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# **Modeling Radiation Sources in MCNP**

# Modeling Radiation Source in MCNP

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Every Radiation Source has:

- **Location**
  - Point, surface, or volume
- **Direction**
  - Isotropic, beam like, or angular distribution
- **Energy**
  - Single energy, multiple discrete lines, distribution
- **Particle type**
  - Neutrons, photons, electrons, or positrons (MCNP5)
  - 34 elementary particles + heavy ions (MCNPX)
- **Time distribution**
  - Constant, radioactive decay

# MCNP Particles

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<u>IPT</u>	<u>Name</u>	<u>Symbol</u>
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1	neutron (n)	n
1	anti-neutron	-n
2	photon (g)	p
3	electron (e-)	e
3	positron (e+)	-e
4	muon- (m-)	l
4	anti-muon- (m+)	-l
6	electron neutrino	u
6	anti-electron neut.	-u
9	proton	h
9	anti-proton	-h

MCNPX only

Both MCNPX & MCNP5

<u>IPT</u>	<u>Name</u>	<u>Symbol</u>
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20	pion+ (p+)	/
20	pion- (p-)	-/
21	neutral pion (p0)	z
22	Kaon+ (K+)	k
22	Kaon- (K-)	-k
23	K0 short	%
24	K0 long	^
31	Deuteron	d
32	Triton	t
33	Helium-3	s
34	Helium-4 (a)	a
35	Heavy ion	#

## MCNP Sources

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MCNP can model the previous physical descriptions in one of four source data cards:

- General      **SDEF**      [This Lecture]
- Criticality      **KSRC**
- Surface      **SSW / SSR**
- User-Supplied (Fortran Routine)

## MCNP Sources

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**For SDEF source types, can use source distribution functions to provide details:**

- SIn      information    about the variable
- SPn      probability    of choosing a particular value
- SBn      probability biasing
- DSn      dependent source distribution

(More on these later.)

## MCNP Sources

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**In this lecture, we will use the SDEF card to work the following examples:**

- A  $^{99m}\text{Tc}$  (monoenergetic) point  $\gamma$  source in lung
- A  $^{99m}\text{Tc}$  spherical  $\gamma$  source in Pb shield
- A  $^{60}\text{Co}$  spherical  $\gamma$  source in Pb shield [optional]
- Two point gamma sources:  $^{99m}\text{Tc}$  bottom,  $^{38}\text{S}$  top
- Two spherical gamma sources:  $^{99m}\text{Tc}$  bottom,  $^{38}\text{S}$  top
- A neutron beam source [optional]

## General Source: SDEF Data Card

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**Form:**            **SDEF**    source\_variable=specification   ....

**Source Variable is an abbreviation for a physical description:**

- **erg**            energy (MeV)
- **pos**            position (location)
- **dir**            cosine of angle
- **vec**            reference vector (direction) for DIR
- **rad**            radial distance of the position
- **ext**            extents (distance or angle)
- **axs**            reference vector for EXT and RAD
- **cel**            cell
- **others**

## SDEF Data Card (cont.)

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**Specification is a value or distribution, in one of three forms:**

1. **explicit value:** **SDEF** **ERG=2.0**

[default values + source energy = 2.0 MeV]

2. **distribution number:** **SDEF** **ERG=D1**

[default values + source energy is a distribution (“D1” notation is explained later)]

3. **as a function of another variable:**

**SDEF** **POS=D1** **ERG=FPOS=D2**

[default values + src position is a distribution + src energy depends on which position]



## SDEF Source

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When a physical description is omitted from the **SDEF** card, a default is assumed

### Defaults:

Energy	[erg]	14.0 MeV
Position	[pos]	0.0 0.0 0.0
Direction	[vec]	Isotropic
Time	[tme]	0.0
Particle Type	[par]	neutrons if mode n, mode n p, or mode n p e photons if mode p or mode p e electrons if mode e or mode e f

The mode data card is a listing of all particles to be used in the simulation.

