

DeltaSpin (Self-Adaptive Spin Constraint in VASP)

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Contents

1	Terms of Use	2
2	Docker Image User Guide	3
3	Getting Started	4
4	INCAR Template	5
5	Selected Multi-step Use Cases	7
6	Strategy	9
7	Compatibility	10
8	Magnetization in VASP	11
9	Disclaimer	13

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2 Docker Image User Guide

For the following commands, replace the values within the angle brackets (<>) with relevant values for your system:

1. Pull the DeltaSpin image from ghcr.io:

```
1 docker pull ghcr.io/caizefeng/deltaspin:latest
```

2. Start a container using this image. Please note that you should allocate at least 16 GB of shared memory when starting the container from this image:

```
1 docker run -d --shm-size=<size of shared memory> \  
2 --name=deltaspin ghcr.io/caizefeng/deltaspin:latest
```

3. Enter the container:

```
1 docker exec -it deltaspin bash
```

4. Navigate to your calculation directory and place the activation code in the same directory as the INCAR, POSCAR, POTCAR, and KPOINTS files:

```
1 cd /root/DeltaSpin/examples/metal/Fe  
2 cp <path to your activation code file> .
```

5. Run vasp_deltaspin:

```
1 mpirun -np <number of threads> /root/DeltaSpin/bin/vasp_deltaspin
```

3 Getting Started

1. *Determine Wigner-Seitz radius (in Å) for each atom type in the system.*

Read the `RWIGS` entry presented in `POTCAR` and use that value for `INCAR`. Note that this value should **remain constant** in every single calculation of this system.

(OPTIONAL) Though tedious and not recommended, you can also determine the Wigner-Seitz radius through a Bader-type **spin density** analysis.

2. *Determine the magnetic ground state of the system.*

Set up `INCAR` with the following parameters:

```
RWIGS = <values from the 1st step>
I_CONSTRAINED_M = 2
SCTYPE = 0
LAMBDA = 0
```

and run `vasp_deltaspin`. When the calculation is finished, the last `MW_current` in `OSZICAR` represents the new ground-state `RWIGS`-defined magnetic moments.

Note that `MAGMOM` should be close to the ground-state.

You can add the following two tags to increase precision in ANY non-collinear calculation.

```
LASPH = .TRUE.
GGA_COMPAT = .FALSE.
```

3. *Constrain the system to the desired magnetic configuration.*

Refer to the `INCAR` template in `DeltaSpin/templates` directory. Set `M_CONSTR` to any value you want to constrain the magnetic moments to.

Run `vasp_deltaspin`.

Note that the ground-state magnetic configuration for reference, throughout all DeltaSpin calculations of this specific material system, is what we obtained in the 2nd step instead of any empirical value or **magnetization** in `OUTCAR` ¹. Different definitions of “magnetization” used in VASP will be discussed in another section.

4. (OPTIONAL) *Non-Self-Consistent (NSCF) calculation.*

Set `INCAR-LAMDA` to the last value of `OSZICAR-lambda` from the same **SCF** calculation that `CHGCAR` is from.

Also, complete the standard procedure for a typical VASP NSCF calculation, like preparing a `CHGCAR` and setting `INCAR-ICHARG` ≥ 10 .

Run `vasp_deltaspin`.

¹This means that if you set `M_CONSTR` to the last `MW_current` in `OSZICAR` from the 2nd step and apply the constraining, the energy will be exactly the same.

