

Willoughby Seago

**Theoretical Physics**

# **Relativity, Nuclear and Particle Physics —Particle Physics**

September 20, 2021

**COURSE NOTES**

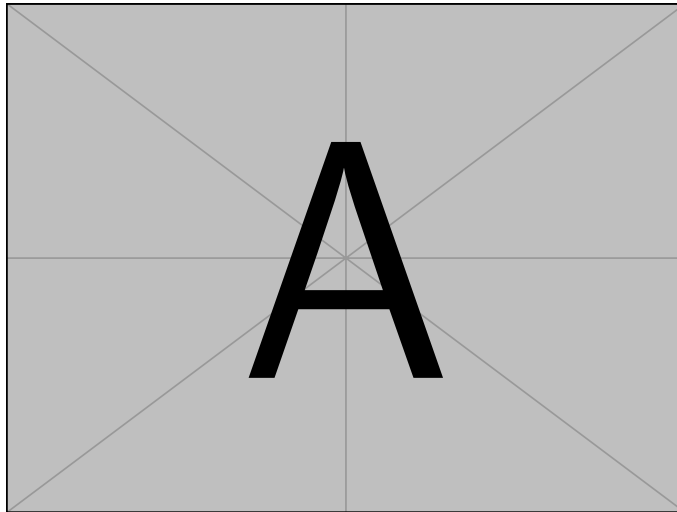
# Relativity, Nuclear and Particle Physics (Particle Physics)

Willoughby Seago

September 20, 2021

These are my notes from the course relativity, nuclear and particle physics (particle physics). I took this course as a part of the theoretical physics degree at the University of Edinburgh.

These notes were last updated at 12:40 on December 20, 2021. For notes on other topics see <https://github.com/WilloughbySeago/Uni-Notes>.



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# One

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## Introduction

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In this course we introduce particle physics. The realm in which the physics of these particles lie is quantum field theory, although this is not the focus of this course. We will develop key ideas and along the way we will meet many particles. We will discuss particle physics experiments and how we measure properties of particles. The focus will be on making sense of experimental observations without getting too bogged down in mathematical formalism.

Part I

The Basics

# Two

## The Standard Model

### 2.1 What is Particle Physics?

Particle physics is about studying the fundamental constituents of matter, and some composite particles too. Such particles are small, and hence necessarily quantum mechanical, and travel fast, and hence necessarily relativistic. The realm in which quantum mechanics and relativity come together is quantum field theory (QFT). While we won't fully develop QFT in this course we will use ideas from QFT as we see fit.

The realm of particle physics is typically restricted to protons, neutrons, other similar composite particles, and fundamental particles. Once we get to a few nucleons we enter the range of nuclear physics, and if we start including electrons in this then we enter the range of atomic physics.

As well as being interested in the building blocks of matter particle physics can also tell us about the initial state of the universe, after the big bang there was so much energy around that only particles could form for a very long time. Neutral atoms didn't form for millions of years after the big bang, so the early universe is the realm of nuclear and particle physics.

### 2.2 Natural Units

The SI units that we are all familiar with are based on a set of base units, in particular length, in metres (m), mass, in kilograms (kg), and time, in seconds (s). However, there is no particular reason why length, mass, and time should be our base units, and there is certainly no reason why we should measure them in kilograms, metres, and second.

One simple alternative is the centimetre-gram-second system (CGS), where we still take length, mass, and time as the base units but now we measure length in centimetres (cm), mass in grams (g), and time in seconds (s). This system has some

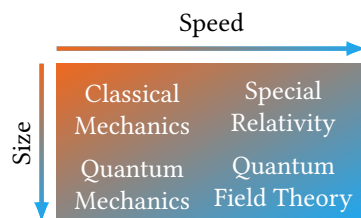


Figure 2.1: Realms of Physics.

use, particularly in electromagnetism, but it's just a rescaling of the SI base units by a couple of factors of 1000, and isn't particularly suitable for particle physics.

In particle physics we like to work with a system called **natural units**. These typically give reasonable sized numbers in particle physics calculations. In this system the base units are energy, velocity, and action<sup>1</sup>. We measure energy in electronvolts (eV), typically working with megaelectronvolts (MeV), or gigaelectronvolts (GeV)<sup>2</sup>. We measure velocity in multiples of the speed of light ( $c$ ). We measure action in multiples of the reduced Planck's constant ( $\hbar$ ).

From these three base units we can define all other quantities of interest. In particular we can measure momentum in electronvolts per speed of light (eV/ $c$ ) and mass in electronvolts per speed of light squared (eV/ $c^2$ ). We can check that this works out dimensionally by comparing to the SI base units:

$$[\text{energy}][\text{velocity}]^{-1} = \frac{[\text{mass}][\text{length}]^2[\text{time}]^{-2}}{[\text{length}][\text{time}]^{-1}} \quad (2.2.1)$$

$$= [\text{mass}][\text{length}][\text{time}]^{-1} = [\text{momentum}] \quad (2.2.2)$$

$$[\text{energy}][\text{velocity}]^{-2} = \frac{[\text{mass}][\text{length}]^2[\text{time}]^{-2}}{([\text{length}][\text{time}]^{-1})^2} = [\text{mass}] \quad (2.2.3)$$

Alternatively, we can check the units for consistency with  $E^2 = m^2c^4 + p^2c^2$ , from which we get the same result if we measure  $E$  in electronvolts and take  $m$  or  $p$  to be zero.

Other dimensions can also be expressed in electronvolts, speeds of light, and Planck's constants, in particular

- $[\text{time}] = [\text{action}][\text{energy}]^{-1}$ , so time can be measured in  $\hbar/\text{GeV}$ .
- $[\text{length}] = [\text{action}][\text{velocity}][\text{energy}]^{-1}$ , so length can be measured in  $\hbar c/\text{eV}$ .
- $[\text{area}] = [\text{length}]^2$ , so area can be measured in  $\hbar^2 c^2/\text{GeV}^2$ .

So far we have only defined the relation between units, not the actual size of units. Of course, electronvolts, the speed of light, and Planck's constant all have well defined sizes, *in SI units*, but we've come this far in abandoning SI units, it would be silly to cling to the values of a few constants now. Instead we usually choose the size of our units such that  $c = \hbar = 1$ . In this case we don't write the factors of  $c$  and  $\hbar$  so we measure energy, mass, and momentum in eV, time and length can be measured in  $\text{eV}^{-1}$ , and area in  $\text{eV}^{-2}$ .

One advantage of this system of units is that time and length are measured in the same units, which is good from a relativistic point of view as we should treat time and space in the same way. It also simplifies a lot of formulas in relativity and quantum mechanics by removing constants. For example

$$E^2 = m^2c^4 + p^2c^2 \rightarrow E^2 = m^2 + p^2, \quad (2.2.4)$$

$$\left( \frac{\hbar^2}{2m} \nabla^2 - V(\mathbf{r}) \right) \psi(\mathbf{r}, t) = i\hbar \frac{\partial \psi}{\partial t} \rightarrow \left( \frac{1}{2m} \nabla^2 - V(\mathbf{r}) \right) \psi(\mathbf{r}, t) = i \frac{\partial \psi}{\partial t}, \quad (2.2.5)$$

$$t = \frac{\tau}{\sqrt{1+v^2/c^2}} \rightarrow t = \frac{\tau}{\sqrt{1-v^2}}. \quad (2.2.6)$$

To convert back to units where  $c, \hbar \neq 1$  we have to do some dimensional analysis, but generally the effort saved by setting  $c = \hbar = 1$  makes this extra work at the end worth while, if we always work in these units then the conversion work is minimised.

<sup>1</sup> $[\text{velocity}] = [\text{length}][\text{time}]^{-1}$ ,  
 $[\text{energy}] = [\text{mass}][\text{length}]^2[\text{time}]^{-2}$ ,  
 $[\text{action}] = [\text{mass}][\text{length}]^2[\text{time}]^{-1}$ .

<sup>2</sup> $1 \text{ GeV} = 10^3 \text{ MeV} = 10^9 \text{ eV}$

■ **Example 2.2.7** A W boson has energy width  $\Gamma = 2 \text{ GeV}$ . What is the lifetime of the particle in SI units? The **lifetime** of a particle is defined as  $\tau := 1/\Gamma = 0.5 \text{ GeV}^{-1}$

Currently we have units of  $[\text{energy}]^{-1} = [E]^{-1}$ , and we need units of  $[\text{time}] = [T]$ . We get that by multiplying by some integer multiples of  $\hbar$  and  $c$ , which have units of  $[\hbar] = [E][T][L]$ , where  $[L] = [\text{length}]$ , and  $[c] = [L][T]^{-1}$ . Hence we need to find  $n, m \in \mathbb{Z}$  such that

$$[E]^{-1} [\hbar]^n [c]^m = [T]. \quad (2.2.8)$$

We will then have  $\tau = \hbar^n c^m / \Gamma$ . We see that

$$[E]^{-1} [\hbar]^n [c]^m = [E]^{-1} [E]^n [T]^n [L]^m [T]^{-m}. \quad (2.2.9)$$

From this we see that  $n = 1$  and  $m = 0$ . Hence  $\tau = \hbar/\Gamma$ , and so for the W boson  $\tau = 0.5 \hbar/\text{GeV}$ .

Natural units as defined here don't uniquely fix all dimensions. For example there are seven SI base units needed to fix all dimensions. We would need a similar number of base units in natural units also. Most of the time this isn't a problem as we aren't that interested in, say, the luminous intensity of a light source, so we don't need an equivalent to the candela (cd). Similarly we deal with a few particles at a time so don't need moles (mol).

One quantity that we will need, which can't be expressed in natural units alone, is charge. In SI units the charge is fixed by defining the ampere as the rate of flow of charge. In particle physics we simply define charge as multiples of the electron charge,  $e$ . Most fundamental particles have an integer multiple of  $e$  as their charge, with the exception of quarks, which have third integer multiples of  $e$ .

## 2.3 Basic Particles

The **leptons** are a family of particles made up of **electrons** ( $e^-$ ) and **electron neutrinos** ( $\nu_e$ ), as well as the heavier **muon** ( $\mu^-$ ) and **muon neutrino** ( $\nu_\mu$ ), and the heavier still **tau** ( $\tau^-$ ) and **tau neutrino** ( $\nu_\tau$ ). These are all **fermions**, i.e. particles with half-integer spin, in particular leptons have spin<sup>3</sup>  $1/2$ . Electrons, muons, and tau have charge  $-e$ . Neutrinos have no charge. The exact masses of the neutrinos aren't known, but it is known that they are very light, but have non-zero mass. We define three quantum numbers, the lepton family numbers, which correspond to electrons, muons, and tau. A lepton and the related neutrino have a lepton family number of  $+1$ , and all other particles have  $0$ . The properties of leptons are given in [Table 2.1](#).

The **quarks** are a family of particles made up of up (u), down (d), charm (c), strange (s), top (t), and bottom (b) quarks. Of these three, u, c, t, have charge  $2e/3$ , and the other three, d, s, b, have charge  $-e/3$ . These are also fermions with spin  $1/2$ . The properties of quarks are given in [Table 2.2](#).

As well as the leptons and quarks there are the antileptons and antiquarks. These have the same mass but their charges and additive quantum numbers are negated. For example, antielectrons, also known as **positrons** ( $e^+$ ) have electron lepton family member  $L_e = -1$ , antiquarks have baryon number  $-1/3$ , and anti strange quarks have strangeness quantum number  $+1$ . For particles with an explicit charge, such as  $\tau^-$ ,

<sup>3</sup> $\hbar/2$  in non-natural units

Table 2.1: Some key properties of leptons, including the generation, charge, mass, and lepton flavour number, specifically the electron, muon, and tau lepton family numbers,  $L_e$ ,  $L_\mu$ , and  $L_\tau$  respectively.

Lepton	Symbol	Generation	Charge/ $e$	Mass/MeV	Lepton Family Number		
					$L_e$	$L_\mu$	$L_\tau$
Electron	$e^-$	1	-1	0.511	1	0	0
Electron Neutrino	$\nu_e$	1	0	$< 2.1 \times 10^{-6}$	1	0	0
Muon	$\mu^-$	2	-1	105.66	0	1	0
Muon Neutrino	$\nu_\mu$	2	0	$< 0.19$	0	1	0
Tau	$\tau^-$	3	-1	1777	0	0	1
Tau Neutrino	$\nu_\tau$	3	0	$< 18$	0	0	1

Table 2.2: Some key properties of quarks, including the generation, charge, mass, isospin, quark flavour number, and baryon number.

Quark	Symbol	Generation	Charge/ $e$	Mass/MeV	Isospin ( $I, I_z$ )	Quark Flavour Quantum Number	Baryon Number, $\mathcal{B}$
Up	u	1	2/3	2.2	(1/2, +1/2)		1/3
Down	d	1	-1/3	4.7	(1/2, -1/2)		1/3
Charm	c	2	2/3	1270		$C = +1$	1/3
Strange	s	2	-1/3	93		$S = -1$	1/3
Top	t	3	2/3	$172.9 \times 10^3$		$T = +1$	1/3
Bottom	b	3	-1/3	4180		$B = -1$	1/3

we denote the antiparticle by negating the charge, so  $\tau^+$ . For particles without an explicit charge, for example, u, we denote the antiparticle with a bar, so  $\bar{u}$ .

### 2.3.1 A Brief History of Particle Discoveries

1887 The electron was discovered by J. J. Thompson.

1920 The proton was discovered by E. Rutherford.

1931 The positron was discovered by C. Anderson. It was originally thought to be a proton.

<sup>4</sup>leave some for everyone else Anderson. 1936 The muon was discovered by C. Anderson<sup>4</sup>.

1956 The electron neutrino was discovered by C. Cowan and F. Reines.

1962 The muon neutrino was discovered by L. Lederman, M. Schwartz, and J. Steinberger.

- 1968 It was discovered that the proton is made of quarks, in particular the proton is uud, this was discovered by J. Friedman *et al.* at the Stanford linear accelerator centre (SLAC).
- 1973 The  $J/\psi$  resonance was discovered, which is the most common resonance of charmonium, a charm-anticharm quark meson ( $c\bar{c}$ ). This was evidence for the charm quark, and was discovered by B. Richter *et al.* SLAC as well as the Brookhaven national laboratory (BNL).
- 1975 The tau was discovered by M. Perl *et al.* at SLAC.
- 1977 The  $\Upsilon$  resonance was discovered, which is a resonance of bottomonium, a bottom-antibottom quark meson ( $b\bar{b}$ ). This was evidence for the bottom quark, and was discovered by L. Lederman<sup>5</sup> *et al.* at Fermilab.
- 1995 The top quark was discovered by the collider detector at Fermilab (CDF) and the DØ experiment (DZero experiment) also at Fermilab.
- 2000 The tau neutrino was discovered by the direct observation of the nu tau experiment, or the DONUT<sup>6</sup> experiment.
- 2021 The Higgs was discovered in 2012 by a toroidal LHC apparatus (ATLAS)<sup>7</sup> and the compact muon solenoid (CMS) experiments at *conseil européen pour la recherche nucléaire* (CERN), with the large hadron collider (LHC).

<sup>5</sup>another greedy particle physicist, discovering all the particles for himself.

<sup>6</sup>they really *really* wanted that acronym to work.

<sup>7</sup>surely the acronym with the word “toroidal” should be the one to spell DONUT

### 2.3.2 Fundamental Forces

There are four fundamental forces:

- The strong force
- Electromagnetism
- The weak force
- Gravity

Of these only the first three are important at the subatomic level. For example, the strong force binds quarks into nucleons, electromagnetism governs the interaction of electrons, and the weak force controls the decay of radioactive nuclei. Gravity is simply too weak to have any noticeable effect on these scales.

The fundamental forces are propagated by **messenger particles**, all of which are **bosons**, i.e. particles with integer spin. In particular they are **vector bosons**, meaning they have spin 1. For example the  $W^-$  boson is the exchange boson in beta decay of a neutron,  $n \rightarrow p e^- \bar{\nu}_e$ . Information on the fundamental forces and their exchange bosons is given in

## 2.4 Space-Time Symmetries

**Theorem 2.4.1 — Noether's Theorem.** Every continuous symmetry of nature corresponds to a conservation law.

Noether's theorem relates continuous symmetries (in particular symmetries under which the Lagrangian is invariant) to conservation laws. The three most common examples being



Table 2.3: Information on the messenger particles for the four fundamental forces. Note that the existence of the graviton has not been experimentally verified.

Force	Boson	Mass/GeV	Relative Strength
Strong	Gluon	0	1
EM	Photon	0	$1 \times 10^{-2}$
Weak	W/Z	80.4, 91.2	$1 \times 10^{-7}$
Gravity	Graviton (???)	0	$1 \times 10^{-40}$

- Invariance under spatial translation giving momentum conservation.
- Invariance under time translation giving conservation of energy.
- Invariance under rotation giving conservation of angular momentum.

Notice that in relativity the first two become one and invariance under translations of the four-position lead to conservation of the four-momentum.

Noether's theorem deals with continuous symmetries, in the sense that we can consider arbitrarily small translations or rotations. There are also symmetries relating to discrete symmetries, and these too lead to conservation laws. There are other conservation laws that we have observed but do not know what symmetry leads to them, if indeed there is a symmetry at all.

These conservation laws are usually phrased in terms of conservation of quantum numbers. Many of these are familiar already.

The **charge** is a quantum number, which measures the strength of interactions with the electromagnetic field.

The **lepton flavour number** is a quantum number, in particular we define the electron lepton flavour number as

$$L_e := N(e^-) + N(\nu_e) - N(e^+) - N(\bar{\nu}_e). \quad (2.4.2)$$

Here  $N(x)$  is the number of  $x$  present in the system. We define the muon and tau lepton flavour numbers analogously. The **lepton number** is then simply the sum of these,

$$L := L_e + L_\mu + L_\tau. \quad (2.4.3)$$

Which is just the number of leptons minus the number of antileptons.

The **quark flavour number** is a quantum number, for up quarks we define it as

$$N_u := N(u) - N(\bar{u}). \quad (2.4.4)$$

We define analogous quantities for the other quark flavours. The **quark number** is simply the total number of quarks minus the number of antiquarks,

$$N_q := N(q) - N(\bar{q}) = N_u + N_d + N_c + N_s + N_t + N_b. \quad (2.4.5)$$

The **baryon number** is three times the quark number, that is

$$\mathcal{B} := 3N_q = 3(N_u + N_d + N_c + N_s + N_t + N_b). \quad (2.4.6)$$

Notice that by this definition quarks all have a baryon number of  $1/3$ , and antiquarks have  $-1/3$ .

- Invariance under gauge transformation<sup>8</sup> in quantum electrodynamics (QED) leads to charge conservation.
- The lepton flavour numbers,  $L_e$ ,  $L_\mu$ , and  $L_\tau$ , are conserved in all known interactions, and hence the lepton number,  $L$ , is conserved. There is no known symmetry leading to this.
- The quark flavour numbers,  $N_u$ ,  $N_d$ ,  $N_c$ ,  $N_s$ ,  $N_t$ , and  $N_b$ , and hence the quark number,  $N_q$ , and baryon number,  $\mathcal{B}$ , are conserved in quantum chromodynamics (QCD), and QED, but not in weak interactions. There is no known symmetry leading to this.
- Parity<sup>9</sup> is conserved in QCD and QED, but not in weak interactions. There is no known symmetry leading to this.

<sup>8</sup>a gauge transformation in this case acting on the electromagnetic potentials is the transformation  $\varphi \rightarrow \varphi - \partial_t \chi$  and  $\mathbf{A} \rightarrow \mathbf{A} + \nabla \chi$ , for scalar and vector potentials  $\varphi$  and  $\mathbf{A}$  and some arbitrary, twice differentiable, function of space and time,  $\chi$ . Relativistically this is the same as  $A_\mu \rightarrow A_\mu - \partial_\mu \chi$  where  $A^\mu = (\varphi, \mathbf{A})$  is the four-potential and  $\partial_\mu = (\partial_t, \nabla)$ .

<sup>9</sup>Recall that the parity of a state,  $\psi$ , is the eigenvalue of the operation  $\psi(\mathbf{r}) \rightarrow \psi(-\mathbf{r})$ , in particular if  $\psi(-\mathbf{r}) = \pm \psi(\mathbf{r})$  we say that the parity is even (+) or odd (-).

## 2.5 Open Questions

There are many open questions in particle theory. For example, the unknown symmetries mentioned in the previous section. Other questions include

- There are three generations of quarks and leptons. Why? Why not 2 or 3, why do we even “need” more than 1?
- The mass difference between generations is large. Why?
- What about the stuff not in the standard model? What is dark matter/dark energy and how does it fit with the standard model?

This last question is fairly large given that approximately 95 % of the universe is dark matter or dark energy.

