

The logo for ExCALIBUR 10, featuring the word 'ExCALIBUR' in white and '10' in white inside an orange circle.

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# NEPTUNE UKAEA TECHNICAL: UQ

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NEPTUNE Workshop (The Cosener's House,  
Abingdon)

6 September 2022

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# Smooth function fitting: GPs vs Multivariate Adaptive Regression Splines (MARS)

- **Introduction:** overview of GPs and MARS, description of test, metrics. Scaling – splines cheap  $O(N)$ ; GP  $O(N^3)$ .
- **Smooth functions**
- **Smooth functions with noise**
- **Conclusion**
- **Additional:**
  - **Nektar++ slot convection**
  - **Smallab**



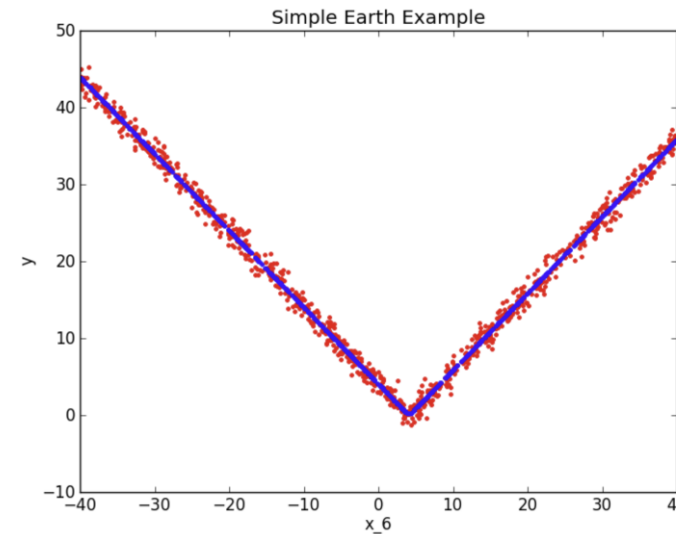
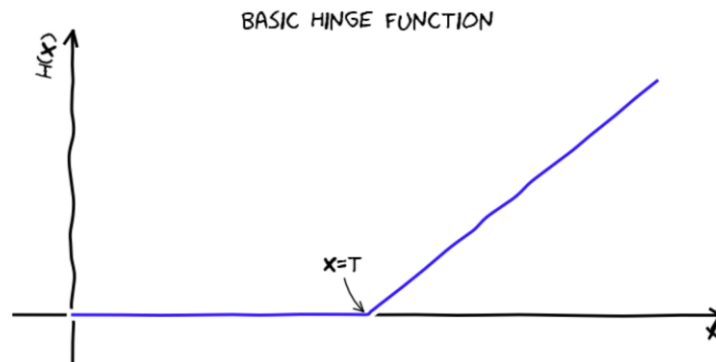
# MARS

J.H. Friedman. Multivariate adaptive regression splines. *The Annals of Statistics*, pages 1–67, 1991.

- **Multivariate Adaptive Regression Splines**

- Hinge functions used in fit – can handle discontinuous derivative.
- Fit is linear combination of global basis functions; each term is product of constant, linear function, hinge functions. (Num terms, product size are params.)

$$h(x - t) = [x - t]_+ = \begin{cases} x - t, & x > t \\ 0, & x \leq t \end{cases}$$