4 INCAR Template

```

SYSTEM = NiO
#GGA = PE

#Electronic minimization
PREC = Accurate          # Accurate is always preferred since non-collinear magnetization is very sensitive to the completeness
                           of FFT grid.
ENCUT = 600
EDIFF = 1E-9
LORBIT = 11
LREAL = .FALSE.
ISTART = 0
NPAR = 4
#ICHARG = 1
VOSKOWN = 1
LWAVE = .FALSE.
LCHARG = .FALSE.
NELMIN = 10
NELM = 500

#IBRION = 2 #OPT
#EDIFFG = -1E-3 #OPT
#ISIF = 3 #OPT
#NSW = 500 #OPT
#POTIM = 0.2 #OPT

#DOS
ISMEAR = -5
SIGMA = 0.2
#NEDOS = 3000 #DOS

#Magnetism
# The DEFINITION of atomic spins which are constrained (M_CONSTR in INCAR, MW in constrmag.F/OSZICAR):
# \vec{M}_{\mathbf{I}} =
# \int_{\Omega_{\mathbf{I}}} \vec{m}(\mathbf{r}) W_{\mathbf{I}}(\mathbf{r}) d\mathbf{r}

#External magnetic field (Zeeman field)
Bfield = 0 0 0          # External field in Cartesian coordinate (X,Y,Z), The unit is Tesla. !IMPORTANT!

#Non-collinear
ISYM = -1
RWIGS = 1.286 0.820     # Wigner-Seitz radius (in angstrom) for each atom type in the system. !IMPORTANT!
GGA_COMPAT = .FALSE.    # Restore the full lattice symmetry for gradient corrected functionals. Recommended.
LASPH = .TRUE.          # Non-spherical contribution to the gradient of the density. Recommended.
ISPIN = 2               # Does not matter in non-collinear calculation.
LWONCOLLINEAR = .TRUE.
LSORBIT = .TRUE.

#SASC
I_CONSTRAINED_M = 2
MAGMOM = 0.75914 0.16146 -1.00688 -0.67954 -0.91284 0.56668 0.00000 0.00000 0.00000 0.00000 0.00000 0.00000
M_CONSTR = 0.75914 0.16146 -1.00688 -0.67954 -0.91284 0.56668 0.00000 0.00000 0.00000 0.00000 0.00000 0.00000
SCTYPE = 1              # Type of SASC calculation 0: noSASC 1: SASC(L) 2: SASC(Q) 3: SASC(QL) !IMPORTANT!

#SASC(L)
IDECOSC = 0             # Whether Lagrangian coefficients are coupled 0: coupled 1: decoupled between atoms (deprecated) 2:
                           decoupled between every Cartesian component (deprecated)
LAMBDA = 12*0           # Initial Lagrangian coefficients for SASC(L)
CONSTRL = 6*1 6*0       # Whether the component is constrained or not 0: not constrained 1: constrained !IMPORTANT!
NSC = 100               # Maximum number of steps for each inner optimization
NSCHIN = 2              # Minimum number of steps for each inner optimization
SCDIFF = 1E-8           # Convergence criterion of iteration (RMS)
SCCONVB_GRAD = 1.9 -1   # Break condition for local gradients of spins w.r.t. Lagrangian coefficients (local field) -1: off
                           for the element

IDECAY_GRAD = 2         # Decay policy of gradient break condition 0: no decay 1: exponential decay 2: step down
SCDECAI_GRAD = 0.9      # Exponential base of decayed gradient break condition (IDECAY_GRAD = 1)
SCGRADB = 0.1           # Lower bound of exponentially decayed gradient break condition
NGRAD = 2               # Number of "steps" the gradient break condition curve includes (IDECAY_GRAD = 2)
NGRAD_STEP = 200 400    # Indices in the outer loop where the gradient break condition discontinuously decay
NGRAD_VALUE = 1.7 -1 \  # SCCONVB_GRAD for corresponding "steps"
1.5 -1

SCDECAI = -1            # Exponential decay of convergence criterion -1: no decay >0,<1: exponential base
SCDIFFB = 1E-7          # Lower bound of criterion decay, not gonna work if SCDECAI = -1
SCCONVB = -1            # Break condition for Lagrangian coefficients difference between two inner steps -1: this condition
                           is off
INISC = 0.01            # Initial trial step size
SCCUT = 3               # Restriction of step size
NELMSCI = 1             # Number of electronic steps BEFORE SASC(L)
NELMSTC = 0             # Number of INTERMEDIATE normal electronic steps
IALGOSC = 1             # Inner optimization algorithm 1: CG(F-R) 2: CG(P-R) 3: CG(H-S) 4: CG(D-Y)
IALGOSC_DIAG = 1        # Inner diagonalization algorithm 1: sub-space rotation 2: Blocked-Davidson 3: sub-space rotation
                           with B-D for the last step
LCUTSC_TRIAL = .TRUE.   # Whether trial step size is updated according to last-step optimal value TRUE: update FALSE: do not
                           update
LDESC = .FALSE.         # Debug mode TRUE: on FALSE: off

#SASC(Q)

