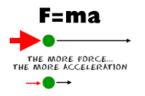
Quantum many-body systems (8.513 fa19) Lecture note 1

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https://stellar.mit.edu/S/course/8/fa19/8.513/index.html

Classical motion of a particle and Newton's Law

The motion of electrons or holes in a semiconductor does not follow Newton's law. They follow a generalized Newton law.





First-order equation of motion and phase-space Lagrangian

• If (x, p) fully characterize the state of a particle, then their equation of motion is first-order:

$$\dot{x} = \partial_p H(x, p), \quad \dot{p} = -\partial_x H(x, p)$$
 Why this form?

which can be obtained via phase-space Lagrangian

$$\mathcal{L}(x,\dot{x},p,\dot{p})=p\dot{x}-H(x,p),\quad S=\int \mathrm{d}t\ \mathcal{L}(x,\dot{x},p,\dot{p}).$$

- A classical system is fully characterized by 1) EOM + Hamiltonian, or by 2) phase-space Lagrangian.
- A phase-space point fully characterises a classical state.
- Phase-space Lagrangian contains only first order time derivative.
- From *S* to first-order equation of motion

$$\delta S = \int dt \, \delta p \underbrace{\left[\dot{x} - \partial_p H(x, p)\right]}_{=0} + \delta x \underbrace{\left[-\dot{p} - \partial_x H(x, p)\right]}_{=0},$$

we got that above equation of motion.

Phase-space Lagrangian description of Shrödinger equation

For a quantum system, its state is fully characterized by a vector ϕ in a Hilbert space \mathcal{V} :



$$|\phi\rangle = \begin{pmatrix} \phi_1 \\ \phi_2 \\ \vdots \end{pmatrix}, \ o \ \ \text{first-order E.O.M} \ \ i\hbar\dot{\phi}_{\textit{m}} = \textit{H}_{\textit{mn}}\phi_{\textit{n}}$$

(Why ϕ_m is complex? Why $|\phi_m|^2$ related to probability?)

Phase-space Lagrangian

$$L = \mathrm{i}\,\hbar\phi_{m}^{*}\dot{\phi}_{m} - \phi_{m}^{*}H_{mn}\phi_{n} = \langle\phi|\mathrm{i}\,\hbar\frac{\mathrm{d}}{\mathrm{d}t} - H|\phi\rangle, \qquad S = \int \mathrm{d}t \ L.$$

• From (Can we have non-linear Shrödinger equation?)

$$\delta S = \int dt \, \delta \phi_m^* [i\hbar \dot{\phi}_m - H_{mn}\phi_n] + \delta \phi_n [-i\hbar \dot{\phi}_m^* - \phi_m^* H_{mn}]$$

we get the equation of motion

$$i\hbar\dot{\phi}_m = H_{mn}\phi_n, \quad -i\hbar\dot{\phi}_n^* = \phi_m^*H_{mn}.$$

Dynamical variational approach

- Given a Hamiltonian H, we can use variational approach to get an approximate ground state, by minimizing $\langle \phi_{\xi^I} | H | \phi_{\xi^I} \rangle$, where ξ^I are the variational parameters \to approximate ground state $|\phi_{\xi^I_0}\rangle$. But how to get the low energy excited states?
- Dynamical variational approach (semi-classical approach):
- we assume the variational parameters has a time-dependence $\xi^{I}(t)$.
- The variational parameters ξ^I fully characterize the state, $ie \ \xi^I$ parametrize a phase-space.
- The dynamics of $\xi'(t)$ is given by the phase-space Lagrangian

$$\mathcal{L}(\xi^I, \dot{\xi}^I) = \langle \phi_{\xi^I(t)} | i\hbar \frac{\mathrm{d}}{\mathrm{d}t} - H | \phi_{\xi^I(t)} \rangle = \hbar a_I(\xi^I) \dot{\xi}^I - \bar{H}(\xi^I)$$

where

$$a_I(\xi^I) \equiv i \langle \phi_{\xi^I} | \partial_{\xi^I} | \phi_{\xi^I} \rangle,$$

which is the vector potential in the phase-space.

Most general phase-space description of classical system

From
$$S = \int dt \ L(\dot{\xi}^I, \xi^I) = \int dt \ [\hbar a_I \dot{\xi}^I - \bar{H}]$$
, we get
$$\delta S = \int dt \ [\hbar (\partial_J a_I) \delta \xi^J \dot{\xi}^I - \hbar \dot{a}_I \delta \xi^I - \delta \xi^I \partial_I \bar{H}(\xi^I)]$$
$$= \int dt \ \delta \xi^I [\hbar (\partial_I a_J) \dot{\xi}^J - \hbar (\partial_J a_I) \dot{\xi}^J - \partial_I \bar{H}] = \int dt \ \delta \xi^I [\hbar b_{IJ} \dot{\xi}^J - \partial_I \bar{H}]$$

and the equation of motion

$$\hbar b_{IJ}\dot{\xi}^J = \frac{\partial \bar{H}}{\partial \xi^I}, \qquad b_{IJ} = \partial_I a_J - \partial_J a_I = \text{"magnetic field" in phase-space}$$

- The above EOM conserve energy $\partial_t \bar{H}(\xi^I(t)) = 0$.
- Gauge redundancy: we may choose an equivalent (redundant) trial wave function $e^{i\theta(\xi^l)}|\psi_{\xi^l}\rangle$. We will get

$$L(\dot{\xi}^I, \xi^I) = \hbar a_I \dot{\xi}^I - \dot{\theta}(\xi^I) - \bar{H}(\xi^I) = \hbar [a_I - \partial_I \theta] \dot{\xi}^I - \bar{H}(\xi^I)$$

which gives rise to the same EOM.

Change the phase space Lagrangian by a total time derivative of any function does not change the EOM.

Gauge "symmetry" and symmetry

Gauge redundancy (also called gauge symmetry by mistake) and **symmetry** (real physical symmetry) in quantum system:

- If we give a single quantum state two names $|a\rangle$ and $|b\rangle$, then $|a\rangle$ and $|b\rangle$ will have the same properties (since $|a\rangle = |b\rangle$). We say there is a gauge redundancy or gauge symmetry, and the theory of $|a\rangle$ and $|b\rangle$ is a gauge theory.
- If two orthogonal states $|a\rangle$ and $|b\rangle$ same properties, then we say there is a symmetry between $|a\rangle$ and $|b\rangle$ (since $\langle a|b\rangle = 0$).

Gauge "symmetry" is indeed a symmetry in classical system

Change of variables

• If we change the variables to $\eta^{l} = \eta^{l}(\xi^{l})$, we get

$$L(\dot{\eta}^I, \eta^I) = \int dt \ [\hbar a_I^{\eta} \dot{\eta}^I - \bar{H}(\eta^I)], \quad \hbar b_{IJ}^{\eta} \dot{\eta}^J = \frac{\partial \bar{H}}{\partial \eta^I}, \quad b_{IJ}^{\eta} = \partial_{\eta^I} a_j^{\eta} - \partial_{\eta^J} a_I^{\eta}$$

where

$$\begin{split} a_{I}^{\eta} &= \mathrm{i} \langle \phi | \partial_{\eta^{I}} | \phi \rangle = \mathrm{i} \langle \phi | \partial_{\xi^{J}} | \phi \rangle \frac{\partial \xi^{J}}{\partial \eta^{I}} = a_{J} \frac{\partial \xi^{J}}{\partial \eta^{I}}. \\ b_{IJ}^{\eta} &= \partial_{\eta^{I}} (a_{K} \frac{\partial \xi^{K}}{\partial \eta^{J}}) - \partial_{\eta^{J}} (a_{K} \frac{\partial \xi^{K}}{\partial \eta^{I}}) = (\partial_{\eta^{I}} a_{K}) \frac{\partial \xi^{K}}{\partial \eta^{J}} - (\partial_{\eta^{J}} a_{K}) \frac{\partial \xi^{K}}{\partial \eta^{I}} \\ &= (\partial_{\xi^{L}} a_{K}) \frac{\partial \xi^{L}}{\partial \eta^{I}} \frac{\partial \xi^{K}}{\partial \eta^{J}} - (\partial_{\xi^{L}} a_{K}) \frac{\partial \xi^{L}}{\partial \eta^{J}} \frac{\partial \xi^{K}}{\partial \eta^{I}} = (\partial_{\xi^{L}} a_{K} - \partial_{\xi^{K}} a_{L}) \frac{\partial \xi^{L}}{\partial \eta^{I}} \frac{\partial \xi^{K}}{\partial \eta^{J}} \\ &= b_{LK} \frac{\partial \xi^{L}}{\partial \eta^{I}} \frac{\partial \xi^{K}}{\partial \eta^{J}} \end{split}$$

Generalized Liouville's theorm

• Consider time evolution from $t \to \tilde{t}$, $\xi^I \to \tilde{\xi}^I$. We have

$$d^n \tilde{\xi}^I = \text{Det}(\hat{J}) d^n \xi^I, \quad J_{IJ} = \frac{\partial \tilde{\xi}^I}{\partial \xi^J}$$

For
$$\tilde{t} = t + \delta t$$
, $\tilde{\xi}^I = \xi^I + b^{IK} \frac{\partial \bar{H}}{\partial \xi^K} \delta t$, where $b_{IJ} b^{JK} = \delta_{IK}$.
$$J_{IJ} = \delta_{IJ} + \partial_J (b^{IK}) \frac{\partial \bar{H}}{\partial \xi^K} \delta t + b^{IK} \frac{\partial^2 \bar{H}}{\partial \xi^K \partial \xi^J} \delta t, \quad \text{Det}(\hat{J}) = 1 + \partial_I (b^{IK}) \frac{\partial \bar{H}}{\partial \xi^K} \delta t$$

• Assume for η^I variable, b_{IJ}^{η} is independent of η^I . Then (Liouville's theorm)

$$\mathrm{d}^n\eta^I = \mathrm{d}^n\tilde{\eta}^I, \quad \sqrt{\mathrm{Det}(b^\eta_{IJ})}\mathrm{d}^n\eta^I = \sqrt{\mathrm{Det}(\tilde{b}^\eta_{IJ})}\mathrm{d}^n\tilde{\eta}^I$$

Change variables (Generalized Liouville's theorem)

$$\sqrt{\operatorname{Det}(b_{IJ})}\operatorname{Det}(\frac{\partial \xi^{I}}{\partial \eta^{J}}) \operatorname{d}^{n} \xi^{I} \operatorname{Det}(\frac{\partial \eta^{I}}{\partial \xi^{J}}) = \sqrt{\operatorname{Det}(\tilde{b}_{IJ}^{\eta})} \operatorname{Det}(\frac{\partial \tilde{\xi}^{I}}{\partial \tilde{\eta}^{J}}) \operatorname{d}^{n} \tilde{\eta}^{I} \operatorname{Det}(\frac{\partial \tilde{\eta}^{I}}{\partial \tilde{\xi}^{J}})
\sqrt{\operatorname{Det}(b_{IJ})} \operatorname{d}^{n} \xi^{I} = \sqrt{\operatorname{Det}(\tilde{b}_{IJ})} \operatorname{d}^{n} \tilde{\xi}^{I} \quad \text{or} \quad \operatorname{Pf}(b_{IJ}) \operatorname{d}^{n} \xi^{I} = \operatorname{Pf}(\tilde{b}_{IJ}) \operatorname{d}^{n} \tilde{\xi}^{I}$$

Phase space volume occupied by a quantum state

- For a classical theory every space-time point represents a distinct state. There is an ∞ number of states for a finite phase space.
- For a quantum system, $|\phi_{\xi^I(t)}\rangle$ and $|\phi_{\tilde{\xi}^I(t)}\rangle$ are orthogonal (ie are different quantum states) only when ξ^I and $\tilde{\xi}^I$ are different enough \to uncertainty of ξ^I . There is a finite number of states for a finite phase space.

• A phase space region D^n contain how many quantum states? We guess

$$N = \int_{D^n} \frac{\mathrm{d}^n \xi^I}{(2\pi)^{n/2}} \mathsf{Pf}(b_{IJ})$$

We will confirm it later.