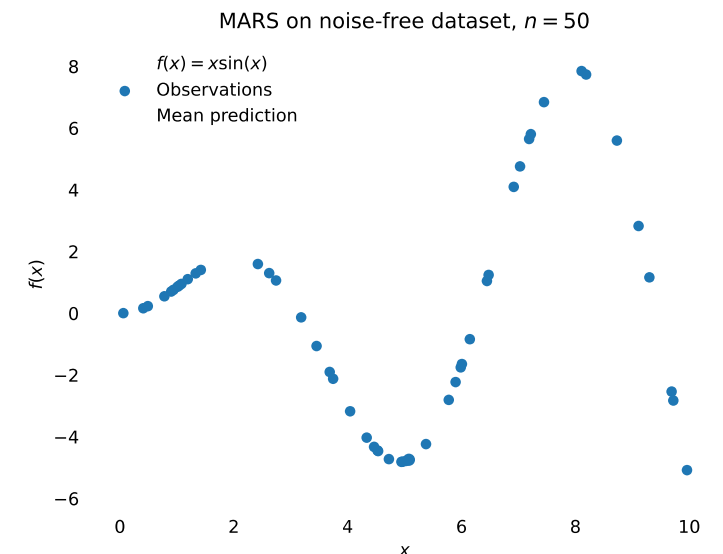
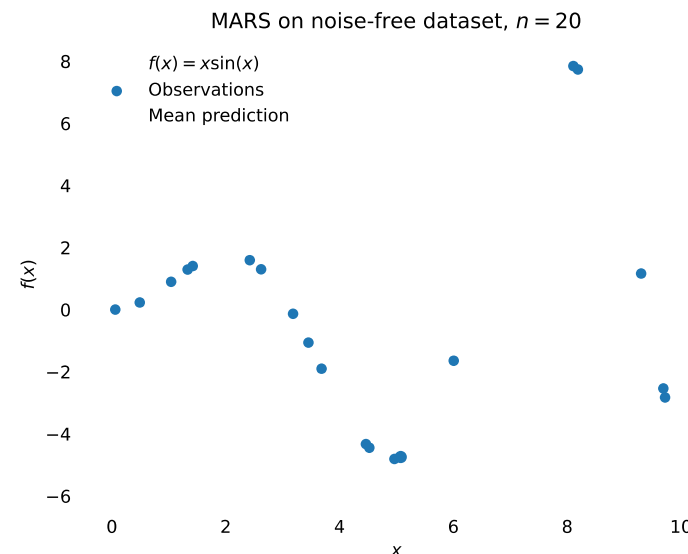
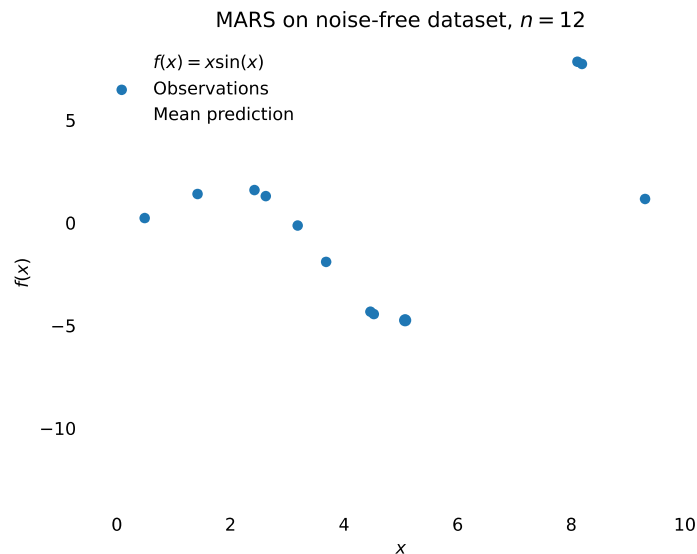
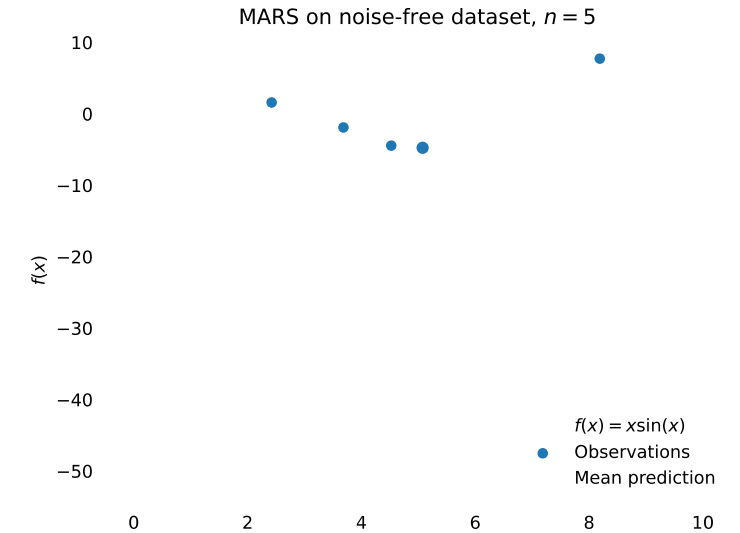
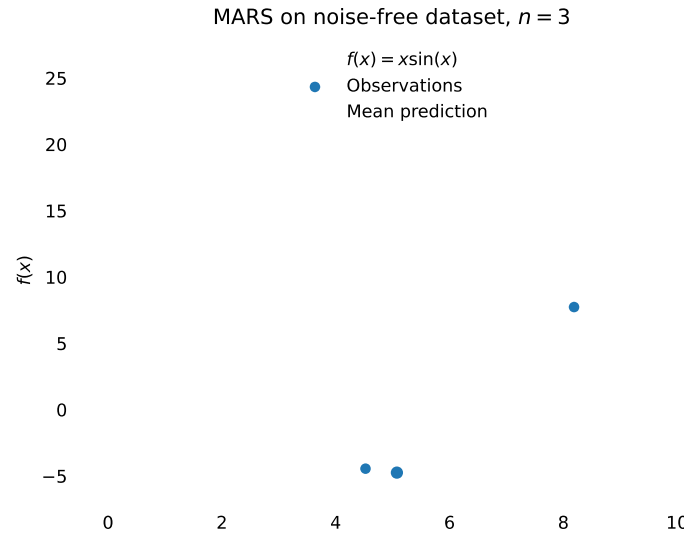


- **See** <https://contrib.scikit-learn.org/py-earth/index.html> (is source of these figures).

# Fitting a smooth function with MARS

- $f(x) = x \sin(x)$ , random sampling points

- $n < 10$ : pretty bad
- $10 \lesssim n$ : reasonable fits, but requires many more points for good fit
- Can be dubious when there are gaps in data (see  $x \approx 2, n = 50$ )