Leptons	Mass Charge Spin	$\approx 2.3 \text{ MeV}$ $2/3$ $1/2$	$\approx 95 \text{ MeV}$ $2/3$ $1/2$	$\approx 173.07 \text{ GeV}$ $2/3$ $1/2$	$80.4 \text{ GeV}$ $\pm 1$ $1$	$125 \text{ GeV}$ $0$ $0$
		<b>u</b>	<b>c</b>	<b>t</b>	<b>W</b>	<b>H</b>
		Up	Charm	Top	W Boson	Higgs Boson
Quarks		$\approx 4.8 \text{ MeV}$ $-1/3$ $1/2$	$\approx 1.275 \text{ GeV}$ $-1/3$ $1/2$	$\approx 4.18 \text{ GeV}$ $-1/3$ $1/2$	$91.2 \text{ GeV}$ $0$ $1$	
		<b>d</b>	<b>s</b>	<b>b</b>	<b>Z</b>	
		Down	Down	Bottom	Z Boson	
Vector Bosons		$511 \text{ keV}$ $-1$ $1/2$	$105.7 \text{ MeV}$ $-1$ $1/2$	$1.777 \text{ GeV}$ $-1$ $1/2$	$0$ $0$ $1$	
		<b>e</b>	<b><math>\mu</math></b>	<b><math>\tau</math></b>	<b><math>\gamma</math></b>	
		Electron	upmuon	uptau	Photon	
Scalar Bosons		$< 2.2 \text{ eV}$ $0$ $1/2$	$< 0.17 \text{ MeV}$ $0$ $1/2$	$< 15.5 \text{ MeV}$ $0$ $1/2$	$0$ $0$ $1$	
		<b><math>\nu_e</math></b>	<b><math>\nu_\mu</math></b>	<b><math>\nu_\tau</math></b>	<b>g</b>	
		Electron Neutrino	upmuon Neutrino	uptau Neutrino	Gluon	

Figure 2.2: Particles of the standard model, with their masses, charges, spins, symbols, and names.

# Three

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## Tools of Particle Physics

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### 3.1 Measurable Properties

There are many properties of particles that we may wish to measure. These broadly separate into two categories, static properties, which don't change, and dynamic properties which do. Examples of static properties includes

- Mass,  $m$ .
- Charge,  $Q$ .
- Spin,  $J$ , and parity,  $P$ , together  $J^P$ .
- Colour charge (see [Section 7.2](#)).
- Weak hypercharge (see ).

Examples of dynamic properties includes

- Energy,  $E$ .
- Momentum,  $p$ .


There are typically two ways that we measure properties. We need some sort of interaction in order to extract information from the system, the common interactions used are particle decay and particle scattering. Particle decays can tell us

- Particle lifetime,  $\tau$ .
- Decay width,  $\Gamma = \hbar/\tau$ .
- Allowed and forbidden decays, which imply conservation laws.

Particle scattering can tell us

- Total cross section,  $\sigma$ .
- Differential cross section,  $d\sigma/d\Omega$ .
- Luminosity,  $\mathcal{L}$  (see [Section 3.5.1](#)).
- Event rate,  $N$ .

### 3.2 Relativity Recap

 See the notes for the relativity part of the course for more details.

Recall that we define  $\beta := v/c$ . In natural units  $\beta = v$  and we use them interchangeably. The Lorentz gamma factor is then

$$\gamma := \frac{1}{\sqrt{1 - \beta^2}}. \quad (3.2.1)$$

The relativistic three-momentum of the particle is defined as  $\mathbf{p} := \gamma m \mathbf{v}$ . The relativistic energy of a particle is defined as

$$E^2 = m^2 c^4 + \mathbf{p}^2 c^2 \quad (3.2.2)$$

where  $p = |\mathbf{p}|$ . The energy is also given by  $E = \gamma m c^2$ . In natural units these equations reduce to

$$E^2 = m^2 + p^2, \quad \text{and} \quad E = \gamma m. \quad (3.2.3)$$

We also have  $\beta = |\mathbf{p}|/E$  and so  $\beta\gamma = |\mathbf{p}|/m$ . The four-momentum of the particle is then  $p^\mu = (E/c, \mathbf{p})$ , or in natural units  $p^\mu = (E, \mathbf{p})$ .

For a massless particle  $v = c$ ,  $\beta = 1$ , and  $E = p$ , the gamma factor,  $\gamma$ , is not defined.

For a particle at rest  $v = \beta = 0$ ,  $E = m$ , and  $\gamma = 1$ .

The invariant mass of a particle is defined as the square of its four momentum,

$$p^2 = p_\mu p^\mu = E^2 - \mathbf{p}^2 = m^2 \quad (3.2.4)$$

where the final equality comes from evaluating the scalar product, which is a Lorentz invariant, in the rest frame of the particle.

### 3.3 Quantum Mechanics Recap

A free particle in three dimensional space has a wave function given by

$$\psi(\mathbf{r}, t) = \exp[i(\mathbf{p} \cdot \mathbf{r} - Et)/\hbar] \quad (3.3.1)$$

where  $\mathbf{r}$  is the position of the particle,  $\mathbf{p}$  is the momentum of the particle,  $E$  is the energy of the particle, and  $t$  is the time.

Define the momentum operator,  $\hat{\mathbf{p}} = -i\hbar\nabla$ , and energy operator,  $\hat{E} := i\hbar\partial_t$ . Using the non-relativistic kinetic energy,  $E = p^2/(2m)$ , as the energy of the particle replacing  $E$  and  $p$  with their operators we get


$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi \quad (3.3.2)$$

which is the Schrödinger equation. In natural units this reduces to

$$i \frac{\partial \psi}{\partial t} = -\frac{1}{2m} \nabla^2 \psi. \quad (3.3.3)$$

The Schrödinger equation is non-relativistic, since we used  $E = p^2/(2m)$ , and it is not Lorentz invariant.

### 3.4 Relativistic Quantum Mechanics

 See Quantum Theory notes for more detail.

If we instead use the full relativistic energy,  $E^2 = m^2 c^4 + p^2 c^2$  we get the equation

$$-\hbar^2 \frac{\partial^2 \psi}{\partial t^2} = -\hbar^2 c^2 \nabla^2 \psi + m^2 c^4 \psi. \quad (3.4.1)$$

This is the **Klein–Gordon** equation, which can be written in natural units as

$$-\frac{\partial^2 \psi}{\partial t^2} = -\nabla^2 \psi + m^2 \psi. \quad (3.4.2)$$

Notice that the relativistic energy equation gives two solutions,  $E = \pm \sqrt{m^2 + p^2}$ . We can understand the negative energy no as corresponding to an antiparticle. Back when Schrödinger was trying to come up with a wave equation for particles he did first try the relativistic momentum equation but was put off by the negative energy solutions and so returned to the non-relativistic energy and found the Schrödinger equation.

The Klein–Gordon equation applies only to spinless particles, which means the only standard model particle it applies to is the Higgs boson. There are other issues such as not being guaranteed to have positive-definite probability for the position. We will see the (non-examinable) Dirac equation later, which is more general.

The Klein–Gordon equation can more compactly be written as

$$(\square^2 + m^2)\psi = 0 \quad (3.4.3)$$

where  $\square^2$  is the d'Alembert operator, also written as  $\square$  or  $\partial^2$ , and is defined to be

$$\square^2 := \partial^\mu \partial_\mu = \frac{\partial^2}{\partial t^2} - \nabla^2 = \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} \quad (3.4.4)$$

#### 3.4.1 The Yukawa Potential

Consider static solutions to the Klein–Gordon equation, that is  $\partial_t \psi = 0$ . The equation then reduces to

$$\nabla^2 \psi = m^2 \psi. \quad (3.4.5)$$

One solution to this is

$$\psi(\mathbf{r}) = -\frac{g^2}{4\pi r} e^{-mr}. \quad (3.4.6)$$

This is called the **Yukawa potential**. Here  $g$  is a constant that characterises the strength of the potential. The range of interactions is determined by the constant  $\alpha := 1/m$ . In the case of  $m = 0$  we have  $\alpha = \infty$  and we get the Coulomb potential (or any other  $-1/r$  potential) with infinite range.

This potential can be used to determine the low energy (long range) behaviour of the strong force. See the notes for the nuclear physics part of this course for more details.

### 3.5 Particle Interactions

#### 3.5.1 Particle Scattering

Consider the interaction  $a + b \rightarrow c + d$ , in which particles,  $a$ , in a beam hit particles,  $b$ , in a target. There are two parameters that characterise the interaction. First the **flux** of the beam, which is the rate of particles crossing the target per unit area perpendicular to the beam direction. This is given by the beam density,  $\rho_a$ , and speed,  $v$ :

$$\text{flux} := \rho_a v. \quad (3.5.1)$$

The units of flux are  $[\text{time}]^{-1}[\text{length}]^{-2}$ .

The other parameter is the rate at which the interaction outputting  $c + d$  occurs. This can be expressed as the product of the incident flux, the number of particles,  $N_b$ , at the target participating in the process and the interaction strength. The rate is then

$$R = \rho_a v N_b \sigma \quad (3.5.2)$$

where  $\sigma$  is the total cross section, which we take to be a constant of proportionality relating the interaction rate to the rate of possible interactions. The units of  $\sigma$  are  $[\text{length}]^2$  and the units of  $R$  are  $[\text{time}]^{-1}$ .

The rate can be split into two factors, the **luminosity**,  $\mathcal{L} := \rho_a v N_b$ , which relates the density of the beam and target and the rate at which the particles come into contact, and the probability that upon “contact” any two particles react to produce the given products, this is characterised by the interaction strength, measured by  $\sigma$ . Cross sections are often measured in units of barns (b), where  $1 \text{ b} = 10^{-28} \text{ m}^2 = 10^{-24} \text{ cm}^2$ . Typical cross sections are about 10 mb for the strong force,  $10^{-2} \text{ mb}$  for the electromagnetic force, and  $10^{-13} \text{ mb}$  for the weak force. The name “barn” comes from the idiom “hit the broad side of a barn”, the idea being that a barn (as in the building on a farm) is large, and as evidenced by these typical cross sections the barn is indeed quite a large unit for what it measures.

In QFT a particle initially described by the state  $|\psi_i\rangle$  can interact with another particle and end up in some final state  $|\psi_f\rangle$ . The interaction is represented by an operator,  $\hat{O}$ , on this state, so  $\hat{O}|\psi_i\rangle = |\psi_f\rangle$ . The transition probability,  $T$ , is given by **Fermi’s golden rule**, which states that

$$T = \frac{2\pi}{\hbar} |\mathcal{M}|^2 \rho \quad (3.5.3)$$

where  $\rho$  is the density of final states and  $\mathcal{M}$  is the matrix element

$$\mathcal{M} = V_{fi} := \langle \psi_f | \hat{O} | \psi_i \rangle \quad (3.5.4)$$

where  $V$  is the interaction potential. This matrix element appears in both scatterings and decays, although of course the potential will be different in both cases, in particular

$$\sigma \propto |\mathcal{M}|^2, \quad \text{and} \quad \tau \propto \frac{1}{|\mathcal{M}|^2}. \quad (3.5.5)$$

### 3.5.2 Particle Decays

Most particles are unstable and quickly decay. We can measure the time taken for decay and the particles produced in the process. These properties relate to the underlying dynamics of the interaction that causes the decay.

Given a system with  $N_0$  initial particles we expect that the decay rate is proportional to the number of particles, that is

$$\frac{dN}{dt} = -\frac{N(t)}{\tau} \quad (3.5.6)$$

where  $N(t)$  is the number of particles at time  $t$  and  $\tau$  is a characteristic time which we call the **lifetime** of the particle. Solving this differential equation we get

$$N(t) = N_0 e^{-t/\tau} = N_0 e^{-\Gamma t} \quad (3.5.7)$$

where we define the **energy width** of the particle to be  $\Gamma := 1/\tau$ . We can see that  $\tau$  is the time taken for the sample to decay to  $1/e \approx 37\%$  of its original count.

Returning to the differential equation we can work out the number of decays per particle per second, which is simply

$$\frac{1}{N} \frac{dN}{dt} = -\frac{1}{\tau} = -\Gamma. \quad (3.5.8)$$

What this means is that the probability that a given particle will decay in a given length of time is constant. This means that elementary particles do not age.

We use  $\tau$  to denote the particle lifetime for a reason. We report lifetimes in the rest frame of the particle, so they are proper times. This allows for unstable particles to travel a surprisingly large distance,  $L = \gamma v \tau = \gamma \beta c \tau$ , before decaying, since in the frame of the particle it can travel a distance (or rather everything else travels a distance past the particle) given by  $v \tau = \beta c \tau$ , which then gives  $L$  by considering the time dilation needed to translate  $\tau$  into a quantity in the frame of the measurement. We can measure this to determine the lifetime. The most famous example of this is cosmic muons reaching Earth's surface when without relativistic effects they would mostly decay before they were even close.

Typical lifetimes of particles depend on the force that mediates their decay. For particles decaying by the strong force a typical lifetime is  $10^{-23}$  s, for electromagnetism  $10^{-20}$  s, and for the weak force  $10^{-8}$  s.

Many particles can decay in more than one way. For example the K-short meson ( $K_S$ ) can commonly decay in one of two ways, either to two charged pions,  $\pi^+ \pi^-$ , or to two neutral pions,  $\pi^0 \pi^0$ . These two decay modes cover at least 99.9% of decays.

$$K_S \rightarrow \pi^+ \pi^-, \quad \text{or} \quad K_S \rightarrow \pi^0 \pi^0. \quad (3.5.9)$$

Each decay has its own value of the matrix element,  $\mathcal{M}$ , appearing in Fermi's golden rule. This is proportional to the partial decay width for each decay mode, which is simply the decay width calculated only for one decay mode.

$$\Gamma(K_S \rightarrow \pi^+ \pi^-) \propto |\mathcal{M}(K_S \rightarrow \pi^+ \pi^-)|^2, \quad (3.5.10)$$

$$\Gamma(K_S \rightarrow \pi^0 \pi^0) \propto |\mathcal{M}(K_S \rightarrow \pi^0 \pi^0)|^2. \quad (3.5.11)$$

The total decay width is simply the sum of partial decay widths,

$$\Gamma(K_S) = \Gamma(K_S \rightarrow \pi^+ \pi^-) + \Gamma(K_S \rightarrow \pi^0 \pi^0). \quad (3.5.12)$$



Ignoring the small fraction of other decay modes in this example.

We define the **branching fraction**,  $B$ , or  $BF$ , is the fraction of times a particle decays to a given final state.

$$BF(K_S \rightarrow \pi^+\pi^-) = \frac{\Gamma(K_S \rightarrow \pi^+\pi^-)}{\Gamma(K_S)}, \quad (3.5.13)$$

$$BF(K_S \rightarrow \pi^0\pi^0) = \frac{\Gamma(K_S \rightarrow \pi^0\pi^0)}{\Gamma(K_S)}. \quad (3.5.14)$$

It isn't always possible to directly measure the particles produced in interactions, as they may decay before we can measure them. However, a small lifetime corresponds to a small uncertainty in the lifetime,  $\Delta t$ , which means that the uncertainty in the energy,  $\Delta E$ , must be quite large, so that we don't violate the uncertainty principle,  $\Delta E \Delta t \geq \hbar/2$ . This means that, with enough data, we can measure  $\Delta E$ , and take this as evidence for the existence of a particle. We call this the resonance of the particle.

## 3.6 Relativistic Kinematics

### 3.6.1 Decay

Consider a particle  $A$  decaying into two particles  $a$  and  $b$ . We consider the decay in the rest frame of  $A$ . The particles have four-momenta  $p_A = (M_A, 0)$ ,  $p_a = (E_a, \mathbf{p}_a)$ , and  $p_b = (E_b, \mathbf{p}_b)$ . Conservation of four-momentum means that  $p_A = p_b + p_a$ , which immediately leads to  $\mathbf{p}_a = -\mathbf{p}_b$ . We can rearrange to get  $p_b = p_A - p_a$ , and so the invariant mass of  $b$  is

$$p_b^2 = m_b^2 = (p_A - p_a)^2 = p_A^2 + p_a^2 - 2p_A \cdot p_a = m_A^2 + m_a^2 - 2m_A E_a. \quad (3.6.1)$$

Rearranging this we see that the energy of  $a$  is

$$E_a = \frac{m_a^2 + m_b^2 - m_A^2}{2m_A}. \quad (3.6.2)$$

■ **Example 3.6.3** Consider the decay of a positive pion,  $\pi^+$ , into an antimuon and a muon neutrino,  $\pi^+ \rightarrow \mu^+ \nu_\mu$ . The pion has rest mass  $m_\pi = 139.6$  MeV, the muon  $m_\mu = 105.7$  MeV and the neutrino's mass is negligible. What is the energy of the antimuon, and what is the magnitude of the three-momenta of the daughter particles?

The energy of the muon is

$$E_\mu = \frac{m_\pi^2 + m_\mu^2}{2m_\pi} = \frac{(139.6 \text{ MeV})^2 + (105.7 \text{ MeV})^2}{2(139.6 \text{ MeV})} = 109.9 \text{ MeV}. \quad (3.6.4)$$

The momentum of the daughter particles is

$$|\mathbf{p}_\mu| = |\mathbf{p}_\nu| = \sqrt{E_\mu^2 - m_\mu^2} = \sqrt{(109.9 \text{ MeV})^2 - (105.7 \text{ MeV})^2} = 29.8 \text{ MeV}. \quad (3.6.5)$$

We can also determine the mass of the particle which decayed if we know the three-momentum, energy and mass of the daughter particles.

■ **Example 3.6.6** A  $K_S$  decays into  $\pi^+\pi^-$ . Reconstruct the mass of the particle assuming the energy, momentum and masses of the pions are known.  
By conservation of momentum

$$p(K_S) = p(\pi^+) + p(\pi^-). \quad (3.6.7)$$

Squaring this we get the invariant mass of the kaon

$$m(K_S)^2 = (p(\pi^+) + p(\pi^-))^2 \quad (3.6.8)$$

$$= p(\pi^+)^2 + p(\pi^-)^2 + 2p(\pi^+) \cdot p(\pi^-) \quad (3.6.9)$$

$$= m(\pi^+)^2 + m(\pi^-)^2 + 2E(\pi^+)E(\pi^-) - 2\mathbf{p}(\pi^+) \cdot \mathbf{p}(\pi^-). \quad (3.6.10)$$

### 3.6.2 Scattering

The most important parameter in particle scattering experiments is the centre of mass energy,  $\sqrt{s}$ , which is the maximum amount of energy available in a collision, for a collision in which the initial particles are annihilated  $\sqrt{s}$  is the maximum mass of a particle that can be produced. If the incoming particles are  $a$  and  $b$  then  $\sqrt{s}$  can be found from the square of the sum of their four-momenta:

$$s := (p_a + p_b)^2 = p_a^2 + p_b^2 + 2p_a \cdot p_b = m_a^2 + m_b^2 + 2(E_a E_b - |\mathbf{p}_a| |\mathbf{p}_b| \cos \vartheta). \quad (3.6.11)$$

There are two main types of scattering experiment. First is fixed target scattering, in which a beam of particles,  $a$ , is accelerated towards a fixed target,  $b$ , which is at rest. Second is a collider experiment in which a beam of  $a$  and a beam of  $b$  are collided against each other, often, but not always, with  $\mathbf{p}_a = -\mathbf{p}_b$ .

For fixed target scattering we have  $\mathbf{p}_b = 0$  and so  $E_b = m_b$ . We typically assume  $E_a \gg m_a, m_b$  also and so

$$s = m_a^2 + m_b^2 + 2E_a m_b \approx 2E_a m_b \implies \sqrt{s} \approx \sqrt{2E_a m_b}. \quad (3.6.12)$$

For collider experiments we typically have  $E_a = E_b = E$  and  $E \gg m_a, m_b$ , and head on collisions, so  $\vartheta = \pi$ , then

$$s = 4E^2 \implies \sqrt{s} = 2E. \quad (3.6.13)$$

This means that we can get much higher collision energies from collider experiments for some fixed beam energy. For this reason collider experiments are often preferred.

■ **Example 3.6.14** What is  $\sqrt{s}$  for a fixed-target experiment and a collider experiment with protons ( $m_p \approx 1 \text{ GeV}$ ) if we use proton beams with energy<sup>a</sup>  $E_p = 1 \text{ TeV}$ ?

For the fixed-target experiment the collision energy is

$$\sqrt{s} = \sqrt{2E_p m_p} = \sqrt{2(10^3 \text{ GeV})(1 \text{ GeV})} = 45 \text{ GeV}. \quad (3.6.15)$$

For the collider experiment the collision energy is

$$\sqrt{s} = 2E_p = 2 \text{ TeV}. \quad (3.6.16)$$

So the collider experiment has about 45 times as much energy available after the collision.

