

# PyDIVIDE User's Guide

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# 1 Introduction

PyDIVIDE is a toolkit that allows the user to quickly plot MAVEN Key Parameter data. This toolkit reads all data into a common structure, which allows comparisons across multiple instruments. PyDIVIDE also includes generic analysis routines for comparing data to existing models of Martian atmosphere.

## 2 Toolkit Installation

### 2.1 System Requirements

The MAVEN PyDIVIDE toolkit currently requires Anaconda 5.0 or above. Anaconda will install Python, as well as numerous software libraries for scientific computing. This toolkit is only compatible with Python 3.0 or above.

### 2.2 Downloading the Toolkit

To install the PyDIVIDE toolkit, type the following command into the local terminal/Anaconda Prompt terminal.

```
>> pip install pydivide
```

The following is necessary for custom bokeh colorbars and timestamps:

```
>> conda install -c bokeh nodejs
```

The PyDIVIDE toolkit can also be downloaded from the MAVEN Science Data Center GitHub page, <https://github.com/MAVENSDC>. This will require manual installation of all dependencies of PyDIVIDE. It is recommended that PyDIVIDE be installed via the pip command above.

### 2.3 Updating the Toolkit

The latest version of PyDIVIDE can be installed by typing the following command into the terminal:

```
>> pip install pydivide --upgrade
```

## 2.4 Mandatory Data Directory Structure

PyDIVIDE requires data files to be stored in an automatically-created directory structure, elaborated upon later in this section. This has a similar format to the SDC and SSL directory structures. The root directory for data storage can be chosen by the user. When first running a `download_files` or `read` procedure, the user will be prompted to select the `root_data_dir`. After the directory is selected, it is saved in `mvn_toolkit_prefs.txt`, and can later be changed manually as desired. After the first selection of the directory, the user will not be prompted by `download_files` or `read` again. `download_files` will place files into the chosen directory structure, and `read` will pull data files from that directory structure. The requisite directory structure is formatted as such:

```
<root_data_dir>/maven/data/sci/kp/insitu/YYYY/MM/  
                /kp/iuvs/YYYY/MM/
```

Level 2 instrument data downloaded via `download_files` will be placed into the following directory structure:

```
<root_data_dir>/maven/data/sci/sta/12/YYYY/MM/  
                /sep/12/YYYY/MM/  
                /swi/12/YYYY/MM/  
                /swe/12/YYYY/MM/  
                /lpw/12/YYYY/MM/  
                /mag/12/YYYY/MM/  
                /iuv/12/YYYY/MM/  
                /ngi/12/YYYY/MM/  
                /euv/12/YYYY/MM/  
                /acc/12/YYYY/MM/
```

Note: For Windows systems, the forward slashes above (/) will instead be back slashes (\).

## 2.5 Starting PyDIVIDE

An IDE is the recommended way to run PyDIVIDE procedures; however, they can also be run from the terminal. To start an interactive session of Python, enter the following commands into the terminal:

```
>> IPython  
>> import pydivide
```

PyDIVIDE function calls can now be entered into the terminal.

## 2.6 Getting More Help

Feel free to express any further problems or questions about installation or operation of the toolkit to the developers at [maven\\_divide@lasp.colorado.edu](mailto:maven_divide@lasp.colorado.edu).

## 3 Function Categories

This section serves to outline the rough order in which the functions should be called in order to plot/list/model/manipulate various sets of data.

### 3.1 Downloading and Reading Data Files

`download_files`: [Section 4.6](#)  
`read_model_results`: [Section 4.14](#)  
`read`: [Section 4.13](#)

### 3.2 Manipulating Key Parameter Data

`bin`: [Section 4.2](#)  
`insitu_search`: [Section 4.7](#)  
`resample`: [Section 4.15](#)

### 3.3 Plotting Key Parameter Data

`altplot`: [Section 4.1](#)  
`corona`: [Section 4.4](#)  
`map2d`: [Section 4.9](#)  
`occultation`: [Section 4.10](#)  
`periapse`: [Section 4.11](#)  
`plot`: [Section 4.12](#)  
`standards`: [Section 4.16](#)

### 3.4 Prediction Models

`create_model_maps`: [Section 4.5](#)  
`interpol_model`: [Section 4.8](#)

### 3.5 Toolkit Utilities

`cleanup_files`: [Section 4.3](#)  
`tplot_varcreate`: [Section 4.17](#)

## 4 PyDIVIDE Functions

All functions within PyDIVIDE are listed in this section in alphabetical order.

### 4.1 altplot

*Function Call:*

```
altplot(kp,
        parameter=None,
        time=None,
        sameplot=True,
        list=False,
        title='Altitude Plot',
        qt=True)
```

*Description:*

Plot the provided data against spacecraft altitude. If time is not provided, plot the entire dataset.

*Required Arguments:*

**kp:** STRUCT  
KP insitu data structure read from file(s).  
**parameter:** INT, LIST, STR  
Parameter(s) to be plotted. Can be provided as integer (by index) or string (by name: inst.obs). List may contain various data types.

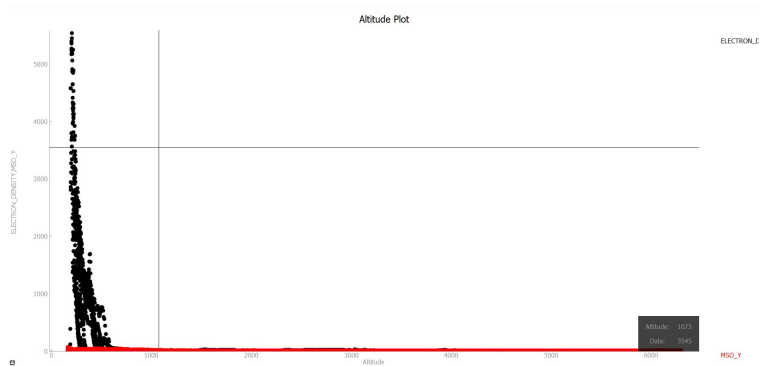
*Optional Arguments:*

**time:** [STR/INT,STR/INT]  
Two-element list of strings or integers indicating the time range to be plotted. Currently, no checks if time range is within data.  
**sameplot:** BOOL  
If True, put all curves on same axes.  
If False, generate new axes for each plot.  
**list:** BOOL  
List all KP parameters instead of plotting.  
**title:** STR  
Sets plot title. Default is 'Altitude Plot'.  
**qt:** BOOL  
If True, plot with PyQtGraph.  
If False, plot with bokeh.

*Examples:*

Plot LPW.ELECTRON\_DENSITY and MAG.MSO\_Y against spacecraft altitude.

```
>> pydivide.altplot(insitu,
                    parameter=['LPW.ELECTRON_DENSITY', 'MAG.MSO_Y'])
```



## 4.2 bin

*Function Call:*

```
bin(kp,
    parameter=None,
    bin_by=None,
    mins=None,
    maxs=None,
    binsize=None,
    avg=False,
    std=False,
    density=False,
    median=False)
```

*Description:*

Bins insitu Key Parameters by up to 8 different parameters, specified within the data structure. Necessary that at least one of **avg**, **std**, **median**, or **density** be specified.

*Required Arguments:*

**kp:** STRUCT  
KP insitu data structure read from file(s).  
**parameter:** STR  
Key Parameter to be binned. Only one may be binned at a time.  
**bin\_by:** INT, STR  
Parameters (index or name) by which to bin the specified Key Parameter.  
**binsize:** INT, LIST  
Bin size for each binning dimension. Number of elements must be equal to those in **bin\_by**.

*Optional Arguments:*

**mins:** INT, LIST  
Minimum value(s) for each binning scheme. Number of elements must be equal to those in **bin\_by**.  
**maxs:** INT, LIST



Maximum value(s) for each binning scheme. Number of elements must be equal to those in `bin_by`.

**avg:** BOOL

Calculate average per bin.

**std:** BOOL

Calculate standard deviation per bin.

**density:** BOOL

Returns number of items in each bin.

**median:** BOOL

Calculate median per bin.

*Returns:*

This procedure outputs up to 4 arrays to user-defined variables, corresponding to `avg`, `std`, `median`, and `density`.