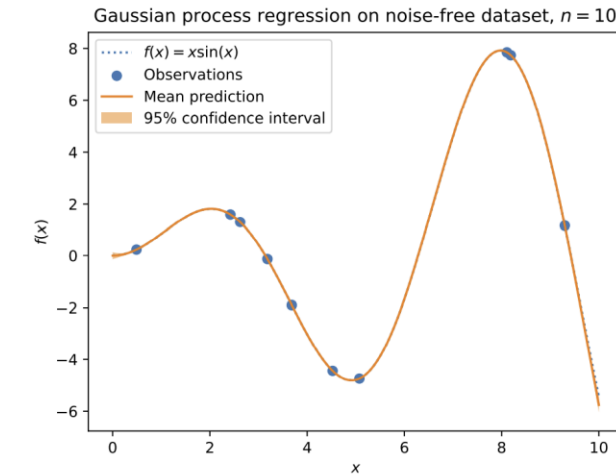
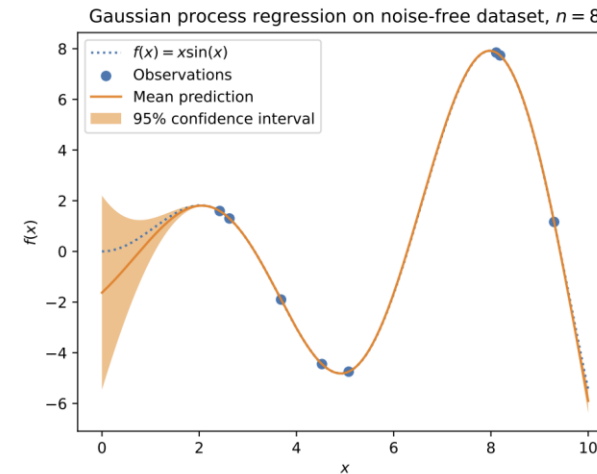
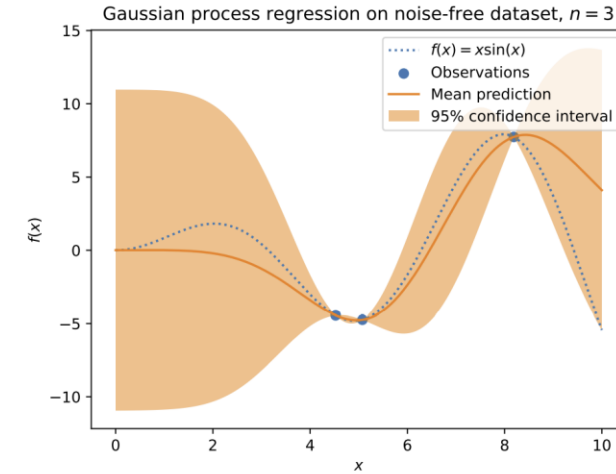
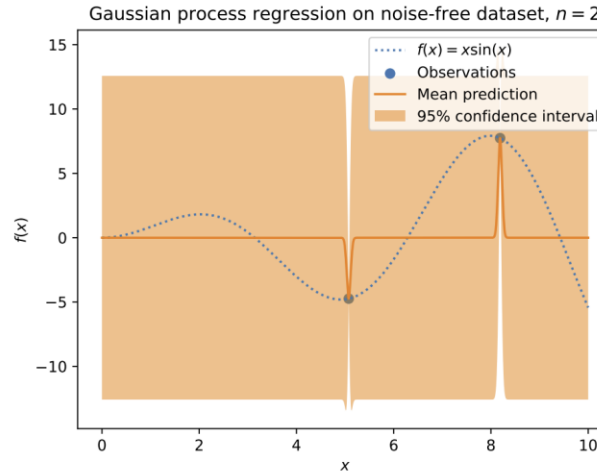


# Fitting a smooth function with GPs

- $f(x) = x \sin(x)$ , random sampling points\*

- $n < 2$ : meaningless...
- $3 \lesssim n \lesssim 7$ : poor but meaningful, bad when extrapolating
- $8 \lesssim n \lesssim 9$ : good within training range
- $10 \lesssim n$ : near-perfect by eye

GP has info about own uncertainty  
... MARS is able to do, but less clear; may be only implemented in R version of package.  
Cheap to do perturbed ensemble ...



\*Test cases adapted from  
SKLearn examples:

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[https://scikit-learn.org/stable/auto\\_examples/gaussian\\_process/plot\\_gpr\\_noisy\\_targets.html](https://scikit-learn.org/stable/auto_examples/gaussian_process/plot_gpr_noisy_targets.html)