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$$^a 1 \text{ TeV} = 10^3 \text{ GeV} = 10^{12} \text{ eV}.$$

## Part II

# Experiments

# Four

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## Accelerators

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### 4.1 Source

All accelerators need a source of particles for the beam. For beams of protons this typically comes as hydrogen gas, which can fairly easily be broken down and ionised into protons and electrons. The actual amount of material needed is very small on human scales. The small bottle that feeds the LHC is only a few litres in volume (although is at high pressure, so contains more than a few litres of gas at 1 atm), this bottle contains enough hydrogen to run the LHC for 200 000 years.

### 4.2 LINAC

Most accelerators start with a **linear accelerator (LINAC)**. This consists of long evacuated pipe containing a set of metal drift tubes. These are shorter tubes through which the particles pass. The tubes are attached to alternate sides of an alternating power supply, which is in the radio frequency (RF), typically about 100 MHz. Particles are produced continuously by the source but only those in phase with the source are accelerated.

The energy a LINAC can accelerate a particle to is limited by the length of the accelerator and the voltage, increasing either increases the power consumption, which is the real limiting factor.

One problem with LINACs is they don't produce high luminosity beams and have a low duty cycle, that is they emit a burst of particles each time the phase of the source reverses, but between these reversals they don't emit any particles.

#### 4.2.1 Examples

LINAC 4 is the first stage of CERN's accelerator complex, it accelerates protons to 160 MeV in several stages.

The largest LINAC ever constructed was the 3 km long SLAC linear collider which accelerated electrons and positrons to 50 GeV with the goal of studying the Z boson resonance. It had a duty cycle producing bursts that lasted 2  $\mu$ s 60 times a second, which means that 99.99 % of the time it wasn't producing any particles.

### 4.3 Synchrotrons

For higher energy collisions LINACs become infeasible and we instead turn to **synchrotrons**. These are essentially LINACS alternated with magnets to curve the path of the beam, so that it goes in a large (almost) circle. The particles can then orbit as

many times as we like without running out of accelerator, fixing one of the main the problems with LINACs.

Particles in the synchrotron bunch according to the frequency used to accelerate them.

A synchrotron of radius  $R$  and magnetic field strength  $B$  can accelerate particles up to a momentum of  $p \propto BR$ . The particles must be injected all at once when  $B$  is relatively low and then as the particles accelerate  $B$  is increased to keep them in the synchrotron. For non-superconducting magnets the strongest the field can get is about 1 T (which is quite high, teslas are a large unit). Superconducting magnets such as those used at the LHC can get up to 8 T.

Once the particles reach their top speed they make approximately  $10^5$  revolutions per second. Each revolution gives them a kick of about 0.1 MeV. It is possible for particles to travel on the order of  $10^6$  km while being accelerated.

As well as accelerating the particles we also need to focus the beams. At injection the beam will fill the vacuum pipe. After the beam has been accelerated and focused the width of the beam will be only a few millimetres. This focusing is done with still more magnets, in particular quadrupoles with alternating polarity are used to squash the beam.

One limitation to synchrotrons is the loss of energy as radiation due to the extra acceleration of charged particles to get them to go in a circle, called **synchrotron radiation**. This energy is given by

$$\Delta E = \frac{q^2 \beta^3 \gamma^4}{3\epsilon_0 R}. \quad (4.3.1)$$

Since  $\gamma = E/m$  the energy loss is proportional to  $1/m^4$ , which means that this problem is  $(m_p/m_e)^4 \approx 10^{13}$  times worse for protons.

Synchrotron radiation means that for light particles we actually need a longer synchrotron than LINAC once we reach a certain energy. For example, the large electron positron collider (LEP) was 27 km in circumference and accelerated electrons and positrons to 200 GeV. The proposed international linear collider (ILC) will be about 20 km in length and will be able to accelerate electrons and positrons up to 1 TeV.

Synchrotron radiation isn't a complete waste though. We can use synchrotrons to create beams of x-rays which are used to study material properties in x-ray diffraction<sup>1</sup>.

<sup>1</sup>see the notes for the introduction to condensed matter physics course for more detail

## 4.4 Beam Losses

The energy stored in a beam is often very large. For example, the LHC beam stores 326 MJ of energy. This is equivalent to the kinetic energy of a 747 airliner travelling at  $140 \text{ km h}^{-1}$ . If all this energy was delivered at once it would be equivalent to setting off about 70 kg of TNT. Clearly this is not good.

It is therefore very important that we don't lose control of the beam during experiments, even if no one is nearby to get hurt particle accelerators tend to be quite expensive, and also by their nature contain a lot of very sensitive equipment which would not fair well if struck by the beam.

When we are finished with a beam it is dumped into special beam dumps, which are large blocks of metal or concrete designed to absorb the beam safely.

## 4.5 Colliders

Colliders are at the forefront of experimental particle physics. Broadly we can separate them into three classes depending on what they collide:

- $e^-e^+$  colliders—precision machines, the interactions are simple as these are fundamental particles and don't interact through the strong force.
- $e^-p$  colliders—mostly used to probe the structure of the proton.
- $pp$  or  $p\bar{p}$  colliders—discovery machines, essentially the idea is if we smash enough particles together at high enough energy something interesting will happen eventually.

Hadron colliders (that is  $pp$  or  $p\bar{p}$  colliders) reach the highest energies and have the highest luminosities. However, the interactions are messy and complicated by the presence of the strong force and the fact that protons aren't fundamental, really we're interacting quarks within the protons.

Electron-positron colliders are good for precision physics, if we know what particles we want to produce, and electrons are fundamental and don't interact by the strong force, which makes the results easier to work with. However, they can't get to as high energy or luminosity as hadron colliders.

### 4.5.1 $Sp\bar{p}S$

The super proton antiproton synchrotron ( $Sp\bar{p}S$ ) was a proton-antiproton collider running at CERN running at ran from 1981 to 1991. It had a centre of mass energy of 400 GeV. It was 6.9 km in diameter and ran two experiments UA1 and UA2 (where UA stands for underground area). The 1984 Nobel prize in physics was awarded to Carlo Rubbia and Simon van der Meer who were working with the  $Sp\bar{p}S$  when they discovered W and Z bosons.

### 4.5.2 Tevatron

The tevatron was a proton-antiproton collider at Fermilab from 1987 to 1995 at an energy of 1.8 TeV, and then from 2000 to 2008 at 1.96 TeV. It had two experiments, CDF and  $D\bar{0}$ , which discovered the top quark. The upgrade in the late nineties lead to approximately a 40 % increase in interesting cross sections, despite the relatively small increase in energy.

### 4.5.3 HERA

The hadron electron ring accelerator (HERA) is the words only electron-proton collider, it ran from 1992 to 2007. It ran at asymmetric beam energies where the electron beam had energy 30 GeV and the proton beam had energy 820 GeV. It was 6.3 km in circumference. There were three experiments, two were general purpose, called ZEUS and H1, and one called HERA-B which was interested in the properties of B mesons. HERA was used to take detailed measurements of the proton's structure.

### 4.5.4 LEP

LEP was the worlds highest energy electron-positron collider. It ran from 1989 to 2000 and was 27 km in circumference. It had four experiments, Aleph, Delph, L3, and Opal.

From 1989 to 1995 it focussed on measurements of Z bosons with resonant production, essentially choosing  $\sqrt{s} = m(Z)$ . It observed approximately four million  $e^-e^+ \rightarrow Z$  events.

From 1997 to 2000 it focused on measurements of W bosons with resonant production at  $\sqrt{s} \geq 2m(W)$ . It observed approximately 8000  $e^-e^+ \rightarrow Z \rightarrow W^+W^-$  events.

During its lifetime scientists precisely measured many properties of W and Z bosons, including their masses, widths, and couplings. The LEP was used to evaluate electroweak theory (see ).

#### 4.5.5 B-factories

Two colliders, the BaBar<sup>2</sup> experiment at SLAC and the Belle experiment at the KEK collider in Japan, were looking into the B meson. Both ran with asymmetric beam energies, 9 GeV and 3.1 GeV, and 8 GeV and 3.5 GeV respectively for electrons and positrons respectively. The KEKB experiment is still running, having recently been upgraded. The BaBar experiment ceased in 2007 but data analysis is ongoing.

<sup>2</sup>named for B and  $\bar{B}$  (B-bar)

#### 4.5.6 LHC

The LHC is the highest energy collider, currently reaching 13 TeV, with plans to upgrade to 13.5 TeV. It is 27 km in diameter (they repurposed the tunnel that used to hold LEP). It is a proton-proton collider. It has 2556 proton bunches at once, separated by only 25 ns.

The LHC runs off of a beam injected from the super proton synchrotron (SPS), which is a repurposed Sp̄pS, it will also need to be upgraded when the LHC is. The injected beam is 450 GeV and is injected 8 bunches at a time.

Once in the LHC the beams are adjusted and accelerated until they reach the desired energy, then called stable beams. This can take several hours. The beams are then used in several experiments. The longest fill lasted about 36 hours, most last for about a day and in full operation hundreds of fills are performed in a year. In 2012 the LHC was used to discover the Higgs boson.

#### 4.5.7 FCC

The future circular collider (FCC) is a proposed collider at CERN. If built it will be 80 km to 100 km in circumference. It is undecided what exactly it will collide. The three options being proton-proton, electron-positron, or electron-proton. It is thought that the proton-proton centre of mass energy would be 100 TeV, whereas various proposals have put the electron-positron centre of mass energy at 91 GeV, 160 GeV, 216 GeV and 350 GeV (these roughly correspond to the mass of Z, twice the mass of  $W^\pm$ , 1.7 times the mass of  $H^0$ , and twice the mass of t, so these different energies would focus on different particles).

If built it will probably use the proposed compact linear collider (CLIC) as the first step in acceleration.



# Five

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## Detectors

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After we collide particles we need a way to see what happened. This is where detectors come in, although their use is not limited to collider experiments, they can also measure the result of fixed-target experiments or particle decays. There are several properties that we might want to measure. Top of the list has to be mass, since this pretty much tells us what the particle is, up to the few particles with equal mass partners with different charge etc. Unfortunately we can't measure mass directly, instead we measure the energy and momentum of particles and from this we can work out their mass. Other properties we may measure include charge, spin, position, and various other quantum numbers.

Frequently particles decay before we can measure them, for example the W and Z bosons, the Higgs boson, the top quark, the list goes on. Instead we are left with the daughter particles, or if they decay quickly the granddaughters and so on. It is then someone's job to track backwards through the decays to try and work out properties of the original particle, mostly relying on conservation laws.

### 5.1 The First Detectors

The first particle detector was a cloud chamber. In this a supersaturated vapour (that is as much vapour as the air can hold) is created. An ionising (i.e. charged, relatively high energy) particle crossing the cloud chamber will cause bubbles of liquid to condense along its path. We see this as a trail. By applying a magnetic field of known strength to a cloud chamber we can curve the paths of the particles and this gives us a way to work out the sign of the charge (based on the direction of curvature) and compare masses of equally charged particles, based on the radius of curvature. For example, a cloud chamber was used to distinguish the light positron from the heavier proton, both of which have the same charge, as the positron track had a tighter curve.

One development on the cloud chamber was the bubble chamber. This used a superheated liquid which would start to boil when an ionising particle passed through. The liquid is maintained at high temperature until the beam of particles reaches it and then the pressure is released and the tracks can be seen. Bubble chambers were used to discover many hadronic states.

Cloud and bubble chambers were limited by the fact that they required a human to analyse the photographs of paths (although there are now ways for computers to do this using image processing and AI). Bubble chambers had another limitation of taking a while to set up and then only working for a brief time while the pressure is being released.

Geiger–Müller tubes consist of an anode wire, surrounded by a cathode cylinder. The cylinder is filled with an inert gas and a large potential applied across the anode and cathode. Particles ionise the gas, knocking off electrons, which then drift toward the anode. Along the way the electrons knock off more electrons, this is known as the avalanche effect, while the initial radiation may only liberate about 100 electrons by the time the electrons reach the anode there will be more like  $10^6$ . Eventually a large number of electrons reach the anode and produce a measurable electrical signal. One problem with Geiger–Müller tubes is that they take time for signals to disappear, so if the rate of irradiation is too high they will miss events. They also can't tell us much more than that there was an ionising event.

## 5.2 Wire and Drift Chambers

The wire chamber is similar to the Geiger–Müller tube in how it works. It was invented by Georges Charpak in the 1960s. It consists of an array of anode wires spaced 5 mm to 10 mm apart. Instead of a cathode tube cathode planes are used. When an ionising event happens the electrons travel to the nearest anode. The precision is limited by the wire spacing,  $d$ , such that, due to the geometry of the setup, the closest position of an event can be resolved to is  $d/\sqrt{12}$ .

A drift chamber works in the same way but also measures the time taken for the electrons to reach nearby wires, which gives an idea of how far from the wire the ionising event occurred. This can improve the spatial resolution to up to 200  $\mu\text{m}$ . It is also possible to use this to track the trajectory of the ionising particle.

## 5.3 Silicon Detectors

The most precise detectors today use silicon semi-conductors in a reverse bias mode, that is a small voltage is applied to stop current flow under normal circumstances. A charged particle passing through then allows a small amount of current to flow, which we can measure. These semi-conductors are placed with a spacing of approximately 100  $\mu\text{m}$ , with a resulting resolution of approximately 20  $\mu\text{m}$ .

For example, the compact muon solenoid (CMS) experiment at CERN is one of the LHC experiments. It has 210  $\text{m}^2$  of silicon strip detectors, with 10 million readout channels. It also has 1  $\text{m}^2$  of pixels, with 65 million readout channels.

## 5.4 Electromagnetic Calorimeters

In high energy experiments, like the LHC, it can be hard to tell the difference between photons and electrons, both are light particles interacting via the electromagnetic force. The main difference is that electrons are charged. This means that they leave trails in the tracking detectors mentioned in the previous section.

The way that photons interact typically depends on their energy. For photons of less than 100 keV the main interaction is the photoelectric effect. For photons of approximately 1 MeV the main interaction is Compton scattering. For photons of more than 10 MeV the main interaction is pair production of electrons and positrons.

A calorimeter is a detector which measures the energy deposited when a particle is absorbed. Typically they are placed behind the tracking detectors. An electromagnetic calorimeter is usually the first calorimeter. It measures photons and electron-positron showers from pair production and **bremsstrahlung**<sup>1</sup>, which is ra-

<sup>1</sup>from the German *bremsen* “to break” and *strahlung* “radiation”

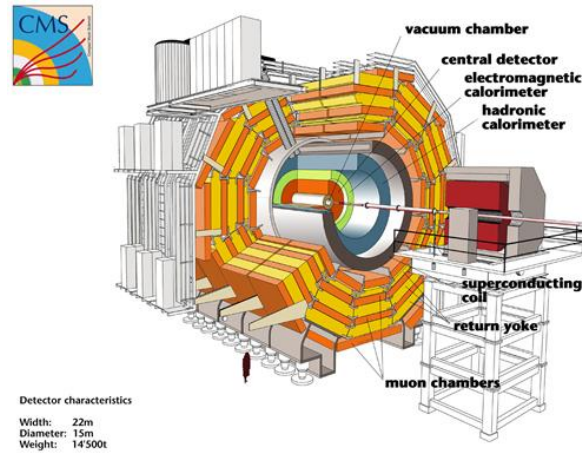


Figure 5.1: The detectors used in the CMS experiment at CERN [1]

diation produced when a charged particle slows down suddenly when interacting with another charged particle, in this case an electron interacting with a nucleus in the detector.

## 5.5 Hadronic Calorimeters

After electromagnetic calorimeters detectors typically have hadronic calorimeters, which measure the result of hadronic showers. These are typically messy and difficult to reconstruct, due to the main interaction being the strong force. For this reason hadronic calorimeters tend to be coarser and have lower resolution than their electromagnetic counter parts.

## 5.6 Muon Detectors

The final stage in most detector setups are muon detectors. These can be very coarse as there are typically only a few muons per event. They are detected with a mixture of scintillating detectors, drift tubes and other simpler detectors.

## 5.7 Particle Identification

There are a variety of methods for identifying particles. Muons are easily detected as they are the only particles to penetrate into the muon detectors, which are the last stage of the detectors. Electrons can be distinguished by causing a shower in the electromagnetic calorimeter, and no corresponding shower in the hadronic calorimeter.

Things get a bit more tricky when it comes to differentiating hadrons, such as pions and kaons. One thing we can do is measure the energy loss in silicon trackers. It can be shown that this depends on  $\beta\gamma = p/m$ . We can also measure time of flight using drift chambers or scintillation counters.

Time of flight detectors are particularly useful for particle identification. Suppose a particle travels a distance  $L$  at speed  $\beta$ , this takes time  $t = L/\beta$ . We have

$$\frac{1}{\beta} = \frac{E}{p} = \frac{\sqrt{p^2 + m^2}}{p} = \sqrt{1 + \frac{m^2}{p^2}} \approx 1 + \frac{m^2}{2p^2}. \quad (5.7.1)$$

We therefore have

$$t = L \left( 1 + \frac{m^2}{2p^2} \right). \quad (5.7.2)$$

Suppose an event has the possibility of producing either a kaon or pion, both with equal momentum. We can differentiate between the two by considering the time taken to reach a given point a distance  $L$  away from the event. In particular the time difference between them arriving is

$$\Delta t = \frac{L}{2p^2} (m_K^2 + m_\pi^2). \quad (5.7.3)$$

For typically experimental values where  $L$  is on the order of a few metres and  $p$  is on the order of a few GeV we find that  $\Delta t$  is on the order of a few hundred pico seconds, which is a measurable time difference, so by measuring momentum and time of flight we can be sure whether we have detected a kaon or pion.

Another effect we can use is **Cherenkov** radiation, which is electromagnetic radiation emitted when a particle travels in a medium faster than the speed of light in that medium, essentially the electromagnetic version of a sonic boom. The radiation spreads out in a cone, and the angle of the cone depends on  $\beta$ . Therefore looking at the particle head on we see a ring of Cherenkov radiation, and the radius of this relates to the particle's speed.

Cherenkov radiation can actually be seen by the human eye, and this was used in early accelerators to align beams, by looking down the beam and aligning the beam so that the Cherenkov radiation forms circles, in which case you knew you were looking at it head on. Needless to say this is a *bad idea* and if you did it with today's higher energy beams you would almost certainly die.

Astronauts in space (and non-astronauts in space) see flashes of light, which it turns out are due to cherenkov radiation being produced *in their eyes* when particles, mostly from the sun, pass through their eye faster than light travels through their eyes. Cherenkov radiation is also seen in water surrounding nuclear detectors, typically the radiation falls in the blue part of the spectrum.

## 5.8 Neutrinos

Neutrinos don't interact in detectors. Instead we can infer their presence by looking for "lost" momentum. However, this method can't tell us what the extra, undetected, particle is, just that there is one. Most likely the particle is a neutrino, but we can't rule out other non-interacting particles which have yet to be discovered.