An example: an anharmonic oscillator

• What is low energy spectrum of (choose $\hbar=1$ unit)

$$H = \frac{k^2}{2} + \frac{1}{2}vx^2 + \frac{1}{4}x^4, \quad k = -i\partial_x$$

• Trial ground state:

$$|\psi_0\rangle = \left(\frac{\alpha}{\pi}\right)^{1/4} e^{-\frac{1}{2}\alpha x^2}$$

The value of α is determined by minimizing the average energy

$$\langle \psi_0^{\alpha} | \hat{H} | \psi_0^{\alpha} \rangle = \frac{3 + 4\alpha^2 + 4\alpha \nu}{16\alpha^2}.$$

We find

$$\alpha = \frac{2 \times 6^{\frac{2}{3}} v + 6^{\frac{1}{3}} \left(27 + \sqrt{729 - 48 v^3}\right)^{\frac{2}{3}}}{6 \left(27 + \sqrt{729 - 48 v^3}\right)^{\frac{1}{3}}} = \sqrt{v} + \frac{3}{4v} + O(1/v^2)$$

$$\langle \hat{H} \rangle = \frac{1}{2} \sqrt{v} + \frac{3}{16v} + O(1/v^2)$$

An anharmonic oscillator

Dynamical trial ground state

$$|\psi_{\xi^I}\rangle = \left(\frac{\alpha}{\pi}\right)^{1/4} e^{i\xi^2 x} e^{-\frac{1}{2}\alpha(x-\xi^1)^2}$$

a state with position $x = \xi^1$ and momentum $k = \xi^2$ fluctuations.

$$\begin{split} L(\dot{\xi}^I,\xi^I) &= \langle \psi_{\xi^I(t)}| \mathrm{i} \, \frac{\mathrm{d}}{\mathrm{d} t} - H |\psi_{\xi^I(t)}\rangle = a_I(\xi^I) \dot{\xi}^I - \bar{H}(\xi^I) \end{split}$$
 where $a_I = \mathrm{i} \, \langle \psi_{\xi^I} | \frac{\partial}{\partial \xi^I} |\psi_{\xi^I}\rangle, \qquad \bar{H}(\xi^I) = \langle \psi_{\xi^I} | \hat{H} |\psi_{\xi^I}\rangle$

The resulting equation of motion is given by

$$b_{IJ}\dot{\xi}^{J} = \frac{\partial \bar{H}}{\partial \xi^{I}}, \quad b_{IJ} = \partial_{I}a_{J} - \partial_{J}a_{I}$$

• Calculate $\mathbf{a}_{l} = \mathrm{i} \langle \psi_{\xi^{l}} | \frac{\partial}{\partial \xi^{l}} | \psi_{\xi^{l}} \rangle$:

$$a_{1} = i \int dx \left(\frac{\alpha}{\pi}\right)^{1/2} e^{-i\xi^{2}x} e^{-\frac{1}{2}\alpha(x-\xi^{1})^{2}} \alpha(x-\xi^{1}) e^{i\xi^{2}x} e^{-\frac{1}{2}\alpha(x-\xi^{1})^{2}} = 0$$

$$a_{2} = i \int dx \left(\frac{\alpha}{\pi}\right)^{1/2} e^{-i\xi^{2}x} e^{-\frac{1}{2}\alpha(x-\xi^{1})^{2}} ix e^{i\xi^{2}x} e^{-\frac{1}{2}\alpha(x-\xi^{1})^{2}} = -\xi^{1}$$

An anharmonic oscillator

We find $b_{IJ} = -\epsilon_{ij}$ and

$$\bar{H}(\xi^I) = \frac{1}{2}(\xi^2)^2 + \frac{1}{2}\nu\Big(1 + \frac{3}{2\alpha\nu}\Big)(\xi^1)^2 + \frac{1}{4}(\xi^1)^4 + \frac{3 + 4\alpha^3 + 4\alpha\nu}{16\alpha^2}$$

 The corresponding equation of motion has a form

$$\dot{\xi}^1 = \xi^2, \quad \dot{\xi}^2 = -v \left(1 + \frac{3}{2\alpha v}\right) \xi^1 - (\xi^1)^3$$

• The number of quantum states

$$N = \int_{D^2} \frac{d\xi^1 d\xi^2}{2\pi} \mathsf{Pf}(b_{IJ}) = \int_{D^2} \frac{d\xi^1 d\xi^2}{2\pi} = \int_{D^2} \frac{dx dk}{2\pi}$$

which is what we expected.

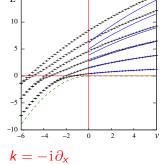
An anharmonic oscillator

- The small motions around the ground state $\xi_0^I \to A$ collection of Harmonic oscillators \rightarrow low energy spectrum.
- This is why for many interacting systems, the low energy excitations are non-interacting (like phonons in interacting crystals).
- This is why semi-classical approach works well for many systems.
- For small motion around the ground state $\xi^1 = 0, \xi^2 = 0$:

$$\dot{\xi}^1 = \xi^2, \quad \dot{\xi}^2 = -v \left(1 + \frac{3}{2\alpha v}\right) \xi^1$$

A harmonic oscillator with mass m = 1, spring constant $K = \frac{3\alpha + 2\alpha^2 v}{2\alpha^2}$, and frequency $\omega = \sqrt{v(1 + \frac{3}{2\alpha v})}$.

ullet Re-quantizing the harmonic oscillator olow energy spectrum for the Hamiltonian



$$H = \frac{k^2}{2} + \frac{1}{2}vx^2 + \frac{1}{4}x^4, \quad k = -i\partial_x$$

$$k = -i\partial_x$$

Geometric phase

- $a_l = i \langle \psi_{\xi^l} | \frac{\partial}{\partial \xi^l} | \psi_{\xi^l} \rangle$ is the so call **geometric phase** (Berry Phase).
- What is the geometric phase? Consider $|\psi_{\xi^l}\rangle$ and $|\psi_{\xi^l+\delta\xi^l}\rangle$, what is the phase different between $|\psi_{\xi^l}\rangle$ and $|\psi_{\xi^l+\delta\xi^l}\rangle$?
- But $|\psi_{\xi^I}\rangle$ and $|\psi_{\xi^I+\delta\xi^I}\rangle$ are not parallel: $|\psi_{\xi^I+\delta\xi^I}\rangle \neq e^{i\delta\phi}|\psi_{\xi^I}\rangle$. They differ by more than a phase.
- But for small $\delta \xi^I$

$$\langle \psi_{\xi'} | \psi_{\xi' + \delta \xi'} \rangle \approx 1 + i O(\delta \xi'), \qquad \langle \psi_{\xi' + \delta \xi'} | \psi_{\xi'} \rangle \approx 1 - i O(\delta \xi')$$

since, to the first order in δ

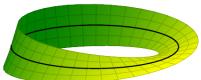
$$\begin{aligned} 0 &= \delta \langle \psi_{\xi^I} | \psi_{\xi^I} \rangle = \left(\langle \psi_{\xi^I + \delta \xi^I} | - \langle \psi_{\xi^I} | \right) | \psi_{\xi^I} \rangle + \langle \psi_{\xi^I} | \left(| \psi_{\xi^I + \delta \xi^I} \rangle - | \psi_{\xi^I} \rangle \right) \\ &= \left[\langle \psi_{\xi^I + \delta \xi^I} | \psi_{\xi^I} \rangle - 1 \right] + \left[\langle \psi_{\xi^I} | \psi_{\xi^I + \delta \xi^I} \rangle - 1 \right] \end{aligned}$$

Therefore
$$\langle \psi_{\xi^I} | \psi_{\xi^I + \delta \xi^I} \rangle \approx e^{i O(\delta \xi)}$$
, or $|\psi_{\xi^I + \delta \xi^I} \rangle = e^{i \delta \phi} |\psi_{\xi^I} \rangle + \#(\delta \xi^I)^2$, $\delta \phi = a_I(\xi^I) \delta \xi^I$

U(1) fiber bundle and global view of geometric phase

The physical states are characterized by a point ξ^i on the phase-space, after we pick the phase of $|\psi(\xi^i)\rangle$. Different choices of phases are equivalent \to the notion of U(1) fiber bundle:

- The phase space ξ^i is the base space. The equivalent normalized quantum states $\mathrm{e}^{\mathrm{i}\,\phi}|\psi(\xi^i)\rangle$ form the fiber, which is S^1 .
- So, a U(1) fiber bundle is (locally) $S^1 \times$ phase-space.
- the ξ^i -labeled quantum states $|\psi(\xi^i)\rangle$ is a cross section of the U(1) bundle. Pick a phase = pick a cross section.
- Trivial U(1) bundle = $S^1 \times \text{base-space}$ (globally).
- Non-trivial U(1) fiber bundle different topology from $S^1 \times \text{base-space}$. No smooth cross section. \rightarrow A different class of classical system.
- An example: Möbius strip is a non-trivial / bundle on base-space S¹
 (I = [0,1] is the fiber)



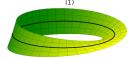
base space

Spin-1/2 example: geometric phase and fiber bundle

• All possible spin-1/2 states (or qubit states)

$$(a+ib)|\uparrow\rangle + (c+id)|\downarrow\rangle = \begin{pmatrix} a+ib\\c+id \end{pmatrix} = z, \ a^2+b^2+c^2+d^2=1$$
 form a 3-dimensional sphere S^3 (a sphere in 4-dimensional space).

- But since $|\psi\rangle \sim \mathrm{e}^{\mathrm{i}\phi}|\psi\rangle$, all possible spin-1/2 states (or qubit states) actually form a 2-dimensional sphere S^2 . $z^{\dagger}\sigma z = n$: a map $S^3 \to S^2 \to |n\rangle$: spin-1/2 in n direction.
- S^3 locally looks like $S^1 \times S^2$: S^3 is a **fiber bundle** with **fiber** S^1 and **base space** S^2 : $pt \rightarrow S^1 \xrightarrow{inj} S^3 \xrightarrow{surj} S^2 \rightarrow pt$
- If we pick a phase ϕ for each $|\mathbf{n}\rangle$, we may get $|\mathbf{n}\rangle = \begin{pmatrix} \mathrm{e}^{-\mathrm{i}\,\varphi/2}\cos(\theta/2) \\ \mathrm{e}^{\mathrm{i}\,\varphi/2}\sin(\theta/2) \end{pmatrix}$ or $|\mathbf{n}\rangle = \begin{pmatrix} \cos(\theta/2) \\ \mathrm{e}^{\mathrm{i}\,\varphi}\sin(\theta/2) \end{pmatrix}$



Möbius strip locally $I \times S^1$

The above correspond to two cross sections of the fiber bundle.

What is the geometric phase?

In the above, the phase ϕ for each $|n\rangle$ is chosen quite arbitarily. Can we make better choice?

• Let us compare the phase of $|\mathbf{n}(\theta,\varphi)\rangle$ and $|\mathbf{n}(\theta+\delta\theta,\varphi+\delta\varphi)\rangle$:

$$\begin{split} &\langle \boldsymbol{n}(\theta,\varphi)|\boldsymbol{n}(\theta+\delta\theta,\varphi+\delta\varphi)\rangle\\ &=1+\underbrace{\langle \boldsymbol{n}(\theta,\varphi)|\frac{\partial}{\partial\theta}|\boldsymbol{n}(\theta,\varphi)\rangle}_{\mathrm{i}\,a_{\theta}}\delta\theta+\underbrace{\langle \boldsymbol{n}(\theta,\varphi)|\frac{\partial}{\partial\varphi}|\boldsymbol{n}(\theta,\varphi)\rangle}_{\mathrm{i}\,a_{\varphi}}\delta\varphi\\ &=1+\mathrm{i}\,a_{\theta}\delta\theta+\mathrm{i}\,a_{\varphi}\delta\varphi\approx\mathrm{e}^{\mathrm{i}\,(a_{\theta}\delta\theta+a_{\varphi}\delta\varphi)}, \end{split}$$

where
$$i a_{\theta} = \langle \mathbf{n}(\theta, \varphi) | \frac{\partial}{\partial \theta} | \mathbf{n}(\theta, \varphi) \rangle$$
 and $i a_{\varphi} = \langle \mathbf{n}(\theta, \varphi) | \frac{\partial}{\partial \varphi} | \mathbf{n}(\theta, \varphi) \rangle$

- $e^{i(a_{\theta}\delta\theta + a_{\varphi}\delta\varphi)} = e^{i\mathbf{a}\cdot\Delta\mathbf{n}}$ is the **geometric phase** as we change $|\mathbf{n}(\theta,\varphi)\rangle$ to $|\mathbf{n}(\theta + \delta\theta,\varphi + \delta\varphi)\rangle = |\mathbf{n} + \Delta\mathbf{n}\rangle$.
- $\mathbf{a} = (a_{\theta}, a_{\varphi})$ is the **connection (vector potential)** of the geometric phase. (Like the vector potential in electromagnetism.)

Is the geometric phase meaningless?

- If $\langle \boldsymbol{n} | \boldsymbol{n} + \Delta \boldsymbol{n} \rangle = \mathrm{e}^{\mathrm{i} \boldsymbol{a} \cdot \Delta \boldsymbol{n}}$, we can always change the phase of $|\boldsymbol{n} + \Delta \boldsymbol{n} \rangle \to |\boldsymbol{n} + \Delta \boldsymbol{n} \rangle_1 = \mathrm{e}^{-\mathrm{i} \boldsymbol{a} \cdot \Delta \boldsymbol{n}} |\boldsymbol{n} + \Delta \boldsymbol{n} \rangle$, so that $\langle \boldsymbol{n} | \boldsymbol{n} + \Delta \boldsymbol{n} \rangle_1 = \mathrm{e}^{-\mathrm{i} \boldsymbol{a} \cdot \Delta \boldsymbol{n}} \mathrm{e}^{\mathrm{i} \boldsymbol{a} \cdot \Delta \boldsymbol{n}} = 1$. We can always make geometric phase = 0, and the geometric phase is meaningless. Wrong!
- As we change the phase of $|\mathbf{n}\rangle$: $|\mathbf{n}\rangle \to \mathrm{e}^{\mathrm{i}f(\theta,\varphi)}|\mathbf{n}\rangle$, the geometric phase (*ie* the connection) also changes:

$$(a_{ heta},a_{arphi})
ightarrow (a_{ heta}+\partial_{ heta}f,a_{arphi}+\partial_{arphi}f)$$

- We can always choose a f to make $(a_{\theta}, a_{\varphi}) = (0, 0)$ at any chosen (θ, φ) , ie to make $|\mathbf{n}\rangle$ and $|\mathbf{n} + \Delta \mathbf{n}\rangle$ to have the same phase.
- But since $S^3 \to S^2$ is not a trivial bundle, we cannot find a f to make $(a_\theta, a_\varphi) = (0, 0)$ for all (θ, φ) , ie to make all $|\mathbf{n}\rangle$'s to have the same phase. There is no crossection such that $|\mathbf{n}\rangle$ all have the same phases $(ie\ (a_\theta, a_\varphi) = (0, 0)$ for all (θ, φ)).

Some part of the geometric connection $\mathbf{a} = (\mathbf{a}_{\theta}, \mathbf{a}_{\varphi})$ is physical, and other part is not.

The notion of the "flux" of the geometric phase

• Consider a loop $|\mathbf{n}(t)\rangle$, $t \in [0,1]$, $\mathbf{n}(0) = \mathbf{n}(1)$. The total geometric phase of the loop

$$e^{i\sum \delta\varphi(t)} = \langle \mathbf{n}(0)|\mathbf{n}(t_1)\rangle\langle\mathbf{n}(t_1)|\mathbf{n}(t_2)\rangle\langle\mathbf{n}(t_2)|\mathbf{n}(t_3)\rangle\cdots\langle\mathbf{n}(t_{N-1})|\mathbf{n}(1)\rangle$$

$$= e^{i\sum \mathbf{a}(t)\cdot\delta\mathbf{n}(t)} = e^{i\int \mathbf{a}(t)\cdot d\mathbf{n}(t)} = e^{i\int \mathbf{a}(t)\cdot \frac{d\mathbf{n}(t)}{dt}dt}$$

- If we change the phase of $|n\rangle$: $|n\rangle \to e^{if(n)}|n\rangle$, the total geometric phase for a loop the **geometric flux** does not change.
- Computing the geometric flux:

$$\oint_C \mathbf{a} \cdot d\mathbf{n} = \oint_C a_\theta d\theta + a_\varphi d\varphi = \iint_D (\partial_\theta a_\varphi - \partial_\varphi a_\theta) d\theta d\varphi$$
 where $C = \partial D$, *ie* the loop C is the boundary of the disk D .