## SDEF Source Scalar Variables

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### Explicit Value Only:

WGT      EFF      NRM

### Explicit Value or Distribution:

SUR      TME      CCC      ARA      TR      X      Y      Z  
CEL      ERG      DIR      RAD      EXT      PAR

**DIR**      cosine of angle between reference vector VEC and u,v,w.  
azimuthal angle always sampled uniformly from 0 to 360.

**NRM**      sign of surface normal

**RAD**      radial distance of the position from POS or AXS vector

**EXT**      cell case: distance from POS along AXS  
surface case: cosine of angle from AXS

## **SDEF Source Vector Variables**

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**Each vector variable has 3 entries: the x,y,z component.**

<b>VEC</b>	<b>reference vector for DIR</b>
<b>POS</b>	<b>reference point for sampling position</b>
<b>AXS</b>	<b>reference vector for EXT and RAD</b>

## EXAMPLE PROBLEMS

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Problem    source1

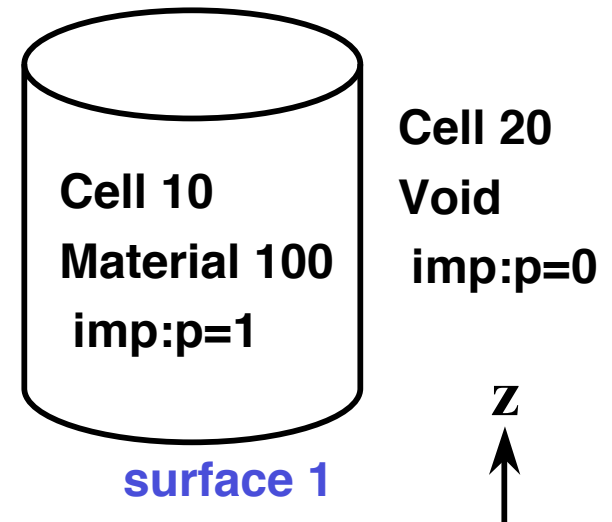
Tc-99m point source in lung

(Source variables are values)

## Tc99m in lung -- using SDEF Sources

- Bare cylinder of (almost) ICRU lung,  $\rho = 1.06 \text{ g/cm}^3$
- Tc99m emits one 0.14 MeV  $\gamma$  per decay
- Tc99m is not a material in the geometry

Nuclide	Mass-fraction	ZAID
– Hydrogen	.103	1001
– Carbon	.105	6000
– Nitrogen-14	.03	7014
– Oxygen-16	.749	8016



(1) Copy file “lung“ to file “source1”

(2) Use macrobodies, with center at (0.0, 0.0, 5.0) height = 10.0  
rad=5.0

cylinder: RCC  $V_x V_y V_z H_x H_y H_z R$  (see next slide)

(3) Add these data cards:

SDEF (Add a point source of photons at center of cylinder)

mode p \$ for photon transport, mode p e for photon and electron transport

nps 100

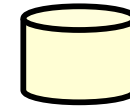
print 10 110 170

# RCC Macrobody

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## Right Circular Cylinder

RCC     $V_x$   $V_y$   $V_z$      $H_x$   $H_y$   $H_z$      $R$



$V_x$   $V_y$   $V_z$  = center of base

$H_x$   $H_y$   $H_z$  = axis of cylinder, magnitude = height

$R$  = radius

# Checking the Source

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## WHAT SHOULD YOU DO?