*Examples:*

Bin STATIC O<sup>+</sup> characteristic energy by spacecraft latitude (1° resolution) and longitude (2° resolution).

```
>> output_avg = pydivide.bin(insitu,
                             parameter='static.oplus_char_energy',
                             bin_by=['spacecraft.geo_latitude',
                                       'spacecraft.geo_longitude'],
                             avg=True, binsize=[2,1])
```

Bin SWIA H<sup>+</sup> density by spacecraft altitude (10km resolution), return average value and standard deviation for each bin.

```
>> output_avg,output_std = pydivide.bin(insitu,
                                         parameter='swia.hplus_density',
                                         bin_by='spacecraft.altitude',
                                         binsize=10, avg=True, std=True)
```

### 4.3 cleanup\_files

*Function Call:*

```
cleanup_files()
```

*Description:*

Searches code directory for \*.tab files, keeps latest versions/revisions, asks to delete old versions/revisions. Will ignore files not ending in .tab and not starting with "mvn\_kp\_insitu" or "mvn\_kp\_iuvs".

*Required Arguments:*

None.

*Optional Arguments:*

None.

*File Requirements:*

All \*.tab files must be named with the following formats:

"mvn\_kp\_insitu\_YYYYMMDD\_vXX\_rXX.tab"

Ex: mvn\_kp\_insitu\_20170619\_v13\_r04.tab  
“mvn\_kp\_iuvs\_ORBIT\_YYYYMMDDTHHMMSS\_vXX\_rXX.tab”  
Ex: mvn\_kp\_iuvs\_02403\_20151225T003727\_v07\_r01.tab

Any extraneous characters or formatting changes to the filename are not compatible with the function regexing.

*Examples:*

Remove all out-of-date insitu and IUVS files from the local directory.

```
>> pydivide.cleanup_files()
```

## 4.4 corona

*Function Call:*

```
corona(iuvs,  
       sameplot=True,  
       density=True,  
       radiance=True,  
       orbit_num=None,  
       species=None,  
       log=False,  
       title='IUVS Corona Observations',  
       qt=True)
```

*Description:*

Create altitude plots of corona limb scans from IUVS KP files, containing radiance and density data of various chemical species.

*Required Arguments:*

**iuvs:** STRUCT

IUVS Key Parameter data structure returned from `read`.

*Optional Arguments:*

**sameplot:** BOOL

If True, will plot everything on one plot.

If False, will create stacked plots.

**density:** BOOL

Plot density.

**radiance:** BOOL

Plot radiance.

**orbit\_num:** INT, LIST

Orbit number(s) to be plotted.

**species:** STR, LIST

Plots only specified chemical species (Appendix A).

**log:** BOOL

Sets logarithmic axes for plotting.

**title:** STR

Sets plot title. Default is 'IUVS Corona Observations'.

**qt:** BOOL

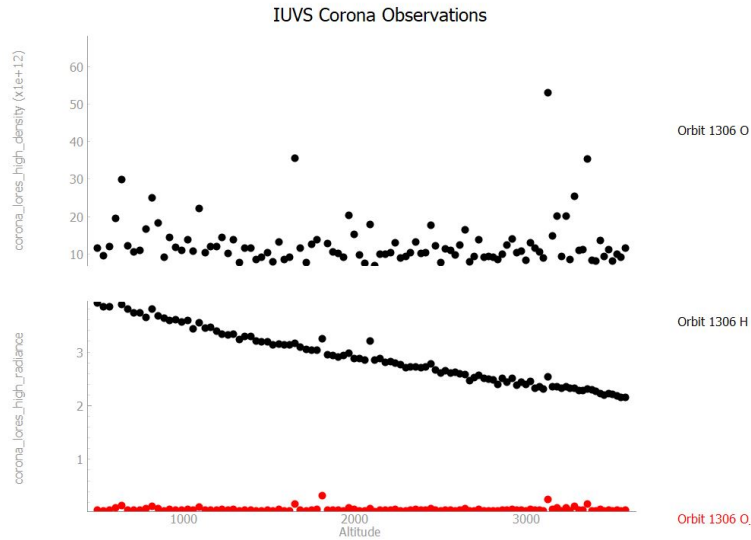
If True, plot with PyQtGraph.

If False, plot with bokeh.

*Examples:*

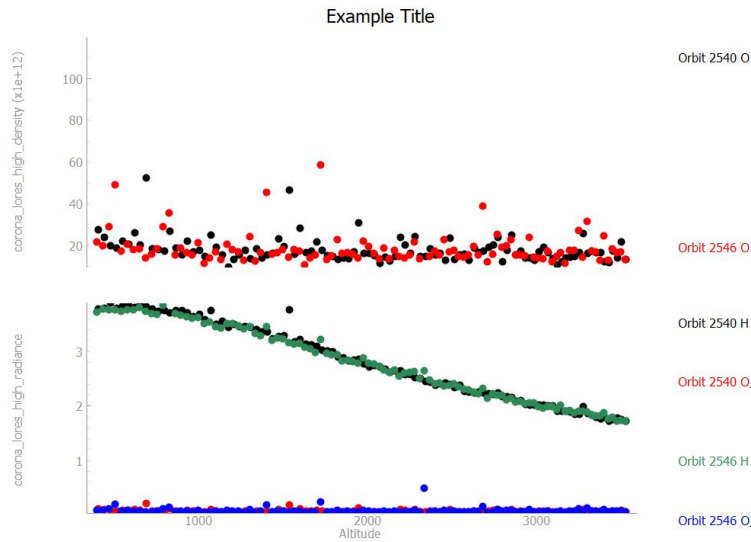
Plot all IUVS radiance and density data.

```
>> insitu,iuvs = pydivide.read('2015-06-02','2015-06-03')
>> pydivide.corona(iuvs)
```



Plot IUVS radiance and density data for H, O, and O<sub>1304</sub> for orbits 2540 and 2456.

```
>> insitu,iuvs = pydivide.read('2016-01-19','2016-01-20')
>> pydivide.corona(iuvs,
                    species=['H','O','O_1304'],
                    orbit_num=[2540,2456],
                    title='Example Title')
```



## 4.5 create\_model\_maps

*Function Call:*

```
create_model_maps(altitude,
                  variable=None,
                  model=None,
                  file=None,
                  numContours=25,
                  fill=False,
                  ct='viridis',
                  transparency=1,
                  nearest=False,
                  linear=True,
                  saveFig=True)
```

*Description:*

Generates a .png contour map of a model at a specific altitude. These can be used as a background in map2d. The models must be downloaded manually from the SDC website:

<https://lasp.colorado.edu/maven/sdc/public/pages/models.html>.

*Required Arguments:*

**altitude:** INT

Specified altitude of output map.

**variable:** STR

Plots specified chemical species (Appendix A).

**model:** DICT

Model variable produced from prior call to `read_model_results`.

**file:** STR

If model *not* provided (produced from `read_model_results`), full path to model can be set and read.

*Optional Arguments:*

**numContours:** INT

Specifies number of contour lines. Default is 25.

**fill:** BOOL

If True, fills in contour levels instead of generating lines.

**ct:** STR

Sets color table. Valid color tables can be found here:

[https://matplotlib.org/examples/color/colormaps\\_\\_reference.html](https://matplotlib.org/examples/color/colormaps__reference.html)

**transparency:** INT, FLOAT

Sets transparency between [0,1] inclusive. 0 is completely transparent, and 1 is completely opaque.

**nearest:** BOOL

If True, instead of interpolating nearby values, this returns the value of the nearest neighbor altitude.

**linear:** BOOL

If True, performs linear interpolation between 2 altitude layers.

**saveFig:** BOOL

If True, saves figure as .png file.

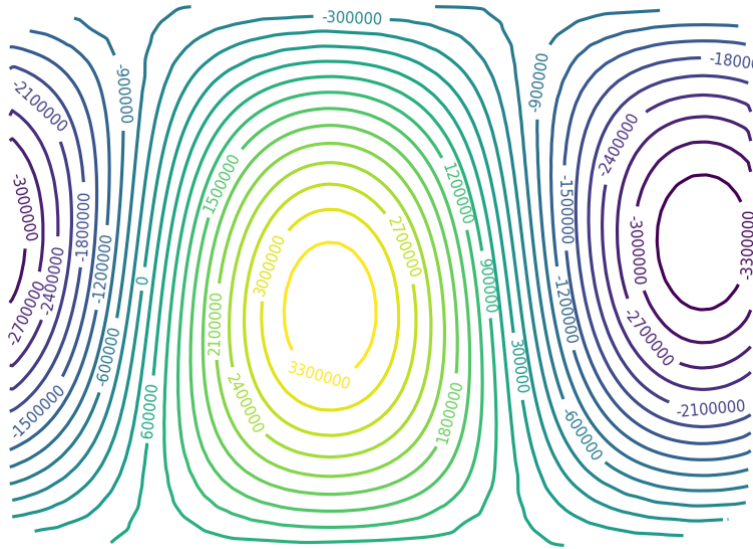
*Returns:*

A .png file will be created in the same directory as the specified input model.