- $b = \partial_{\theta} a_{\varphi} \partial_{\varphi} a_{\theta}$ is called the geometric curvature (magnetic field): $b\Delta\theta\Delta\varphi$ = the total geometric phase for a small loop $(\theta,\varphi) \to (\theta+\Delta\theta,\varphi) \to (\theta+\Delta\theta,\varphi+\Delta\varphi) \to (\theta\theta,\varphi+\Delta\varphi) \to (\theta,\varphi)$.
- The total geometric phase for a loop $\oint_C a \cdot dn$ and the geometric curvature b are meaningful, since they are invariant under the gauge transformation $|n\rangle \rightarrow e^{if(n)}|n\rangle$ and $a \rightarrow a + \partial f$.

The geometric phase (the flux) for spin-1/2

From
$$i a_{\theta} = \langle \boldsymbol{n}(\theta, \varphi) | \frac{\partial}{\partial \theta} | \boldsymbol{n}(\theta, \varphi) \rangle$$
 and $i a_{\varphi} = \langle \boldsymbol{n}(\theta, \varphi) | \frac{\partial}{\partial \varphi} | \boldsymbol{n}(\theta, \varphi) \rangle$ and $|\boldsymbol{n}\rangle = \begin{pmatrix} \cos(\theta/2) \\ e^{i\varphi} \sin(\theta/2) \end{pmatrix}$, we find that $a_{\theta} = 0$, $a_{\varphi} = \sin(\theta/2) \sin(\theta/2) = \frac{1 - \cos(\theta)}{2}$

"Flux" of geometric phase: total geometric phase around a loop

For a loop
$$(\theta, \varphi) \to (\theta + \Delta \theta, \varphi) \to (\theta + \Delta \theta, \varphi + \Delta \varphi) \to (\theta \theta, \varphi + \Delta \varphi) \to (\theta, \varphi)$$
:
$$\oint_{[\Delta \theta, \Delta \varphi]} a_{\theta} d\theta + a_{\varphi} d\varphi = 0 + \frac{1 - \cos(\theta + \Delta \theta)}{2} \Delta \varphi + 0 - \frac{1 - \cos(\theta)}{2} \Delta \varphi$$

$$= \frac{1}{2} \sin(\theta) \Delta \theta \Delta \varphi = \frac{1}{2} \Omega([\Delta \theta, \Delta \varphi]) \quad \to \quad \text{half of the solid angle.}$$

 The total "flux" of the geometric phase on any campact space S² must be quantized

$$\int_{\mathcal{C}^2} \frac{1}{2!} b_{IJ} \mathrm{d} \xi^I \, \mathrm{d} \xi^J = 2\pi \times \mathsf{integer}$$

 $=2\pi \times \text{Chern number}.$

Spin-1/2 has a Chern number 1





The geometric phase of spin-1

• The geometric connection for spin-1/2 $|\mathbf{n}_{S_n=\frac{1}{2}}\rangle$ is

$$(a_{\theta}^{S=\frac{1}{2}}, a_{\varphi}^{S=\frac{1}{2}}) = (0, \frac{1-\cos(\theta)}{2}).$$

• The geometric connection for spin-1 $|n_{S_{n-1}}\rangle$ is

$$(a_{\theta}^{S=1}, a_{\varphi}^{S=1}) = 2(a_{\theta}^{S=\frac{1}{2}}, a_{\varphi}^{S=\frac{1}{2}}) = (0, 1 - \cos(\theta)).$$

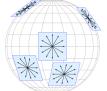
- This is because we may view $|\mathbf{n}_{S_n=1}\rangle = |\mathbf{n}_{S_n=\frac{1}{2}}\rangle \otimes |\mathbf{n}_{S_n=\frac{1}{2}}\rangle$

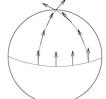
$$e^{i\Delta\phi^{S=1}} = \langle \mathbf{n}_{S_{n}=1} | \mathbf{n}_{S_{n}=1}' \rangle = \langle \mathbf{n}_{S_{n}=\frac{1}{2}} | \mathbf{n}_{S_{n}=\frac{1}{2}}' \rangle \times \langle \mathbf{n}_{S_{n}=\frac{1}{2}} | \mathbf{n}_{S_{n}=\frac{1}{2}}' \rangle = e^{i2\Delta\phi^{S=\frac{1}{2}}}$$

How to visualize the geometric phase of spin-1

Different arrows in the plan at a point n on the sphere correspond to the different phase choices $e^{i\phi}|\mathbf{n}_{S_n=1}\rangle$. We try to choose ϕ for

the spin-1 states along the loop, such that $|n_{S_n=1}\rangle$ all have the same phase. But after going around the loop, the phase miss match is the total geometric phase along the loop.

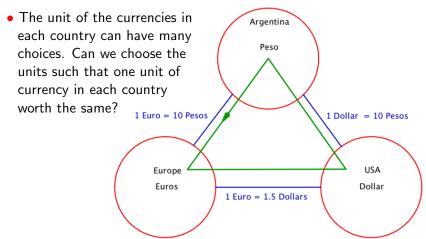




Tangent bundle on a 2-sphere

The geometric phase and currency exchange

• The phase of the state $|\mathbf{n}\rangle$ for a spin in \mathbf{n} direction can have many choices. Can we choose a phase for each $|\mathbf{n}\rangle$ such that all the states $|\mathbf{n}\rangle$ have a same phase? Only when $b_{IJ} = 0$.



Why do we care about such a subtle geometric phase?

The geometric phase is a quantum effect that can affect the equation of motion. Its effect can be real and not subtle in quantum materials.

Classical motion of spin-1/2: two views

The phase-space action

$$S = \int dt \left[\frac{1}{2} (1 - \cos \theta) \dot{\varphi} - V(\theta, \varphi) \right] = \int dt \left[-\frac{1}{2} \cos \theta \dot{\varphi} - V(\theta, \varphi) \right] + \dots$$

- Near the equator, $\cos\theta = \frac{\pi}{2} \theta = L_z$: $S = \int dt [L_z \dot{\varphi} - V(\frac{\pi}{2} - L_z, \varphi)]$
- The uniform phase-space magnetic field \rightarrow the usual canonical coordinate-momentum pair.
- $L = p\dot{x} H(p, x) \rightarrow$ uniform phase-space magnetic field $a_p = 0, a_x = p$ and $b_{px} = \partial_p a_x \partial_x a_p = 1$.
- A particle moving on S^2 with a uniform magnetic field $b_{\theta\varphi}$ of total flux 2π . It is the motion in the lowest Landau level assuming $\hbar\omega_c$ is large. Modified Newton law $F = \mathbf{v} \times \mathbf{B}$ (not $F = \mathbf{ma}$).
- A sphere with a uniform magnetic field of $2\pi N_{\text{Chern}}$ flux \rightarrow lowest Landau level has $N_{\text{Chern}} + 1$ -fold degeneracy \rightarrow spin- $N_{\text{Chern}}/2$.

The motion of a neutron in a non-uniform magnetic field

Consider a spin-1/2 neutron moving in a strong non-uniform spin magnetic field B(x). The neutron magnetic moment is $\mu_n = -1.91304272(45)\mu_N$, where $\mu_N = \frac{e\hbar}{2m_p}$ in SI unit (or $\mu_N = \frac{e\hbar}{2m_pc}$ in CGS unit). The interaction between the magnetic moment and the magnetic field, $-\mu_n B \cdot \sigma$, will force the neutron spin to be anti-parallel to the magnetic field B at low energies.

- What is the quantum Hamiltonian H
 that describes the quantum motion of the above low energy neutron?
- What is the classical equation that describes the motion of the above low energy neutron?

Our first guess:

• $\hat{H} = -\frac{\hbar^2}{2m_n} \partial^2 + V(\mathbf{x})$ where $V(\mathbf{x}) = -|\mu_n \mathbf{B}(\mathbf{x})|$ is the effective potential energy.

Is this guess correct?

Schrödinger equation and coordinate basis

- Schrödinger equation (basis independent): $i\hbar\partial_t|\psi\rangle = \hat{H}(\hat{p},\hat{x})|\psi\rangle$
- In a coordinate basis $|\psi\rangle = \int d\mathbf{x} \ \psi(\mathbf{x})|\mathbf{x}\rangle$, it becomes $i\hbar \partial_t \psi(\mathbf{x},t) = H(\frac{1}{i\hbar}\partial_{\mathbf{x}}\mathbf{x})\psi(\mathbf{x},t) = \left(-\frac{\hbar^2}{2m_p}\partial^2 + V(\mathbf{x})\right)\psi(\mathbf{x},t)$
- In the above, we have assumed that there is no geometric phase for $|\mathbf{x}\rangle$, ie the phase change from $|\mathbf{x}\rangle$ to $|\mathbf{x}+\delta\mathbf{x}\rangle$ is 0.
- But for our neutron problem, the phase from $|x\rangle$ to $|x + \delta x\rangle$ is not 0. How to to compute the phase change?
- For our neutron problem, $|x\rangle$ is actually $|x\rangle \otimes |n(x)\rangle$.
- The phase change from $|x\rangle \otimes |n(x)\rangle$ to $|x + \delta x\rangle \otimes |n(x + \delta x)\rangle$ is given by $\mathbf{a} \cdot \delta \mathbf{x}$:

$$e^{i \mathbf{a}(\mathbf{x}) \cdot \delta \mathbf{x}} = \langle \mathbf{n}(\mathbf{x}) | \mathbf{n}(\mathbf{x} + \delta \mathbf{x}) \rangle \quad \rightarrow \quad i \mathbf{a}(\mathbf{x}) = \langle \mathbf{n}(\mathbf{x}) | \partial | \mathbf{n}(\mathbf{x}) \rangle$$

- If there is a geometric phase for $|\mathbf{x}\rangle$, ie a phase change $\mathrm{e}^{\mathrm{i}\mathbf{a}(\mathbf{x})\cdot\delta\mathbf{x}}$ from $|\mathbf{x}\rangle$ to $|\mathbf{x}+\delta\mathbf{x}\rangle$, what will the Schrödinger equation look like?
- The result $\hat{H} = -\frac{\hbar^2}{2m_n} \partial^2 |\mu_n B(x)|$ is valid only if the direction of B(x) does not change.

How geometric phase affects Schrödinger equation?

- If we choose a new basis $|\mathbf{x}\rangle_{\mathsf{tw}} = \mathrm{e}^{\mathrm{i}\phi(\mathbf{x})}|\mathbf{x}\rangle$. $|\mathbf{x}\rangle_{\mathsf{tw}}$ will have an non-zero geometric phase: The phase change from $|\mathbf{x}\rangle_{\mathsf{tw}}$ to $|\mathbf{x}+\delta\mathbf{x}\rangle_{\mathsf{tw}}$ is $\mathrm{e}^{\mathrm{i}[\phi(\mathbf{x}+\delta\mathbf{x})-\phi(\mathbf{x})]} = \mathrm{e}^{\mathrm{i}\mathbf{a}(\mathbf{x})\cdot\delta\mathbf{x}}$ where $\mathbf{a}=\partial\phi(\mathbf{x})$.
- What is the Schrödinger equation in the new basis $|\psi\rangle = \int \mathrm{d}\mathbf{x} \; \psi(\mathbf{x})|\mathbf{x}\rangle = \int \mathrm{d}\mathbf{x} \; \psi_{\mathsf{tw}}(\mathbf{x})|\mathbf{x}\rangle_{\mathsf{tw}} \; \text{or} \; \mathrm{e}^{\mathrm{i}\phi(\mathbf{x})}\psi_{\mathsf{tw}} = \psi(\mathbf{x})$ $\mathrm{i}\hbar\partial_t\psi(\mathbf{x},t) = \hat{H}\psi(\mathbf{x},t) = \hat{H}\mathrm{e}^{\mathrm{i}\phi(\mathbf{x})}\psi_{\mathsf{tw}}$ $\mathrm{e}^{-\mathrm{i}\phi(\mathbf{x})}\mathrm{i}\hbar\partial_t\psi(\mathbf{x},t) = \mathrm{e}^{-\mathrm{i}\phi(\mathbf{x})}\hat{H}\mathrm{e}^{\mathrm{i}\phi(\mathbf{x})}\psi_{\mathsf{tw}}$ $\mathrm{i}\hbar\partial_t\psi_{\mathsf{tw}}(\mathbf{x},t) = \hat{H}_{\mathsf{tw}}\psi_{\mathsf{tw}}, \quad \hat{H}_{\mathsf{tw}} = \mathrm{e}^{-\mathrm{i}\phi(\mathbf{x})}\hat{H}\mathrm{e}^{\mathrm{i}\phi(\mathbf{x})}.$
- $\hat{H}_{tw}(\partial, \mathbf{x})$ is obtained from $\hat{H}(\partial, \mathbf{x})$ by replacing ∂ in \hat{H} by $e^{-i\phi(\mathbf{x})}\partial e^{i\phi(\mathbf{x})} = \partial + i\partial\phi(\mathbf{x}) = \partial + i\mathbf{a}(\mathbf{x})$.

$$\hat{H}_{\mathsf{tw}} = \hat{H}(\partial + i\mathbf{a}, \mathbf{x}) = -\frac{\hbar^2}{2m_{\mathsf{p}}}(\partial + i\mathbf{a})^2 + V.$$

The above is how geometric phase affects Schrödinger equation.