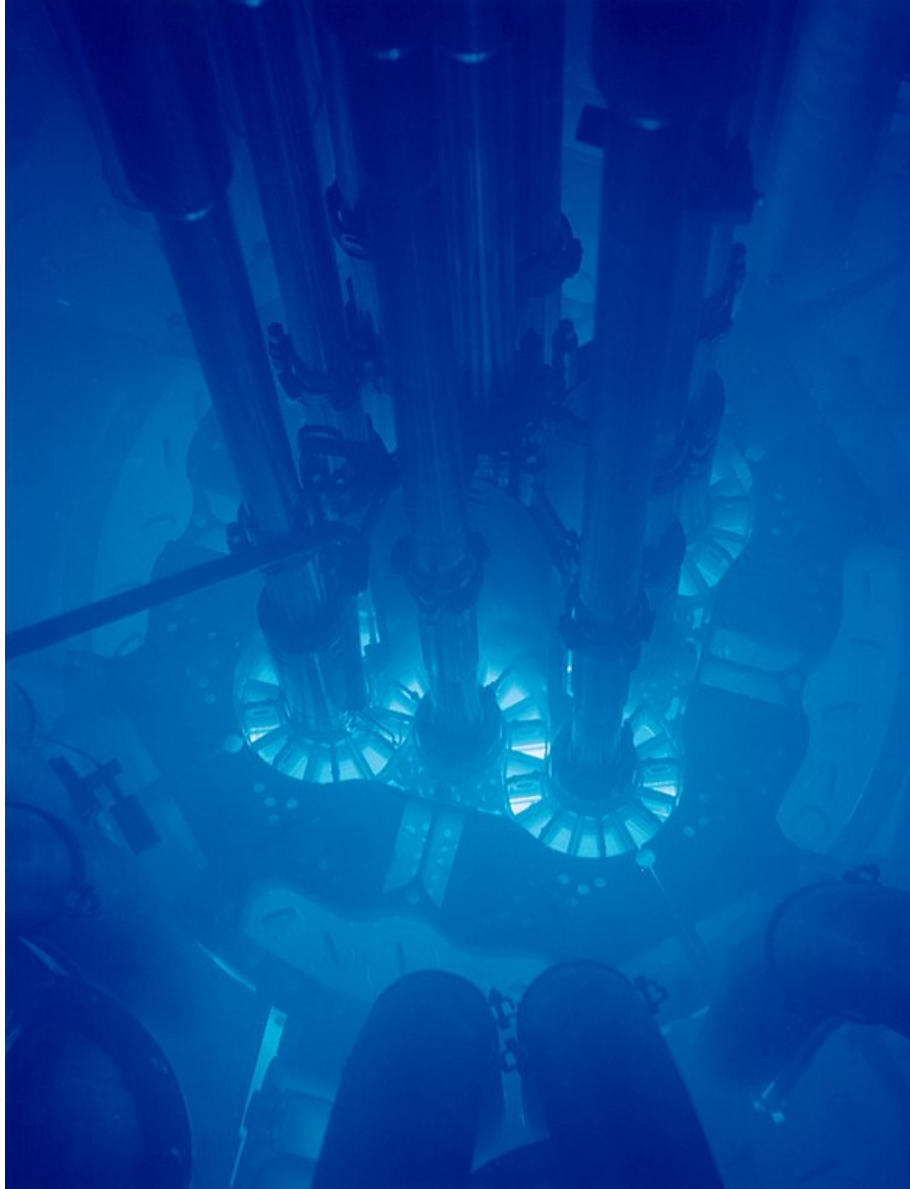


Figure 5.2: Cherenkov radiation in water surrounding a nuclear reactor [2]

## Part III

# Quantum Electrodynamics

## Six

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# Quantum Electrodynamics

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### 6.1 Dirac Equation

Recall that in [Section 3.4](#) we discussed the Klein-Gordon equation,

$$-\frac{\partial^2 \psi}{\partial t^2} = -\nabla^2 \psi + m^2 \psi, \quad (6.1.1)$$

which applies to spin zero particles. We briefly discussed the existence of negative energy solutions, which we determined correspond to antimatter.

We now briefly cover the **Dirac equation**. This was found by Dirac, who was dissatisfied with the Klein-Gordon equation. He started with the Schrödinger-like equation

$$i\hbar \frac{\partial \psi}{\partial t} = \hat{H} \psi \quad (6.1.2)$$

for some Hamiltonian operator  $\hat{H}$ . Dirac then required that the time and position be treated in the same way, as is common in relativity, this means first order spatial derivatives. He concluded that the Hamiltonian must be of the form

$$\hat{H} = -i\hbar c \alpha_i \partial_i + \beta m c^2 \quad (6.1.3)$$

where we employ the Einstein summation convention summing  $i$  from 1 to 3 in the first term. It turns out that  $\alpha_i$ ,  $\beta$ , and  $\psi$  cannot be numbers, but must instead be matrices.

The simplest solution has  $\alpha_i$  and  $\beta$  being  $4 \times 4$  matrices and  $\psi$  being a four component wave-function, known as a spinor. The Dirac equation can be written in the form

$$(i\gamma^0 \partial_t + i\boldsymbol{\gamma} \cdot \nabla - m) \psi = (i\gamma^\mu \partial_\mu - m) \psi = 0 \quad (6.1.4)$$

where  $\gamma^\mu$  are again  $4 \times 4$  constant matrices, which are related to  $\alpha_i$  and  $\beta$  by  $\gamma^0 = \beta$  and  $\gamma^i = \beta \alpha_i$ , and we understand that  $\boldsymbol{\gamma} \cdot \nabla$  is shorthand for

$$\boldsymbol{\gamma} \cdot \nabla = \gamma^i \partial_i = \gamma^1 \partial_1 + \gamma^2 \partial_2 + \gamma^3 \partial_3. \quad (6.1.5)$$

The Dirac equation can be written even more compactly using the Feynman slash notation, where we define  $\not{A} := \gamma^\mu A_\mu$ , so

$$(i\not{\partial} - m) \psi = 0. \quad (6.1.6)$$

For completeness the gamma matrices, in this four-dimensional representation, are given by

$$\gamma^0 := \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}, \quad \gamma^1 = \begin{bmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{bmatrix}, \quad (6.1.7)$$

$$\gamma^2 := \begin{bmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & i & 0 & 0 \\ -i & 0 & 0 & 0 \end{bmatrix}, \quad \gamma^3 = \begin{bmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{bmatrix}. \quad (6.1.8)$$

These can be compactly written as block matrices:

$$\gamma^0 = \begin{bmatrix} I & 0 \\ 0 & -I \end{bmatrix} = \sigma^3 \otimes I, \quad \text{and} \quad \gamma^k = \begin{bmatrix} 0 & \sigma_k \\ -\sigma_k & 0 \end{bmatrix} = i\sigma^2 \otimes \sigma^k \quad (6.1.9)$$

where  $\otimes$  is the tensor product<sup>1</sup>,  $I$  is the  $2 \times 2$  identity matrix and  $\sigma_k$  are the Pauli spin matrices,

<sup>1</sup>see principles of quantum mechanics notes for more details

$$\sigma_1 = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix}, \quad \sigma_2 = \begin{bmatrix} 0 & -i \\ i & 0 \end{bmatrix}, \quad \text{and} \quad \sigma_3 = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix}. \quad (6.1.10)$$

The Dirac equation has plane wave solutions given by

$$\psi(\mathbf{r}, t) = u(\mathbf{p}) \exp \left[ -\frac{i}{\hbar}(Et - \mathbf{r} \cdot \mathbf{p}) \right] \quad (6.1.11)$$

where  $u$  is some function of the momentum,  $\mathbf{p}$ . In four-vector notation and natural units this becomes

$$\psi(x^\mu) = u(p^i) e^{-ix_\mu p^\mu}. \quad (6.1.12)$$

There are still positive and negative energy solutions. In fact, there are four solutions, two of which are positive and two of which are negative, once we include spin.

## 6.2 Antimatter

The negative energy solutions to the Dirac and Klein–Gordon equation gave rise to the prediction of antimatter. There are two common ways to interpret antimatter.

### 6.2.1 Dirac Interpretation

Dirac interpreted the negative energy solutions as being hole states. That is he imagined that the vacuum is filled with negative energy levels, each of which is occupied by two electrons (with opposite spins). This is called the **Dirac sea**. A hole in the sea then appears as a positively charged state with positive energy. From this Dirac predicted the positron in 1928, and it was discovered just a few years later in 1931. Note that the energy required to jump from a negative energy state to the positive energy hole state is  $2m_e c^2$ .



### 6.2.2 Feynman–Stückelberg Interpretation

A more modern view of antimatter is as normal matter moving backwards in time. This interpretation is due to the fact that if we have a state,

$$\psi(\mathbf{r}, t) \propto \exp[-i(Et - \mathbf{r} \cdot \mathbf{p})], \quad (6.2.1)$$

and we reverse it's direction, so  $\mathbf{p} \rightarrow -\mathbf{p}$  and time,  $t \rightarrow -t$ , we get the same state, but moving backwards, so  $\mathbf{r} \rightarrow -\mathbf{r}$ :

$$\psi(-\mathbf{r}, -t) \propto \exp[-i((-E)(-t) - (-\mathbf{r}) \cdot (-\mathbf{p}))] = \exp[-i(Et - \mathbf{r} \cdot \mathbf{p})]. \quad (6.2.2)$$

## 6.3 Feynman Diagrams

**Feynman diagrams** are a pictorial representation of a process at a particular order in perturbation theory. They consist of a directed graph with different types of edges for different types of particles. We view time as going from left to right in a Feynman diagram, although other conventions have time flowing upwards. A fermion is represented as a straight line with an arrow in the direction of time. An antifermion is represented by a straight line with an arrow *backwards* in time. A boson is represented by a wavy (photons), loopy (gluons), or dashed (W and Z bosons) line. This boson line has an arrow only for  $W^\pm$ , where the direction the charge is carried is important.

Each Feynman diagram represents a term in a perturbation expansion for a given process. This means that there is a set of rules for turning a Feynman diagram into maths, in particular for the matrix element of the term in perturbation theory. The set of rules depends on the context of the diagram (QED, QCD, etc.). We give the set of rules for QED here:

- A vertex for the interaction of a charged fermion and a photon, the only type of vertex considered in QED, gives a contribution of the charge of the fermion,  $Q$ . This may also be written in terms of the fine structure constant

$$\alpha = \frac{e^2}{4\pi\epsilon_0\hbar c} \quad (6.3.1)$$

or in natural units, with  $\epsilon_0 = 1$ ,

$$\alpha = \frac{e^2}{4\pi}. \quad (6.3.2)$$

<sup>2</sup>see Section 6.7

This is a dimensionless value which is (at low energies<sup>2</sup>) approximately  $1/137$ . So, for fermions with unit magnitude charge we can instead consider a contribution of  $\sqrt{\alpha}$ , since these are all just proportionality factors, so the  $1/\sqrt{4\pi}$  can be absorbed in the constant of proportionality.

- A photon propagator gives a contribution of  $1/q^2$ , where  $q$  is the transferred four-momentum.

At each vertex charge, momentum, energy, and flavour must be conserved.

A simple first example of a Feynman diagram is an electron scattering off of a muon. The Feynman diagram for this is



The matrix element corresponding to this diagram is

$$\mathcal{M}(e^- \mu^- \rightarrow e^- \mu^-) \propto \frac{1}{q^2} = \frac{e^2}{q^2} \propto \sqrt{\alpha} \frac{1}{q^2} \sqrt{\alpha} = \frac{\alpha}{q^2}. \quad (6.3.4)$$

### 6.3.1 s, p, u-Channel Diagrams

For this section we work with examples of **Bhabha scattering**, which is the electron-positron scattering process

$$e^- e^+ \rightarrow e^- e^+, \quad (6.3.5)$$

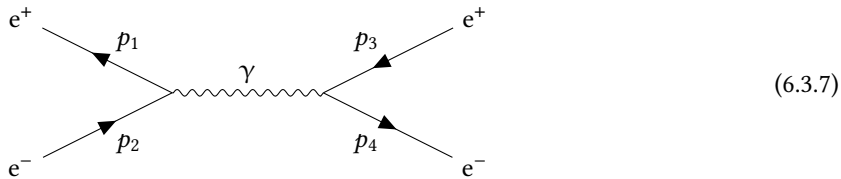
as well as electron-electron scattering,

$$e^- e^- \rightarrow e^- e^-. \quad (6.3.6)$$

To first order there are three different Feynman diagrams to consider. While we refer to the process as scattering this is simply because the initial and final particles are the same. We actually have to allow for things like annihilation and then pair production.

#### 6.3.1.1 s-Channel

Suppose that the initial electron and positron annihilate, producing a photon, and then that photon produces an electron-positron pair. This is shown in the following interaction:



For this process we define the quantity  $s$  to be the invariant total four-momentum squared,

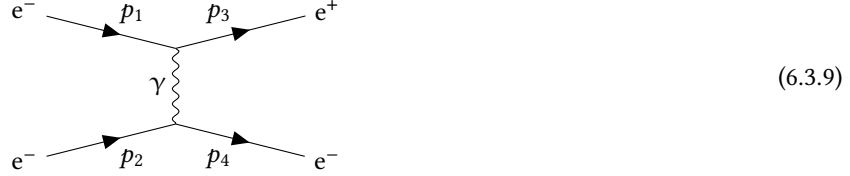
$$s := (p_1 + p_2)^2 = (p_3 + p_4)^2 = q^2, \quad (6.3.8)$$

where  $q$  is the four-momentum of the photon. This quantity,  $s$ , is one of the three **Mandelstam variables**.

This quantity can be defined for any interaction of this type, where two particles combine and then break apart into two new particles. We call this an **s-channel** diagram, where the  $s$  stands for space, as the photon is space-like, travelling horizontally, parallel to the time axis.

### 6.3.1.2 t-Channel

Suppose that two electrons scatter, the most obvious way for this to happen is as represented in the following diagram:



For this process we define the quantity  $t$  to be the invariant change in four-momentum squared, so

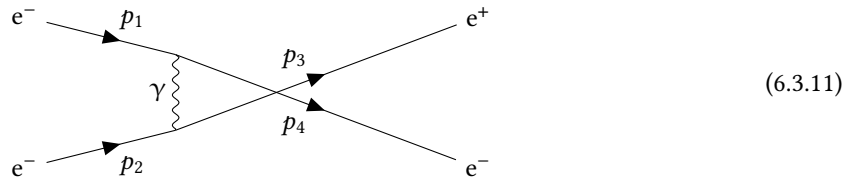
$$t := (p_1 - p_3)^2 = (p_2 - p_4)^2 = q^2, \quad (6.3.10)$$

where  $q$  is the four-momentum of the photon. This is another one of the Mandelstam variables.

We can define this quantity for any interaction of this type, where one particle emits a particle and the other absorbs it. We call this a ***t*-channel** diagram, where the  $t$  stands for time, as the photon is time like, travelling vertically, parallel to the space axis.

### 6.3.1.3 u-Channel

The next process is very similar to the previous process. Since electrons are all identical we can, in theory, swap the two electrons somewhere along the way and the result should be indistinguishable. This is represented by the following diagram:



For this process we define the quantity  $u$  to be the invariant change in four-momentum squared, so

$$u := (p_1 - p_4)^2 = (p_2 - p_3)^2 = q^2 \quad (6.3.12)$$

where  $q$  is the four-momentum of the photon. This is the third Mandelstam variable.

We can define this quantity for any interaction of this type, where two identical particles scatter and swap. We call this a ***u*-Channel** diagram, where the  $u$  stands for “um, well, the photon is time like but we’ve already used  $t$ , I guess  $u$  is next in the alphabet”.

## 6.4 QED Rules

### 6.4.1 QED Vertex

The building block of QED is the QED vertex, for example,



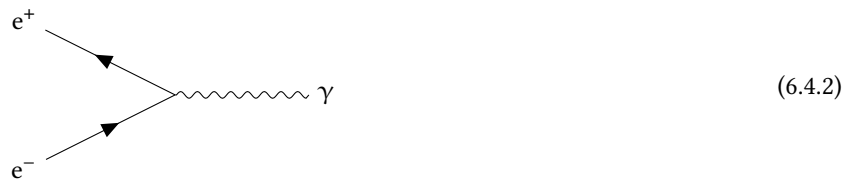
This shows a fermion and antifermion annihilating. In general a QED vertex involves two charged fermions interacting and then absorption or emission of a photon.

The matrix element corresponding to this vertex is proportional to  $\sqrt{\alpha} \propto e$ , where  $\alpha$  is the fine structure constant, and  $e$  is the elementary charge.

At every QED vertex the following must be conserved:

- momentum,
- energy,
- charge, and
- fermion flavour.

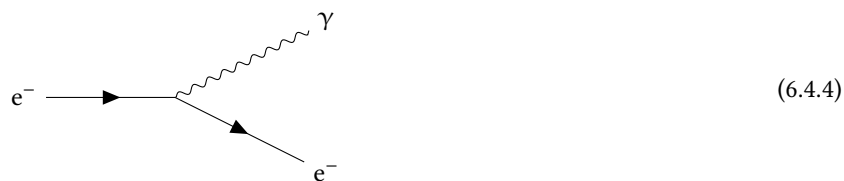
Examples of QED vertices are annihilation,



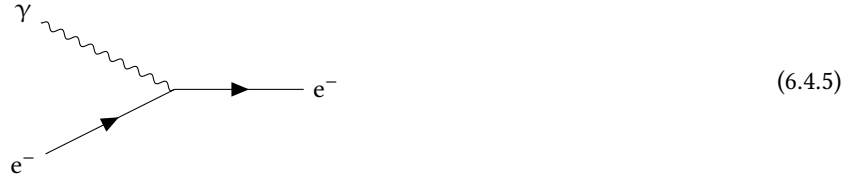
pair production,



photon bremsstrahlung,



and photon absorption,



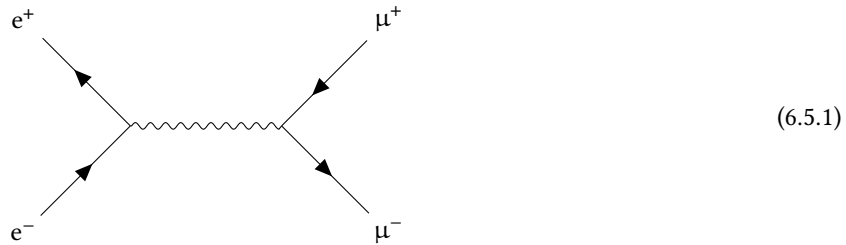
These are all described by the same basic vertex. None of these is a physical process on its own, energy-momentum conservation is violated. We need to combine at least two vertices to get a valid physical process.

## 6.5 Perturbation Theory

**R** For more details on basic perturbation theory see the principles of quantum mechanics notes. For details on perturbation theory with path integrals, as used in QFT, and diagrammatic representations of said integrals see the notes from the quantum theory course.

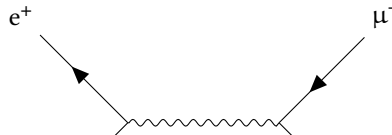
We said that each Feynman diagram represents a term in perturbation theory to a certain order. By this we mean that in a full perturbation theory calculation for the matrix element,  $\mathcal{M}$ , we consider the contribution of all physically allowed diagrams. The order of the diagram is simply the power of  $\alpha$  that appears in the associated matrix element, which is now just one term in the sum giving the overall matrix element.

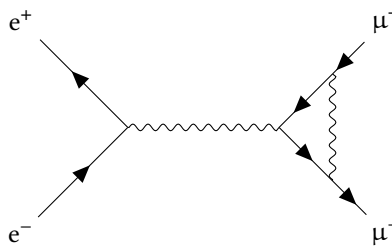
For example, the following diagram is first order in perturbation theory, it has two vertices, each contributing a factor of  $\sqrt{\alpha}$ , meaning  $\mathcal{M} \propto \alpha$ , and so  $|\mathcal{M}|^2 \propto \alpha^2$ .



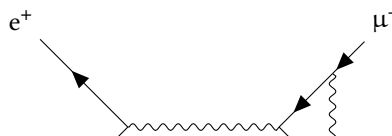
The next two diagrams are both second order in  $\alpha$ , having four QED vertices and so

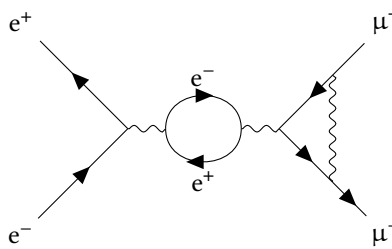
$$|\mathcal{M}|^2 \propto \alpha^4.$$


(6.5.2)


(6.5.3)

The next two diagrams are third order in  $\alpha$ , and so  $|\mathcal{M}|^2 \propto \alpha^6$ .


(6.5.4)


(6.5.5)

In general a graph will have  $|\mathcal{M}|^2 \propto \alpha^{|V|}$ , where  $|V|$  is the number of vertices. It can also be shown that an  $2n$ th order diagram, with  $|\mathcal{M}|^2 \propto \alpha^{2n}$ , will have  $n$  loops.

## 6.6 Experimental Tests of QED

### 6.6.1 Lamb Shift

One of the earliest experimental victories of QED was computing the Lamb shift. For hydrogen if we solve the Dirac equation then we would predict that  $2p_{1/2}$  and  $2s_{1/2}$  states have the same energy. However, if we actually measure the energies of the states we find that they don't have identical energies. QED correctly predicts the small difference in energies if we consider the production of virtual particles, which differs slightly in the two cases as the angular momentums aren't the same.