*Examples:*

Interpolate all model tracers to spacecraft trajectory using nearest neighbor interpolation.

```
>> pydivide.create_model_maps(altitude=170,  
    file = '<dir_path>/MAMPS_LS180_F130_081216.nc',  
    variable='geo_x',  
    saveFig=True)
```



## 4.6 download\_files

*Function Call:*

```
download_files(filenamees=None,
               instruments=None,
               list_files=False,
               level='12',
               insitu=True,
               iuvs=False,
               new_files=False,
               start_date='2014-01-01',
               end_date='2020-01-01',
               update_prefs=False,
               only_update_prefs=False,
               exclude_orbit_file=False,
               local_dir=None)
```

*Description:*

Download insitu and/or IUVS Key Parameter (KP) data files from the MAVEN SDC web server. Also compatible with instrument-specific data downloads. **insitu**, **iuvs**, or at least one **instrument** must be specified.

*Required Arguments:*

**insitu:** BOOL  
Search/download insitu KP data files.

**iuvs:** BOOL  
Search/download IUVS KP data files.

**instruments:** STR, LIST

Search/download data for one or more 3-character instrument abbreviations listed in Section 2.4.

*Optional Arguments:*

**filenames:** STR, ARRAY

Specific filename strings to search/download.

**list\_files:** BOOL

List files available for download based on various parameters.

**level:** STR

Specifies desired data level to download, requires specific instrument argument(s). Options include: 10

11, 11a, 11b, 11c

12, 12a, 12b, 12c

13, 13a, 13b, 13c

By default, KP data is downloaded, at which the data is sufficiently calibrated for science analysis.

**new\_files:** BOOL

Search/download files on the SDC server that do not exist locally.

**start\_date:** 'YYYY-MM-DD'

Search/download data from start\_date to present, inclusive.

**end\_date:** 'YYYY-MM-DD'

Search/download data prior to end\_date, inclusive.

**update\_prefs:** BOOL

Before searching/downloading data, open window to allow user to update root\_data\_dir in mvn\_toolkit\_prefs.txt. Once the new path is selected, the search/download will proceed according to the remaining arguments.

**only\_update\_prefs:** BOOL

Like update\_prefs, but will not attempt to search/download files.

**exclude\_orbit\_file:** BOOL

Do not download new version of orbit number file from <https://naif.jpl.nasa.gov/naif/>.

**local\_dir:** STR

Specify a directory for file download. Overrides (but does not overwrite) root\_data\_dir in mvn\_toolkit\_prefs.txt.

*Examples:*

Download all available insitu data between 2015-01-01 and 2015-01-31, inclusive:

```
>> pydivide.download_files(start_date='2015-01-01',
                             end_date='2015-01-31',
                             insitu=True)
```

List all available CDF insitu KP files on the server:

```
>> pydivide.download_files(insitu=True,
                             list_files=True)
```

Download all *new* IUVS files from 6 April 2015 not found in the local directory.

```
>> pydivide.download_files(iuvs=True,
                             new_files=True,
                             end_date='2015-04-06')
```

List all available Level 2 data files for SWIA.

```
>> pydivide.download_files(instruments='swi',
                             list_files=True,
                             level='L2')
```

List all available Level 2 data files for SWIA for the month of January 2015.

```
>> pydivide.download_files(start_date='2015-01-01',
                             end_date='2015-01-31',
                             instruments='swi',
                             list_files=True,
                             level='L2')
```

Download all *new* Level 2 data files for NGIMS, STATIC, and EUV.

```
>> pydivide.download_files(instruments=['ngi', 'sta', 'euv'],
                             new_files=True)
```

## 4.7 insitu\_search

*Function Call:*

```
insitu_search(kp,
               parameter,
               min=None,
               max=None,
               list=False)
```

*Description:*

Search existing insitu data structure for specified data, output new structure. containing only the specified data.

*Required Arguments:*

**kp:** STRUCT  
KP insitu data structure read from file(s).

*Optional Arguments:*

**parameter:** STR, INT, LIST  
Name or index of Key Parameter to be searched. May also be entered as list of strings or indices to search for multiple parameters.

**min:** INT, LIST  
Minimum value for specified search criteria. If excluded, minimum is assumed to be  $-\infty$ . If multiple minima specified, they will



be applied respectively to specified parameters.

**max:** INT, LIST

Maximum value for specified search criteria. If excluded, maximum is assumed to be  $\infty$ . If multiple maxima specified, they will be applied respectively to specified parameters.

**list:** BOOL

Display ordered list of all parameters in **insitu** input.

If list is called, no other arguments will be executed, and no output data structure will be returned.

*Examples:*

Find all STATIC O<sup>+</sup> density data greater than 3000 cm<sup>-3</sup> and less than 1000000 cm<sup>-3</sup>, store results in **insitu\_new**.

```
>> insitu_new = pydivide.insitu_search(insitu,
                                     parameter='static.oplus_density',
                                     min=3000,
                                     max=1000000)
```

## 4.8 interpol\_model

*Function Call:*

```
interpol_model(kp,
               model=None,
               file=None,
               nearest=False)
```

*Description:*

Reads in MAVEN's position from insitu, and determines the value of the models at those points.

*Required Arguments:*

**kp:** STRUCT

KP insitu data structure read from file(s).

**model:** STR

Source of simulation data to be interpolated.

**file:** STR

If model not provided, can specify the full path to the model.

*Optional Arguments:*

**nearest:** BOOL

If True, instead of interpolating nearby values, this returns the value of the nearest neighbor altitude.

*Returns:*

Returns array of data representative of what the spacecraft would have measured if it were traveling through the model.

*Examples:*

Interpolate all model tracers to the spacecraft trajectory using nearest neighbor interpolation.

```
>> results = pydivide.interpol_model(insitu,
                                     file='<dir_path>/Elew_18_06_14_t00600.nc',
                                     nearest=True)
```

## 4.9 map2d

*Function Call:*

```
map2d(kp,
      parameter=None,
      time=None,
      list=False,
      color_table=None,
      subsolar=False,
      mso=False,
      map_limit=None,
      basemap=None,
      alpha=None,
      title='MAVEN Mars',
      qt=True)
```

*Description:*

Produces a 2D map of Mars, either in the planetocentric or MSO coordinate system, with the MAVEN orbital projection and a variety of basemaps. Spacecraft orbital path may be colored by a given insitu KP data value.

*Required Arguments:*

**kp:** STRUCT  
KP insitu data structure read from file(s).  
**parameter:** STR  
Insitu Key Parameter for setting spacecraft trajectory color.

*Optional Arguments:*

**time:** LIST  
Plots subset of insitu KP data. **time** must be expressed in the format 'YYYY-MM-DD HH:MM:SS'.  
**list:** BOOL  
Display list of all parameters in data structure. All other keywords will be ignored if set.  
**color\_table:** STR, LIST  
Specifies color table(s) to use for plotting.  
**subsolar:** BOOL  
Plot path of subsolar point, not valid for MSO coordinates.  
**mso:** BOOL  
Plot using MSO map projection.  
**map\_limit:** LIST  
Sets the bounding box on the map in lat/lon coordinates: [x0,y0,x1,y1].  
**basemap:** STR

Name of basemap on which the spacecraft data will be overlaid.  
Choices include:

- 'mdim': Mars Digital Image Model
- 'mola': Mars Topography (color)
- 'mola\_bw': Mars Topography (black and white)
- 'mag': Mars Crustal Magnetism
- '<dir\_path>/file.png': User-defined basemap

**alpha:** INT, FLOAT

Sets trajectory transparency, valid between [0,1] inclusive.

**title:** STR

Sets plot title.

**qt:** BOOL

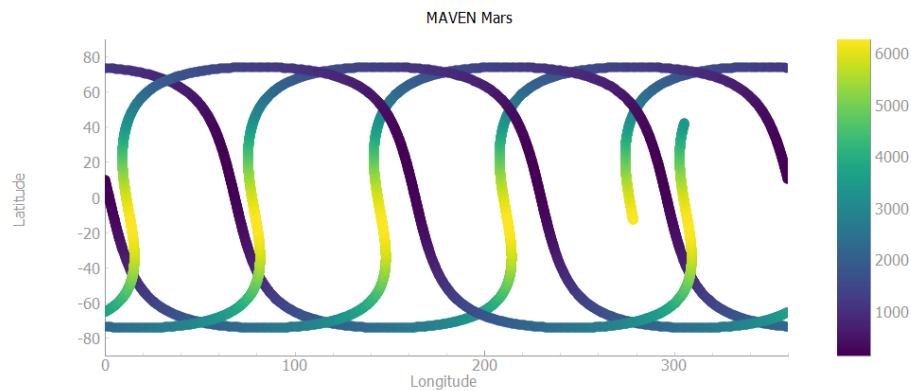
If True, plot with PyQtGraph.