Effective Hamiltonian for neutron in spin magnetic field

$$\hat{H}_{\text{eff}} = -\frac{\hbar^2}{2m_n}(\boldsymbol{\partial} + \mathrm{i}\,\boldsymbol{a})^2 + V$$

where

$$\mathrm{i}\, a(x) = \langle n(x)|\partial |n(x)\rangle, \quad n = -\frac{B(x)}{|B(x)|}, \quad V(x) = -|\mu_n B(x)|.$$

- a(x) comes from geometric phase and V(x) is potential energy.
- V(x) generates a potential force $F = -\partial V$ on the particle.
- We will see that $\mathbf{a}(\mathbf{x})$ generates a Lorentz force $\mathbf{F} \propto \mathbf{v} \times \mathbf{b}$ on the particle, as if there is a "orbital magnetic field" $\mathbf{b} = \mathbf{\partial} \times \mathbf{a}$.

The geometric phase gives rise to an effective orbital magnetic field.

Obtain classical equation of motion

 Consider wavepacket with space-time dependent spin

avepacket with dependent spin
$$|\psi_{\mathbf{x}_0,\mathbf{k}_0}\rangle = \left(\frac{\alpha}{\pi}\right)^{1/4} \mathrm{e}^{\mathrm{i}\,\mathbf{k}_0\mathbf{x}} \mathrm{e}^{-\frac{1}{2}\alpha(\mathbf{x}-\mathbf{x}_0)^2} |\mathbf{n}(\mathbf{x}_0)\rangle$$

Phase space Lagrangian

$$\mathcal{L} = \langle \psi_{\mathbf{x}_{0}(t),\mathbf{k}_{0}(t)} | i \hbar \frac{\mathrm{d}}{\mathrm{d}t} - H | \psi_{\mathbf{x}_{0}(t),\mathbf{k}_{0}(t)} \rangle$$

$$= \hbar \underbrace{\mathbf{a}'}_{=0} \cdot \dot{\mathbf{x}}_{0} + \hbar \underbrace{\mathbf{a}''}_{-\mathbf{x}_{0}} \cdot \dot{\mathbf{k}}_{0} + \hbar \underbrace{\mathbf{a}(\mathbf{x}_{0})}_{-i \langle \mathbf{n} | \partial_{\mathbf{x}_{0}} | \mathbf{n} \rangle} \cdot \dot{\mathbf{x}}_{0} - \frac{\hbar^{2} \mathbf{k}_{0}^{2}}{2m_{n}} - |\mu_{n} \mathbf{B}(\mathbf{x}_{0})|$$

$$= -\hbar \mathbf{x}_{0} \cdot \dot{\mathbf{k}}_{0} + \hbar \mathbf{a}(\mathbf{x}_{0}) \cdot \dot{\mathbf{x}}_{0} - \frac{\hbar^{2} \mathbf{k}_{0}^{2}}{2m_{n}} + |\mu_{n} \mathbf{B}(\mathbf{x}_{0})|$$

$$\approx \mathbf{p}_{0} \cdot \dot{\mathbf{x}}_{0} + \hbar \mathbf{a}(\mathbf{x}_{0}) \cdot \dot{\mathbf{x}}_{0} - \frac{\mathbf{p}_{0}^{2}}{2m} - V(\mathbf{x}_{0})$$

Obtain classical equation of motion

For
$$S = \int dt \left[\mathbf{p} \cdot \dot{\mathbf{x}} + \hbar \mathbf{a}(\mathbf{x}) \cdot \dot{\mathbf{x}} - \frac{\mathbf{p}^2}{2m_n} - V(\mathbf{x}) \right]$$

From $\int dt \, \hbar \delta(\mathbf{a}_i(\mathbf{x})\dot{\mathbf{x}}^i) = \int dt \, \hbar \left[\delta x^j (\partial_j \mathbf{a}_i) \dot{\mathbf{x}}^i - \dot{\mathbf{a}}_i(\mathbf{x}) \delta x^i \right]$
 $\delta S = \int dt \, \delta p_i \left[\dot{\mathbf{x}}^i - \frac{p_i}{m_n} \right] + \delta x^i \left[-\dot{p}_i + \hbar (\partial_i \mathbf{a}_J) \dot{\mathbf{x}}^j - \hbar (\partial_j \mathbf{a}_i) \dot{\mathbf{x}}^j - \partial_i V \right]$

we obtain the phase space equation of motion

$$\dot{x}^{i} = \frac{p_{i}}{m_{n}}, \qquad \dot{p}_{i} = \underbrace{\hbar(\partial_{i}a_{J} - \partial_{j}a_{i})\dot{x}^{j}}_{\text{Lorentz force}} - \partial_{i}V = \hbar b_{IJ}\dot{x}^{j} - \partial_{i}V$$

Spin twist gives rise to simulated vector potential $\mathbf{a}(\mathbf{x}) = -\mathrm{i} \langle \mathbf{n}(\mathbf{x}) | \partial | \mathbf{n}(\mathbf{x}) \rangle \to \text{simulated magnetic field.}$

Geometric phase = orbital magnetic field

- Equation of motion for $x^3 = z$

$$m_n\ddot{z} = -\partial_z V + \dot{x}[\partial_z \hbar a_x - \partial_x \hbar a_z] + \dot{y}[\partial_z \hbar a_y - \partial_y \hbar a_z]$$

- Compare with the equation of motion in a magnetic field B

$$m_{n}\ddot{z} = -\partial_{z}V + \frac{e}{c}(\dot{x}B_{y} - \dot{y}B_{x})$$

$$= -\partial_{z}V + \dot{x}(\partial_{z}\frac{e}{c}A_{x} - \partial_{x}\frac{e}{c}A_{z}) - \dot{y}(\partial_{y}\frac{e}{c}A_{z} - \partial_{z}\frac{e}{c}A_{y}).$$

- We find that $\hbar a = \frac{e}{c} A$
- The geometric meaning of magnetic field

of flux quanta =
$$\int_{S} d\mathbf{S} \cdot \mathbf{B} / \frac{hc}{e} = \oint_{\partial S} d\mathbf{x} \cdot \frac{e}{hc} \mathbf{A} = \frac{1}{2\pi} \oint_{\partial S} d\mathbf{x} \cdot \mathbf{a}$$

= geometric phase around a loop/ 2π

Simulate orbital magnetic field by twisted spin

When an electron move in a background twisted spins, the electron spin may following the direction of the background twisted spins \rightarrow geometric phase = simulated magnetic field.

The geometric phase around a loop/ 2π = The number of flux quanta of the simulated magnetic field through the loop.

- Note that $hc/e = 4.135667516 \times 10^{-15} \text{T m}^2$.
- If there is one flux quantum per $(10^{-8}m)^2$, then $B=4.135667516\times 10^{-15}/(10^{-8})^2=41 \mathrm{T}$ (About the highest static magnetic field produced)



- For electron hoping in a non-coplannar magnet, the geometric phase from the spin-twist is of order 1 per unit cell: There is one flux quantum per $(10^{-9}m)^2$, or the simulated magnetic field by the spin-twist geometric phase is $B_{\rm spin} = 4.135667516 \times 10^{-15}/(10^{-9})^2 = 4100 {\rm T}$

Energy bands in a crystal

• Hopping Hamiltonian

$$H_{\boldsymbol{m}\alpha;\boldsymbol{n}\beta} = \sum_{\Delta \boldsymbol{n}} -t_{\alpha\beta}^{\Delta \boldsymbol{n}} \delta_{\boldsymbol{m},\boldsymbol{n}+\Delta \boldsymbol{n}},$$

n lable unit cell, α, β label orbitals



Si



• Plane wave state $(x_n = n_1 a_1 + n_2 a_2 + n_3 a_3)_{(a)}$

$$\psi_{\mathbf{k}}(\mathbf{n},\beta) = \psi_{\beta}(\mathbf{k}) e^{\mathrm{i} \mathbf{k} \cdot \mathbf{x}_{\mathbf{n}}}, \qquad \sum_{\mathbf{n},\beta} H_{\mathbf{m}\alpha;\mathbf{n}\beta} \ \psi_{\mathbf{k}}(\mathbf{n},\beta) = \epsilon_{\mathbf{k}} \psi_{\mathbf{k}}(\mathbf{m},\alpha)$$

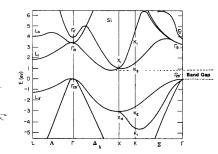
• The energy bands $\epsilon_{\pmb{k}}$ are eigenvalues of $M_{\alpha\beta}(\pmb{k})$

Si bands

$$\sum_{\beta} M_{\alpha\beta}(\mathbf{k}) \psi_{\beta}(\mathbf{k}) = \epsilon_{\mathbf{k}} \psi_{\alpha}(\mathbf{k}),$$

$$M_{\alpha\beta}(\mathbf{k}) = -\sum_{\Delta \mathbf{n}} t_{\alpha\beta}^{\Delta \mathbf{n}} e^{-i \mathbf{x}_{\Delta \mathbf{n}} \cdot \mathbf{k}}$$

 Number of bands = number of orbitals in a unit cell.



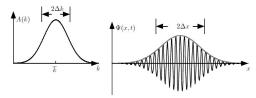
Dynamics of an electron in semiconductor

The standard theory

• Quantum dynamics: $H(\hat{\pmb{p}}) = \epsilon(\hat{\pmb{p}}), \ \hat{p} = -i\hbar\partial \rightarrow$ A plane wave $e^{i\pmb{k}\cdot\pmb{x}}\psi_{\alpha}(\pmb{k}) = e^{i\pmb{k}\cdot\pmb{x}}|\psi(\pmb{k})\rangle$ evolves as $e^{i\pmb{k}\cdot\pmb{x}}e^{-i\frac{\epsilon(\hbar\pmb{k})}{\hbar}t}|\psi(\pmb{k})\rangle$.

With potential term, the Hamiltonian is changed to $H(\hat{\boldsymbol{p}}, \hat{\boldsymbol{x}}) = \epsilon(\hat{\boldsymbol{p}}) + V(\hat{\boldsymbol{x}})$, where $[\hat{\boldsymbol{p}}^i, \hat{\boldsymbol{x}}^j] = -i\hbar\delta_{ij}$.

• Classical dynamics: $\frac{\mathrm{d}}{\mathrm{d}t}\langle\hat{O}\rangle = \frac{\mathrm{i}}{\hbar}\langle[H,\hat{O}]\rangle \rightarrow$ $\dot{\boldsymbol{p}} = -\frac{\partial H(\boldsymbol{p},\boldsymbol{x})}{\partial \boldsymbol{x}}, \quad \dot{\boldsymbol{x}} = \frac{\partial H(\boldsymbol{p},\boldsymbol{x})}{\partial \boldsymbol{p}}.$



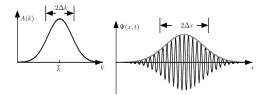
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With potential term, the Hamiltonian is changed to $H(\hat{\boldsymbol{p}}, \hat{\boldsymbol{x}}) = \epsilon(\hat{\boldsymbol{p}}) + V(\hat{\boldsymbol{x}})$, where $[\hat{\boldsymbol{p}}^i, \hat{\boldsymbol{x}}^j] = -i\hbar\delta_{ij}$.

• Classical dynamics: $\frac{\mathrm{d}}{\mathrm{d}t}\langle\hat{O}\rangle = \frac{\mathrm{i}}{\hbar}\langle[H,\hat{O}]\rangle \rightarrow$ $\dot{\boldsymbol{p}} = -\frac{\partial H(\boldsymbol{p},\boldsymbol{x})}{\partial \boldsymbol{x}}, \quad \dot{\boldsymbol{x}} = \frac{\partial H(\boldsymbol{p},\boldsymbol{x})}{\partial \boldsymbol{p}}.$



• The standard theory is wrong. $V(\hat{x})$ is wrong

Obtain classical EOM of an electron in a band

 Consider wavepacket with space-time dependent spin

avepacket with dependent spin
$$|\psi_{\mathbf{x}_0,\mathbf{k}_0}\rangle = \left(\frac{\alpha}{\pi}\right)^{1/4} \mathrm{e}^{\mathrm{i}\,\mathbf{k}_0\mathbf{x}} \mathrm{e}^{-\frac{1}{2}\alpha(\mathbf{x}-\mathbf{x}_0)^2} |\psi(\mathbf{k}_0)\rangle$$

Phase space Lagrangian

$$\mathcal{L} = \langle \psi_{\mathbf{x}_{0}(t),\mathbf{k}_{0}(t)} | i\hbar \frac{\mathrm{d}}{\mathrm{d}t} - H | \psi_{\mathbf{x}_{0}(t),\mathbf{k}_{0}(t)} \rangle$$

$$= \hbar \underbrace{\mathbf{a}'}_{=0} \cdot \dot{\mathbf{x}}_{0} + \hbar \underbrace{\mathbf{a}''}_{-\mathbf{x}_{0}} \cdot \dot{\mathbf{k}}_{0} + \hbar \underbrace{\mathbf{a}(\mathbf{k}_{0})}_{-i \langle \psi | \partial_{\mathbf{k}_{0}} | \psi \rangle} \cdot \dot{\mathbf{k}}_{0} - \frac{\hbar^{2} \mathbf{k}_{0}^{2}}{2m_{n}} - |\mu_{n} \mathbf{B}(\mathbf{x}_{0})|$$

$$= -\hbar \mathbf{x}_{0} \cdot \dot{\mathbf{k}}_{0} + \hbar \tilde{\mathbf{a}}(\mathbf{k}_{0}) \cdot \dot{\mathbf{k}}_{0} - \frac{\hbar^{2} \mathbf{k}_{0}^{2}}{2m_{n}} + |\mu_{n} \mathbf{B}(\mathbf{x}_{0})|$$

$$\approx \mathbf{p}_{0} \cdot \dot{\mathbf{x}}_{0} + \tilde{\mathbf{a}}(\mathbf{p}_{0}/\hbar) \cdot \dot{\mathbf{p}}_{0} - \frac{\mathbf{p}_{0}^{2}}{2m} - V(\mathbf{x}_{0})$$

Obtain classical EOM of an electron in a band

• The **k**-space connection (vector potential) in Brillouin zone.

$$i\tilde{\mathbf{a}}(\mathbf{k}) = \langle \psi(\mathbf{k}) | \partial_{\mathbf{k}} | \psi(\mathbf{k}) \rangle$$

• For $S = \int dt \left[\boldsymbol{p} \cdot \dot{\boldsymbol{x}} + \tilde{\boldsymbol{a}}(\boldsymbol{p}/\hbar) \cdot \dot{\boldsymbol{p}} - \frac{\boldsymbol{p}^2}{2m_n} - V(\boldsymbol{x}) \right]$ From $\int dt \, \delta(\tilde{\boldsymbol{a}}_i(\boldsymbol{p}/\hbar)\dot{\boldsymbol{p}}^i) = \int dt \left[\delta p^j (\partial_{p_j} \tilde{\boldsymbol{a}}_i) \dot{\boldsymbol{p}}^i - \dot{\tilde{\boldsymbol{a}}}_i(\boldsymbol{p}/\hbar) \delta p^i \right]$ $\delta S = \int dt \, \delta p_i \left[\dot{\boldsymbol{x}}^i - \frac{p_i}{m_n} + \hbar^{-1} (\partial_{k_i} \tilde{\boldsymbol{a}}_j) \dot{\boldsymbol{p}}^j - \hbar^{-1} (\partial_{k_j} \tilde{\boldsymbol{a}}_i) \dot{\boldsymbol{p}}^j \right] + \delta \boldsymbol{x}^i \left[-\dot{p}_i - \partial_i V \right]$

we obtain the phase space equation of motion

$$\dot{x}^{i} = \frac{p_{i}}{m_{n}} - \underbrace{\hbar^{-1}(\partial_{k_{i}}\tilde{a}_{j} - \partial_{k_{j}}\tilde{a}_{i})\dot{p}^{j}}_{\text{Velocity correction}} = \frac{p_{i}}{m_{n}} - \hbar^{-1}\tilde{b}_{IJ}\dot{p}^{j}, \qquad \dot{p}_{i} = -\partial_{i}V$$

where $\tilde{b}_{IJ} = \partial_{k_i} \tilde{a}_j - \partial_{k_j} \tilde{a}_i$ is the **k**-space "magnetic" field (geometric curvature).