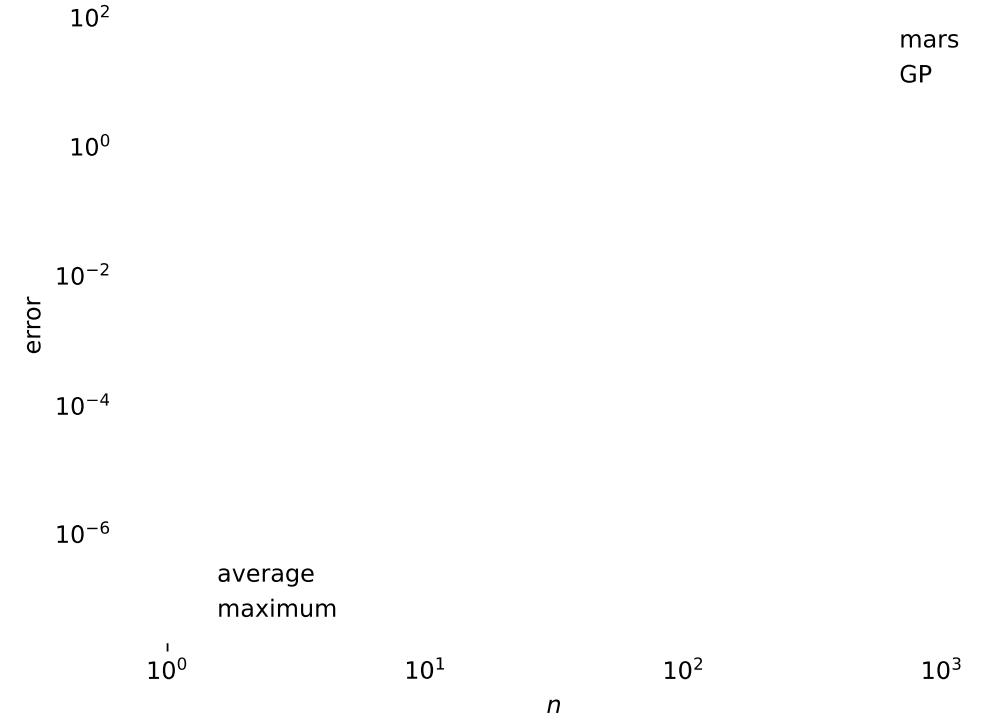
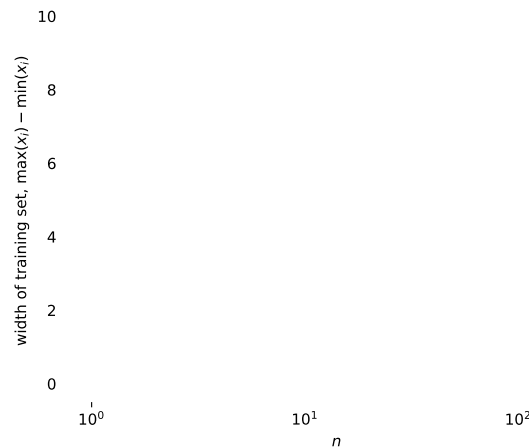
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# Comparative error: GPs vs MARS

- $f(x) = x \sin(x)$ , random sampling points\*
- GPs have smaller error and converge more quickly than MARS
- Jumps in convergence correspond to making the training set wider (more interpolation, less extrapolation)
- Can we make convergence quicker by choosing training points adaptively?



\*Test cases adapted from  
SKLearn examples:

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[https://scikit-learn.org/stable/auto\\_examples/gaussian\\_process/plot\\_gpr\\_noisy\\_targets.html](https://scikit-learn.org/stable/auto_examples/gaussian_process/plot_gpr_noisy_targets.html)