If False, plot with bokeh.

*Examples:*

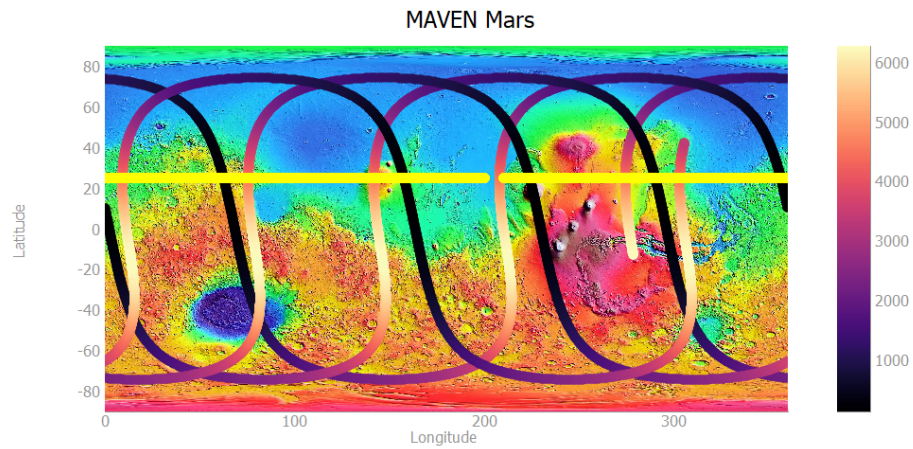
Plot spacecraft altitude along MAVEN surface orbital track.

```
>> pydivide.map2d(insitu,  
                  'spacecraft.altitude')
```



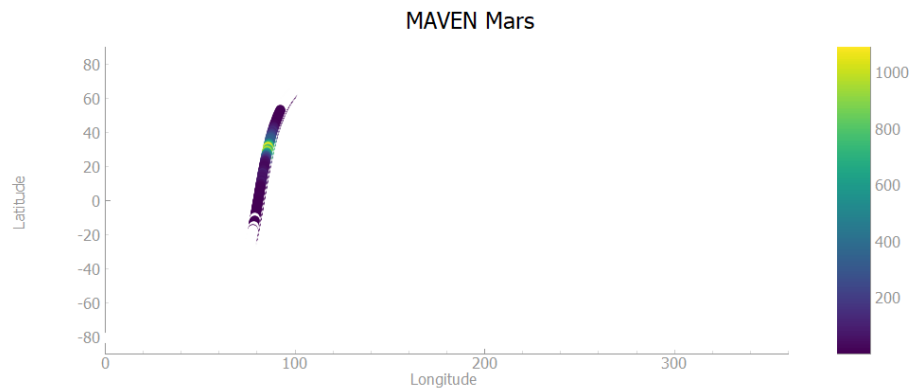
Plot spacecraft altitude along MAVEN surface orbital track using MOLA altimetry basemap; plot subsolar point path.

```
>> pydivide.map2d(insitu,  
                  'spacecraft.altitude',  
                  basemap='mola',  
                  subsolar=True)
```



Plot NGIMS  $\text{CO}_2^+$  density along MAVEN surface orbital track. Limit map to  $\pm 60^\circ$  latitude and  $90^\circ$  to  $270^\circ$  longitude in MSO coordinates.

```
>> pydivide.map2d(insitu,
                   'ngims.co2plus_density',
                   map_limit=[-60,90,60,270],
                   mso=True)
```



## 4.10 occultation

*Function Call:*

```
occultation(iuvs,
            sameplot=True,
            orbit_num=None,
            species=None,
            log=False,
            title='IUVS Occultation Observations',
            qt=True):
```

*Description:*

Plots altitude vs. various species' densities.

*Required Arguments:*

**iuvs:** STRUCT

IUVS Key Parameter data structure returned from `read`.

*Optional Arguments:*

**sameplot:** BOOL

If True, will plot everything on one plot.

If False, will create stacked plots.

**orbit\_num:** INT, LIST

Orbit number(s) to be plotted.

**species:** STR, LIST

Plots only specified chemical species (Appendix A).

**log:** BOOL

Sets logarithmic axes for plotting.

**title:** STR

Sets plot title. Default is 'IUVS Occultation Observations'.

**qt:** BOOL

If True, plot with PyQtGraph.

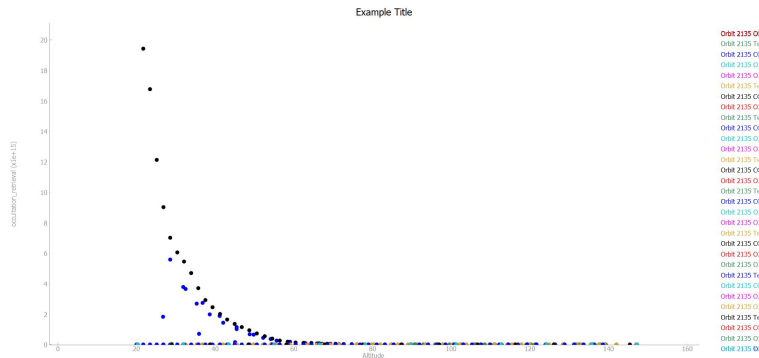
If False, plot with bokeh.

*Examples:*

Plot IUVS data for orbit 2135.

```
>> insitu,iuvs = pydivide.read('2015-11-03','2015-11-04')
```

```
>> pydivide.occultation(iuvs,
                        orbit_num=[2135],
                        title='Example Title')
```



#### 4.11 periapse

*Function Call:*

```
periapse(iuvs,
        sameplot=True,
        density=True,
        radiance=True,
        orbit_num=None,
        species=None,
        obs_num=None,
        log=False,
        title='IUVS Periapse Observations',
        qt=True)
```

*Description:*

Create altitude plots of periapse limb scans from IUVS KP files. These scans contain radiance and density data of various species.

*Required Arguments:*

**iuvs:** STRUCT  
IUVS Key Parameter data structure returned from `read`.

*Optional Arguments:*

**sameplot:** BOOL  
If True, will plot everything on one plot.  
If False, will create stacked plots.

**density:** BOOL  
Plot density.

**radiance:** BOOL  
Plot radiance.

**orbit\_num:** INT, LIST  
Orbit number(s) to be plotted.

**species:** STR, LIST  
Plots only specified chemical species (Appendix A).

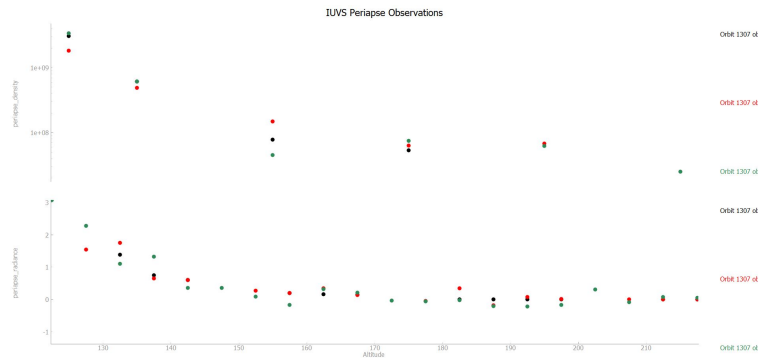
**obs\_num:** INT, LIST  
Observation number(s) to be plotted. There are up to 3 periapse observations

per orbit.  
**log**: BOOL  
 Sets logarithmic axes for plotting.  
**title**: STR  
 Sets plot title. Default is 'IUVS Periapse Observations'.  
**qt**: BOOL  
 If True, plot with PyQtGraph.  
 If False, plot with bokeh.

*Examples:*

Plot Orbit 1307 N<sub>2</sub> density and radiance data.

```
>> pydivide.periapse(iuvs,
                      log=True,
                      species='N2',
                      orbit_num=1307)
```



## 4.12 plot

*Function Call:*

```
plot(kp,
     parameter=None,
     time=None,
     sameplot=True,
     list=False,
     title = '',
     qt=True)
```

*Description:*

Plot time-series data from **insitu** data structure.

*Required Arguments:*

**kp**: STRUCT

KP insitu data structure read from file(s).

**parameter**: INT, STR

Name or index of Key Parameter(s) to be plotted.

*Optional Arguments:*

**time:** [STR,STR]

Plots subset of insitu KP data. **time** must be expressed in the format  
[‘YYYY-MM-DD HH:MM:SS’,‘YYYY-MM-DD HH:MM:SS’].

**sameplot:** BOOL

If True, will plot everything on one plot.

If False, will create stacked plots.

