Parton Showers

Alexander Huss

July 26, 2023

Contents

1	Introduction	1
2	Emission probability and the Sudakov form factor	1
3	Implementation	2

1 Introduction

We will investigate the emission probability of gluons off quarks and gluons and use that to implement a **very** simplified parton shower (only final state primary branching, leading double-log, only virtuality q^2 and not proper kinematics, ...).

2 Emission probability and the Sudakov form factor

In the leading double-log approximation (soft and collinear emission), we have seen in the lecture that the emission propability is given as

$$d\omega_{X\to X+g} = 2\frac{\alpha_s}{\pi} C_X \frac{dE}{E} \frac{d\theta}{\theta}, \qquad (1)$$

where E denotes the energy of the emitted gluon and θ the angle w.r.t. the parent particle. We denote the emitting particle by "X" and C_X is the associated colour factor. For quarks, $C_X = C_F = \frac{4}{3}$ and for gluons $C_X = C_A = 3$.

For any parton shower, we first need to fix the evolution variable w.r.t. which we want to generate emissions. To this end, we choose the virtuality q^2 associated with the emission for which we find

$$d\mathcal{P} = \frac{\alpha_s}{\pi} C_X \frac{\mathrm{d}q^2}{q^2} \ln\left(\frac{q^2}{Q_0^2}\right) \tag{2}$$

where Q_0 denotes a cutoff below which emissions are considered unresolved. We next define the so-called Sudakov form factor $\Delta(Q^2, q^2)$, which is the probability for no resolved

emissions to happen between $Q^2 \to q^2$. It saitisfies a differential equation reminiscent of radiative decay with a simple solution

$$\frac{d\Delta(Q^2, q^2)}{dq^2} = \Delta(Q^2, q^2) \frac{d\mathcal{P}}{dq^2}, \quad \Delta(Q^2) \equiv \Delta(Q^2, Q_0^2) = \exp\left\{-\frac{\alpha_s C_X}{2\pi} \ln^2\left(\frac{q^2}{Q_0^2}\right)\right\}$$
(3)

3 Implementation

With the Sudakov form factor at hand, we can easily iterate the sampling of emissions using these steps:

- 1. set $Q = Q_{\text{start}}$
- 2. draw a uniform random number r in the range [0, 1]
- 3. if $r < \Delta(Q^2)$, no resolvable emission can be generated ($< Q_0$): Terminate loop.
- 4. solve $r = \Delta(Q^2)/\Delta(Q_{\text{new}}^2)$ for Q_{new} , which is the new emission scale.
- 5. set $Q = Q_{\text{new}}$ and go back to step 2.

```
#!/usr/bin/env python
import math
import random
import sys
random.seed(42)
alphas = 0.118
def generate_event(Q2_start: float, Q2_cutoff: float, CX: float):
 sudakov = 1. # initialize Sudakov to the starting scale
  fac = alphas*CX/(2.*math.pi)
  Qlist = []
  while True:
   r = random.uniform(0.,1.)
    \#> sudakov = exp(-[alphas*CX/(2.*pi)]*log^2[Q2/Q2_start])
    #> determine Q2 from the associated sudakov
   L2 = - math.log(sudakov) / fac
   Q2 = Q2\_start * math.exp(-math.sqrt(L2))
   if Q2 < Q2_cutoff:
     break
   Qlist.append( math.sqrt(Q2) )
   print("#summary2 {} {} {} {} {}".format(len(Qlist),sum(Qlist),Qlist[0],Qlist))
if __name__ == "__main__":
 if len(sys.argv) < 3:</pre>
   raise RuntimeError("I expect at least two arguments: Q_start [g|q]")
  Q_start = float(sys.argv[1]) # the hard scale
  Q_cutoff = 1 # shower cutoff (PS stops -> hand over to hadronization)
  if sys.argv[2] == "q":
   CX = 4./3. # quark
  elif sys.argv[2] == "g":
```

```
CX = 3.  # gluon
else:
    raise RuntimeError("unrecognised parton: {}".format(sys.argv[2]))
if len(sys.argv) >= 4:
    alphas = float(sys.argv[3])
if len(sys.argv) >= 5:
    nevents = int(sys.argv[4])
else:
    nevents = 1000
for i in range(nevents):
    print("# event {} {{} {{} {} {} {}} {{} {}} {{}}".format(i,Q_start,sys.argv[2],CX,alphas,nevents))
    generate_event(Q_start**2, Q_cutoff**2, CX)
```

Let's use the implementation to generate some "events"

We can see that the all-order description damps the divergent behaviour of a pure fixed-order prediction for $Q \to 0$. Given $C_A > C_F$, we also see how a gluon generates more emissions than quarks. This property can be exploited to try and discriminate between "quark jets" and "gluon jets".