### 6.6.2 $g - 2$

Elementary particles have intrinsic angular momentum through their spin. One way of measuring this is to measure the magnetic moment related to the spin. We predict that it is given by

$$\boldsymbol{\mu} = -g \frac{e}{2m} \mathbf{L}, \quad (6.6.1)$$

where  $\mathbf{L}$  is the electron's angular momentum,  $e$  is the elementary charge, and  $m$  is the electron's mass. Classically one would expect that  $g = 1$ . If we take into account spin, in its most basic form, as in the Stern–Gerlach experiment<sup>3</sup>, then we find that  $g = 2$ . Actually measuring the magnetic moment we find that for an electron

$$g_e = 2.00238(6), \quad (6.6.2)$$

which differs from 2 by

$$a_e = \frac{g_e - 2}{2} \approx 0.1 \%. \quad (6.6.3)$$

To first order in QED we instead predict

$$g_e = 2 \left( 1 + \frac{\alpha}{2\pi} \right) \approx 2.00232. \quad (6.6.4)$$

A fuller calculation with QED using up to five loops, which corresponds to 12672 diagrams makes predictions that are in remarkable agreement with the measured values:

$$\text{predicted:} \quad a_e = 115965218161(23) \times 10^{-12}, \quad (6.6.5)$$

$$\text{measured:} \quad a_e = 115965218073(28) \times 10^{-12}. \quad (6.6.6)$$

These agree to within  $2.5\sigma$ . Feynman famously said

To give you a feeling for the accuracy of these numbers, it comes out something like this: If you were to measure the distance from Los Angeles to New York to this accuracy, it would be exact to the thickness of a human hair. [3]

This was with the numbers as they were in 1985, using the newer numbers above we have about a factor of 10 improvement, meaning that an equivalent statement now would be measuring the distance from New York to Los Angeles to within the width of a red blood cell. Alternatively this is approximately equivalent to measuring the circumference of Earth to within the width of a human hair.

## 6.7 Running of $\alpha$

The free electron is surrounded by a cloud of virtual particles, including  $e^-e^+$  pairs, due to quantum fluctuations in the vacuum. These provide shielding for the real electron. This is somewhat analogous to how a positive charge in an otherwise neutral medium induces a charge separation that surrounds it with negative charge, reducing its effective charge from the outside.

At large distances, which correspond to low energies, an electron is shielded. At smaller distances, and higher energies, the shielding is reduced. The result is that

<sup>3</sup>For details on the Stern–Gerlach experiment see the notes from principles of quantum mechanics

the effective coupling constant,  $\alpha$ , is not actually constant<sup>4</sup>. We say that the value of  $\alpha$  runs.

Consider the interaction in Equation (6.5.1). This is first order in  $\alpha$ , since there are two vertices, both contributing a factor of  $\sqrt{\alpha}$ . The momentum of the photon,  $q$ , is how we measure the energy of the interaction. When  $q^2 = 0$  we have  $\alpha = 1/137$ . When  $q^2 = (100 \text{ GeV})^2$  we measure  $\alpha = 1/128$ .

The value of  $\alpha$  increases with energy, until at some point  $\alpha$  is greater than one. At this point we have a problem as the perturbation series that QED calculations are full of no longer converge. Fortunately this doesn't happen at any reasonable energy, and it is predicted that gravitational effects will become important at lower energies that the series failing to converge so QED isn't even valid at these energies without modification anyway.

<sup>4</sup>the value  $e^2/(4\pi)$  is constant, it's just that at low energies this no longer effectively measures the strength of QED interactions



## Part IV

# Quantum Chromodynamics

# Seven

## Introduction to QCD

Quantum chromodynamics is the quantum theory of the strong force. The strong force is mediated by massless bosons, known as gluons,  $g$ . The strong force equivalent of charge is **colour**, which is a property possessed only by quarks and gluons.

### 7.1 Feynman Diagrams

#### 7.1.1 QCD Vertex

The QCD vertex is very similar to that of QED, except that instead of charged fermions we have quarks, and instead of photons we have gluons, which we represent with a loopy line. The basic QCD vertex involves a quark and a gluon, for example



(7.1.1)

#### 7.1.2 Feynman Rules

The Feynman rules for QCD diagrams are almost identical to the QED Feynman rules. In QCD the propagator is the gluon, and each propagator of momentum  $q$  leads to a factor of  $1/q^2$  in the matrix element. The coupling constant is  $\alpha_s$ , which in analogy with QED we define to be

$$\alpha_s = \frac{g_s^2}{4\pi} \approx 0.2 \quad (7.1.2)$$

where  $g_s$  is a constant which is somewhat analogous to the charge in QED.

Notice that  $\alpha_s \approx 0.2$  is significantly larger than the QED equivalent of  $\alpha \approx 0.007$ . This is because the strong force is stronger than electromagnetism.

#### 7.1.3 Examples

Since QCD involve quarks, which we only ever see in bound states (see ), Feynman diagrams tend to be slightly more complicated in QCD. For example, the  $\phi$  meson, which is a bound strange and antistrange quark,  $s\bar{s}$ , can decay in one of two ways:

$$\phi \rightarrow K^+ K^-, \quad \text{or} \quad \phi \rightarrow K^0 \bar{K}^0, \quad (7.1.3)$$

which in terms of quark content is

$$s\bar{s} \rightarrow (u\bar{s})(\bar{u}s), \quad \text{or} \quad s\bar{s} \rightarrow (d\bar{s})(\bar{d}s). \quad (7.1.4)$$

These two decays are represented by the following diagrams,

$$\varphi \left\{ \begin{array}{l} s \\ \bar{s} \end{array} \right\} \rightarrow \left\{ \begin{array}{l} s \\ \bar{u} \end{array} \right\} K^- + \left\{ \begin{array}{l} u \\ \bar{s} \end{array} \right\} K^+ \quad (7.1.5)$$

and

$$\varphi \left\{ \begin{array}{l} s \\ \bar{s} \end{array} \right\} \rightarrow \left\{ \begin{array}{l} s \\ \bar{d} \end{array} \right\} \bar{K} + \left\{ \begin{array}{l} d \\ \bar{s} \end{array} \right\} K^0 \quad (7.1.6)$$

Notice that in both the antistrange quark isn't actually partaking in the reaction. For this reason we may sometimes draw this diagram with just the strange quark, and the two (anti)up quarks, with the implicit second quark in the initial state, since we can't have a single quark on its own.

Another example would be the decay of a  $\Delta^0$  hadron, which has quark content  $udd$ , this is essentially an excited neutron with spin  $3/2$ . This decays to a proton, with quark content  $uud$ , and a negative pion, with quark content  $d\bar{u}$ . This is represented by the diagram

$$\Delta^0 \left\{ \begin{array}{l} u \\ d \\ d \end{array} \right\} \rightarrow \left\{ \begin{array}{l} u \\ d \end{array} \right\} p + \left\{ \begin{array}{l} u \\ \bar{d} \end{array} \right\} \pi^0 \quad (7.1.7)$$

## 7.2 Colour

**Colour** is the charge of QCD. Colour is a conserved quantum number, with three possible values, which we call red (**r**), green (**g**), and blue (**b**). Quarks carry colour, and antiquarks carry anticolour,  $\bar{r}$ ,  $\bar{g}$ , and  $\bar{b}$ . Leptons don't carry colour, and therefore

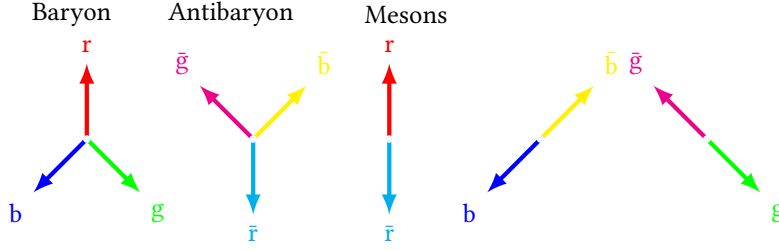


Figure 7.1: An intuitive, and completely non-physical, picture explaining why “colour-anticolour”, and “all three (anti)colours” are both considered colourless states.

don’t interact via the strong force. Colour is not like electric charge in that any coloured particle can have posses any colour, it’s more like spin in this nature, it’s also like spin in the fact that spin has nothing to do with the particle actually rotating, and colour has nothing to do with the particle having a colour in the everyday meaning.

A **hadron** is any particle made of quarks. They always form such that they are colourless. They can do this in one of two ways, firstly they can form **mesons**, which are quark-antiquark pairs, in this case the antiquark must have the anticolour of the quark’s colour, this gives three options for colours in mesons:

$$r\bar{r}, \quad g\bar{g}, \quad \text{and} \quad b\bar{b}. \quad (7.2.1)$$

Secondly, they can form **(anti)baryons**, which are made of three (anti)quarks, all with different colours, so for baryons the colours are  $rgb$ , and for antibaryons the colours are  $\bar{r}\bar{g}\bar{b}$ .

### 7.3 Gluons

The force carriers of the strong force are **gluons**. These are massless, electrically neutral, spin 1 bosons, in this regard they behave a lot like photons. However, gluons carry colour charge, this means that they also interact through the strong force. Gluons carry equal parts colour and anticolour. Naively one might expect that as there are 3 colours, and 3 anticouls, we would have 9 combinations:

$$\begin{array}{lll} r\bar{r}, & r\bar{g}, & r\bar{b}, \\ g\bar{r}, & g\bar{g}, & g\bar{b}, \\ b\bar{r}, & b\bar{g}, & b\bar{b}. \end{array} \quad (7.3.1)$$

However, three of these,  $r\bar{r}$ ,  $g\bar{g}$ , and  $b\bar{b}$ , are colourless. Instead we see that gluons have colour charge given by linear combinations of these states, for example

$$r\bar{g}, g\bar{r}, r\bar{b}, b\bar{r}, b\bar{g}, g\bar{b}, \frac{\sqrt{2}}{2}(r\bar{r} - g\bar{g}), \frac{\sqrt{6}}{6}(r\bar{r} + g\bar{g} - 2b\bar{b}). \quad (7.3.2)$$

We may also expect that we would find a colourless gluon like

$$\frac{\sqrt{3}}{3}(r\bar{r} + g\bar{g} + b\bar{b}) \quad (7.3.3)$$

but this isn’t the case. If this existed it would be a massless, colourless, chargeless, vector boson, and so would behave a lot like photon, in that it wouldn’t interact via

QCD or QED and would be free, this would give it an infinite range, which we do not observe with the strong force, so we can be certain that this doesn't exist.

From a more abstract point of view QCD is a gauge theory with  $SU(3)$  gauge symmetry. Gluons are just one representation of this Lie group, so what we mean when we take linear combinations of colours is linear combinations of this representation, which means linear combinations of elements of the general linear group.

There is not just one set of colours that gluons can have, instead we can think of them as a basis for the representation, viewing the general linear group as a vector space, this happens to be an 8-dimensional vector space for this particular representation, which in turn is due to the fact that there are 8 generators of the Lie algebra  $\mathfrak{su}(3)$ , hence there are 8 gluon colour states. Importantly the gluon colours we pick must be linearly independent, and must be such that no combination gives one of the forbidden singlet states, which are  $r\bar{r}$ ,  $g\bar{g}$ ,  $b\bar{b}$ , and  $(r\bar{r} + b\bar{b} + g\bar{g})/\sqrt{3}$ .

The requirement that these singlet states be excluded is not one that falls out of the theory, but something that we see in experiments. We therefore demand it in our theory. If we did see singlet states then we would instead work with the group  $U(3)$ , which has one more degree of freedom, and hence an extra state, which would be one of these singlet states, and the others can then be achieved as linear combinations of the 9 basis states.

In a Feynman diagram we represent gluons as loopy lines if we aren't being specific about colour, if we need to specify colours we can draw a diagram doing so, for example the quark-quark scattering process



may have the colour change as



Notice that the quark flavour doesn't change, just the colour. If we see the gluon as travelling upwards then it is a  $b\bar{r}$  gluon, whereas if we see it as travelling downwards then it is a  $r\bar{b}$  gluon.

### 7.3.1 Self-Interaction

Since gluons carry colour they can interact via the strong force. For example we can get a three gluon vertex:



We can also get four gluon vertices.

The existence of these gluon self-interactions makes QCD very different from QED, as we shall soon see.

## 7.4 Colour Confinement

In QED we have a potential given by

$$V_{\text{QED}} = -\frac{\alpha}{r}, \quad (7.4.1)$$

which is the Coulomb potential in units where  $\epsilon_0 = 1$ . At short distances in QCD we have a similar potential,

$$V_{\text{QCD}} = -\frac{4}{3} \frac{\alpha_s}{r}. \quad (7.4.2)$$

However, at larger distances self-interactions become more important and the potential actually increases linearly with distance, meaning that at large distances we have a spring like potential

$$V_{\text{QCD}} = kr. \quad (7.4.3)$$

This means that we need infinite energy to separate two quarks. We call this **colour confinement**. We can take the potential to be

$$V_{\text{QCD}} = -\frac{4}{3} \frac{\alpha_s}{r} + kr \quad (7.4.4)$$

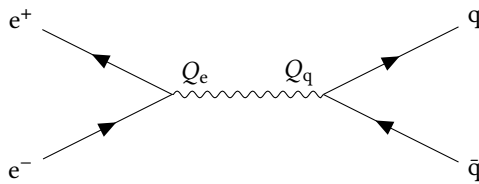
and calculate the force needed to separate quarks:

$$F_{\text{QCD}} = -\frac{dV_{\text{QCD}}}{dr} = \frac{4}{3} \frac{\alpha_s}{r} + k. \quad (7.4.5)$$

At large distances we have  $F_{\text{QCD}} \approx k$  and we have measured  $k \approx 1 \text{ GeV fm}^{-1} \approx 160\,000 \text{ N}$ .

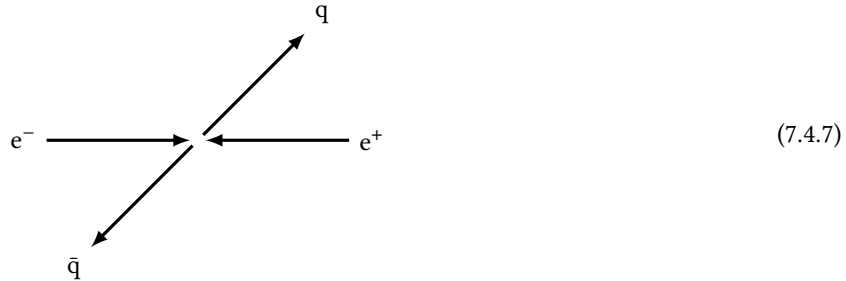
In reality what happens when we try to separate two quarks is eventually we put in enough energy that a quark-antiquark pair can be created. These will then pair up with the two quarks we were initially trying to separate and the potential energy decreases. Quark-antiquark pairs will keep being produced until there is no longer sufficient energy, and we will be left with more quarks, but they will not be separated. This process is referred to as hadronisation.

We can use this process to jets of hadrons. To do so we need to start with a high energy pair of quarks. One way to do this is to start with an electron-positron annihilation:



The diagram shows an electron-positron annihilation process. On the left, a positron ( $e^+$ ) and an electron ( $e^-$ ) are shown as incoming particles with arrows pointing towards the interaction vertex. They annihilate into a virtual photon, represented by a wavy line. This photon then splits into a quark ( $q$ ) and an antiquark ( $\bar{q}$ ) at a second vertex. The quark and antiquark are shown as outgoing particles with arrows pointing away from the vertex. The vertices are labeled  $Q_e$  and  $Q_q$  respectively. The entire process is labeled (7.4.6).

This interaction is drawn in time and space. In just space it looks more like



## 7.5 Running of $\alpha_s$

For QED we saw that the coupling constant increases with energy due to screening. In QCD we also have screening from spontaneous quark-antiquark pair production. We also have the effect of gluon pair production. This effect provides anti-screening. It can be shown that the anti-screening effect of the gluons beats the screening effect of the quarks and  $\alpha_s$  decreases with distance.

At low energies  $\alpha_s$  is large and we can't use perturbation theory as the series don't converge. At higher energies  $\alpha_s$  decreases, and eventually it is effectively zero. We call this **asymptotic freedom**, at sufficiently high energies QCD effects become negligible.

It just so happens that at about 1 fm we have  $\alpha_s \approx 1$ , which means that on the scales of nuclear physics QCD is non-perturbative, and calculations in QCD are very difficult. This is why we need high energies, such as those produced in colliders, to study QCD.

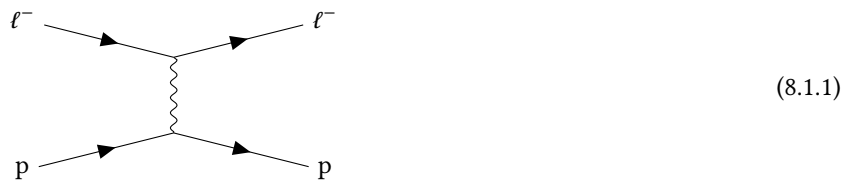
# Eight

## Tests of QCD

### 8.1 Partons

In the 1960s many were uncomfortable with the idea of quarks that couldn't be individually observed. Feynman proposed a model that served as a stepping stone to the full theory of QCD with quarks and gluons, his idea was that hadrons are made of point like constituent particles, but he made no claim as to what they were. He called these particles **partons**.

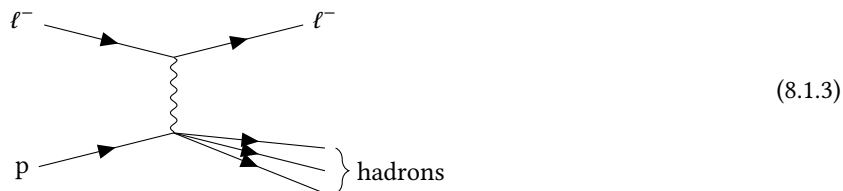
Consider a lepton scattering off of a proton, as in the diagram



If the particles involved in the scattering are point-like then the cross section is given by the Rutherford scattering formula:

$$\frac{d\sigma}{d\Omega} = \frac{e^4}{4E^2} \frac{\cos^2(\vartheta/2)}{\sin^4(\vartheta/2)}. \quad (8.1.2)$$

We can therefore use deviations from this result to probe the inner structure of the proton. One example being inelastic scattering, such as



Experimentally we find that the cross section decreases with the momentum transferred, and that the momentum transferred is not constant. This is evidence that the proton is not point-like, but has some internal structure. Inelastic scattering is particularly useful for this since the result doesn't depend very strongly on the momentum transferred. Therefore we can consider very high energy scattering, which we call **deep inelastic scattering** and in this regime we see point like scattering,



but not off of the proton, off of its point-like constituents, which were thought of as partons in the early experiments.

At relatively low energies the lepton scatters off of the valence quarks, that is the uud quarks that make up the proton. At higher energies, which correspond to shorter distances, scattering also occurs from the sea of virtual quarks and gluons that are produced in loops.

We have now performed lepton-proton scattering at wide range of energies and the results are consistent with predictions from QCD. In particular we can measure the parton density function, which tells us the probability of finding a parton with given kinematics.

## 8.2 Evidence for Colour

### 8.2.1 $\Delta^{++}$ Baryon

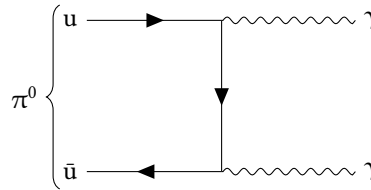
The  $\Delta^{++}$  baryon is made of three up quarks, uuu, and has spin 3/2, so all three quarks have the same spin. The wave function for  $\Delta^{++}$  is symmetric under interchange of identical quarks, since all quarks are the same flavour and spin, which means that the flavour,  $\psi_{\text{flavour}}$ , and spin,  $\psi_{\text{spin}}$ , factors in the wave function are symmetric. However, since  $\Delta^{++}$  is a fermion it must have an antisymmetric wave function. This suggests that we are missing a term in the wave function. Indeed, we can write the wave function as

$$\psi = \psi_{\text{space}} \psi_{\text{flavour}} \psi_{\text{spin}} \psi_{\text{colour}} \quad (8.2.1)$$

where  $\psi_{\text{space}}$ , the spatial component of the wave function, is symmetric since the ground state of the  $\Delta^{++}$  baryon has no orbital angular momentum, and so must be spatially symmetric. The solution is that the  $\psi_{\text{colour}}$  factor is antisymmetric, which is to say that the quarks all have different colour.

### 8.2.2 $\pi^0 \rightarrow \gamma\gamma$

Consider the interaction  $\pi^0 \rightarrow \gamma\gamma$ , recall that the  $\pi^0$  is made of an up-antiup pair,  $u\bar{u}$ . This decay is then



$$\pi^0 \left\{ \begin{array}{l} u \longrightarrow \gamma \\ \bar{u} \longleftarrow \gamma \end{array} \right. \quad (8.2.2)$$

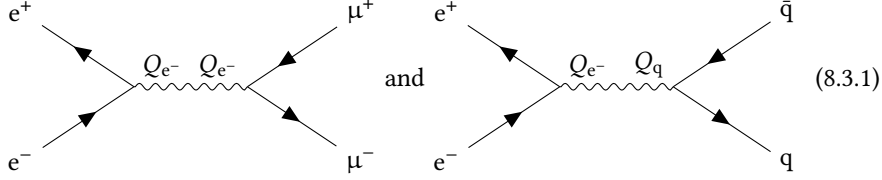
The width of this interaction is proportional to the number of colours,  $N_c$ , squared. This comes from considering all possible colours of the two quarks. Measurements of  $\Gamma(\pi^0 \rightarrow \gamma\gamma) \propto N_c^2$  gives a value of

$$N_c = 2.99 \pm 0.12. \quad (8.2.3)$$

Which is in excellent agreement with the expected value of  $N_c = 3$ .