**list:** BOOL

List all KP parameters.

**title:** STR

Sets plot title.

**qt:** BOOL

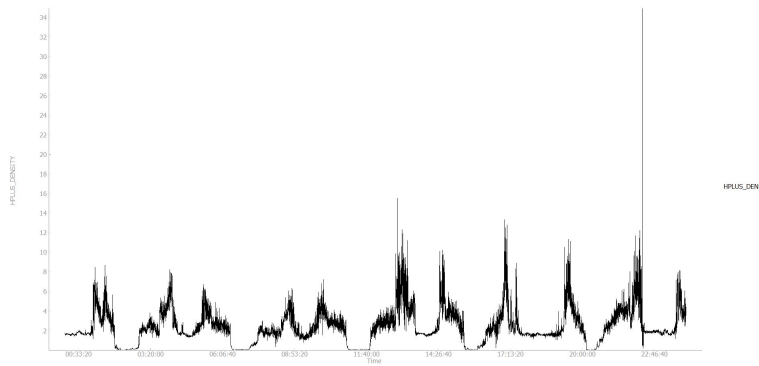
If True, plot with PyQtGraph.

If False, plot with bokeh.

*Examples:*

Plot SWIA H<sup>+</sup> density.

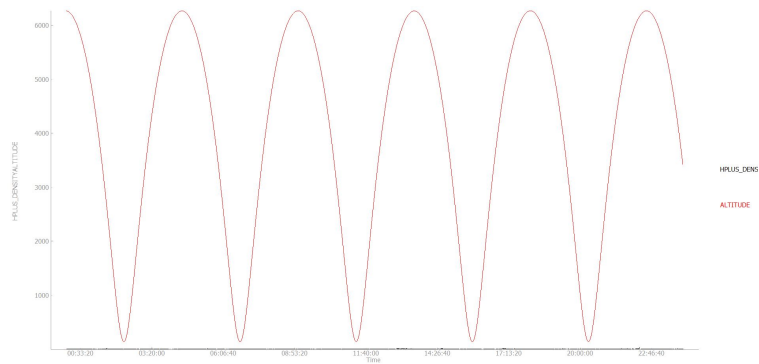
```
>> pydivide.plot(insitu,parameter='swia.hplus_density')
```



Plot SWIA H<sup>+</sup> density and altitude in the same window.

```
>> pydivide.plot(insitu,parameter=['swia.hplus_density',  
                                   'spacecraft.altitude'],sameplot=True)
```



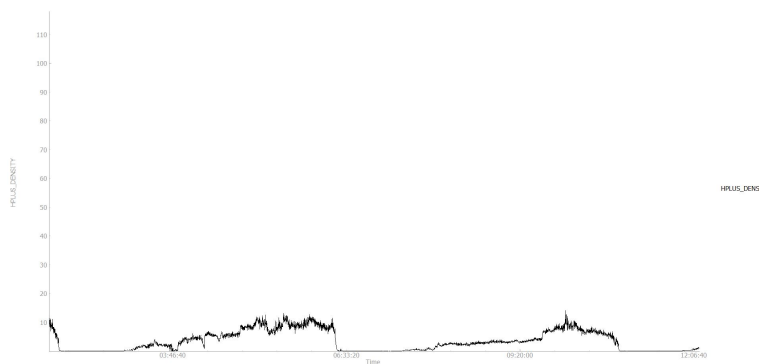


List all KP data parameters.

```
>> pydivide.plot(insitu,list=True)
```

Plot SWIA H<sup>+</sup> density between 02:00 and 12:00 UTC on 2016-04-10.

```
>> pydivide.plot(insitu,parameter='swia.hplus_density',
time=['2016-04-10 02:00:00','2016-04-10 12:00:00'])
```



### 4.13 read

*Function Call:*

```
read(input_time,  
     instruments=None,  
     insitu_only=False)
```

*Description:*

Ingests a subset of local KP data into one or more data structures. These data structures are the primary inputs to various functions in the PyDIVIDE and PyTplot toolkits. Upon first calling the `read` function, the user will be prompted to choose the `root_data_dir`.

*Required Arguments:*

**input\_time:** DATE, [START DATE, END DATE].

The user must provide a time constraint for data retrieval. These constraints must be provided in either of the following formats:

‘YYYY-MM-DD’

‘YYYY-MM-DD HH:MM:SS’

For a single time string, the function will return data for the default time period (1 day = 86400 s), beginning at the specified time. The user may also enter a two-element list, corresponding to beginning and end times.

*Optional Arguments:*

**instruments:** STR, LIST

One or more 3-character instrument abbreviations for instrument-specific data retrieval.

**insitu\_only:** BOOL

Will only read insitu instrument data.

*Returns:*

**insitu:** STRUCT

Contains insitu instrument KP data, as well as spacecraft position and orientation information.

**iuvs:** STRUCT

Contains IUVS instrument KP data.

*Examples:*

Retrieve insitu and IUVS data for LPW and MAG on 2015-12-26.

```
>> insitu,iuvs = pydivide.read('2015-12-26',  
                               instruments=['lpw','mag'])
```

Retrieve only insitu data for all instruments on 2017-06-19.

```
>> insitu = pydivide.read('2017-06-19',  
                           insitu_only=True)
```

## 4.14 read\_model\_results

*Function Call:*

```
read_model_results(file)
```

*Description:*

Reads results of specified simulation into a dictionary object containing sub-directories for metadata, dimension information, and model tracers.

This function can read any of the models currently on the MAVEN SDC website with the .nc extension, which can be found here:

<https://lasp.colorado.edu/maven/sdc/public/pages/models.html>

The desired model must be downloaded prior to running this procedure.

*Required Arguments:*

**file:** STR

Simulation result filename to be read.

*Optional Arguments:*

None.

*Returns:*

Dictionary object of simulation results, structured roughly as follows:

```
output
  meta
    longsubsol
    ls
    ...
  dim
    lat/x
    lon/y
    alt/z
  var1
    dim_order (x,y,z; z,y,x; etc)
    data
  var2
    dim_order
    data
  ...
```

*Examples:*

Read the University of Michigan group's ionospheric model for mean solar activity ( $F_{10.7} = 130$ ).

```
>> model = pydivide.read_model_results(
    '<dir_path>/MGITM_LS270_F130_150519.nc')
```

Read the LATMOS group's ionospheric model for solar maximum levels.

```
>> model = pydivide.read_model_results(
    '<dir_path>/Heliosares_Ionos_Ls270_SolMax1_26_01_15.nc')
```

## 4.15 resample

*Function Call:*

```
resample(kp,  
         time)
```

*Description:*

Modifies KP structure index to user specified time via interpolation.

*Required Arguments:*

**kp:** STRUCT

KP insitu data structure read from file(s).

**time:** LIST

Specifies subset of insitu KP data for resampling. **time** must be expressed in the format 'YYYY-MM-DD HH:MM:SS'.

*Optional Arguments:*

None.

*Examples:*

Resample insitu time to 2016-06-20 coarse survey 3D file time.

```
>> swi_cdf = cdflib.CDF(  
        '<dir_path>/mvn_swi_l2_coarsesvy3d_20160620_v01_r00.cdf')  
>> newtime = swi_cdf.varget('time_unix')  
>> insitu_resampled = pydivide.resample(insitu,  
                                         newtime)
```

## 4.16 standards

*Function Call:*

```
standards(kp,  
          list_plots=False,  
          all_plots=False,  
          euv=False,  
          mag_mso=False,  
          mag_geo=False,  
          mag_cone=False,  
          mag_dir=False,  
          ngims_neutral=False,  
          ngims_ions=False,  
          eph_angle=False,  
          eph_geo=False,  
          eph_mso=False,  
          swea=False,  
          sep_ion=False,  
          sep_electron=False,  
          wave=False,  
          plasma_den=False,
```

```

plasma_temp=False,
swia_h_vel=False,
static_h_vel=False,
static_o2_vel=False,
static_flux=False,
static_energy=False,
sun_bar=False,
solar_wind=False,
ionosphere=False,
sc_pot=False,
altitude=False,
title='Standard Plots',
qt=True)

```

*Description:*

Generate all or a subset of 25 standardized plots, created from insitu KP data on the MAVEN SDC website.

*Required Arguments:*

**kp:** STRUCT

KP insitu data structure read from file(s).

**keyword:** BOOL

At least one or more of the following arguments are required; they may be used in conjunction with one another.