The k-space connection (ie the k-space magnetic field) also modifies the equation of motion



The correct classical EOM of an electron in a band

$$L = \mathbf{p} \cdot \dot{\mathbf{x}} + \frac{e}{c} \mathbf{A}(\mathbf{x}) \cdot \dot{\mathbf{x}} + \tilde{\mathbf{a}}(\mathbf{p}/\hbar) \cdot \dot{\mathbf{p}} - \frac{\mathbf{p}^2}{2m_n} - V(\mathbf{x})$$
$$= \hbar [\mathbf{k} \cdot \dot{\mathbf{x}} + \mathbf{a}(\mathbf{x}) \cdot \dot{\mathbf{x}} + \tilde{\mathbf{a}}(\mathbf{k}) \cdot \dot{\mathbf{k}}] - \frac{\mathbf{p}^2}{2m_n} - V(\mathbf{x})$$

The real equation of motion in semiconductor

$$\dot{p}_i = -\frac{\partial V}{\partial x^i} + \frac{e}{c}B_{ij}\dot{x}^j = F_i, \quad \dot{x}_i = \frac{\partial \epsilon}{\partial p_i} - \hbar^{-1}\tilde{b}_{ij}(\mathbf{k})\dot{p}_j.$$

*F*_i include both potential force and Lorentz force.

• Phase space curvature
$$(I = x^1, x^2, x^3, k^1, k^2, k^3)$$

$$(b_{IJ}) = \begin{pmatrix} b_{ij} & \delta_{ij} \\ -\delta_{ij} & \tilde{b}_{ij} \end{pmatrix}, \quad \begin{pmatrix} 0 & -\delta_{ij} \\ \delta_{ij} & 0 \end{pmatrix} \begin{pmatrix} b_{ij} & \delta_{ij} \\ -\delta_{ij} & \tilde{b}_{ij} \end{pmatrix} = \begin{pmatrix} \delta_{ij} & \tilde{b}_{ij} \\ b_{ij} & \delta_{ij} \end{pmatrix}$$

$$\log \operatorname{Det} \begin{pmatrix} \delta_{ij} & \tilde{b}_{ij} \\ b_{ij} & \delta_{ij} \end{pmatrix} = \operatorname{Tr} \log \begin{pmatrix} \delta_{ij} & \tilde{b}_{ij} \\ b_{ij} & \delta_{ij} \end{pmatrix} = 2b_{ij}\tilde{b}_{ji} + O(b_{ik}\tilde{b}_{kj})^2$$

$$\operatorname{Pf} \begin{pmatrix} b_{ij} & \delta_{ij} \\ -\delta_{ij} & \tilde{b}_{ij} \end{pmatrix} \equiv \operatorname{Pf}(b, \tilde{b}) = 1 + b_{ij}\tilde{b}_{ji} + O(b_{ik}\tilde{b}_{kj})^2.$$

Compare with Newton's law

From the EOM

$$\begin{split} \dot{k}_i &= \hbar^{-1} F_i, \quad \dot{x}_i = \hbar^{-1} \frac{\partial \epsilon}{\partial k_i} - \tilde{b}_{ij}(\boldsymbol{k}) \dot{k}_j = \hbar^{-1} \frac{\partial \epsilon}{\partial k_i} - \hbar^{-1} \tilde{b}_{ij}(\boldsymbol{k}) F_j \\ \text{and assume } H &= \frac{\hbar^2 k^2}{2m} + V(\boldsymbol{x}), \ \tilde{b}_{ij} = \tilde{b}_{ij}(\boldsymbol{k}), \ \text{we obtain} \\ \ddot{x}^i &= \hbar^{-2} (\partial_{k_i} \partial_{k_j} H) F_j - \hbar^{-1} \tilde{b}_{ij} \dot{F}_j - \hbar^{-2} \partial_{k_l} \tilde{b}_{ij} F_j F_l \\ \text{or} \quad \ddot{x}^i &= (\partial_{p_i} \partial_{p_j} H) F_j - D_{ij} \dot{F}_j - \partial_{k_l} D_{ij} F_j F_l \\ &= m^{-1} F_i - D_{ij} \dot{F}_j - \partial_{k_l} D_{ij} F_j F_l \end{split}$$

where $p_i = \hbar k_i$, $D_{ij} = \hbar^{-1} \tilde{b}_{ij}$.

We obtain correction to the Newton law $-D_{ij}\dot{F}_j - \partial_{p_l}D_{ij}F_jF_l$.

$$\frac{p^2}{2m} \rightarrow \sqrt{m^2c^4 + c^2p^2}$$
 is the relativistic correction.

AC conductivity (from classical Drude model)

First way to include a friction force

$$F_i \rightarrow F_i - \gamma \dot{x}^i$$

We obtain

$$\ddot{x}^i = m^{-1}(F_i - \gamma \dot{x}^i) - D_{ij}(\dot{F_j} - \gamma \ddot{x}^i) - \partial_{P_l}D_{ij}(F_j - \gamma \dot{x}^j)(F_l - \gamma \dot{x}^l)$$

- Assume $\partial_{P_l} D_{ij} = 0$ and go to ω -space ${\it x} = {\it x}_\omega \, {\rm e}^{-{\rm i}\,\omega t}$:

$$[-\omega^{2}(\delta_{ij} - \gamma D_{ij}) - i\omega\gamma m^{-1}\delta_{ij}]x_{\omega}^{j} = [m^{-1}\delta_{ij} + iD_{ij}]F_{j}$$

$$\mathbf{x}_{\omega} = [-\omega^{2}(m - \gamma mD) - i\omega\gamma]^{-1}(1 + i\omega mD)\mathbf{F}_{\omega}$$

$$\mathbf{v}_{\omega} = [\gamma - i\omega m(1 - \gamma D)]^{-1}(1 + i\omega mD)\mathbf{F}_{\omega}$$

Effect of Dii disappear for DC conductance.

AC conductivity (from classical Drude model)

Second way to include a friction force

$$F_i \to F_i - \gamma \partial_{p_i} H = F_i - \gamma m^{-1} p_i$$

Still assume $\partial_{p_l} D_{ij} = 0$:

$$\dot{\mathbf{x}} = \partial_{\mathbf{p}} H - D(\mathbf{F} - \gamma m^{-1} \mathbf{p}) = (1 + \gamma D) m^{-1} \mathbf{p} - D\mathbf{F}$$
$$\dot{\mathbf{p}} = \mathbf{F} - \gamma m^{-1} \mathbf{p}.$$

- Go to
$$\omega$$
-space $\mathbf{x} = \mathbf{x}_{\omega} e^{-\mathrm{i}\omega t}$: $-\mathrm{i}\omega \mathbf{p}_{\omega} = \mathbf{F}_{\omega} - \gamma m^{-1} \mathbf{p}_{\omega}$

$$\mathbf{v}_{\omega} = -\mathrm{i}\omega \mathbf{x}_{\omega} = (1 + \gamma D) m^{-1} \mathbf{p}_{\omega} - D \mathbf{F}_{\omega}$$

$$= (1 + \gamma D) m^{-1} \frac{1}{\gamma m^{-1} - \mathrm{i}\omega} \mathbf{F}_{\omega} - D \mathbf{F}_{\omega}$$

$$= (1 + \gamma D) \frac{1}{\gamma - \mathrm{i}\omega m} \mathbf{F}_{\omega} - D \mathbf{F}_{\omega}$$

$$= (1 + \mathrm{i}\omega D m) (\gamma - \mathrm{i}\omega m)^{-1} \mathbf{F}_{\omega}$$

Effect of D_{ij} still disappear for DC conductance, the result is different from the first one.

Transport: Boltzmann equation (fluid equ. in phase space)

In fact, D_{ij} has effect on DC conductance, at least for quantum Fermi gas.

• Phase space is parametrized by $\xi^I = x^1, x^2, x^3, k^1, k^2, k^3$

$$L(\dot{\xi}^I, \xi^I) = \hbar a_I \dot{\xi}^I - H, \qquad \hbar b_{IJ} \dot{\xi}^J = \frac{\partial H}{\partial \xi^I}, \qquad b_{IJ} = \partial_I a_J - \partial_J a_I$$

• We introduce phase space density distribution

$$\mathrm{d}N = g(\xi^I) \mathsf{Pf}[b(\xi^I)] \frac{\mathrm{d}^n \xi^I}{(2\pi)^{n/2}}$$

g is the number per orbital.

• Local equilibrium distribution

$$\begin{split} g_0(\xi^I) &= \frac{1}{\mathrm{e}^{\beta(\xi^I)[H(\xi^I) - \mu]} + 1}, \qquad \text{for fermions} \\ g_0(\xi^I) &= \frac{1}{\mathrm{e}^{\beta(\xi^I)[H(\xi^I) - \mu]} - 1}, \qquad \text{for bosons} \\ g_0(\xi^I) &= \mathrm{e}^{-\beta(\xi^I)[H(\xi^I) - \mu]}, \qquad \text{for classical particles} \end{split}$$

Hydrodynamic equation of motion

ullet Consider a small cluster of gas, that evolve from time t to $ilde{t}$

$$\mathrm{d}N = \mathrm{d}\tilde{N}$$
 or $g(\xi^I)\mathrm{Pf}[b(\xi^I)]\frac{\mathrm{d}^n\xi^I}{(2\pi)^{n/2}} = g(\tilde{\xi}^I)\mathrm{Pf}[b(\tilde{\xi}^I)]\frac{\mathrm{d}^n\tilde{\xi}^I}{(2\pi)^{n/2}}$

Due to Liouville's theorm $Pf[b(\xi^I)]d^n\xi^I = Pf[b(\tilde{\xi}^I)]d^n\tilde{\xi}^I$, we have

$$g(\xi^I) = g(\tilde{\xi}^I)$$
 or $\frac{\mathrm{d}}{\mathrm{d}t}g[\xi^I(t)] = 0$

We obtain hydrodynamic equation

$$\frac{\mathrm{d}}{\mathrm{d}t}g[\xi'(t)] = 0 \quad \to \quad \frac{\partial g}{\partial t} + \dot{\xi'}\partial_I g = \frac{\partial g}{\partial t} + \hbar b^{IJ}\partial_J H \partial_I g = 0$$

• Consistent with the conservation of particle number ($b_{IJ} = const.$):

$$\begin{split} &\frac{\partial g}{\partial t} + \partial_I \mathcal{J}^I = 0, \quad \text{current: } \mathcal{J}^I = g\dot{\xi}^I = \hbar g \ b^{IJ}\partial_J H \\ &0 = \frac{\partial g}{\partial t} + \partial_I \mathcal{J}^I = \frac{\partial g}{\partial t} + \hbar b^{IJ}\partial_I g\partial_J H + \hbar b^{IJ} g\partial_I \partial_J H \\ &= \frac{\partial g}{\partial t} + \hbar b^{IJ}\partial_I g\partial_J H \end{split}$$

The conservation of particle number for $b_{IJ} \neq const$.

Assume for phase space coordinates $\tilde{\xi}^I$, $\tilde{b}_{IJ} = const.$

• Hydrodynamic EOM and conservation equation:

$$\begin{split} &\frac{\partial \tilde{g}}{\partial t} + \dot{\tilde{\xi}}^I \tilde{\partial}_I \tilde{g} = \frac{\partial \tilde{g}}{\partial t} + \hbar \tilde{b}^{IJ} \tilde{\partial}_J H \tilde{\partial}_I \tilde{g} = 0 \\ &\frac{\partial \tilde{g}}{\partial t} + \tilde{\partial}_I \tilde{\mathcal{J}}^I = 0, \quad \tilde{\mathcal{J}}^I = \tilde{g} \dot{\tilde{\xi}}^I, \qquad \dot{\tilde{\xi}}^I = \hbar \tilde{b}^{IJ} \tilde{\partial}_J H \end{split}$$

• Change of coordinates: $\xi^I = \xi^I(\tilde{\xi}^I)$

$$g(\xi^{I}) = \tilde{g}(\tilde{\xi}^{I}), \quad \partial_{I} = \frac{\partial \tilde{\xi}^{J}}{\partial \xi^{I}} \tilde{\partial}_{J}, \quad \dot{\xi}^{I} = \frac{\partial \xi^{I}}{\partial \tilde{\xi}^{J}} \dot{\tilde{\xi}}^{J}, \quad \mathcal{J}^{I} = \frac{\partial \xi^{I}}{\partial \tilde{\xi}^{J}} \tilde{\mathcal{J}}^{J},$$

$$b_{IJ} = \frac{\partial \tilde{\xi}^{K}}{\partial \xi^{I}} \frac{\partial \tilde{\xi}^{L}}{\partial \xi^{J}} \tilde{b}_{KL}, \qquad b^{IJ} = \frac{\partial \xi^{I}}{\partial \tilde{\xi}^{K}} \frac{\partial \xi^{J}}{\partial \tilde{\xi}^{L}} \tilde{b}^{KL}$$

- The subscript and superscript indecate how the quantity transforms under the coordinate transformation.
- The form of the hydrodynamic EOM remain unchanged:

$$\frac{\partial g}{\partial t} + \dot{\xi}^I \partial_I g = \frac{\partial g}{\partial t} + \hbar b^{IJ} \partial_J H \partial_I g = 0$$

The conservation of particle number for $b_{IJ} \neq const$.