```

```
LAMBDA_Q   = 10 10 10 10 10 10 # Initial Lagrangian coefficients for SASC(Q)
CONSTRL_Q  = 1 1 1 1 1 1        # Whether the component is constrained or not 0: not constrained 1: constrained
NSC_Q      = 500                 # Maximum number of steps for SASC(Q) iteration
SCDIFF_Q   = 3E-4                # Convergence criterion of iteration (+Inf-Norm)
EDIFF_Q    = 1E-6                # Global break condition (EDIFF) for SASC(Q)
INISC_Q    = 0.1                 # Initial trial step size
LDESC_Q    = .FALSE.             # Debug mode TRUE: on FALSE: off

#Orbital moments
#LORBMOM   = .TRUE.

#Mixer
AMIX       = 0.2
BMIX       = 0.0001
AMIX_MAG   = 0.8
BMIX_MAG   = 0.0001

#LSDA+U
LDAU       = .TRUE.
LDAUTYPE   = 2
LDAUL      = 2 -1
LDAUU      = 5.3 0.0
LDAUJ      = 0.0 0.0
LDAUPRINT  = 2
LMAXMIX    = 4
```

5 Selected Multi-step Use Cases

5.1 Calculate the effective field of the original ground-state configuration under Zeeman field

1. Relax the system to reach the ground-state magnetic configuration without an external field.

```
Bfield = 0 0 0
I_CONSTRAINED_M = 2
SCTYPE = 0
LAMBDA = 0
```

2. Apply a finite external field B. Constrain the size and direct of the magnetic moments using DeltaSpin.

```
Bfield = <B_x> <B_y> <B_z>
I_CONSTRAINED_M = 2
SCTYPE = 1
LAMBDA = <number of atoms>*0
M_CONSTR = <the last "M_current" from step 1>
```

5.2 Calculate the effective field of a direction-constrained magnetic configuration

1. Constrain the direction of the magnetic moments using the original VASP algorithm that is compatible with DeltaSpin.

```
I_CONSTRAINED_M = 1
SCTYPE = 0
LAMBDA = <some positive real number>
```

2. Constrain the size and direct of the magnetic moments using DeltaSpin.

```
I_CONSTRAINED_M = 2
SCTYPE = 1
LAMBDA = <number of atoms>*0
M_CONSTR = <the last "M_current" from step 1>
```

5.3 Calculate the electronic structure of a fully-constrained configuration

1. Self-consistently constrain the size and direct of the magnetic moments using DeltaSpin.

```
ISTART = 0
ICHARG = 2
LAMBDA = <number of atoms>*0
LWAVE = .TRUE.
LCHARG = .TRUE.
```

2. Change to a desired k-point grid/path. Run DeltaSpin non-self-consistently.

```
ISTART = 1  
ICHARG = 11  
LAMBDA = <the last "lambda" from step 1>
```


6 Strategy

When the energy difference (dE) keeps fluctuating around a value above the normal EDIFF (approximately $1\text{E}+02\text{ eV} \Leftrightarrow 1\text{E}-02\text{ eV}$), consider the following steps:

1. TEST FOR A SUITABLE GRADIENT CRITERION, i.e. `SCCONVB_GRAD`. Modifying this value might have a significant impact on convergence. Some values of `SCCONVB_GRAD` may result in a complete corruption of the diagonalization process.
Note that the same element in different systems may have different optimal gradient criteria. For example, the optimal `SCCONVB_GRAD` for chromium (Cr) is 0.5 (μ_B^2/eV) in a monolayer CrI_3 , but it is 2.0 in CrB_2 . Normally, a simple GRID SEARCH is sufficient to find a decent value for `SCCONVB_GRAD`.
2. SLOW DOWN THE MIXING. Decrease the values of `AMIX` and `AMIX_MAG`, and usually, it is necessary to increase `NELM` at the same time. To avoid out-of-memory issues caused by the increasing number of electronic steps, you may also need to allocate more memory when specifying the computing resources.