**all\_plots:** generate all 25 plots

**euv:** EUV irradiance in each of 3 bands

**mag\_mso:** magnetic field, MSO coordinates

**mag\_geo:** magnetic field, geographic coordinates

**mag\_cone:** magnetic clock and cone angles, MSO coordinates

**mag\_dir:** magnetic field, radial/horizontal/northward/eastward components

**ngims\_neutral:** neutral atmospheric component densities

**ngims\_ions:** ionized atmospheric component densities

**eph\_angle:** spacecraft ephemeris information

**eph\_geo:** spacecraft position, geographic coordinates

**eph\_mso:** spacecraft position, MSO coordinates

**swea:** electron parallel/anti-parallel fluxes

**sep\_ion:** ion energy flux

**sep\_electron:** electron energy flux

**wave:** electric field wave power

**plasma\_den:** plasma density

**plasma\_temp:** plasma temperature

**swia\_h\_vel:** H<sup>+</sup> flow velocity, SWIA MSO coordinates

**static\_h\_vel:** H<sup>+</sup> flow velocity, STATIC MSO coordinates

**static\_o2\_vel:** O<sub>2</sub><sup>+</sup> flow velocity, STATIC MSO coordinates

**static\_flux:** H<sup>+</sup>/H<sup>++</sup> and pick-up ion omni-directional flux

**static\_energy:** H<sup>+</sup>/H<sup>++</sup> and pick-up ion characteristic energy

**sun\_bar:** MAVEN sunlight indicator

**solar\_wind:** solar wind dynamic pressure

**ionosphere:** electron spectrum shape parameter  
**altitude:** spacecraft altitude  
**sc\_pot:** spacecraft potential

*Optional Arguments:*

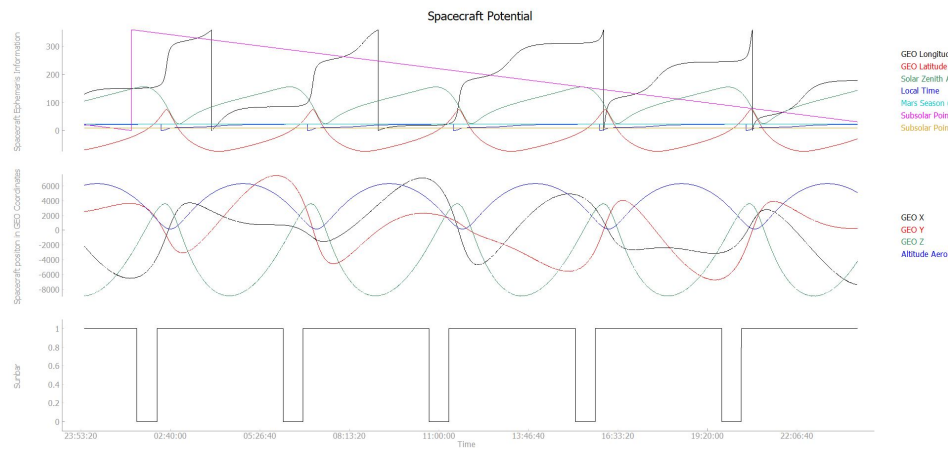
**list\_plots:** BOOL  
 Display list and description of all available plots.  
**title:** STR  
 Title name for all plots.  
**qt:** BOOL  
 If True, plot with PyQtGraph.  
 If False, plot with bokeh.

*Examples:*

Plot all 25 standard plots. **Note:** not recommended for general use as this will generate 25 very narrow plots; may be useful for a quick glance.

```

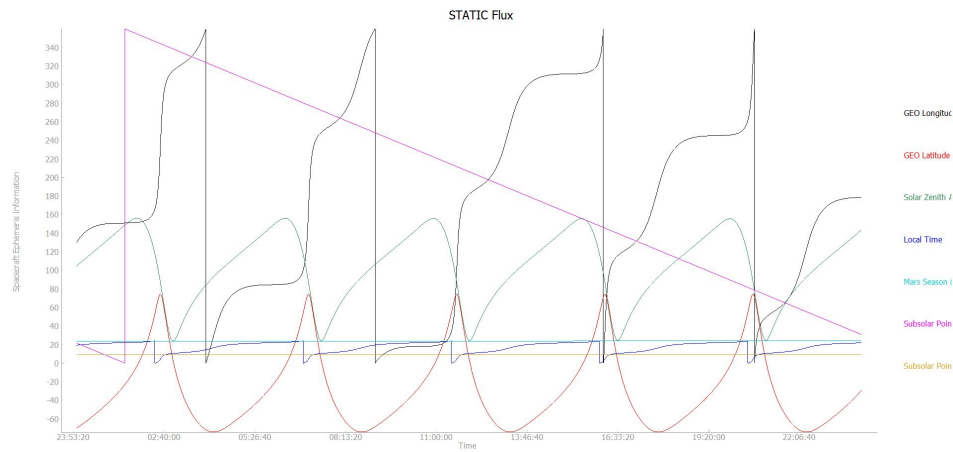
>> insitu,iuvs = pydivide.read('2017-06-19','2017-06-20')
>> pydivide.standards(insitu,
                        all_plots=True)
  
```



Generate figure containing 3 plots: magnetic field standard plot in Mars Solar Orbital coordinates (x, y, z, magnitude), standard spacecraft ephemeris information (sub-spacecraft lat/lon, subsolar lat/lon, local solar time, solar zenith angle, Mars season), and STATIC H<sup>+</sup>/H<sup>++</sup> and pick-up ion omni-directional flux.

```

>> insitu,iuvs = pydivide.read('2017-06-19','2017-06-20')
>> pydivide.standards(insitu,
                        mag_mso=True,
                        eph_angle=True,
                        static_flux=True,
                        title='Example Title')
  
```



#### 4.17 tplot\_varcreate

*Function Call:*

```
tplot_varcreate(kp)
```

*Description:*

Creates TVars (tplot variables) for each instrument, named with the format: `mvn_kp::instrument::observation`.

*Required Arguments:*

**kp:** STRUCT

KP insitu data structure read from file(s).

*Optional Arguments:*

None.

*Examples:*

Create TVars for all instruments: EUV, LPW, STATIC, SWEA, SWIA, MAG, SEP, and NGIMS, with data from 2017-06-19.

```
>> insitu = pydivide.read('2017-06-19')
>> pydivide.tplot_varcreate(insitu)
```

## A Chemical Species

- C
- C\_1561
- C\_1657
- CO Cameron
- CO2
- CO2+
- CO2pUVD
- H
- N
- N2
- N\_1493
- NO
- O
- O2
- O3
- O\_1304
- O\_1356
- O\_2972

## B KP Data Structures

Insitu parameters are callable either by name or number listed below.  
IUVS parameters are only callable by name.

### B.1 INSITU

1. TimeString
2. Time
3. Orbit
4. IOflag



### **B.1.1 LPW**

5. LPW.ELECTRON\_DENSITY
6. LPW.ELECTRON\_DENSITY\_QUAL\_MIN
7. LPW.ELECTRON\_DENSITY\_QUAL\_MAX
8. LPW.ELECTRON\_TEMPERATURE
9. LPW.ELECTRON\_TEMPERATURE\_QUAL\_MIN
10. LPW.ELECTRON\_TEMPERATURE\_QUAL\_MAX
11. LPW.SPACECRAFT\_POTENTIAL
12. LPW.SPACECRAFT\_POTENTIAL\_QUAL\_MIN
13. LPW.SPACECRAFT\_POTENTIAL\_QUAL\_MAX
14. LPW.EWAVE\_LOW\_FREQ
15. LPW.EWAVE\_LOW\_FREQ\_QUAL
16. LPW.EWAVE\_MID\_FREQ
17. LPW.EWAVE\_MID\_FREQ\_QUAL
18. LPW.EWAVE\_HIGH\_FREQ
19. LPW.EWAVE\_HIGH\_FREQ\_QUAL

### **B.1.2 EUV**

20. EUV.IRRADIANCE\_LOW
21. EUV.IRRADIANCE\_LOW\_QUAL
22. EUV.IRRADIANCE\_MID
23. EUV.IRRADIANCE\_MID\_QUAL
24. EUV.IRRADIANCE\_LYMAN
25. EUV.IRRADIANCE\_LYMAN\_QUAL

