having added a passive tracer to our initial conditions, and defined the tracer in the model so that it will be transported by the model, it's not much additional work to define sources and sinks for the tracer. Here, we will define time-invariant sources and sinks that are set up when the model begins running, but we could in principle define time-varying sources and sinks that are periodically updated from an input file, for example.

To add time-invariant sources and sinks, we can change to our MPAS-Model source directory and edit the src/core_atmosphere/Registry.xml file with the editor of our choice. Search for the section of the file that begins the definition of the "tend" var_struct; around line 1712 we should see the following:

Just below the tendencies for prognostic variables comment line, let's add a definition for our passive tracer's sources/sinks:

```
<var name="mytracer1_sources" type="real"
   dimensions="nVertLevels nCells Time"
   units="kg m^{-2} s^{-1}"
   description="Our first passive tracer areal production rate" />
```

As in the lecture on adding passive tracers, our example tracer will be produced and destroyed at the surface of the model.

Next, we can add code to set the "mytracer1_sources" field when the model starts up. In the src/core_atmosphere/mpas_atm_core.F in the atm_mpas_init_block subroutine, we will add declarations of two new array pointers at the end of the variable declarations as in the highlighted code, below:

```
integer, pointer :: nThreads
integer, dimension(:), pointer :: cellThreadStart, cellThreadEnd
integer, dimension(:), pointer :: cellSolveThreadStart, cellSolveThreadEnd
integer, dimension(:), pointer :: celgeThreadStart, edgeThreadEnd
integer, dimension(:), pointer :: edgeSolveThreadStart, edgeSolveThreadEnd
integer, dimension(:), pointer :: vertexThreadStart, vertexThreadEnd
integer, dimension(:), pointer :: vertexSolveThreadStart, vertexSolveThreadEnd
logical, pointer :: config_do_restart, config_do_DAcycling
real (kind=RKIND), dimension(:,:), pointer :: mytracerl_sources
integer, dimension(:), pointer :: ivgtyp
call atm_compute_signs(mesh)
```

Then, near the end of the atm_mpas_init_block subroutine, we can add code to set the production/destruction of our tracer based on whether a model cell is land or water:

```
call mpas_pool_get_array(tend, 'mytracer1_sources', mytracer1_sources)
call mpas_pool_get_array(sfc_input, 'ivgtyp', ivgtyp)

mytracer1_sources(:,:) = 0.0_RKIND
do iCell=1,nCells
    ! Assume that vegetation category 17 is water
    ! (true for MODIS land use, not true for USGS)
    if (ivgtyp(iCell) == 17) then
        ! 2 g/m^2/day destruction over water
        mytracer1_sources(1,iCell) = -0.002 / 86400.0
    else
        ! 4 g/m^2/day production over land
        mytracer1_sources(1,iCell) = 0.004 / 86400.0
    end if
end do
end subroutine atm_mpas_init_block
```

Finally, we can add our source terms in to the tendency array for our new tracer. In

src/core_atmosphere/physics/mpas_atmphys_todynamics.F at the bottom of the physics_get_tend
subroutine, we can add the following variable declarations:

Then, near the bottom of the subroutine before several arrays are deallocated, we can add code to include our source terms in the tendency for our passive tracer:

Since we have modified the Registry.xml file, we will need to clean the code before re-compiling the model:

```
$ make clean CORE=atmosphere
$ make -j4 gfortran CORE=atmosphere PRECISION=single
```

After the model has been successfully re-compiled, we can change directories to our run directory, remove any existing model output files, and re-run the model with 4 MPI tasks:

```
$ cd ${HOME}/passive_tracer
$ rm -f diag*nc history*nc
$ mpiexec -n 4 ./atmosphere model
```

Once the model has started to run and has produced several history files — which should now contain our new "mytracer1" field — we can use, e.g., convert_mpas to interpolate the model history files to a lat-lon grid for quick visualization of our new tracer with sources and sinks in the lowest model layer!

7. Computing new diagnostic field

In this final session, we will add a new diagnostic field to the model: an approximation to the radiative energy balance computed as the integrated difference between the outgoing and incoming radiation fluxes (both long-wave and short-wave) at the top of the model. The integration will occur between output times of the new diagnostic field.

Disclaimer: this exercise was developed by a software engineer; if the diagnostic described here makes no physical sense or otherwise contains errors, please don't let this negatively affect your views of MPAS-Atmosphere as a scientifically credible model!

7.1 Defining the diagnostic field and adding place-holder code

The MPAS-Atmosphere model contains a framework for organizing new diagnostic computations into their own modules, which contain routines to perform different stages of the computation of the diagnostic (e.g., accumulation, averaging, and resetting). To separate the complications of computing the diagnostic from those of implementing the diagnostic in the MPAS-Atmosphere code, we can begin by simply defining a variable for our new energy balance diagnostic along with some placeholder code with print statements. We can then verify that the model compiles correctly and prints the expected lines to the log.atmosphere.0000.out file, demonstrating that at a technical level we have successfully connected our new diagnostic module with the MPAS-Atmosphere diagnostic framework. In the next section, then, we will be free to focus on the correct computation of the diagnostic itself.