### 8.3 The R-Ratio

Consider the two interactions



For the first interaction  $\mathcal{M} \propto \alpha/q^2$ . For the second interaction  $\mathcal{M} \propto \alpha Q_q/q^2$ , where  $Q_q = 2/3, -1/3$ . We then define the quantity

$$R := \frac{\sigma(e^-e^+ \rightarrow \text{hadrons})}{\sigma(e^-e^+ \rightarrow \mu^-\mu^+)} = N_c \sum_i Q_i^2 \quad (8.3.2)$$

where  $N_c$  is the number of colours and  $Q_i$  is the energy of the quarks. Importantly we sum over only the energetically accessible quarks. For an experiment with  $\sqrt{s} \approx 1$  GeV the accessible quarks are u, d, and s. Therefore

$$R = 3 \left( \frac{4}{9} + \frac{1}{9} + \frac{1}{9} \right) = 2. \quad (8.3.3)$$

For  $\sqrt{s} \approx 4$  GeV the charm quark is also accessible and so

$$R = 3 \left( \frac{4}{9} + \frac{1}{9} + \frac{1}{9} + \frac{4}{9} \right) = \frac{10}{3}. \quad (8.3.4)$$

For  $\sqrt{s} \approx 10$  GeV the bottom quark becomes available and

$$R = 3 \left( \frac{4}{9} + \frac{1}{9} + \frac{1}{9} + \frac{4}{9} + \frac{1}{9} \right) = \frac{11}{3}. \quad (8.3.5)$$

Finally at  $\sqrt{s} \approx 350$  GeV the top quark becomes available and

$$R = 3 \left( \frac{4}{9} + \frac{1}{9} + \frac{1}{9} + \frac{4}{9} + \frac{1}{9} + \frac{4}{9} \right) = 5. \quad (8.3.6)$$

We therefore expect that measurements of  $R$  should have discontinuous jumps at energies where new quarks become accessible. Further the size of the jumps corresponds to the number of colours and also the charge of the quark that has just become available. Indeed, this is exactly what we observe. However, the results we observe at lower energies are actually higher than we would predict. We will see the reason for this in [Section 8.5](#).

### 8.4 Evidence for Gluons

High energy quarks are likely to emit gluons. Like high energy quarks high energy gluons will hadronise and form jets. In an interaction producing two quarks we will therefore occasionally see three jets instead of two, if the quarks emit a gluon which creates the third jet. The resulting jet pattern is called the **Mercedes configuration**, since the three pronged jets resemble the car manufactures' logo.

These jets are observed in particle colliders fairly often and are some of the earliest evidence that gluons exist.

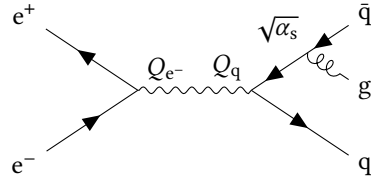
Further, by studying the angles between the highest energy jet and the two other jets it is possible to deduce that the spin of the gluon is 1.

### 8.5 Running of $\alpha_s$

QCD predicts that  $\alpha_s$  should decrease with energy. We can measure  $\alpha_s$  with the  $R$ -ratio of [Section 8.3](#). Recall that this quantity is defined as

$$R := \frac{\sigma(e^-e^+ \rightarrow \text{hadrons})}{\sigma(e^-e^+ \rightarrow \mu^-\mu^+)}. \quad (8.5.1)$$

In the case where we create hadrons we have to account for these hadrons producing a gluon jet in a three jet event like in [Section 8.4](#).



$$(8.5.2)$$

If we account for these then we find that

$$R = N_c \sum_i Q_i^2 \left( 1 + \frac{\alpha_s}{\pi} \right). \quad (8.5.3)$$

The extra factor in the sum is

$$1 + \frac{\alpha_s}{\pi} \approx \frac{3.85}{3.66} \approx 1.05 \quad (8.5.4)$$

at 25 GeV.

We see this in the data at lower energies. We find that  $R$  is larger than the theoretical prediction from first order diagrams. At higher energies  $\alpha_s$  decreases and the data comes in line with the theory as the second order diagrams become less important.

Another useful ratio is the ratio of cross sections for 3-jet events to 2-jet events

$$R_3 := \frac{\sigma(e^-e^+ \rightarrow 3\text{-jets})}{\sigma(e^-e^+ \rightarrow 2\text{-jets})} \propto \alpha_s. \quad (8.5.5)$$

This also decreases with energy, as expected.

# Nine

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## Isospin and Hadron Spectroscopy

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In the 1950s and 1960s many new particle states were discovered. We now understand these to be simply combinations of fundamental particles, such as quarks. But it is still advantageous to study these particle states with the concepts that were discovered at the same time. Two of these concepts being parity and isospin.

### 9.1 Parity

The parity operator,  $P$ , has the action of replacing  $\mathbf{r}$  with  $-\mathbf{r}$  in the wave function:

$$P|\psi(\mathbf{r}, t)\rangle = |\psi(-\mathbf{r}, t)\rangle. \quad (9.1.1)$$

The **parity** of a state is the eigenvalue associated with this operator. The parity of a proton is 1 and the parity of a neutron is  $-1$ . An antifermion will have the opposite parity to its fermion counterpart. Bosons and antibosons will have the same parity.

Ground states have an intrinsic parity, which is the parity mentioned above for the proton and neutron. For states with angular momentum,  $\ell$ , there is also an additional term  $(-1)^\ell$ .

We classify particles into one of four types based on their angular momentum,  $J$ , and parity,  $P$ , which we denote  $\pm$  for  $\pm 1$ .

**Scalar**  $J^P = 0^+$ ,

**Pseudoscalar**  $J^P = 0^-$ ,

**Vector**  $J^P = 1^-$ , and

**Pseudovector**  $J^P = 1^+$ .

A meson will have overall parity  $P_q P_{\bar{q}} (-1)^\ell$ , where  $P_q$  and  $P_{\bar{q}}$  are the parities of the quark and antiquark and  $\ell$  is the angular momentum.

Parity is conserved in the strong and electromagnetic interactions, but not in the weak interaction.

### 9.2 Isospin

The concept of isospin was first introduced by Heisenberg to explain why protons and neutrons are interchangeable in strong and weak interactions. It was later expanded to other concepts. Since it was proposed before we knew of quarks it has a

somewhat clunky relation to more modern concepts. We define the **isospin**,  $|I, I_3\rangle$ , of a particle to be such that

$$I_3 := \frac{1}{2}(N_u - N_d) \quad (9.2.1)$$

and  $I$  is the maximum value of  $I_3$  in the same multiplet. Recall that  $N_q$  is the number of quarks of type  $q$  and that antiquarks of type  $\bar{q}$  count as negative. Multiplets themselves are an idea that is beyond the scope of the course but we can think of them as ways of grouping particles. For example the pions,  $\pi^+$ ,  $\pi^0$ , and  $\pi^-$ , form a triplet:

$$\pi^+(\bar{u}d) : |I, I_3\rangle = |11\rangle, \quad \pi^0(u\bar{u}) : |I, I_3\rangle = |10\rangle, \quad (9.2.2)$$

$$\text{and} \quad \pi^-(\bar{u}d) : |I, I_3\rangle = |1, -1\rangle. \quad (9.2.3)$$

We also have a duplet of the proton and neutron:

$$p(uud) : |I, I_3\rangle = \left| \frac{1}{2}, \frac{1}{2} \right\rangle, \quad \text{and} \quad n(udd) : |I, I_3\rangle = \left| \frac{1}{2}, -\frac{1}{2} \right\rangle. \quad (9.2.4)$$

We consider a quadruplet of delta baryons:

$$\Delta^{++}(uuu) : |I, I_3\rangle = \left| \frac{3}{2}, \frac{3}{2} \right\rangle, \quad \Delta^+(uud) : |I, I_3\rangle = \left| \frac{3}{2}, \frac{1}{2} \right\rangle, \quad (9.2.5)$$

$$\Delta^0(udd) : |I, I_3\rangle = \left| \frac{3}{2}, -\frac{1}{2} \right\rangle, \quad \Delta^-(ddd) : |I, I_3\rangle = \left| \frac{3}{2}, -\frac{3}{2} \right\rangle. \quad (9.2.6)$$

We can define the **hypercharge** as  $Y := \mathcal{B} + S$  where  $\mathcal{B}$  is the baryon number and  $S = N_s - N_{\bar{s}}$  is the strangeness quantum number. It then turns out that  $I_3 = Q - Y/2$  where  $Q$  is the charge.

The concept of isospin is useful because it allows us to predict properties. For example the particle  $\Sigma^+(1189)$  particle, which we now know has quark content  $uus$ , was first observed in the strong interaction

$$K^- + p \rightarrow \pi^- + \Sigma^+(1189). \quad (9.2.7)$$

This particle decays by one of two modes

$$\Sigma^+(1189) \rightarrow \pi^+ + n, \pi^0 + p. \quad (9.2.8)$$

Since the strong interaction conserves strangeness and baryon number, and on the left in the interaction we have  $\mathcal{B} = 1$  and  $S = -1$ , it follows that  $Y = 0$  and  $I_3 = Q = 1$ . Since  $I_3 \neq 0$  we can predict the existence of other particles in the same isospin multiplet with  $I_3 = Q = 0, -1$ .

One of the early victories of particle physics was finding a way to group particles such that we can predict the existence of missing particles. For example one way of arranging hadrons in a way consistent with the underlying symmetries is given in [Figures 9.1](#) and [9.2](#). The grouping in [Figure 9.2](#) allowed Murray Gell-Mann to predict the existence of the  $\Omega^-$  baryon before it was found experimentally.

## 9.3 Charmonium

### 9.3.1 The November Revolution

In the early 1970s the idea of quarks was starting to be accepted, and the notion of colour and QCD was starting to be developed, although it was far from being

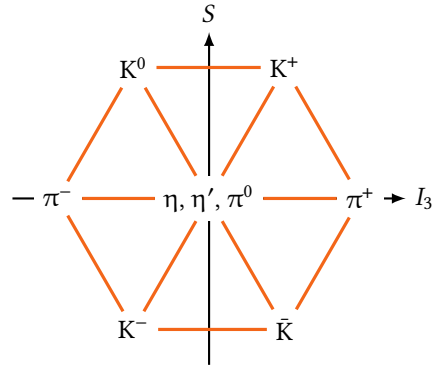
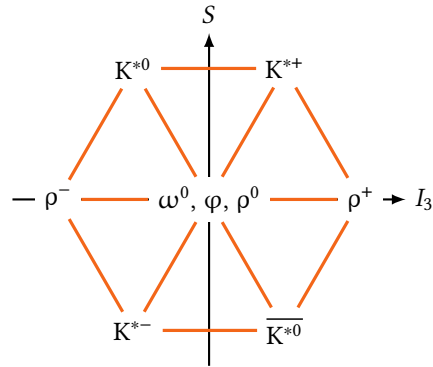
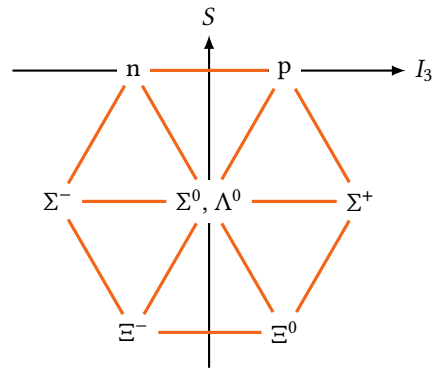
(a) The pseudoscalar ( $J^P = 0^-$ ) nonet.(b) The vector ( $J^P = 1^-$ ) nonet.(c) The  $J^P = \frac{1}{2}^+$  octet.

Figure 9.1: Grouping hadrons.

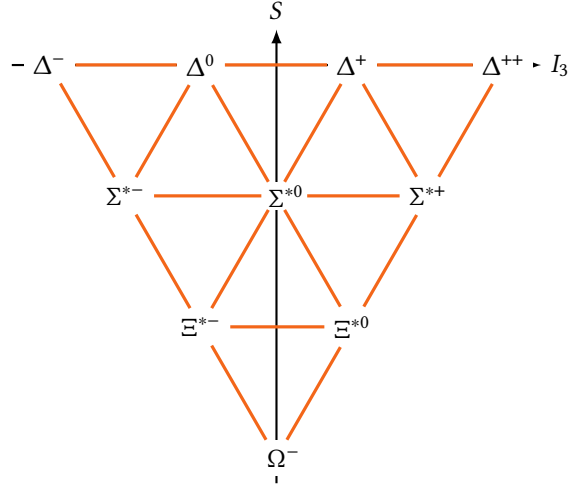


Figure 9.2: More hadron grouping, the  $J^P = \frac{3}{2}^+$  decuplet.

accepted as the theory of the strong force. In 1974 the Mark I detector at SLAC the Stanford's positron electron accelerating ring collider (SPEAR), while studying the  $R$  ratio, of [Section 8.3](#), in the previously unexplored region around  $\sqrt{s} = 3$  GeV a signal was discovered. First discovered in a broad scan they decided to go back and look more carefully. They found that the signal appeared at a beam energy of 1.56 GeV, and vanished when the beam energy was reduced to 1.55 GeV, reappearing at 1.555 GeV. This very narrow resonance corresponds to a new meson.

### 9.3.2 The $J/\psi$

The particle that they had observed is now known as the  $J/\psi$ , the reason for the odd name being that the same particle was discovered at Brookhaven around the same time and both experiments gave it a different name, Brookhaven going for  $J$ , and SLAC for  $\psi$ .

We now know that the  $J/\psi$  is a bound charmonium state, that is a charm-anticharm,  $c\bar{c}$ . It has an angular momentum of 1, since it is produced from the decay of a photon. Our current best measurement of its mass is 3.0969 GeV. The observed width, of approximately 3 MeV, is mostly due to limited experimental resolution. The resonance is actually extremely narrow and is predicted to have width  $\Gamma_{J/\psi} \approx 93$  keV.

The  $J/\psi$  can decay either to hadrons or to leptons, with branching fractions

$$\text{BF}(J/\psi \rightarrow \text{hadrons}) = 0.88, \quad (9.3.1)$$

$$\text{BF}(J/\psi \rightarrow e^-e^+) \approx \text{BF}(J/\psi \rightarrow \mu^-\mu^+) \approx 0.06. \quad (9.3.2)$$

Note that  $2m_\tau = 3.55$  GeV  $>$   $m_{J/\psi}$ , so  $J/\psi \rightarrow \tau^-\tau^+$  cannot occur.

While  $J/\psi$  has an angular momentum of 1, due to its method of production, there are other bound  $c\bar{c}$  states with different quantum numbers. We find these at slightly different masses in a range from approximately 3 GeV to 4.5 GeV.

The  $J/\psi$  is very narrow, the reason for this is it would “like” to decay to  $D^+D^-$ ,

which are mesons consisting of  $c\bar{d}$  and  $d\bar{c}$  respectively,

$$J/\psi \left\{ \begin{array}{l} c \longrightarrow c \\ \bar{c} \longrightarrow \bar{c} \end{array} \right. \begin{array}{l} \longrightarrow \bar{d} \\ \longrightarrow d \end{array} \left. \begin{array}{l} D^+ \\ D^- \end{array} \right. \quad (9.3.3)$$

however, this is kinematically forbidden as

$$2m_D = 3.739 \text{ GeV} > m_{J/\psi}. \quad (9.3.4)$$

Above a certain energy, known as the open charm threshold, this decay becomes accessible, for example, the  $\psi(3770)$  state can decay this way (3370 here being the mass in MeV). This state has a much wider resonance.

Instead of this decay the  $J/\psi$  must decay to lighter mesons, for example an allowed decay is

$$J/\psi \left\{ \begin{array}{l} c \\ \bar{c} \end{array} \right. \rightarrow \left\{ \begin{array}{l} u \\ \bar{d} \end{array} \right\} \pi^+ + \left\{ \begin{array}{l} d \\ \bar{d} \end{array} \right\} \pi^0 + \left\{ \begin{array}{l} d \\ u \end{array} \right\} \pi^- \quad (9.3.5)$$

This is suppressed by the **Zweig rule**, which states that decays in which the Feynman diagram separates into two disconnected components upon removing the gluon lines such that the initial particles are disconnected from the final particles are suppressed. We need three gluons to conserve colour, and so  $\mathcal{M} \propto \alpha_s^3$ . Further the gluons have to be massive, as they need to carry away the approximately 3 GeV, of mass from the quarks. This further suppresses this decay as  $\alpha_s$  runs, and therefore decreases. The combined effect of this suppression is a very narrow resonance.

## 9.4 Other QCD States

Since the 1970s when charmonium was first observed we have also found bottomonium states, which are bound  $b\bar{b}$  states. Measurements of these states agree well with predictions of QCD.

More interestingly we have also observed states with more than 3 quarks. This includes tetraquarks, which are made of two quarks and two antiquarks, essentially forming bound meson-meson pairs, and pentaquarks, which are made of four quarks and an antiquark, essentially forming a bound meson-baryon pair. The existence of hexaquarks, made from either 6 quarks in a bound dibaryon state, or 3 quarks and 3 antiquarks, in a bound 3 meson state, has also been hypothesised.

Another prediction of QCD is the existence of glueballs, which are bound states consisting solely of gluons, with no quarks. It's predicted that such states could be detected at current collider energies but identifying these states would be incredibly difficult. There are some candidate particles that it is thought might be glueballs but this has not yet been confirmed.



We are still actively finding new hadronic resonances regularly. A recent example being  $T_{cc}^+$ , with quark content  $cc\bar{u}\bar{d}$ , which decays to  $D^0 D^0 \pi^+$ , where  $D^0$  has quark content  $c\bar{u}$ . It is not yet known whether  $T_{cc}^+$  is a tetra quark or a bound  $D^0 D^{*+}$  pair of mesons,  $D^{*+}$  has quark content  $cd$ .

Part V

## The Weak Force

## The Weak Interaction

### 10.1 The Basics

The weak interaction is responsible for

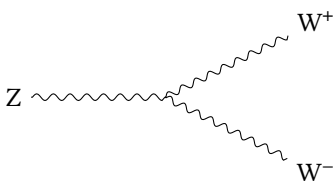
- muon and tau decay,
- light hadron decay,
- nuclear fission and fusion,
- neutrino interactions.

The weak force is, as the name suggests, weak. Cross sections for weak processes are small, on the order of  $10^{-13}$  mb, and lifetimes are relatively long, on the order of  $10^{-13}$  s– $10^3$  s.

There are two types of weak interactions. Neutral current processes are mediated by Z bosons, which have mass  $m_Z = 91.2$  GeV. Charged current processes are mediated by W bosons, which have charge  $\pm 1$ , and mass  $m_W = 80.4$  GeV.

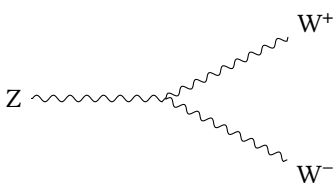
The Z boson interacts with all quarks and all leptons, including neutrinos. It doesn't change fermion flavour. The W boson interacts with all quarks and leptons, including neutrinos, and it can change fermion flavour. In general in a process involving a photon we can replace the photon with a Z boson, although not vice-versa as photons don't interact with neutrinos.

Z and W can also interact with each other, for example,



$$\text{Z} \rightsquigarrow \text{W}^+ \text{W}^- \quad (10.1.1)$$

another way of creating two W bosons is through a photon decaying:



$$\text{Z} \rightsquigarrow \text{W}^+ \text{W}^- \quad (10.1.2)$$

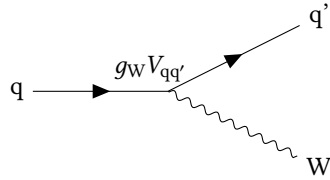
## 10.2 Feynman Rules

The Feynman rules for the weak interaction are more complicated than for QED or QCD since there are multiple massive bosons. The coupling strength depends on the boson, with each weak vertex contributing  $g_W \propto \sqrt{\alpha_W}$  for charged weak currents, or  $g_Z \propto \sqrt{\alpha_Z}$  for neutral weak currents. Each propagator then contributes

$$\frac{1}{q^2 - m_W^2}, \quad \text{or} \quad \frac{1}{q^2 - m_Z^2} \quad (10.2.1)$$

for charged and neutral weak currents respectively. Compare this to the propagator term in QED and QCD of  $1/q^2$ , since the exchange particles are massless in these theories.