### B.1.3 SWEA

- 26. SWEA.SOLAR\_WIND\_ELECTRON\_DENSITY
- 27. SWEA.SOLAR\_WIND\_ELECTRON\_DENSITY\_QUAL
- 28. SWEA.SOLAR\_WIND\_ELECTRON\_TEMPERATURE
- 29. SWEA.SOLAR\_WIND\_ELECTRON\_TEMPERATURE\_QUAL
- 30. SWEA.ELECTRON\_PARALLEL\_FLUX\_LOW
- 31. SWEA.ELECTRON\_PARALLEL\_FLUX\_LOW\_QUAL
- 32. SWEA.ELECTRON\_PARALLEL\_FLUX\_MID
- 33. SWEA.ELECTRON\_PARALLEL\_FLUX\_MID\_QUAL
- 34. SWEA.ELECTRON\_PARALLEL\_FLUX\_HIGH
- 35. SWEA.ELECTRON\_PARALLEL\_FLUX\_HIGH\_QUAL
- 36. SWEA.ELECTRON\_ANTI\_PARALLEL\_FLUX\_LOW
- 37. SWEA.ELECTRON\_ANTI\_PARALLEL\_FLUX\_LOW\_QUAL
- 38. SWEA.ELECTRON\_ANTI\_PARALLEL\_FLUX\_MID
- 39. SWEA.ELECTRON\_ANTI\_PARALLEL\_FLUX\_MID\_QUAL
- 40. SWEA.ELECTRON\_ANTI\_PARALLEL\_FLUX\_HIGH
- 41. SWEA.ELECTRON\_ANTI\_PARALLEL\_FLUX\_HIGH\_QUAL
- 42. SWEA.ELECTRON\_SPECTRUM\_SHAPE\_PARAMETER
- 43. SWEA.ELECTRON\_SPECTRUM\_SHAPE\_PARAMETER\_QUAL

### B.1.4 SWIA

- 44. SWIA.HPLUS\_DENSITY
- 45. SWIA.HPLUS\_DENSITY\_QUAL
- 46. SWIA.HPLUS\_FLOW\_VELOCITY\_MSO\_X
- 47. SWIA.HPLUS\_FLOW\_VELOCITY\_MSO\_X\_QUAL
- 48. SWIA.HPLUS\_FLOW\_VELOCITY\_MSO\_Y
- 49. SWIA.HPLUS\_FLOW\_VELOCITY\_MSO\_Y\_QUAL
- 50. SWIA.HPLUS\_FLOW\_VELOCITY\_MSO\_Z

- 51. SWIA.HPLUS\_FLOW\_VELOCITY\_MSO\_Z\_QUAL
- 52. SWIA.HPLUS\_TEMPERATURE
- 53. SWIA.HPLUS\_TEMPERATURE\_QUAL
- 54. SWIA.SOLAR\_WIND\_DYNAMIC\_PRESSURE
- 55. SWIA.SOLAR\_WIND\_DYNAMIC\_PRESSURE\_QUAL

#### **B.1.5 STATIC**

- 56. STATIC.STATIC\_QUALITY\_FLAG
- 57. STATIC.HPLUS\_DENSITY
- 58. STATIC.HPLUS\_DENSITY\_QUAL
- 59. STATIC.OPLUS\_DENSITY
- 60. STATIC.OPLUS\_DENSITY\_QUAL
- 61. STATIC.O2PLUS\_DENSITY
- 62. STATIC.O2PLUS\_DENSITY\_QUAL
- 63. STATIC.HPLUS\_TEMPERATURE
- 64. STATIC.HPLUS\_TEMPERATURE\_QUAL
- 65. STATIC.OPLUS\_TEMPERATURE
- 66. STATIC.OPLUS\_TEMPERATURE\_QUAL
- 67. STATIC.O2PLUS\_TEMPERATURE
- 68. STATIC.O2PLUS\_TEMPERATURE\_QUAL
- 69. STATIC.O2PLUS\_FLOW\_VELOCITY\_MAVEN\_APP\_X
- 70. STATIC.O2PLUS\_FLOW\_VELOCITY\_MAVEN\_APP\_X\_QUAL
- 71. STATIC.O2PLUS\_FLOW\_VELOCITY\_MAVEN\_APP\_Y
- 72. STATIC.O2PLUS\_FLOW\_VELOCITY\_MAVEN\_APP\_Y\_QUAL
- 73. STATIC.O2PLUS\_FLOW\_VELOCITY\_MAVEN\_APP\_Zz
- 74. STATIC.O2PLUS\_FLOW\_VELOCITY\_MAVEN\_APP\_Z\_QUAL
- 75. STATIC.O2PLUS\_FLOW\_VELOCITY\_MSO\_X
- 76. STATIC.O2PLUS\_FLOW\_VELOCITY\_MSO\_X\_QUAL

77. STATIC.O2PLUS\_FLOW\_VELOCITY\_MSO\_Y  
78. STATIC.O2PLUS\_FLOW\_VELOCITY\_MSO\_Y\_QUAL  
79. STATIC.O2PLUS\_FLOW\_VELOCITY\_MSO\_Z  
80. STATIC.O2PLUS\_FLOW\_VELOCITY\_MSO\_Z\_QUAL  
81. STATIC.HPLUS\_OMNI\_DIRECTIONAL\_FLUX  
82. STATIC.HPLUS\_CHARACTERISTIC\_ENERGY  
83. STATIC.HPLUS\_CHARACTERISTIC\_ENERGY\_QUAL  
84. STATIC.HEPLUS\_OMNI\_DIRECTIONAL\_FLUX  
85. STATIC.HEPLUS\_CHARACTERISTIC\_ENERGY  
86. HSTATIC.HEPLUS\_CHARACTERISTIC\_ENERGY\_QUAL  
87. STATIC.OPLUS\_OMNI\_DIRECTIONAL\_FLUX  
88. STATIC.OPLUS\_CHARACTERISTIC\_ENERGY  
89. STATIC.OPLUS\_CHARACTERISTIC\_ENERGY\_QUAL  
90. STATIC.O2PLUS\_OMNI\_DIRECTIONAL\_FLUX  
91. STATIC.O2PLUS\_CHARACTERISTIC\_ENERGY  
92. STATIC.O2PLUS\_CHARACTERISTIC\_ENERGY\_QUAL  
93. STATIC.HPLUS\_CHARACTERISTIC\_DIRECTION\_MSO\_X  
94. STATIC.HPLUS\_CHARACTERISTIC\_DIRECTION\_MSO\_Y  
95. STATIC.HPLUS\_CHARACTERISTIC\_DIRECTION\_MSO\_Z  
96. STATIC.HPLUS\_CHARACTERISTIC\_ANGULAR\_WIDTH  
97. STATIC.HPLUS\_CHARACTERISTIC\_ANGULAR\_WIDTH\_QUAL  
98. STATIC.DOMINANT\_PICKUP\_ION\_CHARACTERISTIC\_DIRECTION\_MSO\_X  
99. STATIC.DOMINANT\_PICKUP\_ION\_CHARACTERISTIC\_DIRECTION\_MSO\_Y  
100. STATIC.DOMINANT\_PICKUP\_ION\_CHARACTERISTIC\_DIRECTION\_MSO\_Z  
101. STATIC.DOMINANT\_PICKUP\_ION\_CHARACTERISTIC\_ANGULAR\_WIDTH  
102. STATIC.DOMINANT\_PICKUP\_ION\_CHARACTERISTIC\_ANGULAR\_WIDTH\_QUAL

### B.1.6 SEP

- 103. SEP.ION\_ENERGY\_FLUX\_\_FOV\_1\_F
- 104. SEP.ION\_ENERGY\_FLUX\_\_FOV\_1\_F\_QUAL
- 105. SEP.ION\_ENERGY\_FLUX\_\_FOV\_1\_R
- 106. SEP.ION\_ENERGY\_FLUX\_\_FOV\_1\_R\_QUAL
- 107. SEP.ION\_ENERGY\_FLUX\_\_FOV\_2\_F
- 108. SEP.ION\_ENERGY\_FLUX\_\_FOV\_2\_F\_QUAL
- 109. SEP.ION\_ENERGY\_FLUX\_\_FOV\_2\_R
- 110. SEP.ION\_ENERGY\_FLUX\_\_FOV\_2\_R\_QUAL
- 111. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_1\_F
- 112. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_1\_F\_QUAL
- 113. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_1\_R
- 114. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_1\_R\_QUAL
- 115. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_2\_F
- 116. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_2\_F\_QUAL
- 117. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_2\_R
- 118. SEP.ELECTRON\_ENERGY\_FLUX\_\_\_FOV\_2\_R\_QUAL
- 119. SEP.LOOK\_DIRECTION\_1\_F\_MSO\_X
- 120. SEP.LOOK\_DIRECTION\_1\_F\_MSO\_Y
- 121. SEP.LOOK\_DIRECTION\_1\_F\_MSO\_Z
- 122. SEP.LOOK\_DIRECTION\_1\_R\_MSO\_X
- 123. SEP.LOOK\_DIRECTION\_1\_R\_MSO\_Y
- 124. SEP.LOOK\_DIRECTION\_1\_R\_MSO\_Z
- 125. SEP.LOOK\_DIRECTION\_2\_F\_MSO\_X
- 126. SEP.LOOK\_DIRECTION\_2\_F\_MSO\_Y
- 127. SEP.LOOK\_DIRECTION\_2\_F\_MSO\_Z
- 128. SEP.LOOK\_DIRECTION\_2\_R\_MSO\_X
- 129. SEP.LOOK\_DIRECTION\_2\_R\_MSO\_Y
- 130. SEP.LOOK\_DIRECTION\_2\_R\_MSO\_Z