• The form of the conservation equation is changed:

$$0 = \frac{\partial g}{\partial t} + \frac{\partial \xi^{K}}{\partial \tilde{\xi}^{I}} \left(\partial_{K} \frac{\partial \tilde{\xi}^{I}}{\partial \xi^{L}} \mathcal{J}^{L} \right) = \frac{\partial g}{\partial t} + \partial_{I} \mathcal{J}^{I} + \frac{\partial \xi^{K}}{\partial \tilde{\xi}^{I}} \left(\partial_{K} \frac{\partial \tilde{\xi}^{I}}{\partial \xi^{L}} \right) \mathcal{J}^{L}$$

$$= \frac{\partial g}{\partial t} + \partial_{I} \mathcal{J}^{I} + \frac{\partial \xi^{K}}{\partial \tilde{\xi}^{I}} \left(\partial_{L} \frac{\partial \tilde{\xi}^{I}}{\partial \xi^{K}} \right) \mathcal{J}^{L}$$
In fact: $\frac{\partial \xi^{K}}{\partial \tilde{\xi}^{I}} \left(\partial_{L} \frac{\partial \tilde{\xi}^{I}}{\partial \xi^{K}} \right) = \operatorname{Det}^{1/2}(b^{IJ}) \partial_{K} \operatorname{Det}^{1/2}(b_{IJ}), \text{ since the RHS}$

$$= \operatorname{Det} \left(\frac{\partial \xi^{J}}{\partial \tilde{\xi}^{I}} \right) \operatorname{Det}^{1/2}(\tilde{b}^{IJ}) \partial_{K} \left[\operatorname{Det} \left(\frac{\partial \tilde{\xi}^{I}}{\partial \xi^{J}} \right) \operatorname{Det}^{1/2}(\tilde{b}_{IJ}) \right] = \operatorname{Det} \left(\frac{\partial \xi^{J}}{\partial \tilde{\xi}^{I}} \right) \partial_{K} \operatorname{Det} \left(\frac{\partial \tilde{\xi}^{I}}{\partial \xi^{J}} \right)$$
We also have (let $M_{IJ} = \frac{\partial \tilde{\xi}^{I}}{\partial \xi^{J}}$)
$$= \operatorname{Det} \left(M_{IJ}^{IJ} \right) \partial_{L} \operatorname{Det} \left(M_{IJ} \right) = \operatorname{Det} \left(M_{IJ}^{IJ} \right) \partial_{L} \operatorname{Det} \left(M_{IJ} \right) + \delta M_{IJ} \right) - 1$$

$$= \operatorname{Det} \left(\delta_{IJ} + M^{IK} \delta M_{KJ} \right) - 1 = M^{IK} \delta M_{KI}$$

Conservation equation: (not just
$$\frac{\partial g}{\partial t} + \partial_I \mathcal{J}^I = 0$$
)

$$\frac{\partial g}{\partial t} + \partial_I \mathcal{J}^I + \frac{1}{\mathsf{Pf}(\hat{b})} [\partial_I \mathsf{Pf}(\hat{b})] \mathcal{J}^I = \frac{\partial g}{\partial t} + \frac{1}{\mathsf{Pf}(\hat{b})} \partial_I [\mathsf{Pf}(\hat{b}) \mathcal{J}^I] = 0$$

Conservation equation = Hydrodynamic equation

$$0 = \frac{\partial g}{\partial t} + \frac{1}{\mathsf{Pf}(\hat{b})} \partial_I \big[\mathsf{Pf}(\hat{b}) \mathcal{J}^I \big] = \frac{\partial g}{\partial t} + \frac{1}{\mathsf{Pf}(\hat{b})} \partial_I \big[\mathsf{Pf}(\hat{b}) \ g \ \hbar b^{IJ} \partial_J H \big]$$
$$= \frac{\partial g}{\partial t} + \hbar b^{IJ} \partial_I g \partial_J H + \hbar g \partial_J H \underbrace{\frac{1}{\mathsf{Pf}(\hat{b})} \partial_I \big[\mathsf{Pf}(\hat{b}) b^{IJ} \big]}_{\hat{p}}$$

We first note that $0 = \partial_M(b^{IK}b_{KL}) = (\partial_M b^{IK})b_{KL} + b^{IK}(\partial_M b_{KL})$ $\rightarrow 0 = \partial_M b^{IJ} + b^{IK}(\partial_M b_{KL})b^{LJ}$

This allows us to obtain

$$\begin{split} &\frac{\partial_{I}\left[\mathsf{Pf}(\hat{b})b^{IJ}\right]}{\mathsf{Pf}(\hat{b})} = \frac{b^{KL}\partial_{I}b_{LK}}{2}b^{IJ} + \partial_{I}b^{IJ} = \frac{b^{KL}b^{IJ}\partial_{I}b_{LK}}{2} - b^{IK}(\partial_{I}b_{KL})b^{LJ} \\ &= \frac{b^{KL}b^{IJ}\partial_{I}(\partial_{L}a_{K} - \partial_{K}a_{L})}{2} - b^{IK}b^{LJ}\partial_{I}(\partial_{K}a_{L} - \partial_{L}a_{K}) \\ &= b^{KL}b^{IJ}\partial_{I}\partial_{L}a_{K} + b^{IK}b^{LJ}\partial_{I}\partial_{L}a_{K} = b^{KL}b^{IJ}\partial_{I}\partial_{L}a_{K} + b^{LK}b^{IJ}\partial_{L}\partial_{I}a_{K} = 0 \end{split}$$

We recover the hydrodynamic equation $\frac{\partial g}{\partial t} + \hbar b^{IJ} \partial_I g \partial_J H = 0$.

Adding dissipation – difffusion in phase space

The environmental influence only change ξ^I slightly each time. Diffusion current

$$\mathcal{J}_{\text{diff}}^{I} = \gamma^{IJ} \frac{\partial g}{\partial \xi^{J}} = -\gamma^{IJ} \partial_{J} g. \qquad \text{(Should } \gamma^{IJ} \text{ be symmetric?)}$$

New EOM (new continuity equation)

$$\frac{\partial g}{\partial t} + \frac{1}{\mathsf{Pf}(\hat{b})} \partial_{I} \left[\mathsf{Pf}(\hat{b}) \ g \dot{\xi}^{I} \right] - \frac{1}{\mathsf{Pf}(\hat{b})} \partial_{I} \left[\mathsf{Pf}(\hat{b}) \mathcal{J}_{\mathsf{diff}}^{I} \right] = 0$$
or
$$\frac{\partial g}{\partial t} + \dot{\xi}^{I} \partial_{I} g = \frac{1}{\mathsf{Pf}(\hat{b})} \partial_{I} \left[\mathsf{Pf}(\hat{b}) \gamma^{IJ} \partial_{J} g \right]$$

- But the above difusion model does not satisfy detail balance. It assume the transition rates caused by environmntal influence between two states A, B to be the same in either direction: $t_{A \to B} = t_{B \to A}$. Such a transition rates give rise to equilibrium probability distribution that satisfies $P_A = P_B$ regardless the energy difference $E_A - E_B$ of the two states. This coresponds to $T = \infty$ case. Indeed the above diffusion model tends to make g to be uniform in phase space, which is the $T = \infty$ case.

Adding dissipation – difffusion in phase space

How to find a diffussion model that satisfy detail balance? How to find a diffussion model that make g to evolve into the equilibrium distributions for a finite temperature T:

$$g_0(\xi^I) = rac{1}{\mathrm{e}^{eta[H(\xi^I)-\mu]}+1},$$
 for fermions $g_0(\xi^I) = rac{1}{\mathrm{e}^{eta[H(\xi^I)-\mu]}-1},$ for bosons $g_0(\xi^I) = \mathrm{e}^{-eta[H(\xi^I)-\mu]},$ for classical particles

Diffusion current

$$\begin{split} \mathcal{J}_{\mathrm{diff}}^{I} &= -\gamma^{IJ} g \partial_{J} (\log g + \beta H), & \text{for classical particles} \\ \mathcal{J}_{\mathrm{diff}}^{I} &= -\gamma^{IJ} g (1-g) \partial_{J} [-\log (g^{-1}-1) + \beta H], & \text{for fermions} \\ \mathcal{J}_{\mathrm{diff}}^{I} &= -\gamma^{IJ} g (1+g) \partial_{J} [-\log (g^{-1}+1) + \beta H], & \text{for bosons} \end{split}$$

Hydrodynamics in phase space with dissipation

For classical particles (high temperature limit)

$$\frac{\partial g}{\partial t} + \dot{\xi}^I \partial_I g = \frac{1}{\mathsf{Pf}(\hat{b})} \partial_I \big[\mathsf{Pf}(\hat{b}) \gamma^{IJ} g \partial_J (\log g + \beta H) \big]$$

For fermions

$$\frac{\partial g}{\partial t} + \dot{\xi}^I \partial_I g = \frac{1}{\mathsf{Pf}(\hat{b})} \partial_I \big[\mathsf{Pf}(\hat{b}) \gamma^{IJ} g (1-g) \partial_J (\log \frac{g}{1-g} + \beta H) \big]$$

For bosons

$$\frac{\partial g}{\partial t} + \dot{\xi}^I \partial_I g = \frac{1}{\mathsf{Pf}(\hat{b})} \partial_I \big[\mathsf{Pf}(\hat{b}) \gamma^{IJ} g (1+g) \partial_J (\log \frac{g}{1+g} + \beta H) \big]$$

- The equilibrium distribution go satisfies the above EOM.
- The above diffusion term only incorporates the particle number conservation, not energy conservation, since we consider an open system and assume *T* to be fixed.

How to include energy conservation for a closed system?

Go to $\xi^I = x, k$ phase space

But the scattering by impurities and other particles usually cause large Δk . Diffusion only happen in x space. In k-space, we have have something more dramatic.

$$L = \hbar[\mathbf{k} \cdot \dot{\mathbf{x}} + \mathbf{a}(\mathbf{x}) \cdot \dot{\mathbf{x}} + \tilde{\mathbf{a}}(\mathbf{k}) \cdot \dot{\mathbf{k}}] - E(\mathbf{k}, \mathbf{x}), \quad E(\mathbf{k}, \mathbf{x}) = \epsilon(\mathbf{k}) + V(\mathbf{x})$$

$$\hbar \dot{k}_{i} = -\frac{\partial E}{\partial x^{i}} + \underbrace{\hbar b_{ij}}_{e \, P} \dot{x}^{j}, \qquad \hbar \dot{x}_{i} = \frac{\partial E}{\partial k_{i}} - \hbar \tilde{b}_{ij}(\mathbf{k}) \dot{k}_{j}.$$

 \bullet (x, k)-density distribution function

$$g(\mathbf{x}, \mathbf{k}, t) : dN = g(\mathbf{x}, \mathbf{k}, t) \operatorname{Pf}(b, \tilde{b}) \frac{d^3 \mathbf{x} d^3 \mathbf{k}}{(2\pi)^3}$$

g is the number per orbital, and $Pf(b, \tilde{b}) = 1 + b_{ii}\tilde{b}_{ii}' + \cdots$

Local equilibrium distribution

$$g_0(\mathbf{x}, \mathbf{k}) = \frac{1}{\mathrm{e}^{\beta(\mathbf{x})[E(\mathbf{k}, \mathbf{x}) - \mu(\mathbf{x})]} + 1},$$
 for fermions $g_0(\mathbf{x}, \mathbf{k}) = \frac{1}{\mathrm{e}^{\beta(\mathbf{x})[E(\mathbf{k}, \mathbf{x}) - \mu(\mathbf{x})]} - 1},$ for bosons $g_0(\mathbf{x}, \mathbf{k}) = \mathrm{e}^{-\beta(\mathbf{x})[E(\mathbf{k}, \mathbf{x}) - \mu(\mathbf{x})]},$ for classical particles

Adding dissipation - relaxationtime approximation

• We will model large Δk redistribution in k-space by

$$\frac{\partial g}{\partial t} + \dot{\xi}^I \partial_I g = \frac{\partial g}{\partial t} + \dot{x} \cdot \frac{\partial g}{\partial x} + \dot{k} \cdot \frac{\partial g}{\partial k} = -\frac{1}{\tau} (g - g_0)$$

- $\frac{\mathrm{d}g}{\mathrm{d}t} = \frac{1}{\tau}(g-g_0)$ corresponds to the change of g caused by scattering process in k space. It should conserve the x-space particle density $n(x) = \int \mathsf{Pf}(b, \tilde{b}) \frac{\mathrm{d}^3 k}{(2\pi)^3} g$. Thus the chemical potential $\mu(x)$ in g_0 is chosen to make g_0 to satisfy

$$\delta n(\mathbf{x}) = \int \mathsf{Pf}(b, \tilde{b}) \mathrm{d}^3 \mathbf{k} \ (g - g_0) = 0.$$

- No particle diffusion in x-space.
- To conserve the energy density in \mathbf{x} -space $n_E(\mathbf{x}) = \int \mathsf{Pf}(b, \tilde{b}) \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} \; E(\mathbf{x}, \mathbf{k}) g$, we choose temperature $T(\mathbf{x})$ such that $\delta n_E(\mathbf{x}) = \int \mathsf{Pf}(b, \tilde{b}) \, \mathrm{d}^3 \mathbf{k} \; E(\mathbf{x}, \mathbf{k}) (g g_0) = 0.$