For charged interactions involving quarks there is one more complexity, the coupling strength depends on the quarks. We will see this more later, for now in the interaction

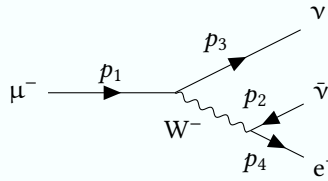


$$(10.2.2)$$

we simply include an extra term  $V_{qq'}$  which depends on the types of the two quarks.

Charged weak currents conserve lepton number, but not quark flavour number. On the other hand neutral weak currents conserve all flavour quantum numbers. Strictly the coupling does depend on the fermion flavour but we won't consider this in this course.

■ **Example 10.2.3** Consider the decay of a muon in the process  $\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$ . The Feynman diagram for this is



$$(10.2.4)$$

The four-momentum transferred is  $q = p_3 - p_1 = p_4 - p_2$ , so the propagator term is  $1/(q^2 - m_W^2)$ , and there are two vertex terms, each contributing a factor of  $g_W$ , for an overall contribution of

$$\mathcal{M} \propto \frac{g_W^2}{q^2 - m_W^2}. \quad (10.2.5)$$

## 10.3 Fermi Theory

At low energies we the momentum transferred is such that  $q^2 \ll m_W^2$ . The propagator term is then approximately constant and equal to  $1/m_W^2$ . We drop the negative

term since we are only interested in proportionality. Low energies includes most processes in nuclear physics, and muon decay.

In this low energy regime we can treat the weak interaction with **Fermi theory**. Fermi originally formulated his theory using the constant

$$G_F = \frac{g_W^2 \sqrt{2}}{8m_W^2} = 1.166\,37(1) \times 10^{-5} \text{ GeV}^{-2}. \quad (10.3.1)$$

Since the  $W$  and  $Z$  bosons are massive this results in a short range force, in particular the range is approximately

$$R_{\text{weak}} = \frac{\hbar}{m_W c} \approx 2 \times 10^{-18} \text{ m} = 2 \times 10^{-3} \text{ fm}. \quad (10.3.2)$$

This incredibly short range is why the weak force seems so weak, the coupling constant actually has a value of about  $g_W \approx 0.66$ , meaning  $\alpha_W = g_W^2/(4\pi) \approx 1/29$ , compared to QED's much smaller  $\alpha \approx 1/137$ , but QED has an infinite range.

### 10.3.1 Muon Decay

The dominant decay mode for a muon is  $\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$ . Electromagnetic decay,  $\mu^- \rightarrow e^- \gamma$  is forbidden as it violates conservation of lepton flavour numbers. The muon is relatively light,  $m_\mu = 105.7 \text{ MeV}$ , and so there aren't really any other decay modes at reasonable energies. See [Example 10.2.3](#) for a calculation of the matrix element. We find that

$$\Gamma(\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu) \propto |\mathcal{M}|^2 \propto \frac{g_W^4}{(q^2 - m_W^2)^2} \approx \frac{g_W^4}{m_W^4} \propto G_F^2. \quad (10.3.3)$$

Since there aren't other viable decay modes this is approximately equal to the full decay width,  $\Gamma_\mu$ , and so we can use the muon lifetime to measure the Fermi constant,  $G_F$ .

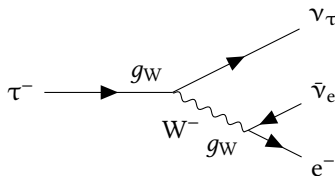
Since  $G_F$  has dimensions of  $[\text{energy}]^{-2}$  and  $\Gamma_\mu$  has dimensions of  $[\text{energy}]$  we posit that  $\Gamma_\mu \approx k G_F^2 m_\mu^5$  with  $k$  being some dimensionless constant. A full calculation gives

$$\Gamma_\mu = \frac{G_F^2 m_\mu^5}{192\pi^3}. \quad (10.3.4)$$

Measurements give  $\tau_\mu = \hbar/\Gamma_\mu = 2.197\,034(21) \times 10^{-6} \text{ s}$ , which gives the value of  $G_F$  from earlier.

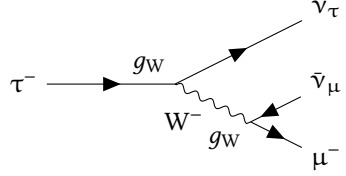
### 10.4 Tau Decay

Tau decay is more complicated than muon decay as the tauon is heavy enough that it has other decay modes. In particular it can decay into  $e^- \bar{\nu}_e \nu_\tau$ ,  $\mu^- \bar{\nu}_\mu \nu_\tau$ , or into hadrons and  $\nu_\tau$ . The first of these is given by



$$\mathcal{M} \propto \frac{g_W^2}{q^2 - m_W^2} \quad (10.4.1)$$

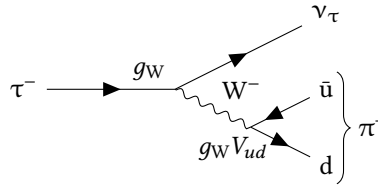
the second decay mode is



The diagram shows a tau lepton line entering from the left, meeting a vertex with a  $W^-$  boson. The  $W^-$  boson then splits into a tau neutrino and a muon. The vertices are labeled with  $g_W$ . The equation to the right is:

$$\mathcal{M} \propto \frac{g_W^2}{q^2 - m_W^2} \quad (10.4.2)$$

and an example of the third decay mode is



The diagram shows a tau lepton line entering from the left, meeting a vertex with a  $W^-$  boson. The  $W^-$  boson then splits into an anti-up quark and a down quark. The vertices are labeled with  $g_W$ . The quark pair is grouped in a bracket and labeled  $\pi^-$ . The equation to the right is:

$$\mathcal{M} \propto \frac{g_W^2}{q^2 - m_W^2} \quad (10.4.3)$$

The branching fractions for these are

$$\text{BF}(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) = 0.178, \quad (10.4.4)$$

$$\text{BF}(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) = 0.174, \quad (10.4.5)$$

$$\text{BF}(\tau^- \rightarrow \text{hadrons} + \nu_\tau) = 0.647. \quad (10.4.6)$$

Notice that the branching fraction to decay to hadrons is approximately three times the branching fraction to decay to each type of lepton. This is because there are three colours and  $V_{ud}$  is sufficiently close to 1 that the matrix elements are pretty much the same for all three decay modes.

Recall that  $\Gamma_\mu = k G_F m_\mu^5$ . By the same argument  $\Gamma_{\tau^-} = 0.178 k G_F m_\tau^5$ , where the extra factor of 0.178 is simply the branching ratio. We therefore find that

$$\frac{\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)}{\Gamma(\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu)} = \frac{0.178 \Gamma_\tau}{\Gamma_\mu} = \frac{m_\tau^5}{m_\mu^5}, \quad (10.4.7)$$

so we can measure the ratio of the masses of the muon and tau from the ratio of the decay widths of their decay to an electron.

Similar logic leads to the prediction that

$$\tau_\tau = \text{BF}(\tau \rightarrow e^- \bar{\nu}_e \nu_\tau) \tau_\mu \frac{m_\tau^5}{m_\mu^5} = 2.91 \times 10^{-13} \text{ s}. \quad (10.4.8)$$

Experimental measurements agree remarkably well with this, giving  $\tau_{\tau^-} = (2.906 \pm 0.110) \times 10^{-13} \text{ s}$ . This is a demonstration of lepton universality, where the only noticeable difference between leptons is their mass, which in this case is accounted for when we include the ratio  $m_\tau^5/m_\mu^5$ , and account for the extra decay modes of  $\tau^-$ .

## 10.5 Weak Interaction of Quarks

A charged current interaction can change the flavour of quarks, in particular charge conservation means that an up-like quark can produce a  $W^+$  and a down-like quark, or a down-like quark can produce a  $W^-$  and an up-like quark. While this violates conservation of quark flavour the total number of quarks is still conserved.

The coupling for these interactions depends on the flavours of quarks involved. The coupling constant for  $q \rightarrow Wq'$  is  $V_{qq'}g_W$ . Here  $V_{qq'}$  are the elements of a matrix known as the **Cabibbo–Kobayashi–Maskawa** matrix, or CKMatrix. This relates the possible flavour changes by

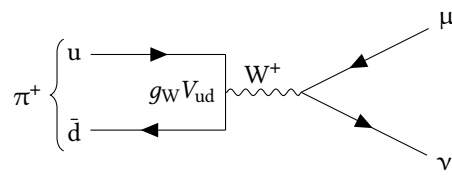
$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix} \quad (10.5.1)$$

Our best experimental results give

$$\begin{array}{lll} V_{ud} = 0.97427 \pm 0.00015 & V_{us} = 0.22534 \pm 0.00065 & V_{ub} = 0.00351^{+0.00015}_{-0.00014} \\ V_{cd} = 0.22520 \pm 0.00065 & V_{cs} = 0.97344 \pm 0.00016 & V_{cb} = 0.0412^{+0.0011}_{-0.0005} \\ V_{td} = 0.00867^{+0.00029}_{-0.00031} & V_{ts} = 0.0404^{+0.0011}_{-0.0005} & V_{tb} = 0.999146^{+0.000021}_{-0.000046} \end{array}$$

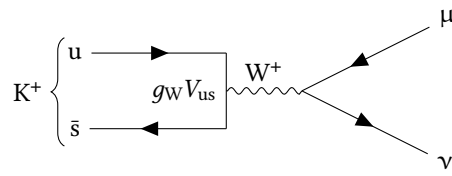
Notice that within a generation  $V_{qq'} \approx 1$ , for example  $V_{ud} = 0.974$ . Interactions crossing all generations, so involving generation one and three quarks, are the most suppressed, for example  $V_{ub} = 0.00351$  and  $V_{td} = 0.00867$ .

One example of the effect of this is that the decay  $\pi^+(u\bar{d}) \rightarrow \mu^+\nu_\mu$ , which proceeds via creation of a  $W^+$ :



$$\pi^+ \left\{ \begin{array}{l} u \\ \bar{d} \end{array} \right\} \xrightarrow{g_W V_{ud}} W^+ \rightarrow \mu^+ \nu_\mu \quad (10.5.2)$$

Another very similar decay is that of the  $K^+(u\bar{s}) \rightarrow \mu^+\nu_\mu$  via the same mechanism:



$$K^+ \left\{ \begin{array}{l} u \\ \bar{s} \end{array} \right\} \xrightarrow{g_W V_{us}} W^+ \rightarrow \mu^+ \nu_\mu \quad (10.5.3)$$

Both of these decays are observed in nature, however, the decay of  $K^+ \rightarrow \mu^+\nu_\mu$  occurs approximately 0.005 times as often as the  $\pi^+ \rightarrow \mu^+\nu_\mu$ . This is because  $V_{us} \approx 0.23 \approx 0.005 \cdot 0.97 \approx V_{ud}$ . We say that the quark mixing causes this decay mode to be suppressed for the kaon.

If we consider only the first two generations of quarks then the mixing can be given by the matrix

$$\begin{pmatrix} d' \\ s' \end{pmatrix} = \begin{pmatrix} \cos \vartheta_c & \sin \vartheta_c \\ -\sin \vartheta_c & \cos \vartheta_c \end{pmatrix} \begin{pmatrix} d \\ s \end{pmatrix}. \quad (10.5.4)$$

Here we have a single parameter,  $\vartheta_c$ , known as the **Cabbibo angle**. It's value is  $\vartheta_c \approx 0.2272$ .

## 10.6 Parity

Recall that the parity operator is defined to have the action  $\mathbf{r} \rightarrow -\mathbf{r}$ . This means that  $\mathbf{p} \rightarrow -\mathbf{p}$ . However, as pseudo vectors, the angular momentum,

$$\mathbf{L} = \mathbf{r} \times \mathbf{p} \rightarrow (-\mathbf{r}) \times (-\mathbf{p}) = \mathbf{L}, \quad (10.6.1)$$

doesn't change under the parity operator, and neither does the magnetic moment,  $\mu$ , for similar reasons. Parity is conserved by the electromagnetic and strong forces. It is not conserved by the weak force.

The first test of parity conservation for the weak force involved an unstable isotope,  $^{60}\text{Co}$ , which decays to the stable isotope  $^{60}\text{Ni}$  by beta decay:



In the experiment  $^{60}\text{Co}$  was cooled to 0.01 K and a magnetic field was applied such that the atomic spins aligned. The decay rate was then studied with respect to the field direction. Note that  $^{60}\text{Co}$  has total angular momentum  $J = 5$ , whereas  $^{60}\text{Ni}$  has total angular momentum  $J = 4$ , with the electron and neutrino each carrying away spin 1/2. We would expect that the electrons are emitted in all directions with uniform probability. What we find is that the angular distribution of electrons is not symmetric. If the direction of the magnetic field is reversed then the favoured direction for electron emission also reverses. It seems like electrons "prefer" to be emitted in the opposite direction to the atomic spin.

We find that not only is parity not conserved but that the violation of parity is as large as it could possibly be. We say that parity is maximally violated.



# Eleven

## Neutrinos

Neutrinos are nearly massless, spin half, neutral, leptons. There are three flavours,  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ . There are similarly three antineutrinos,  $\bar{\nu}_e$ ,  $\bar{\nu}_\mu$ , and  $\bar{\nu}_\tau$ . These carry the associated lepton (flavour) number of +1 for the neutrinos and -1 for the antineutrino.

As neutral leptons the only force (other than gravity) through which neutrinos interact is the weak force. This means that neutrinos hardly ever interact, the only reason we can detect them is because there are so many.

In one second passing through an area the size of your hand there are approximately

- $5 \times 10^{12}$  neutrinos from the sun,
- 20 neutrinos from cosmic rays in the atmosphere,
- $10^9$  neutrinos from the Earth,
- $10^{12}$  neutrinos from supernovae, and
- $2 \times 10^{12}$  neutrinos from the big bang.

The general method for detecting neutrinos involves setting up huge detectors in the hopes that the size and ubiquity of neutrinos allows us to detect some.

The existence of neutrinos was first posited based on the energy spectrum of electrons produced in beta decay. In a two-body interaction decay, which is what people thought beta decay was at the time, the energy of the particles should be the same every time. This is not what was observed so Pauli posited that beta decay must be a three-body interaction, with some, as yet, undetected particle. We now know this to be the neutrino.

### 11.1 Neutrino Interactions

Neutrinos are produced in beta decay processes, both  $\beta^-$  and  $\beta^+$ . Rotating the Feynman diagrams for these interactions we get the inverse beta decay processes:

$$n + \nu_e \rightarrow p + e^-, \quad \text{and} \quad p + \bar{\nu}_e \rightarrow n + e^+. \quad (11.1.1)$$

These processes are incredibly unlikely, for example the cross section for the second reaction is

$$\sigma(p + \bar{\nu}_e \rightarrow n + e^+) \approx 5 \times 10^{-44} \left( \frac{E_{\bar{\nu}}}{\text{MeV}} \right)^2 \text{ cm}^2. \quad (11.1.2)$$

One common detection method for neutrinos is to just have a large tank of water and wait for this interaction and measure the positrons produced. The mean free path for an antineutrino is  $\lambda = 1/(n_p\sigma)$ , where  $n_p$  is the proton density. The mean free path will be very large, meaning we need very large tanks to achieve a reasonable detection rate. We can calculate the proton density,  $n_p$ , using the density of water,  $\rho = 1 \text{ g mL}^{-1}$ , and the molecular mass of water,  $18 \text{ g mol}^{-1}$ . The number of protons per mole of water is then  $10N_A$ , where we've used the fact that water,  $\text{H}_2\text{O}$ , has 10 protons, 2 from the hydrogens and eight from the oxygen. For a given volume of water,  $V$ , with one mole of water we have

$$n_p = \frac{10N_A}{V}. \quad (11.1.3)$$

The density of the water is  $\rho = 18 \text{ g/V}$ , and so

$$n_p = \frac{10N_A\rho}{18 \text{ g}} \approx 3.34 \times 10^{23} \text{ cm}^{-3}. \quad (11.1.4)$$

Which gives

$$\lambda \approx 6 \times 10^{19} \text{ cm} \approx 60 \text{ light years}. \quad (11.1.5)$$

To give another example of just how little neutrinos interact we expect that 1 in  $10^{11}$  neutrinos will interact when passing through the Earth.

As a final example a neutrino travelling through interstellar space has a mean free path of approximately  $\lambda \approx 10^{44} \text{ cm} \approx 10^{26} \text{ light years}$ .

The only reason we have detected any neutrinos is because there are so many of them. In a detector of 1000 kg of water we would expect there to be approximately  $10^{11}$  neutrinos passing through each square centimetre per second. We would therefore expect to have a few interaction events per day.

### 11.1.1 Detecting Neutrinos

The first suggestion for detection of neutrinos was to use an atomic bomb test to produce a large number of neutrinos. Remarkably the US government initially agreed to this before someone realised it was a terrible idea.

The actual method used to detect neutrinos for the first time was the reaction  $p + \bar{\nu}_e \rightarrow n + e^+$ , using Cd as a source of protons. After this reaction occurs the positron annihilates with an atomic electron producing photons. Approximately  $20 \mu\text{s}$  after this the neutron is captured and another photon produced. This whole procedure produces a characteristic double flash which allows us to be fairly certain that we have detected a neutrino.

## 11.2 Neutrino Flavours

The process  $\mu^- \rightarrow e^- \gamma$  is forbidden. This was the first evidence that made people suspect that muon neutrinos are different to electron neutrinos. In order to test this very high energy neutrinos, in the GeV range, were used, since Fermi theory predicts that the cross section increases with energy, so interactions become more likely at high energy.

A source of muon neutrinos is pion decay,  $\pi^+ \rightarrow \mu^+ \nu_\mu$ . Placing a detector far away from a beam of pions we can then detect the interaction  $n + \nu_\mu \rightarrow \mu^- + p$ , by

detecting the muons. The reason for the large distance is so that most non-neutrinos interact before reaching the detector. In the detector we detect mostly interactions producing muons, rather than electrons, suggesting that muon neutrinos are different to electron neutrinos.

### 11.3 Neutrino Oscillations

Between 1967 and the 1990s an experiment was run in the Homestake mine in South Dakota. The experiment was done underground in a mine to block background radiation. The experiment involved 615 tons of cleaning fluid, used as a source of chlorine, for the reaction

$$^{37}\text{Ce} + \nu_e \rightarrow ^{37}\text{Ar} + e^-. \quad (11.3.1)$$

It was possible to extract the argon and count it, giving a measure of the number of interactions. It was predicted that, given a rate of production of solar neutrinos predicted from knowing the processes that occur in the sun, as well as the luminosity of the sun, which is a measure of the rate of these processes, that there would be approximately 1.5 argon atoms produced per day. In fact they actually only found about 0.5 argon atoms per day. This may seem a small difference but the factor of three made people think.

This interaction can only detect electron neutrinos. If neutrinos can change flavour, and do so fairly often, we would expect that by the time they reach the Earth there would be an equal proportion of all three flavours of neutrinos, even though they all started as electron neutrinos. Indeed this is what happens.

#### 11.3.1 Atmospheric Neutrinos

One of the first experiments designed to detect flavour changing used atmospheric neutrinos. These are produced in cosmic rays by high energy proton showers, which produce pions. These then decay via the weak interaction into muons, which can further decay into electrons, both steps produce neutrinos:

$$\pi^+ \rightarrow \mu^+ \nu_\mu, \quad \text{and} \quad \mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu. \quad (11.3.2)$$

Given this reaction scheme we would expect to measure twice as many muon neutrinos as electron neutrinos.

This was tested in the Super-Kamiokande, or SuperK experiment in the 1990s. This was a large water Cherenkov detector. It consisted of a large tank of water surrounded by photomultipliers. Neutrinos were detected indirectly via charge current interactions with nuclei:

$$N + \nu_e \rightarrow X + e^-, \quad \text{and} \quad N + \bar{\nu}_\mu \rightarrow X + \mu^+. \quad (11.3.3)$$

The electrons and muons produced travel faster than the speed of light in water and so produce Cherenkov radiation, which is detected by the photomultipliers.

It was found that fewer muon neutrinos were detected than expected:

$$\frac{\nu_\mu \text{ observed}}{\nu_\mu \text{ expected}} = 0.45 \pm 0.02. \quad (11.3.4)$$

This result can be interpreted as muon neutrinos oscillating to tau neutrinos.

### 11.3.2 Oscillation

For neutrinos to oscillate they must have nonzero mass. This means that we can describe a neutrino as a superposition of either flavour or mass eigenstates. So far we have discussed the flavour, or weak, eigenstates,  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ , which are the particles produced in weak interactions. The mass eigenstates,  $\nu_1$ ,  $\nu_2$ , and  $\nu_3$ , have definite mass values.

Neutrinos propagate as mass mass eigenstates,

$$|\nu_i(t)\rangle = e^{ip \cdot x} |\nu_i\rangle. \quad (11.3.5)$$

The weak eigenstates are then a linear combination of these mass eigenstates. Therefore as the neutrino propagates the superposition changes and the flavour can change, even if the neutrino starts in an initially pure flavour state.