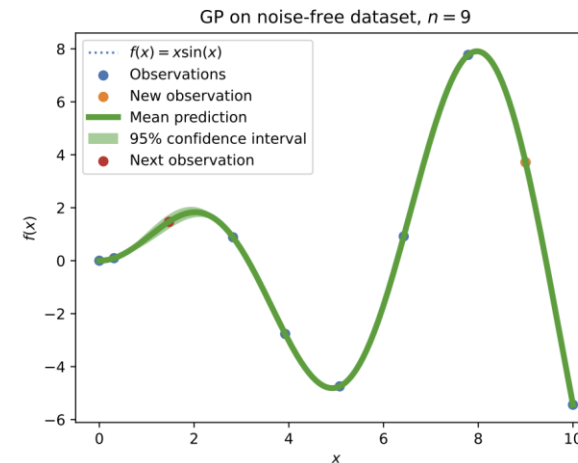
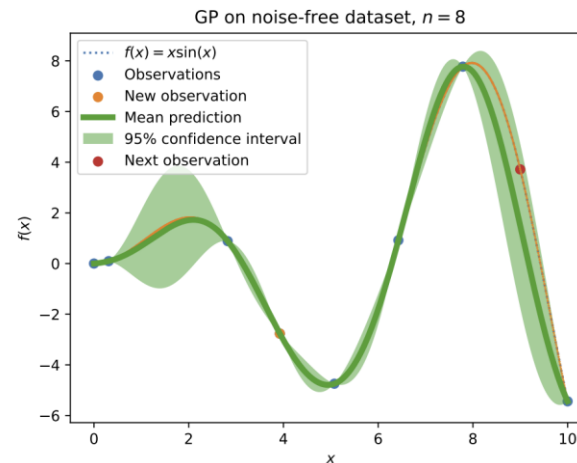
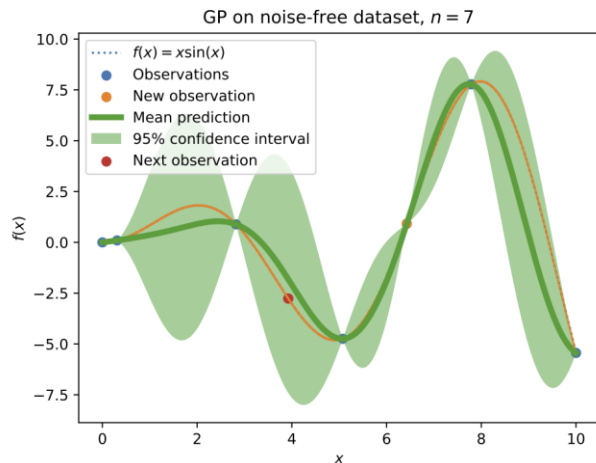
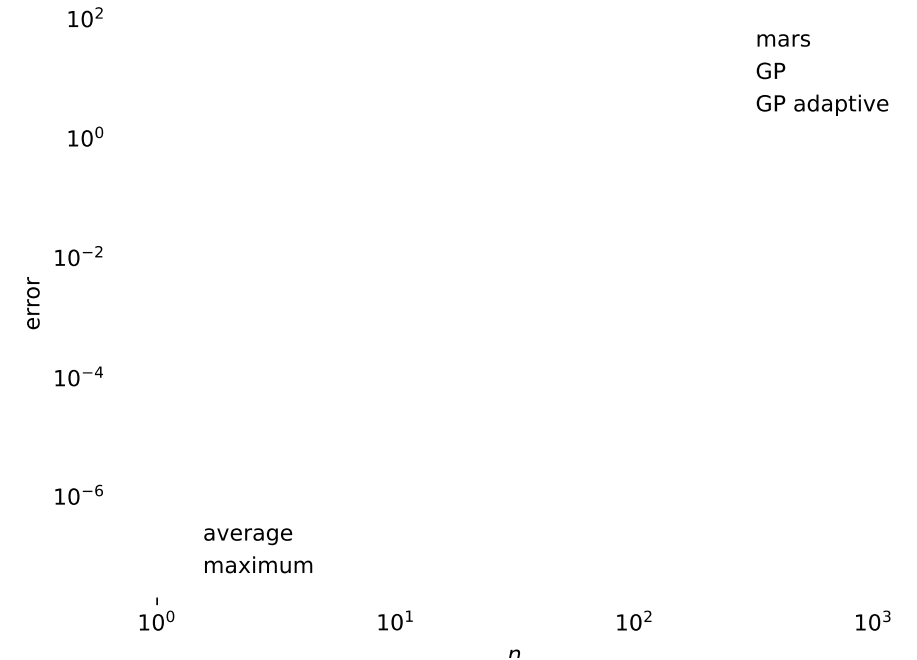
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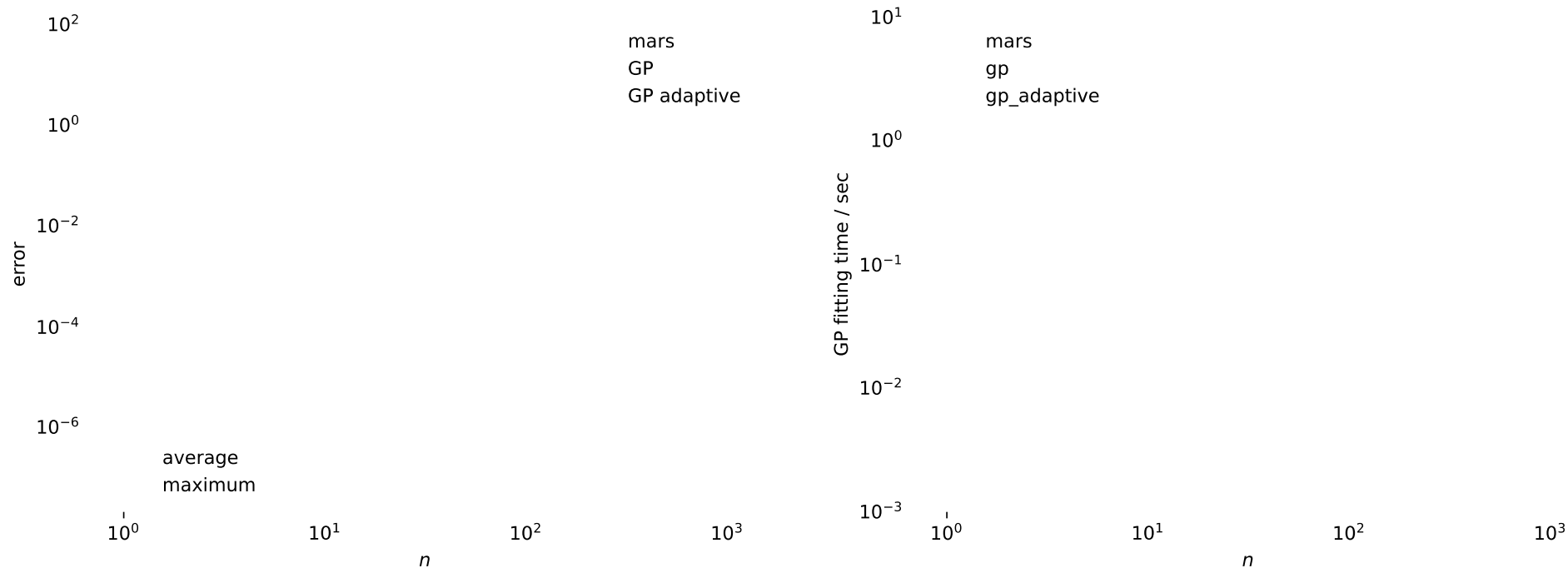
# GPs with adaptive sampling

- $f(x) = x \sin(x)$ , adaptive sampling
- GPs give confidence intervals of fit, choose sampling point where CI is largest
- Speeds convergence, though effect not very large with smooth functions
- Lower bound of error  $\Rightarrow$  no point using more than  $\sim 20$  points



# Comparative run times

- $f(x) = x \sin(x)$ , adaptive sampling
- GPs converge very quickly, but take longer to run
- This is not a concern when GP requires few points
- How does noisy data change this picture?