Linear responce in steady state

- Steady state: $\frac{\partial g}{\partial t} = 0$ or $\dot{\mathbf{x}} \cdot \frac{\partial g}{\partial \mathbf{x}} + \dot{\mathbf{k}} \cdot \frac{\partial g}{\partial \mathbf{k}} = -\frac{1}{\tau} (g g_0)$ with EOM for particles $\hbar \dot{k}_i = -\frac{\partial V}{\partial x^i} + \hbar b_{ij} \dot{x}^j$, $\hbar \dot{x}_i = \frac{\partial \epsilon}{\partial k_i} \hbar \tilde{b}_{ij} (\mathbf{k}) \dot{k}_j$ and $g_0(\mathbf{x}, \mathbf{k}) = 1/(e^{\beta(\mathbf{x})[\epsilon(\mathbf{k}) + V(\mathbf{x}) \mu(\mathbf{x})]} + 1)$
- When $\partial_{\mathbf{x}}V = 0$, $b_{ij} = 0$, $\partial_{\mathbf{x}}\mu = 0$, $\partial_{\mathbf{x}}\beta(\mathbf{x}) = 0$, g_0 satisfies the EOM, since $\dot{\mathbf{k}} = 0$, $\frac{\partial g_0}{\partial \mathbf{x}} = \frac{\partial g_0}{\partial t} = 0$
- Linear responce: first order in

$$\dot{\mathbf{k}} \sim \partial_{\mathbf{x}} V, b_{ij}, \qquad \partial_{\mathbf{x}} g_0 \sim \partial_{\mathbf{x}} \underbrace{(V - \mu)}_{-\bar{\mu}}, \partial_{\mathbf{x}} \beta \qquad \delta g = g - g_0.$$

• Linear response for steady state

$$\begin{split} \delta g + \tau \hbar^{-1} \partial_{k_i} \epsilon \partial_{x_i} \delta g &= -\tau [\hbar^{-1} \partial_{k_i} \epsilon \partial_{x_i} g_0 + \dot{k}_i \partial_{k_i} g_0] \\ \text{or} \quad \delta g + \tau v^i \partial_{x_i} \delta g &= -\tau [v^i \partial_{x_i} g_0 + \dot{k}_i \partial_{k_i} g_0], \quad v^i = \hbar^{-1} \partial_{k_i} \epsilon. \end{split}$$

Assume $\tau v^i \partial_{x_i} \delta g \ll \delta g$ and since $\hbar \dot{k}_i = e E_i + \hbar b_{ii} v^j$:

$$\delta g = -\tau v^i \partial_{x_i} g_0 + \frac{\tau}{\hbar} (e E_i + \hbar b_{ij} v^j) \partial_{k_i} g_0, \quad g_0 = \frac{1}{\mathrm{e}^{\beta(\mathbf{x})[\epsilon(\mathbf{k}) - \bar{\mu}(\mathbf{x})]} + 1}$$

2D conductivity from $m{k}$ -space "magnetic" field $ilde{b}_{ij}$

Assume real space magnetic field $b_{ij} = 0$ and T(x), $\bar{\mu}(x)$ are independent of x:

$$\delta g = \tau e E_i \frac{\partial \epsilon}{\hbar \partial k_i} \frac{\partial g_0}{\partial \epsilon} = \tau e E_i v^i \frac{\partial g_0}{\partial \epsilon}$$

The current $(\mathsf{Pf}(b_{ij}, \tilde{b}_{ij}) = \mathsf{Pf}(0, \tilde{b}_{ij}) = 1)$

$$J^i = \int \frac{\mathrm{d}^3 \boldsymbol{k}}{(2\pi)^3} e \dot{x}^i g = \int \frac{\mathrm{d}^3 \boldsymbol{k}}{(2\pi)^3} (e v^i - e \tilde{b}_{ij} \ \hbar^{-1} e E_j) (g_0 + \tau e E_i v^i \frac{\partial g_0}{\partial \epsilon})$$

Note that

$$\int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} e v^i g_0 = \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} e^{\frac{\partial \epsilon(\mathbf{k})}{\partial k_i}} g_0(\epsilon) = \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} e^{\frac{\partial G_0[\epsilon(\mathbf{k})]}{\partial k_i}} = 0$$

where $\partial G(\epsilon)/\partial \epsilon = g_0(\epsilon)$.

$$J^i = \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} e \dot{x}^i g = \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} \Big[-\frac{e^2}{\hbar} \tilde{b}_{ij} g_0 + \tau e^2 v^j v^i \frac{\partial g_0}{\partial \epsilon} \Big] E_j$$

Conductivity:

$$\sigma_{ij} = \int \frac{\mathrm{d}^3 \mathbf{k}}{(2\pi)^3} \left[-\frac{e^2}{\hbar} \tilde{b}_{ij} g_0 + \tau e^2 v^j v^i \frac{\partial g_0}{\partial \epsilon} \right]$$

Quantized Hall conductance in 2D

For a filled band, $g_0 = 1$

$$\sigma^H_{ij} = -\int rac{\mathrm{d}^2 \mathbf{k}}{(2\pi)^2} rac{e^2}{\hbar} \tilde{b}_{ij} g_0 = -\epsilon_{ij} n_{\mathsf{Chern}} rac{e^2}{\hbar}$$



where (let $\tilde{b}_{ij} = \epsilon_{ij}\tilde{b}$)

Thouless

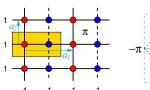
$$\begin{split} n_{\text{Chern}} &= \int_{B.Z.} \frac{\mathrm{d}^2 k}{2\pi} \tilde{b} = \int_{B.Z.} \frac{\mathrm{d}^2 k}{2\pi} \Big(\frac{\partial \tilde{a}_x}{\partial k_y} - \frac{\partial \tilde{a}_y}{\partial k_x} \Big) = \text{integer}, \\ &\mathrm{i} \, \tilde{a}_i = \langle \psi(\mathbf{k}) | \partial_{k_i} | \psi(\mathbf{k}) \rangle. \end{split}$$

We see that we have a quantized Hall conductance. n_{Chern} is Chern number.

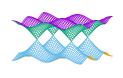
We have a Chern insulator if the total Chern number of the filled bands is non-zero.

• How to make a Chern insulator?

π -flux, Dirac fermion, and its geometric connection $\tilde{a}(k)$







Hopping matrix in k-space ($a_1 = 2x$, $a_2 = y$):

$$M(\mathbf{k}) = \begin{pmatrix} 2t\cos(\mathbf{a}_2 \cdot \mathbf{k}) & t + te^{-i\mathbf{a}_1 \cdot \mathbf{k}} \\ t + te^{i\mathbf{a}_1 \cdot \mathbf{k}} & -2t\cos(\mathbf{a}_2 \cdot \mathbf{k}) \end{pmatrix} = \begin{pmatrix} 2t\cos k_y & t + te^{-2ik_x} \\ t + te^{2ik_x} & -2t\cos k_y \end{pmatrix}$$

• $M(k) = \mathbf{v}(\mathbf{k}) \cdot \mathbf{\sigma}$: $\epsilon = \pm |\mathbf{v}(\mathbf{k})|$. The vector field $\mathbf{v}(\mathbf{k})$ on B.Z.:

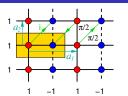
$$v_x = t + t\cos(2k_x), \quad v_y = t\sin(2k_x), \quad v_z = 2t\cos(k_y).$$

 $|\mathbf{v}| = t\sqrt{2 + 2\cos(2k_x) + 4\cos^2(k_y)} = t\sqrt{4\cos^2(k_x) + 4\cos^2(k_y)}.$

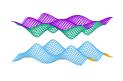
• Eigenstate in conduction band $|\boldsymbol{n}(\boldsymbol{k})\rangle$, $\boldsymbol{n}(\boldsymbol{k}) = \boldsymbol{v}(\boldsymbol{k})/|\boldsymbol{v}(\boldsymbol{k})|$, has geometric connection i $\tilde{a}_i(\boldsymbol{k}) = \langle \boldsymbol{n}(\boldsymbol{k})|\partial_{k_i}|\boldsymbol{n}(\boldsymbol{k})\rangle$: $\tilde{b}_{xy} = \partial_{k_x}\tilde{a}_y - \partial_{k_y}\tilde{a}_x \neq 0$ $\oint_K d\boldsymbol{k} \cdot \tilde{\boldsymbol{a}} = \pi$, $\oint_{K'} d\boldsymbol{k} \cdot \tilde{\boldsymbol{a}} = \pi \rightarrow \text{two } \pi\text{-flux tubes.}$



$\pi/2$ -flux state: complex hopping \to Chern insulator



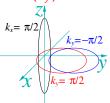




Hopping matrix in k-space $(a_1 = 2x, a_2 = y)$: M(k) =

$$\begin{pmatrix} 2t\cos(\textbf{\textit{a}}_2\cdot\textbf{\textit{k}}) & t+t\,\mathrm{e}^{-\mathrm{i}\,\textbf{\textit{a}}_1\cdot\textbf{\textit{k}}}+\mathrm{i}\,t'\,\mathrm{e}^{\mathrm{i}\,\textbf{\textit{a}}_2\cdot\textbf{\textit{k}}}+\mathrm{i}\,t'\,\mathrm{e}^{-\mathrm{i}\,(\textbf{\textit{a}}_2\cdot\textbf{\textit{k}}+\textbf{\textit{a}}_1\cdot\textbf{\textit{k}})}\\ t+t\,\mathrm{e}^{\mathrm{i}\,\textbf{\textit{a}}_1\cdot\textbf{\textit{k}}}-\mathrm{i}\,t'\,\mathrm{e}^{-\mathrm{i}\,\textbf{\textit{a}}_2\cdot\textbf{\textit{k}}}-\mathrm{i}\,t'\,\mathrm{e}^{\mathrm{i}\,(\textbf{\textit{a}}_2\cdot\textbf{\textit{k}}+\textbf{\textit{a}}_1\cdot\textbf{\textit{k}})} & -2t\cos(\textbf{\textit{a}}_2\cdot\textbf{\textit{k}}) \end{pmatrix}$$

- $M(k) = \mathbf{v}(\mathbf{k}) \cdot \mathbf{\sigma}$: $\epsilon = \pm |\mathbf{v}(\mathbf{k})|$. The vector field $\mathbf{v}(\mathbf{k})$ on B.Z.: $v_x = t + t\cos(2k_x) t'\sin(k_y) + t'\sin(k_y + 2k_x)$, $v_y = t\sin(2k_x) t'\cos(k_y) t'\cos(k_y + 2k_x)$, $v_z = 2t\cos(k_y)$.
- Eigenstate in conduction band $|\mathbf{n}(\mathbf{k})\rangle$, $\mathbf{n}(\mathbf{k}) = \mathbf{v}(\mathbf{k})/|\mathbf{v}(\mathbf{k})|$, has geometric connection i $\tilde{a}_i(\mathbf{k}) = \langle \mathbf{n}(\mathbf{k})|\partial_{k_i}|\mathbf{n}(\mathbf{k})\rangle$: $\tilde{b}_{xy} = \partial_{k_x}\tilde{a}_y \partial_{k_y}\tilde{a}_x \neq 0$ \rightarrow The wrapping number (Chern number) = 1 Chern insulator (IQH state)



Geometric phase from geometric connection $\tilde{a}(k)$

• Geometric phase $\phi = \oint_{\partial D} d\mathbf{k} \cdot \tilde{\mathbf{a}}(\mathbf{k}) = \frac{1}{2}\Omega$ $\phi = \oint_{\partial B, Z} d\mathbf{k} \cdot \tilde{\mathbf{a}}(\mathbf{k}) = 2\pi \times \text{wraping num.}$





• Geometric curvature $\tilde{B} = \partial_{k_x} \tilde{a}_y - \partial_{k_y} \tilde{a}_x$.

$$\phi = \oint_{\partial D} \mathrm{d} \textbf{\textit{k}} \cdot \tilde{\textbf{\textit{a}}}(\textbf{\textit{k}}) = \int_{D} \mathrm{d}^{2} k \tilde{B},$$

$$\int_{B.Z.} \mathrm{d}^{2} k \tilde{B} = 2\pi \times \text{Chern number}$$

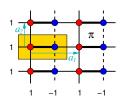
• Compute geometric curvature:

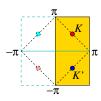
$$\tilde{B}\delta k_{x}\delta k_{y} = \frac{1}{2}\boldsymbol{n}\cdot\left(\left[\boldsymbol{n}(\boldsymbol{k}+\delta k_{x}\boldsymbol{x})-\boldsymbol{n}(\boldsymbol{k})\right]\times\left[\boldsymbol{n}(\boldsymbol{k}+\delta k_{y}\boldsymbol{y})-\boldsymbol{n}(\boldsymbol{k})\right]\right)$$
$$\tilde{B}(\boldsymbol{k}) = \frac{1}{2}\boldsymbol{n}\cdot\left[\partial_{k_{x}}\boldsymbol{n}(\boldsymbol{k})\times\partial_{k_{y}}\boldsymbol{n}(\boldsymbol{k})\right]$$

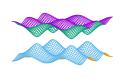
• Compute Chern number (the wrapping number):

$$(4\pi)^{-1}\int_{B_{z}} d^{2}k \ \boldsymbol{n} \cdot [\partial_{k_{x}}\boldsymbol{n}(\boldsymbol{k}) \times \partial_{k_{y}}\boldsymbol{n}(\boldsymbol{k})] = \text{Chern number}$$

Dimmer state





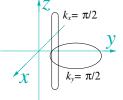


Hopping matrix in k-space ($a_1 = 2x$, $a_2 = y$):

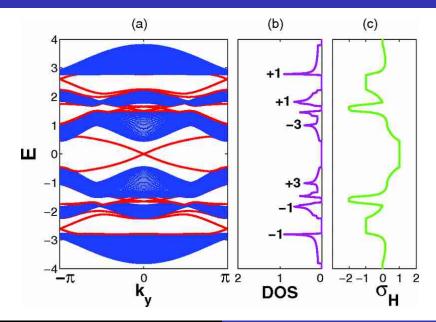
$$M(\mathbf{k}) = \begin{pmatrix} 2t\cos(\mathbf{a}_2 \cdot \mathbf{k}) & t' + te^{-i\mathbf{a}_1 \cdot \mathbf{k}} \\ t' + te^{i\mathbf{a}_1 \cdot \mathbf{k}} & -2t\cos(\mathbf{a}_2 \cdot \mathbf{k}) \end{pmatrix}$$