For simplicity we consider the case of two neutrino flavours. We then have

$$\begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} \cos \vartheta & -\sin \vartheta \\ \sin \vartheta & \cos \vartheta \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}, \quad \text{or} \quad \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \vartheta & \sin \vartheta \\ -\sin \vartheta & \cos \vartheta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}. \quad (11.3.6)$$

Here  $\vartheta$  is some parameter called the mixing angle.

Since the neutrino mass is small

$$E_i \approx |\mathbf{p}| + \frac{m_i^2}{2|\mathbf{p}|} \implies \Delta E = E_2 - E_1 = \frac{\Delta m^2}{2E}, \quad (11.3.7)$$

so if the neutrinos have different masses then the phase differs between the two components, which is what allows flavour oscillation to occur.

Starting with a neutrino that is in a pure muon neutrino state after some time the probability of interacting as an electron neutrino is given by

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2(2\vartheta) \sin^2 \left( \frac{1.27 \text{ eV} L \Delta m^2}{E} \right) \quad (11.3.8)$$

where  $L$  is the distance travelled by the neutrino in kilometres, and  $E$  is the energy in GeV.

The results of SuperK gives

$$\Delta m^2 = 2.3_{-0.08}^{+0.12} \times 10^{-3} \text{ eV}^2. \quad (11.3.9)$$

The latest results on neutrino mixing relate the mass and flavour eigenstates as

$$|\nu_3\rangle \approx \frac{\sqrt{2}}{2} (|\nu_\mu + \nu_\tau\rangle), \quad (11.3.10)$$

$$|\nu_2\rangle \approx 0.53|\nu_e\rangle + 0.60(|\nu_\mu\rangle - |\nu_\tau\rangle), \quad (11.3.11)$$

$$|\nu_3\rangle \approx 0.85|\nu_e\rangle - 0.37(|\nu_\mu\rangle - |\nu_\tau\rangle). \quad (11.3.12)$$

$$(11.3.13)$$

This predicts a relatively large mixing angle. The mass difference is then predicted to be  $\Delta m_{32}^2 \approx 2.5 \times 10^{-3} \text{ eV}^2$ , and  $\Delta m_{21}^2 \approx 8 \times 10^{-5} \text{ eV}^2$ .

We *don't* know the exact neutrino masses, just how they relate to each other and upper bounds, for example the mass of the electron neutrino is known to be less than 1 eV from the beta decay of tritium.

## Part VI

# Electroweak Interaction

# Twelve

## Electroweak Interaction

### 12.1 Weak Force Recap

The weak force is responsible for radioactive decay as well as muon and tau decay. It is propagated by the massive  $W^\pm$ , and  $Z$  bosons. Weak interactions have (relatively) long lifetimes,  $10^{-13}$  s to  $10^3$  s. The cross section for weak processes is also small,  $10^{-13}$  mb. The  $W$  and  $Z$  bosons can interact with each other through the weak force, and  $W^\pm$  can interact through the electromagnetic force as well.

### 12.2 Beyond Fermi Theory

Fermi theory is an effective theory. It breaks down as the total cross section increases with energy, so eventually we start finding probabilities greater than one. At low energy the cross section in Fermi theory is given by

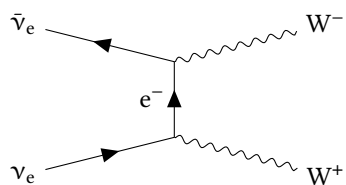
$$\sigma = \frac{G_F^2 s}{\pi}. \quad (12.2.1)$$

Here  $s$  is the centre of mass energy, which will eventually be large enough that we get non-physical probabilities.

The solution to this is that the  $W$  boson is massive, and so we need to include its mass in the propagator term:

$$\frac{1}{q^2 - m_W^2}. \quad (12.2.2)$$

This works in many situations to remove divergences. However, some processes still have divergences. One example of a process that is still divergent is  $\nu_e \bar{\nu}_e \rightarrow W^+ W^-$ . This interaction occurs via the  $t$ -channel diagram



The diagram shows a  $t$ -channel exchange of an electron. On the left, an incoming electron neutrino ( $\nu_e$ ) and an incoming electron antineutrino ( $\bar{\nu}_e$ ) meet at a vertex. A vertical line representing an electron ( $e^-$ ) propagator connects this vertex to another vertex on the right. From the right vertex, a  $W^+$  boson and a  $W^-$  boson emerge. The  $W^+$  boson is represented by a wavy line, and the  $W^-$  boson is also represented by a wavy line. The entire diagram is labeled (12.2.3).

We can solve this problem by including other diagrams which cancel out the divergences, at least to first order. The simplest diagram that does this is

$$(12.2.4)$$

This lead to the prediction of neutral current weak interactions. Another very similar diagram is

$$(12.2.5)$$

The similarity between these last two diagrams lead people to wonder if the electromagnetic and weak interactions are related.

### 12.3 The Electroweak Model

The most common model relating the electromagnetic and weak force was posited by Glashow, Weinberg, and Salam. In this model there are four massless bosons,  $W^+$ ,  $W^-$ ,  $W^0$ , and  $B^0$ . These then mix to produce the physical bosons we observe,  $W^+$ ,  $W^-$ ,  $Z^0$ , and  $\gamma$ . This mixing happens via

$$\begin{pmatrix} \gamma \\ Z^0 \end{pmatrix} = \begin{pmatrix} \cos \vartheta_W & \sin \vartheta_W \\ -\sin \vartheta_W & \cos \vartheta_W \end{pmatrix} \begin{pmatrix} B^0 \\ W^0 \end{pmatrix}. \quad (12.3.1)$$

This mixing has a single parameter,  $\vartheta_W$ , known as the **weak mixing angle** or **Weinberg angle**. Experimentally this has been measured to give  $\sin^2 \vartheta_W = 0.23$ .

Of these two “new” bosons the photon,  $\gamma$ , is massless and the mass of the  $Z^0$  is related to the mass of  $W$  via

$$m_Z^2 = \frac{m_W^2}{\cos^2 \vartheta_W}. \quad (12.3.2)$$

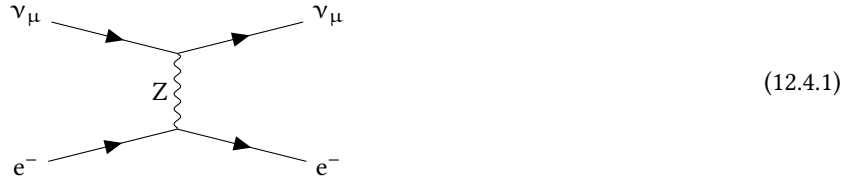
The interaction of these four bosons,  $W^\pm$ ,  $Z^0$ , and  $\gamma$ , is governed by three parameters, namely

$$\alpha_{\text{EM}} = \frac{e^2}{4\pi}, \quad G_F = \frac{g_W^2 \sqrt{2}}{8m_W^2}, \sin \vartheta_W. \quad (12.3.3)$$

Ultimately the fact that the electromagnetic and weak interactions can be described together is due to the fact that there is a single Lagrangian that describes both QED and the weak force. The sudden introduction of massive bosons occurs through the interaction with the Higgs field, which is the topic of discussion of the next part.

## 12.4 Experimental Evidence

In 1973 Gargamelle bubble chamber at CERN observed neutral currents for the first time in the interaction



Since  $\nu_\mu$  and Z can't be seen in the bubble chamber this looks like the electron just changes direction for no reason.

In 1976 Carlo Rubbia and Simon van der Meer proposed turning CERN's SPS fixed target proton synchrotron into a proton-antiproton collider to find W and Z bosons. This was done creating the Sp̄pS collider. Two experiments were run here, underground area one and two, or UA1 and UA2 for short. The main challenge of this experiment was producing an antiproton beam. This was done by hitting a fixed target with a proton beam. The problem was that this produced antiprotons with a wide range of energies and angles. The solution was to inject the antiprotons into an accumulator ring and measure beam properties in the ring and then make small adjustments for them on the opposite side of the ring with finely controlled magnets. This only worked because the signal to adjust the magnets could pass along the diameter of the ring and the antiprotons had to go halfway around the circumference. Making many small adjustments eventually the beam was acceptable for experiments to begin.

One of the other problems is how to detect W and Z once they have been created. This was done by looking for characteristic signatures. The signature of Z is an electron-antielectron pair with high momentum transverse to the beam, called high pt electrons. The signature for the W boson was a single high pt electron and missing transverse energy from the undetected neutrino.

## 12.5 Precision Electroweak Measurements at LEP

Between 1989 and 2000 the LEP collider at CERN made precise measurements of properties predicted by electroweak theory. LEP was an electron-positron collider with a 27 km circumference ring with four experiments. While it was running data was collected on 16 million Z events and 30 thousand  $W^\pm$  events.

During its operation the LEP was run at energies ranging from 90 GeV, which is the mass of Z, to 209 GeV, which is more than twice the mass of W.

Many decay products of Z were observed, including  $e^-e^+$ ,  $\mu^-\mu^+$ , and  $q\bar{q}$  as well as gluon jets emitted by these hadronic products.

At energies below the mass of the Z boson electron-antielectron annihilation is dominated by photon exchange, leading to cross section proportional to

$$\left| \begin{array}{c} e^+ \\ \searrow \\ \gamma \\ \nearrow \\ e^- \end{array} \begin{array}{c} \nearrow \\ f \\ \searrow \\ \bar{f} \end{array} \right|^2. \quad (12.5.1)$$



At slightly higher energies, even before we reach the mass of Z the interaction via bosons becomes important, giving matrix elements proportional to

$$\left| \begin{array}{c} e^+ \\ e^- \end{array} \rightarrow \begin{array}{c} \gamma \\ Z \end{array} \begin{array}{c} \bar{f} \\ f \end{array} \right|^2. \quad (12.5.2)$$

At approximately the mass of the Z boson Z resonance is achieved and most annihilations produce Z, so the matrix element is proportional to

$$\left| \begin{array}{c} e^+ \\ e^- \end{array} \rightarrow Z \begin{array}{c} \bar{f} \\ f \end{array} \right|^2. \quad (12.5.3)$$

At higher energies, approaching twice the mass of W, interactions producing  $W^\pm$  bosons become dominant and the matrix element is proportional to

$$\left| \begin{array}{c} e^+ \\ e^- \end{array} \rightarrow \begin{array}{c} \gamma \\ Z \end{array} \begin{array}{c} W^- \\ W^+ \end{array} + \begin{array}{c} e^+ \\ e^- \end{array} \rightarrow Z \begin{array}{c} W^- \\ W^+ \end{array} + \begin{array}{c} e^+ \\ e^- \end{array} \rightarrow \begin{array}{c} \nu_e \\ \bar{\nu}_e \end{array} \begin{array}{c} W^- \\ W^+ \end{array} \right|^2.$$

The experiments at LEP were remarkably precise. In order to verify results they needed to know the beam energies to an incredibly high precision, so precise that they had to account for

- The gravity of the moon causing tidal distortion in the ring size. An effect which caused approximately a 1 mm change over the 27 km circumference, leading to an approximate change of 10 MeV in the beam energy.
- The amount of water in the nearby lake Geneva.
- Current leak from a nearby railway.

Another experiment that was run at LEP was determining the number of neutrino species through Z- $\nu$  coupling. We can measure that

$$\Gamma(Z \rightarrow q\bar{q}) = (1744.0 \pm 2.0) \text{ MeV}. \quad (12.5.4)$$

As well as

$$\Gamma(Z \rightarrow e^-e^+) \approx \Gamma(Z \rightarrow \mu^-\mu^+) \approx \Gamma(Z \rightarrow \tau^-\tau^+) = (83.984 \pm 0.086) \text{ MeV}. \quad (12.5.5)$$

We predict that

$$\Gamma(Z \rightarrow \nu_e\bar{\nu}_e) = \Gamma(Z \rightarrow \nu_\mu\bar{\nu}_\mu) = \Gamma(Z \rightarrow \nu_\tau\bar{\nu}_\tau) = 167 \text{ MeV}. \quad (12.5.6)$$

The total width for Z is then

$$\Gamma_Z = N_q \Gamma(Z \rightarrow q\bar{q}) + N_\ell \Gamma(Z \rightarrow \ell\bar{\ell}) + N_\nu \Gamma(Z \rightarrow \nu\bar{\nu}), \quad (12.5.7)$$

where  $N_q$  is the number of quark flavours,  $N_\ell$  the number of non-neutrino lepton flavours, and  $N_\nu$  the number of neutrino flavours. Measurements of  $\Gamma_Z$  agree with the prediction that  $N_\nu = 3$ , with the results above we get  $N_\nu = 2.9963(74)$ .

Part VII

Higgs

# Thirteen

## Higgs

### 13.1 Higgs Mechanism

The electroweak theory relates the bosons  $W^\pm$ ,  $Z^0$ , and  $\gamma$ . We find that  $W^\pm$  and  $Z^0$  are massive, while  $\gamma$  is massless. This has profound consequences for the nature of the weak and electromagnetic forces. In the electroweak theory we start with four massless bosons. We then require something breaks the symmetry allowing massive bosons. This process is called the Higgs mechanism.

The solution, which was proposed separately by Higgs, Brout and Englert, Guralnik, and, Hagen and Kibble, was the existence of an additional field,  $\varphi$ , which gives rise to the potential

$$V(\varphi) = -\mu^2|\varphi|^2 + \lambda|\varphi|^4. \quad (13.1.1)$$

This is shown in [Figure 13.1](#).

The reason that this potential works as required is that the minimum is not at  $\varphi = 0$ , but at  $|\varphi_0| = \sqrt{\mu^2/\lambda} =: v$ , where the value is  $V(\varphi_0) = \lambda v^4/4$ .

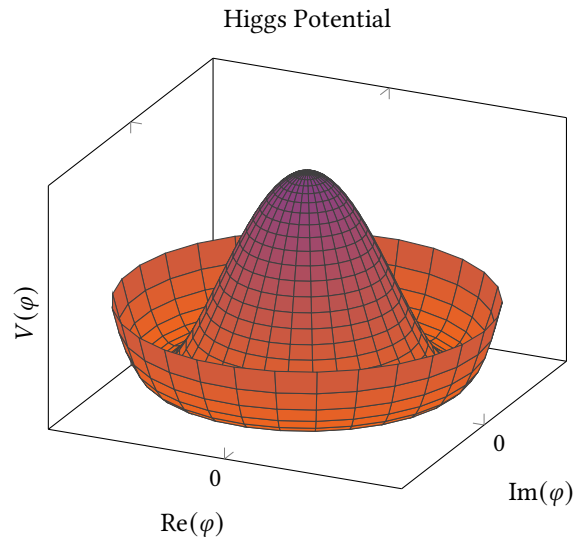


Figure 13.1: Higgs potential, also known as the sombrero potential.

The way that this potential works is that at high energies, such as in the early universe, the electroweak interaction is unified and the high energies mean that the  $\varphi = 0$  state can be achieved easily. At lower energies, later in the universe, the  $\varphi = 0$  state is no longer accessible. It is necessary to pick a direction in which to increase  $\varphi$  in order to move to a lower energy state. This introduces asymmetry, we say that spontaneous symmetry breaking occurs. This allows the W and Z bosons to have mass.

As well as allowing mass for W and Z the field  $\varphi$  actually allows all particles to have mass. To go along with this Peter Higgs predicted a new boson, the **Higgs boson**,  $H^0$ . The Higgs boson is itself massive, and has mass  $m_H = \mu = v\sqrt{\lambda}$ .

The vacuum expectation value,  $v$ , of the Higgs field is related to the mass of the W boson by

$$v = \frac{2m_W}{g_W} = \frac{1}{\sqrt{2}G_F} \approx 246 \text{ GeV}. \quad (13.1.2)$$

The value of  $\lambda$  is a free parameter in the standard model, this means that  $m_H$  is a free parameter also, although we can place certain bounds on it, for example in order to have unitary self interaction, that is have probabilities in  $[0, 1]$ , in the standard model we must have  $m_H < 1 \text{ TeV}$ . The width of the Higgs is

$$\Gamma_H = \sum_f \Gamma(H \rightarrow f\bar{f}) + \Gamma(H \rightarrow W^+W^-) + \Gamma(H \rightarrow ZZ) + \Gamma(H \rightarrow HH). \quad (13.1.3)$$

For heavy Higgs the final, self interaction, term dominates. This leads to a bound on the Higgs mass by requiring that  $\Gamma_H < m_H$ , which implies

$$G_F m_H^3 < m_H \implies m_H < \frac{1}{\sqrt{G_F}} = 296 \text{ GeV}. \quad (13.1.4)$$

Allowing for radiative corrections we give a reasonable upper bound of 1 TeV.

The Higgs couples to all massive particles. For fermions the coupling is proportional to the mass of the fermions. This means interactions involving, for example, electrons, are much less likely than those involving muons.

## 13.2 Search for the Higgs

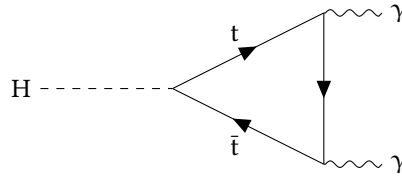
Between 1989 and 2000 the LEP searched for the Higgs. Since the Higgs couples to mass the s-channel process  $e^-e^+ \rightarrow H \rightarrow e^-e^+$  is incredibly unlikely. Instead the **Higgstrahlung** process,  $e^-e^+ \rightarrow Z^* \rightarrow ZH$ , was searched for. With a maximum centre of mass energy of  $\sqrt{s} = 207 \text{ GeV}$  LEP was able to detect the Higgs only if its mass was less than 116 GeV. During its run LEP set a lower bound on the Higgs mass of 114 GeV.

There were signals at about 115 GeV that some argued were the Higgs. To determine whether this was true would have required delaying the building of the LHC, which it was decided was not worth it. Instead they went with the “no lose theorem”, that in building the LHC they would either discover the LHC at 115 GeV, or at a higher energy, or they would find *something* worth finding.

The LHC came online at an energy of  $\sqrt{s} = 7 \text{ TeV}$  in 2010. In 2012 the teams at CERN working on the various experiments announced that they had found the Higgs. A variety of possible production and decay processes were considered, all

of which depend on the mass of the Higgs. The predicted cross sections for these interactions ranged from 1 pb to 30 pb, depending on the mass of the Higgs. For an integrated luminosity of  $100 \text{ fb}^{-1}$  per year we would expect to see between 10 000 and 300 000 Higgs per year.

In the end the process that lead to the discovery of the Higgs was  $H \rightarrow \gamma\gamma$ , where a small, but significant, peak was found at 125 GeV. Note that the Higgs cannot couple directly to photons, since they are massless, instead there are intermediary virtual particles involved in the process. For example,



The diagram shows a Higgs boson (H) represented by a dashed line on the left. It splits into two fermion lines, a top quark (t) and an anti-top quark ( $\bar{t}$ ), which form a loop. The top quark line has an arrow pointing right, and the anti-top quark line has an arrow pointing left. The loop is closed by a vertical line with a downward-pointing arrow. From the top and bottom of this vertical line, two wavy lines representing photons ( $\gamma$ ) emerge to the right.

(13.2.1)

This measurement was later verified by measurements of a peak at the same location in the process  $H \rightarrow Z^*Z \rightarrow \mu^- \mu^+ \mu^- \mu^+$ .

Measurements of the Higgs and its interactions are ongoing.

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# Acronyms

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## B

BNL: Brookhaven national laboratory [7](#)

## C

CABIBBO–KOBAYASHI–MASKAWA: m [62](#)

CDF: collider detector at Fermilab [7](#)

CERN: *conseil européen pour la recherche nucléaire* [7](#)

CLIC: compact linear collider [23](#)

CMS: compact muon solenoid [7](#)

## D

DONUT: direct observation of the nu tau [7](#)

## F

FCC: future circular collider [23](#)

## H

HERA: hadron electron ring accelerator [22](#)

## I

ILC: international linear collider [21](#)

## L

LEP: large electron positron [21](#)

LHC: large hadron collider [7](#)

LINAC: linear accelerator [20](#)

## P

PT: transverse momentum [71](#)

## Q

QCD: quantum chromodynamics [9](#)  
QED: quantum electrodynamics [9](#)  
QFT: quantum field theory [3](#)

**R**

RF: radio frequency [20](#)

**S**

Spp̄S: super proton antiproton synchrotron [22](#)  
SI: le système International d'Unités [3](#)  
SLAC: Stanford linear accelerator laboratory [7](#)  
SPEAR: Stanford's positron electron accelerating ring collider [54](#)  
SPS: super proton synchrotron [23](#)  
SUPERK: Super-Kamiokande [66](#)

**U**

UA: underground area [22](#)



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