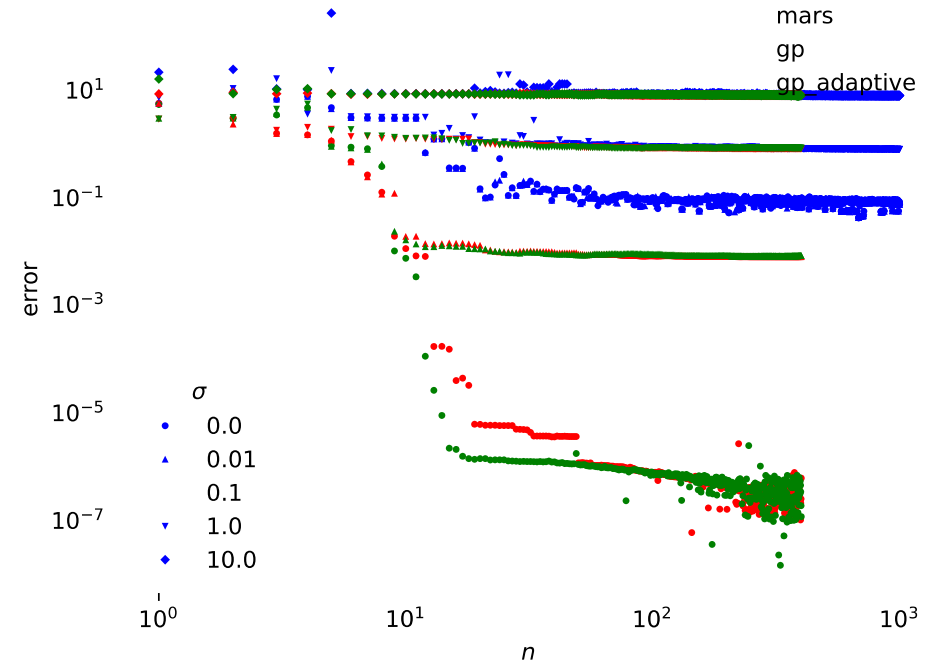
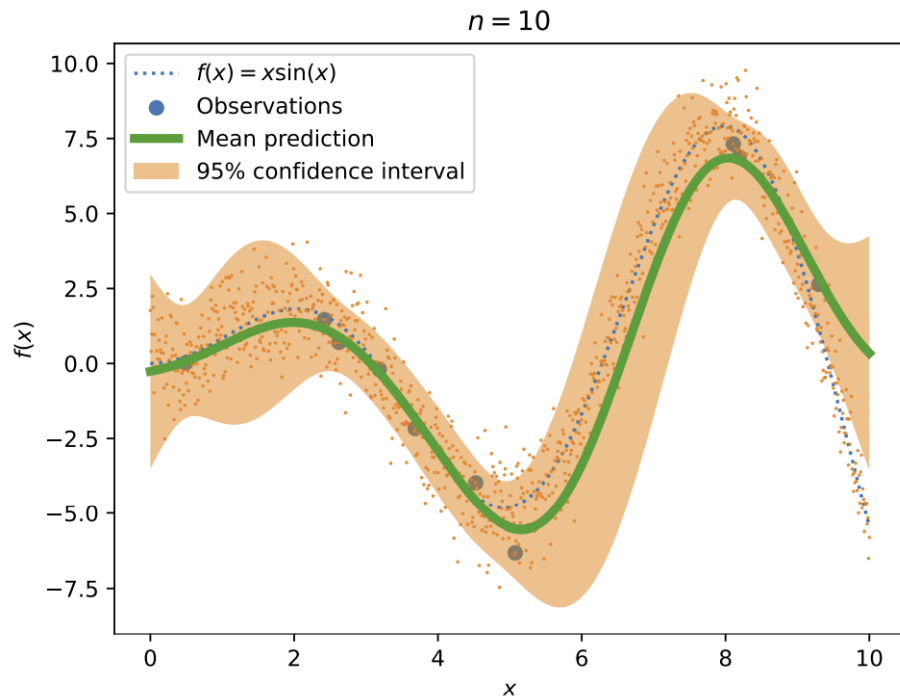




# Fitting a smooth function with noise

- $f(x) = x \sin(x) + X$ , where  $X \sim N(0, \sigma^2)$ , random sampling points

- Do NOT allow GPs to interpolate! Set kernel parameter  $\alpha = \sigma^2$ , i.e. generally need to approximate variance of data's random noise.
- No change to MARS (it doesn't interpolate)



- Noise level sets a lower bound for achievable error
- For noise  $\sigma \gtrsim 0.1$  achieves good fit... use MARS as cheaper
- Caveat: single performance metric studied.

# Conclusion

## • Gaussian Processes vs. MARS

### GPs:

- Work:  $O(N_{samples}^3)$
- Excellent fit to smooth functions
- Noise and bad fits around jumps
- Provides error estimates for fit, makes adaptive sampling easy
- Handle discontinuities with manual detection of jumps and one GP for each continuous section
- With adaptive sampling and jump detection, get very fast convergence

- See UKAEA report Selection of Techniques for Uncertainty Quantification, ([Documents/CD-EXCALIBUR-FMS0063-M5.2.pdf at main](#) · [ExCALIBUR-NEPTUNE/Documents \(github.com\)](#)).