- $M(k) = \mathbf{v}(\mathbf{k}) \cdot \mathbf{\sigma}$: $\epsilon = \pm |\mathbf{v}(\mathbf{k})|$. The vector field $\mathbf{v}(\mathbf{k})$ on B.Z.: $v_x = t' + t\cos(2k_x)$, $v_y = t\sin(2k_x)$, $v_z = 2t\cos(k_y)$.
- Eigenstate in conduction band $|\mathbf{n}(\mathbf{k})\rangle$, $\mathbf{n}(\mathbf{k}) = \mathbf{v}(\mathbf{k})/|\mathbf{v}(\mathbf{k})|$, has geometric connection i $\tilde{a}_i(\mathbf{k}) = \langle \mathbf{n}(\mathbf{k})|\partial_{k_i}|\mathbf{n}(\mathbf{k})\rangle$: $\tilde{b}_{xy} = \partial_{k_x}\tilde{a}_y \partial_{k_y}\tilde{a}_x \neq 0$ \rightarrow The wrapping number (Chern number) = 0

 Atomic insulator



Chern number of the bands



Continuum limit and Dirac equation (Dirac fermion)

• Near K-point $(\frac{\pi}{2}, \frac{\pi}{2})$

$$M(\delta \mathbf{k} + K) = -2t\delta k_x \sigma^y - 2t\delta k_y \sigma^z + \Delta_K \sigma^x$$

The **k**-space magnetic field \tilde{b}_{xy} has a peak near K-point with total flux π . The sign of \tilde{b}_{xy} is determined by the handness of $(-2t\sigma^y, -2t\sigma^z, \Delta_K\sigma^x)$

• **k**-space Shrödinger equation: $i\hbar\dot{\psi}_{\delta \mathbf{k}} = M(\delta \mathbf{k} + K)\psi_{\delta \mathbf{k}}$ **x**-space Shrödinger equation $(\delta \mathbf{k} \to -i\partial)$:

$$i\hbar\dot{\psi}(\mathbf{x}) = 2it(\sigma^y\partial_x + \sigma^z\partial_y + \Delta_K\sigma^X)\psi(\mathbf{x})$$

Dirac equation (multiply σ^{\times} on both sides):

$$\begin{split} &(\mathrm{i}\hbar\sigma^{\mathsf{x}}\partial_{t}+2t\sigma^{\mathsf{z}}\partial_{\mathsf{x}}-2t\sigma^{\mathsf{y}}\partial_{\mathsf{y}}+\Delta_{K})\psi(\boldsymbol{x})=0\\ &(\gamma_{K}^{t}\partial_{t}+v\gamma_{K}^{\mathsf{x}}\partial_{\mathsf{x}}+v\gamma_{K}^{\mathsf{y}}\partial_{\mathsf{y}}+\hbar^{-1}\Delta_{K})\psi(\boldsymbol{x})=0\\ &\gamma_{K}^{t}=\mathrm{i}\sigma^{\mathsf{x}},\ \gamma_{K}^{\mathsf{x}}=\sigma^{\mathsf{z}},\ \gamma_{K}^{\mathsf{y}}=-\sigma^{\mathsf{y}},\ \ \{\gamma_{K}^{\mu},\gamma_{K}^{\nu}\}=2\eta^{\mu\nu},\ \mu,\nu=t,\mathsf{x},\mathsf{y}. \end{split}$$

Continuum limit and Dirac equation (Dirac fermion)

• Near K'-point $(\frac{\pi}{2}, \frac{\pi}{2})$ $M(\delta \mathbf{k} + K') = -2t\delta k_x \sigma^y + 2t\delta k_y \sigma^z + \Delta_{K'} \sigma^x$

The k-space magnetic field \tilde{b}_{xy} has a peak near K-point with total flux π . The sign of \tilde{b}_{xy} is determined by the handness of $(-2t\sigma^y, 2t\sigma^z, \Delta_{K'}\sigma^x)$

• Dirac equation:

$$(\gamma_{K'}^t \partial_t + v \gamma_{K'}^x \partial_x + v \gamma_{K'}^y \partial_y + \hbar^{-1} \Delta_{K'}) \psi(\mathbf{x}) = 0$$

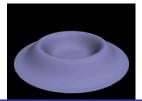
$$\gamma_{K'}^t = i \sigma^x, \ \gamma_{K'}^x = \sigma^z, \ \gamma_{K'}^y = \sigma^y, \quad \{\gamma_{K'}^\mu, \gamma_{K'}^\nu\} = 2\eta^{\mu\nu}, \ \mu, \nu = t, x, y.$$

- For the dimmer phase $\Delta_K = \Delta_{K'}$. The handness of $\left(-2t\sigma^y, -2t\sigma^z, \Delta_K\sigma^x\right)$ and $\left(-2t\sigma^y, 2t\sigma^z, \Delta_{K'}\sigma^x\right)$ are opposite. The k-space flux from K and K' cancel. The Chern number =0.
- For the $\frac{\pi}{2}$ -flux phase $\Delta_K = -\Delta_{K'}$. The handness of $\left(-2t\sigma^y, -2t\sigma^z, \Delta_K\sigma^x\right)$ and $\left(-2t\sigma^y, 2t\sigma^z, \Delta_{K'}\sigma^x\right)$ are the same. The k-space flux from K and K' add up to 2π . The Chern number =1.

Quantum Hall states and their edge excitations

The $\frac{\pi}{2}$ -flux phase (the Chern insulator), although is fully gapped in the bulk, has a gapless boundary which is **topological**. To see this more easily, we turn off the lattice potential.

- The Hamiltonian for a 2D electron in a uniform magnetic field B: $H = -\sum \frac{1}{2m} (\partial_i A_i)^2 = -\sum \frac{1}{2m} (\partial_z \frac{B}{4} z^*) (\partial_{z^*} + \frac{B}{4} z) + const.$ in complex coordinate z = x + iy (in $\hbar = c = e = 1$ unit).
- The lowest energy eigenstates: $P(z)e^{-\frac{1}{4l_B^2}|z|^2}$, where $P(z)=\sum a_Iz^I$, $B=\frac{1}{l_B^2}$, since $e^{\frac{1}{4l_B^2}zz^*}(i\partial_z-i\frac{B}{4}z^*)(i\partial_{z^*}+i\frac{B}{4}z)e^{-\frac{1}{4l_B^2}zz^*}=(i\partial_z-i\frac{B}{2}z^*)i\partial_{z^*}$
- Basis of first Landau level states: angular momentum *I*-orbital $z^I e^{-\frac{1}{4I_B^2}|z|^2}$ with a ring shape with $R_I = \sqrt{2I_B^2I}$ that enclose *I*-unit of magnetic flux, since $\pi R_I^2 B = 2\pi I$.



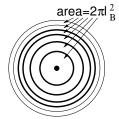
Integer quantum Hall (IQH) state

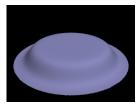
Many-body wave function of the IQH state

$$\Psi = P_1(z_1, \cdots, z_N) e^{-\frac{1}{4l_B^2} \sum_{i=1}^N z_i z_i^*}, \qquad P_1 = \prod_{i < i} (z_i - z_j).$$

- Let S_N is the power of z's in $P(z_1, \dots, z_N)$, the total angular momentum of the N-electron state: $S_N = \frac{N(N-1)}{2}$
- Let $I_N = S_N S_{N-1} \rightarrow$ the N^{th} electron is added to the angular momentum I_N -orbital: $I_N = N 1$
- The shape of quantum Hall wave function

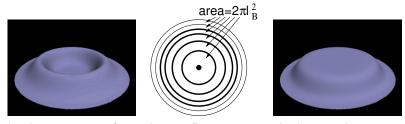




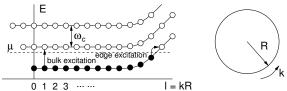


 $\rightarrow P_1$ is a filling fraction $\nu = 1$ IQH state.

Edge excitations of IQH state: chiral edge

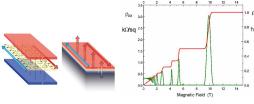


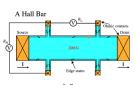
In the presence of circular confining potential, the angular momentum I-orbital has a non-zero energy E(I). Since I = kR (R = radius of the droplet), edge electron dispersion relation $\epsilon(k) = E(kR)$. The velocity of edge excitation is v = c|E|/|B|.



All excitations on the edge move in the same direction.

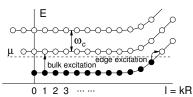
Physics of chiral edge states: Quantized Hall effect





- $I = ev_{\text{edge}} \Delta n_{\text{1D}} = ev_{\text{edge}} \Delta k / 2\pi = e \frac{eV_{\text{cross}}}{\hbar}$ = $\frac{e^2}{h} V_{\text{cross}} \rightarrow \sigma_{xy} = \frac{I}{V_{\text{cross}}} = \frac{e^2}{h}$.
- $\begin{array}{c|c}
 L & R & eV_{cross} \\
 \hline
 \Delta k & k
 \end{array}$

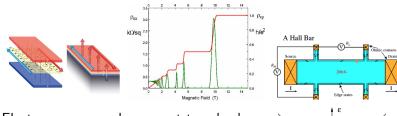
$$\sigma_{xy} = \frac{1}{V_{\text{cross}}} = \text{Wrapping-number} \times \frac{e^2}{h} = \text{Chern-number} \times \frac{e^2}{h}.$$



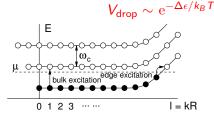




Physics of chiral edge states: Perfect conducting channel

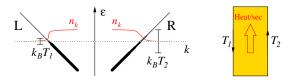


- Electrons on one edge cannot turn back, have to go forward $\rightarrow V_{drop} = 0$, $R_{xx} = 0$
- The non-zero voltage drop $V_{\text{drop}} \neq 0$ can only come from interedge tunneling (back scattering):





Universal thermal Hall conductance κ





- Heat flux = $\kappa \Delta T$. Heat flux = entropy flux $\times T$
- Entropy for one electron $S = -k_B n_k \log n_k$. $n_k = \frac{1}{e^{\epsilon_k/k_BT}+1}$.
- Entropy density

$$s = \int_{-\infty}^{+\infty} \frac{\mathrm{d}k}{2\pi} \left(-k_B n_k \log n_k \right) = \int_{-\infty}^{+\infty} \frac{\mathrm{d}k}{2\pi} k_B \frac{\log(\mathrm{e}^{\epsilon_k/k_B T} + 1)}{\mathrm{e}^{\epsilon_k/k_B T} + 1}$$

• Heat flux = $\int_{-\infty}^{+\infty} \frac{\mathrm{d}k}{2\pi} v T k_B \frac{\log(e^{\hbar v k/k_B T} + 1)}{e^{\hbar v k/k_B T} + 1} = \frac{(k_B T)^2}{\hbar} \int_{-\infty}^{+\infty} \frac{\mathrm{d}x}{2\pi} \frac{\log(e^x + 1)}{e^x + 1}$ $\int dx \frac{\log(e^x + 1)}{e^x + 1} = \frac{\pi^2}{6}$

Heat flux
$$=\frac{(k_BT)^2\pi}{12\hbar}$$
, Total heat flux $=\frac{k_B^2\pi}{12\hbar}(T_2^2-T_1^2)=\frac{\pi k_B^2T}{6\hbar}\Delta T$

- Thermal Hall conductance $\kappa = \frac{\pi}{6} \frac{k_B^2 T}{h}$
- In general $\kappa = c \frac{\pi}{6} \frac{k_B^2 T}{\hbar}$, c = # of right-modes # of left-modes

Stability of Chern insulator against U(1) symm. breaking

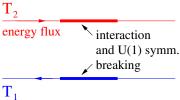
- Chern insulator: non-interacting fermions with translation and U(1) symmetry, characterized by the non-zero Chern number of the filled band in the Brillouin zone.
- If we break the U(1) symmetry (by attaching a superconductor), there will be no Hall conductance nor U(1)-Chern-Simons term. Is Chern insulator still well defined, ie still different from the trivial insulator?

Robust c = 1 chiral gapless edge state

• The thermal Hall conductance $\kappa = \frac{\pi}{6} \frac{k_B^2 T}{h}$ is always quantized, since the time-independent interaction and U(1)-symmetry breaking terms cannot change the energy flux, and the energy flux cannot flow backwards

and cross the bulk.

The 2D topological phase for fermions



• The thermal Hall conductance $\kappa = \frac{\pi}{6} \frac{k_B^2 T}{\hbar}$ is robust against U(1) symm. breaking and interaction. \rightarrow Chern insulator (IQH state) is robust against U(1) symm. breaking and interaction, although it is constructed for non-interacting fermions with U(1) symm.

Chern insulator (IQH state) has a non-trivial topological order (*ie* cannot smoothly deform into trivial insulator via any paths that can have interaction and break any symmetries)

Summary: a more general example

- n layers of $N_{\text{Chern}} = 1$ Chern insulators of charge-1 fermions + m layers of $N_{\text{Chern}} = -1$ Chern insulators of charge-0 fermions.
- Hall conductance $\sigma_{xy} = \frac{n}{2\pi}$.

```
Thermal Hall conductance \kappa = (n - m) \frac{\pi}{6} \frac{k_B^2 T}{\hbar} (ie chiral central charge c = n - m).
```

- Different phases of 2D gapped fermion systems without any symmetry and with non-degenerate ground state on torus are labeled by (at least) one integer c = n m.
 Non-zero c → non-trivial topological order (Cannot be deformed into product state)
- Different phases of 2D gapped fermion systems with U(1) symmetry and non-degenerate ground state on torus are labeled by (at least) two integers (n, m).
 - → symmetry enriched topological order (Cannot be deformed into product state if we break the symmetry. Cannot be deformed into each other if we preserve the symmetry.)