Let's begin by changing to the directory containing our MPAS-Model source code:

```
$ cd ${HOME}/MPAS-Model
```

All of the changes for our energy balance diagnostic will be confined to the src/core_atmosphere/diagnostics subdirectory; we can change to that sub-directory now:

```
$ cd src/core_atmosphere/diagnostics
```

There are just three files in this directory that serve as the interface between the rest of the MPAS-

Atmosphere model and our new diagnostic:

- Makefile changes here will ensure that our new Fortran code will be used when we recompile MPAS
- Registry_diagnostics.xml changes here will allow any new variables defined by our new
 diagnostic to be available to the rest of the MPAS code (in particular, to the stream manager, which
 allows our diagnostic to be written to output files)
- mpas_atm_diagnostics_manager.F changes here will ensure that the different stages of computation required by our new diagnostic are called at the correct points in each model timestep

Before modifying the three files listed above, we can first define a new variable in an XML file for our diagnostic. With the editor of your choice, create a new file named Registry_energy_balance.xml with the following contents (click the light blue header):

Registry_energy_balance.xml

We can also create a new Fortran module with place-holder code, in which each subroutine simply prints a message to the log file. With the editor of your choice, create a new file named energy_balance_diagnostics.F with the following contents (click the light blue header):

energy_balance_diagnostics.F

Having defined a variable to store our diagnostic as well as placeholder code in a new Fortran module, we can edit the Makefile, Registry_diagnostics.xml, and mpas_atm_diagnostics_manager.F files to connect our diagnostic. The required changes to each of these files are highlighted in the code, below.

Makefile

Registry_diagnostics.xml

mpas_atm_diagnostics_manager.F

Now, we can try re-compiling MPAS-Atmosphere. When making changes to diagnostic Registry files, it's generally necessary to clean the code before recompiling. Let's change to the top-level MPAS-Model directory, and clean and re-compile with the following commands:

```
$ cd ../../..
$ make clean CORE=atmosphere
$ make -j4 gfortran CORE=atmosphere PRECISION=single
```

After the model has been successfully compiled, we can create a new directory to run a short test simulation, through which we can verify that our diagnostic is being invoked by the model by looking for our print statements in the log file. Let's first make a new sub-directory in \${HOME} for our new diagnostic simulations:

```
$ mkdir -p ${HOME}/energy_balance
$ cd ${HOME}/energy_balance
```

Rather than create new initial conditions, we can simply re-run the 240-km quasi-uniform simulation, this time with our new diagnostic. We can therefore symbolically link the initial conditions, surface update, and mesh partition file from that simulation into our new run directory, and copy the namelist.atmosphere, streams.atmosphere, and stream_list.atmosphere.* files, too:

```
$ ln -s ${HOME}/240km_uniform/x1.10242.init.nc .
$ ln -s ${HOME}/240km_uniform/x1.10242.sfc_update.nc .
$ ln -s ${HOME}/240km_uniform/x1.10242.graph.info.part.4 .
$ cp ${HOME}/240km_uniform/namelist.atmosphere .
$ cp ${HOME}/240km_uniform/streams.atmosphere .
$ cp ${HOME}/240km_uniform/stream_list.atmosphere .* .
```

Additionally, we will need to symbolically link to our newly recompiled atmosphere_model executable as well as the physics look-up tables with the following commands:

```
$ ln -s ${HOME}/MPAS-Model/atmosphere_model
$ ln -s ${HOME}/MPAS-Model/src/core_atmosphere/physics/physics_wrf/files/* .
```

Lastly, we will edit the namelist.atmosphere file to run a non-restart (i.e., a "cold-start") simulation from 2019-09-10 0000 UTC for just one hour. Our only goal with this initial simulation is to verify that our place-holder diagnostic code is being called!

```
config_start_time = '2014-09-10_00:00:00'
config_run_duration = '1:00:00'
config_do_restart = false
```

After verifying that the variables above are set correctly in the namelist.atmosphere file, we can run our one-hour test simulation with:

```
$ mpiexec -n 4 ./atmosphere_model
```

This simulation should take just a minute or so to run. After the model completes, you can check the log.atmosphere.0000.out file for the print statements from our dignostic routines. Do you see lines like the following?

```
min/max of meshScalingDel2 = 1.00000 1.00000
min/max of meshScalingDel4 = 1.00000 1.00000
Sounding location file 'sounding_locations.txt' not found.
TESTING: energy_balance_setup
TESTING: energy_balance_reset
TESTING: energy_balance_update
TESTING: energy_balance_compute
Timing for diagnostic computation: 6.03111 s
Timing for stream output: 1.10848 s
TESTING: energy_balance_reset
```

If our diagnostic prints appear in the log file, we can proceed to the next section, were we will focus on the computation of the diagnostic without having to worry about the software details that were dealt with in this section.