### MARS:

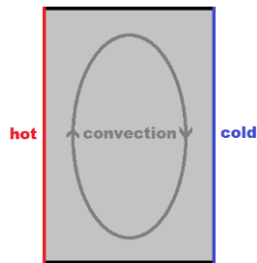
- Work:  $O(N_{samples})$
- Reasonable fit to smooth functions, but has an error floor, never converges to zero error
- No error estimate for fit, no adaptive sampling studied

### Noise:

- Imposes an error floor,  $E \sim \sigma$
- Neither GPs nor MARS can converge
- ... but MARS is much faster

# Nektar++ slot convection (reprise)

- 2D heat transfer problem:

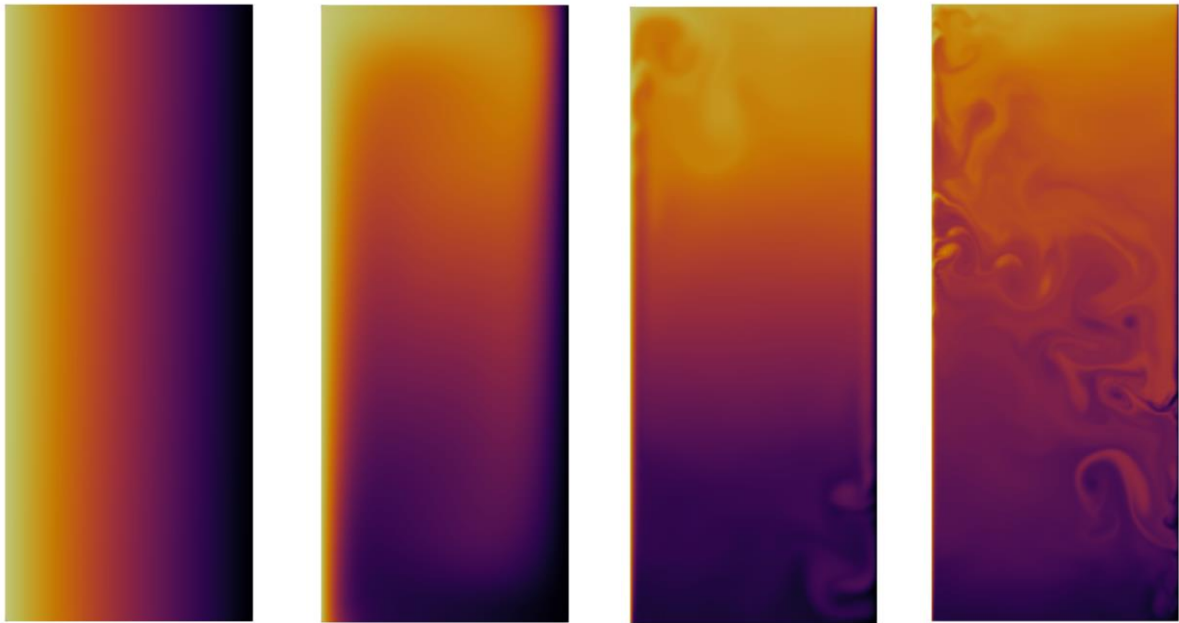


$$\begin{aligned} \frac{1}{Pr} \left( \frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) &= -\nabla p + Ra \, T \, \hat{\mathbf{y}} + \nabla^2 \mathbf{u} \\ \left( \frac{\partial T}{\partial t} + \mathbf{u} \cdot \nabla T \right) &= \nabla^2 T \\ \nabla \cdot \mathbf{u} &= 0. \end{aligned}$$

- Parameter: Rayleigh number  $Ra$ : strength of buoyancy force. ( $Pr = 0.71$ ; air.)
- Phenomenology: increased  $Ra$  gives increased heat transfer rate ... diffusion, then steady convection, then 2D turbulence driven by large intermittent boundary fluctuations (cf. tokamak plasma edge).
- Goals: study of system as turbulent heat transport proxy; also investigate capability of Nektar++.