7 Compatibility

Using `SCTYPE=0`, *DeltaSpin* is now compatible with almost all functionalities from the original VASP, including

1. Constrained local moments approach (direction, `I_CONSTRAINED_M = 1`)
2. Constrained local moments approach (size and direction, `I_CONSTRAINED_M = 2`)
3. Spin spiral (`LSPIRAL=.TRUE.`)

8 Magnetization in VASP

OUTCAR magnetization (`magnetization (x,y,z)` in `OUTCAR`), is the integral of spin density in an area $\Omega_{\mathbf{R}}^{\text{AE}}$, where AE partial wave has integer occupancy, which can be slightly different from the augmentation sphere $\Omega_{\mathbf{R}}$.

$$\mathbf{M}_{\text{OUTCAR}} = \int_{\Omega_{\mathbf{R}}^{\text{AE}}} \rho^1(r) \quad (1)$$

$$\rho^1(r) = \sum_{n,(i,j)} f_n \langle \tilde{\Psi}_n | \tilde{p}_i \rangle \langle \phi_i | r \rangle \sigma \langle r | \phi_j \rangle \langle \tilde{p}_j | \tilde{\Psi}_n \rangle \quad (2)$$

where

$$\int_{\Omega_{\mathbf{R}}^{\text{AE}}} \langle \phi_i | r \rangle \langle r | \phi_j \rangle = \delta_{ij} \quad (3)$$

Because

$$\begin{aligned} \rho(r) &= \tilde{\rho}(r) + \rho^1(r) - \tilde{\rho}^1(r) \\ &= \sum_n f_n \langle \tilde{\Psi}_n | r \rangle \sigma \langle r | \tilde{\Psi}_n \rangle + \rho^1(r) - \sum_{n,(i,j)} f_n \langle \tilde{\Psi}_n | \tilde{p}_i \rangle \langle \tilde{\phi}_i | r \rangle \langle r | \tilde{\phi}_j \rangle \sigma \langle \tilde{p}_j | \tilde{\Psi}_n \rangle \end{aligned} \quad (4)$$

and

$$\sum_i |\tilde{\phi}_i \rangle \langle \tilde{p}_i| = 1 \text{ within } \Omega_{\mathbf{R}} \quad (5)$$

If we assume $\Omega_{\mathbf{R}}^{\text{AE}} \approx \Omega_{\mathbf{R}}$, we have

$$\mathbf{M}_{\text{OUTCAR}} = \int_{\Omega_{\mathbf{R}}^{\text{AE}}} \rho^1(r) \approx \int_{\Omega_{\mathbf{R}}} \rho(r) \quad (6)$$

OSZICAR magnetization (`M_current` in `OSZICAR`), is the integral of spin density in a user-defined real-space sphere.

$$\mathbf{M}_{\text{OSZICAR}} = \int_{\Omega'_{\mathbf{R}}} \rho(r) \quad (7)$$

Weighted OSZICAR magnetization (`MW_current` in `OSZICAR`), is the weighted integral of spin density in a user-defined real-space sphere. Prevalent weight functions are Bessel functions.

$$\mathbf{M}_{\text{OSZICAR}}^{\text{W}} = \int_{\Omega'_{\mathbf{R}}} W(r) \rho(r) \quad (8)$$

If we define $\Omega'_{\mathbf{R}}$ using WS radii (RWIGS) in `POTCAR`, that is $\Omega'_{\mathbf{R}} \approx \Omega_{\mathbf{R}}$, “*OUTCAR magnetization*” and “*OSZICAR magnetization*” will be closer (closer than partial core radius `RPCOR` and outmost cutoff radius `RCORE`), but still have a notable difference. That’s because these two have totally different origins.

Most importantly, both of them are based on the muffin-tin approximation and the summation of all atomic magnetic moments won't be equal to total magnetization. It's because there always exist some gaps in between those spheres.

The only well-defined quantity in DFT is “total magnetization”, the all-space integral of charge density, which is `mag` in `OSZICAR`.

To get access to a new definition of atomic magnetization, the summation of which is exactly the all-space integral, one could apply Bader analysis on spin density.

9 Disclaimer

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