7.2 Adding code to compute the diagnostic

In order to compute the accumulated net radiative energy between output times for our 'accradflux' field, we will need to implement three functions in our new energy_balance_diagnostics.F module. Let's begin by changing directories to our MPAS-Model source directory, and then to the diagnostics subdirectory:

```
$ cd ${HOME}/MPAS-Model
$ cd src/core_atmosphere/diagnostics
```

Conceptually, we will need to implement changes to the following routines:

- energy_balance_setup here we need to set the initial accradflux field to zero
- energy_balance_update here we need to add in the net energy flux for the current model time
 step
- energy_balance_reset here we need to reset the accradflux field after it has been written to an output stream

The expandable headers, below, highlight the changes that we need to make in each of these three subroutines.

subroutine energy_balance_setup

subroutine energy_balance_update

```
subroutine energy_balance_reset
```

After editing the file, we can change back to the top-level MPAS-Model directory and recompile the model. Because we did not change any Registry XML files, there is no need to "make clean", first.

```
$ cd ../../..
$ make -j4 gfortran CORE=atmosphere PRECISION=single
```

If the model compilation was successful, we are ready to try a model simulation with our new accradflux diagnostic in the next section!

7.3 Testing the diagnostic with a short simulation

To test our our diagnostic field, we can use a short, 12-hour simulation in which we write our "accradflux" field every six hours to its own output stream. All code changes have been completed, so we can change directories to \${HOME}/energy_balance:

```
$ cd ${HOME}/energy_balance
```

Let's begin by adding a new stream like the following to our streams.atmosphere file:

Rembemer: the new stream must be defined between the opening <streams> and closing </streams> tags!

Next, we can set the duration of our model simulation to just 12 hours by setting config_run_duration to '12:00:00' in our namelist.atmosphere file:

```
config_run_duration = '12:00:00'
```

Finally, we can remove any existing output files from our earlier test run with

```
$ rm diag*nc history*nc surface_winds*nc
```

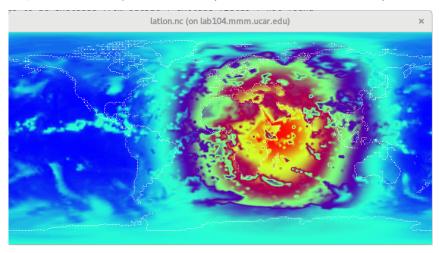
before running our short simulation with 4 MPI tasks:

```
$ mpiexec -n 4 ./atmosphere_model
```

This simulation should take just a couple of minutes to complete, and when it has finished, we should have "accradflux" files in our run directory:

```
$ ls -1 accradflux*
-rw-r--r- 1 class104 cbet 46088 Sep 10 17:56 accradflux.2014-09-10_00.00.00.nc
-rw-r--r- 1 class104 cbet 46088 Sep 10 17:57 accradflux.2014-09-10_06.00.00.nc
-rw-r--r- 1 class104 cbet 46088 Sep 10 17:58 accradflux.2014-09-10_12.00.00.nc
```

If you interpolate the "accradflux" files to a 0.5-degree latitude-longitude grid with the convert_mpas tool, does the last time period for the field (valid 12 hours after our initial time) look like the following?



Now that we have the top-of-the-atmosphere accumulated net radiative energy flux integrated over two six-hour periods from our 12-hour simulation, we can compute a radiation budget (in W / m^2) for these time periods with a Python script. Using an editor of your choice, create a file named $compute_budget.py$ with the following contents:

```
import numpy as np
from netCDF4 import Dataset

f = Dataset('accradflux.2014-09-10_12.00.00.nc')
accradflux = f.variables['accradflux'][:]

g = Dataset('x1.10242.init.nc')
areaCell = g.variables['areaCell'][:]

print(np.sum(accradflux * areaCell) / np.sum(areaCell) / 21600.0)
```

We can run this script with:

```
$ python compute_budget.py
```

You can try changing the time perdiod for which this radiation budget is computed by changing 'accradflux.2014-09-10_12.00.00.nc' to 'accradflux.2014-09-10_06.00.00.nc' in the script.

 $Positive \ values \ indicate \ more \ incoming \ radiation \ at \ the \ top \ of \ the \ atmosphere \ than \ outgoing \ radiation.$

With the remaining time in this session, you may like to try one of the following:

- Run a multi-day simulation, writing the "toa_energy_balance" stream only once per day; then, compute the radiation budget for 24-hour periods using a modified version of the Python script described above
- Run a two-day simulation, writing the "toa_energy_balance" stream every three hours; then, compute
 the radiation budget for these three-hour periods and note how the resulting values change during the
 simulation.

As an example of processing multiple netCDF files in a loop in Python, you can use the following script:

```
import numpy as np
from netCDF4 import Dataset

g = Dataset('x1.10242.init.nc')
areaCell = g.variables['areaCell'][:]
g.close()

for filename in ['accradflux.2014-09-10_00.00.00.nc',
```