heat flux results vs. literature

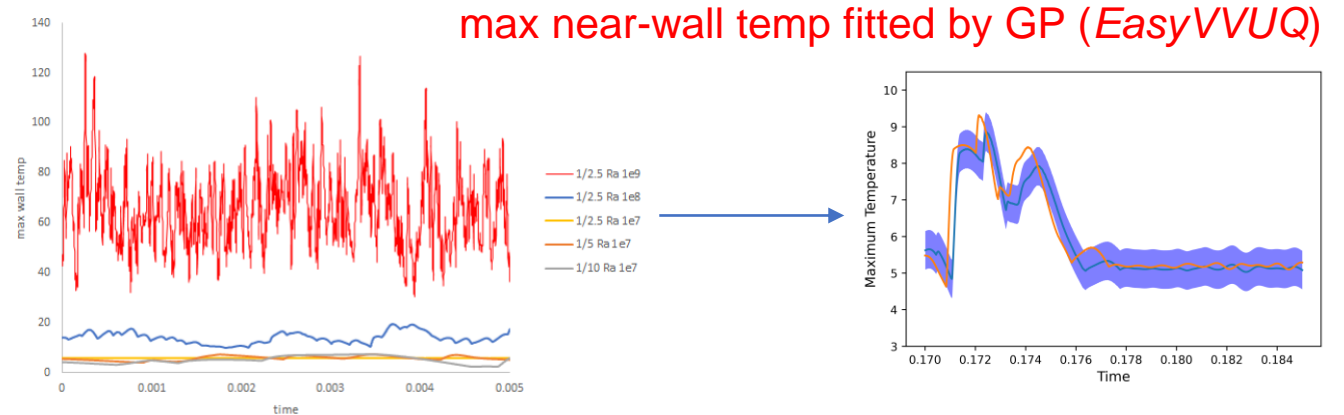
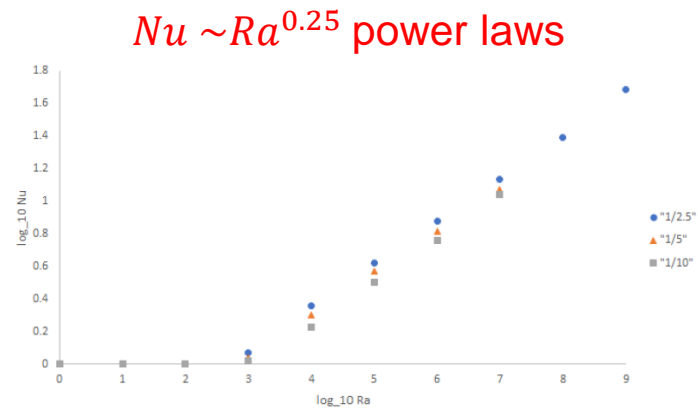
$Ra$	Barletta <i>et al</i> (2005)	Lo <i>et al</i> (2005)	Nektar++ (2021)
$10^3$	1.118	1.118	1.118
$10^4$	2.245	2.243	2.245
$10^5$	4.520/4.522	4.519	4.522
$10^6$	—	8.823	8.826
$10^7$	—	16.641	16.532



temperature fields for  $Ra = 10^2, 10^5, 10^8, 10^9$

# Nektar++ slot convection (reprise)

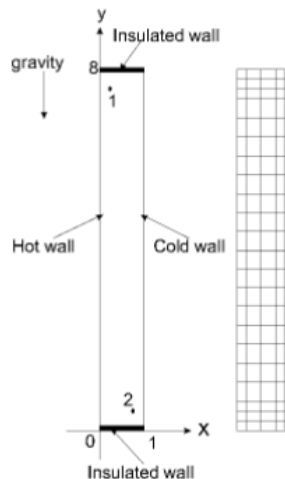
- Small modifications to Nektar++ to compute heat flux, and properties of wall-wave instability.
- Heat flux (Nusselt number): power laws (three different cavity aspect ratios)...



- Wall-wave instability: time series for maximum local temperature and position of hottest point (these were used with *EasyVVUQ* – constructed GP surrogate during hackathons).
- Main conclusions:  $p$ -refinement advised in preference to  $h$ -refinement (more efficient). Large- $Ra$  dynamics mandates small time step in solver (also tried out different schemes for SpectralVanishingViscosity solver option – did not see a big improvement).

# Smallab

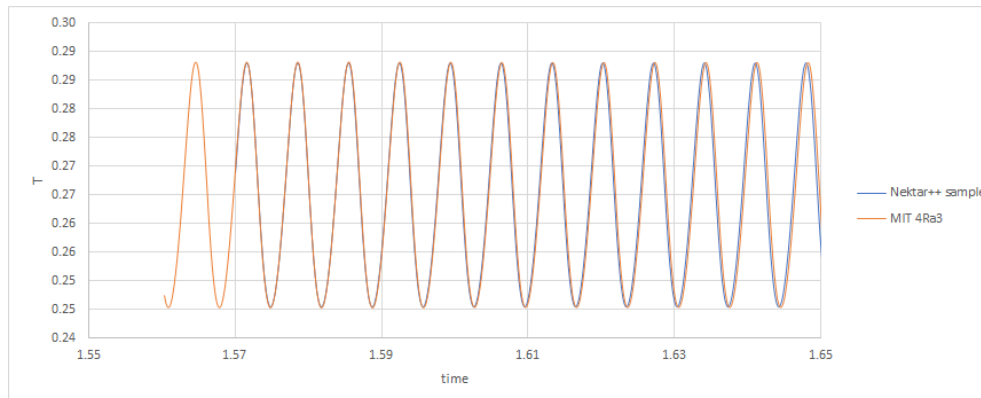
- Interesting problem for DA, GP-ROM surrogate ...
- Nektar++ convection problems are relevant (interesting features e.g. transition from time-periodic behaviour to turbulent chaos).
- MIT benchmark problem (2D 8:1 ratio cavity, Rayleigh numbers close to turbulent transition), then move on to 3D examples based on Smallab experiments (Dom Buta, Wayne Arter) - implemented in FabNEPTUNE. 2D case easy as accessible to single PC – turbulent 3D not. Video of Dom Buta initial results presentation available.



## (QoIs) 2D MIT benchmark

1. The x-velocity at point 1
2. The temperature at point 1
3. The Nusselt number along both sides of the wall in an integral valued function,

$$Nu(t) = \frac{1}{H} \int_0^H \left| \frac{\partial \Theta}{\partial x} \right|_{x=0, W} dx$$



## Smallab apparatus (Buta)



From Elder, *Laminar free convection in a vertical slot*, J. Fluid Mech (1965) vol.23, pt.1