### **B.1.7 MAG**

- 131. MAG.MSO\_X
- 132. MAG.MSO\_X\_QUAL
- 133. MAG.MSO\_Y
- 134. MAG.MSO\_Y\_QUAL
- 135. MAG.MSO\_Z
- 136. MAG.MSO\_Z\_QUAL
- 137. MAG.GEO\_X
- 138. MAG.GEO\_X\_QUAL
- 139. MAG.GEO\_Y
- 140. MAG.GEO\_Y\_QUAL
- 141. MAG.GEO\_Z
- 142. MAG.GEO\_Z\_QUAL
- 143. MAG.RMS\_DEVIATION
- 144. MAG.RMS\_DEVIATION\_QUAL

### **B.1.8 NGIMS**

- 145. NGIMS.HE\_DENSITY
- 146. NGIMS.HE\_DENSITY\_PRECISION
- 147. NGIMS.HE\_DENSITY\_QUAL
- 148. NGIMS.O\_DENSITY
- 149. NGIMS.O\_DENSITY\_PRECISION
- 150. NGIMS.O\_DENSITY\_QUAL
- 151. NGIMS.CO\_DENSITY
- 152. NGIMS.CO\_DENSITY\_PRECISION
- 153. NGIMS.CO\_DENSITY\_QUAL
- 154. NGIMS.N2\_DENSITY
- 155. NGIMS.N2\_DENSITY\_PRECISION

156. NGIMS.N2\_DENSITY\_QUAL  
157. NGIMS.NO\_DENSITY  
158. NGIMS.NO\_DENSITY\_PRECISION  
159. NGIMS.NO\_DENSITY\_QUAL  
160. NGIMS.AR\_DENSITY  
161. NGIMS.AR\_DENSITY\_PRECISION  
162. NGIMS.AR\_DENSITY\_QUAL  
163. NGIMS.CO2\_DENSITY  
164. NGIMS.CO2\_DENSITY\_PRECISION  
165. NGIMS.CO2\_DENSITY\_QUAL  
166. NGIMS.O2PLUS\_DENSITY  
167. NGIMS.O2PLUS\_DENSITY\_PRECISION  
168. NGIMS.O2PLUS\_DENSITY\_QUAL  
169. NGIMS.CO2PLUS\_DENSITY  
170. NGIMS.CO2PLUS\_DENSITY\_PRECISION  
171. NGIMS.CO2PLUS\_DENSITY\_QUAL  
172. NGIMS.NOPLUS\_DENSITY  
173. NGIMS.NOPLUS\_DENSITY\_PRECISION  
174. NGIMS.NOPLUS\_DENSITY\_QUAL  
175. NGIMS.OPLUS\_DENSITY  
176. NGIMS.OPLUS\_DENSITY\_PRECISION  
177. NGIMS.OPLUS\_DENSITY\_QUAL  
178. NGIMS.CO2PLUS\_N2PLUS\_DENSITY  
179. NGIMS.CO2PLUS\_N2PLUS\_DENSITY\_PRECISION  
180. NGIMS.CO2PLUS\_N2PLUS\_DENSITY\_QUAL  
181. NGIMS.CPLUS\_DENSITY  
182. NGIMS.CPLUS\_DENSITY\_PRECISION  
183. NGIMS.CPLUS\_DENSITY\_QUAL

- 184. NGIMS.OHPLUS\_DENSITY
- 185. NGIMS.OHPLUS\_DENSITY\_PRECISION
- 186. NGIMS.OHPLUS\_DENSITY\_QUAL
- 187. NGIMS.NPLUS\_DENSITY
- 188. NGIMS.NPLUS\_DENSITY\_PRECISION
- 189. NGIMS.NPLUS\_DENSITY\_QUAL

#### **B.1.9 APP**

- 190. APP.ATTITUDE\_GEO\_X
- 191. APP.ATTITUDE\_GEO\_Y
- 192. APP.ATTITUDE\_GEO\_Z
- 193. APP.ATTITUDE\_MSO\_X
- 194. APP.ATTITUDE\_MSO\_Y
- 195. APP.ATTITUDE\_MSO\_Z

#### **B.1.10 SPACECRAFT**

- 196. SPACECRAFT.GEO\_X
- 197. SPACECRAFT.GEO\_Y
- 198. SPACECRAFT.GEO\_Z
- 199. SPACECRAFT.MSO\_X
- 200. SPACECRAFT.MSO\_Y
- 201. SPACECRAFT.MSO\_Z
- 202. SPACECRAFT.SUB\_SC\_LONGITUDE
- 203. SPACECRAFT.SUB\_SC\_LATITUDE
- 204. SPACECRAFT.SZA
- 205. SPACECRAFT.LOCAL\_TIME
- 206. SPACECRAFT.ALTITUDE
- 207. SPACECRAFT.ATTITUDE\_GEO\_X



208. SPACECRAFT.ATTITUDE\_GEO\_Y  
209. SPACECRAFT.ATTITUDE\_GEO\_Z  
210. SPACECRAFT.ATTITUDE\_MSO\_X  
211. SPACECRAFT.ATTITUDE\_MSO\_Y  
212. SPACECRAFT.ATTITUDE\_MSO\_Z  
213. SPACECRAFT.MARS\_SEASON  
214. SPACECRAFT.MARS\_SUN\_DISTANCE  
215. SPACECRAFT.SUBSOLAR\_POINT\_GEO\_LONGITUDE  
216. SPACECRAFT.SUBSOLAR\_POINT\_GEO\_LATITUDE  
217. SPACECRAFT.SUBMARS\_POINT\_SOLAR\_LONGITUDE  
218. SPACECRAFT.SUBMARS\_POINT\_SOLAR\_LATITUDE  
219. SPACECRAFT.T11  
220. SPACECRAFT.T12  
221. SPACECRAFT.T13  
222. SPACECRAFT.T21  
223. SPACECRAFT.T22  
224. SPACECRAFT.T23  
225. SPACECRAFT.T31  
226. SPACECRAFT.T32  
227. SPACECRAFT.T33  
228. SPACECRAFT.SPACECRAFT\_T11  
229. SPACECRAFT.SPACECRAFT\_T12  
230. SPACECRAFT.SPACECRAFT\_T13  
231. SPACECRAFT.SPACECRAFT\_T21  
232. SPACECRAFT.SPACECRAFT\_T22  
233. SPACECRAFT.SPACECRAFT\_T23  
234. SPACECRAFT.SPACECRAFT\_T31  
235. SPACECRAFT.SPACECRAFT\_T32  
236. SPACECRAFT.SPACECRAFT\_T33

## B.2 IUVS

### B.2.1 ALL

- TIME\_START
- TIME\_STOP
- SZA
- LOCAL\_TIME
- LAT
- LON
- LAT\_MSO
- LON\_MSO
- ORBIT\_NUMBER
- MARS\_SEASON\_LS
- SPACECRAFT\_GEO
- SPACECRAFT\_MSO
- SUN\_GEO
- SPACECRAFT\_GEO\_LONGITUDE
- SPACECRAFT\_GEO\_LATITUDE
- SPACECRAFT\_MSO\_LONGITUDE
- SPACECRAFT\_MSO\_LATITUDE
- SUBSOLAR\_POINT\_GEO\_LONGITUDE
- SUBSOLAR\_POINT\_GEO\_LATITUDE
- SPACECRAFT\_SZA
- SPACECRAFT\_LOCAL\_TIME
- SPACECRAFT\_ALTITUDE
- MARS\_SUN\_DISTANCE

### **B.2.2 PERIAPSE1/PERIAPSE2/PERIAPSE3**

- SCALE\_HEIGHT
- DENSITY
- RADIANCE
- TEMPERATURE
- ALT

### **B.2.3 APOAPSE**

- OZONE\_DEPTH
- AURORAL\_INDEX
- DUST\_DEPTH
- RADIANCE
- SZA\_BP
- LOCAL\_TIME\_BP
- LON\_BINS
- LAT\_BINS

### **B.2.4 CORONA\_LORES\_HIGH**

- HALF\_INT\_DISTANCE
- TEMPERATURE
- DENSITY
- RADIANCE
- ALT

### **B.2.5 OCCULTATION**

- CO2
- O2
- O3
- TEMPERATURE