- 1) Examine Print Tables in the output file

**Table 10**     **source input information**

**Table 110**   **info about 1<sup>st</sup> 50 source particles**

**x y z   cell   surface u v w time wgt erg**

**Table 170**   **source distribution frequency after run**

- 2) Examine the summary tables in the output file

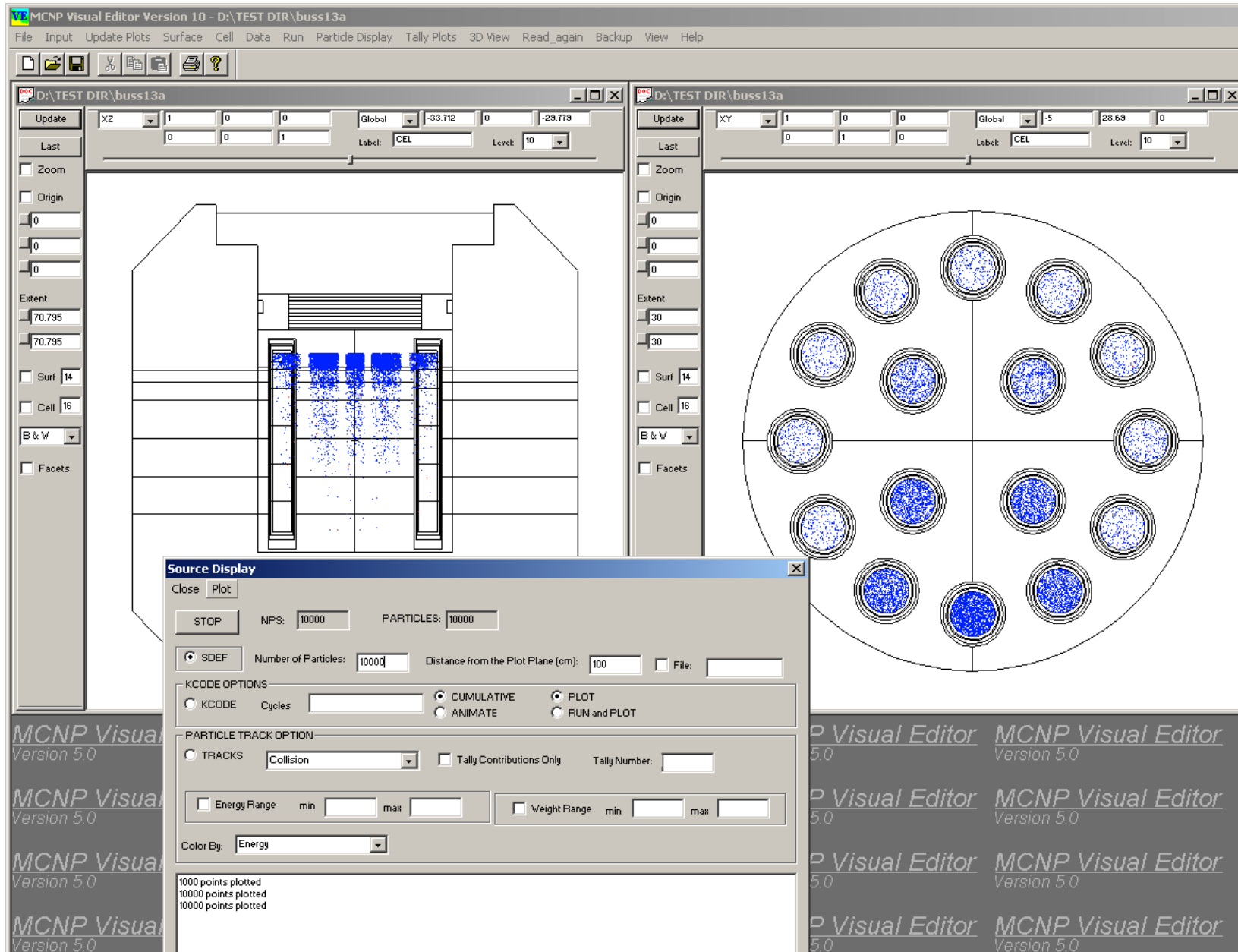
**weight started should be ~total source strength**

**energy started should be ~average source energy**

- 3) [OPTIONAL] Use Visual Editor to plot source points  
(under particle display)

**But for a point source, plot is not interesting.**

# Visual Editor





## SI, SP, SB, and DS Cards

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Often, source variables are not single values.

The following cards are used in conjunction with the SDEF card to describe distributions in location, direction, energy, etc.

<b>SI</b>	<b><u>information</u> about the variable</b> <b>bins, discrete values, distribution numbers</b>
<b>SP</b>	<b><u>probability</u> of choosing particular value</b> <b>true probabilities, built-in functions</b>
<b>SB</b>	<b><u>biased</u> probabilities</b>
<b>DS</b>	<b><u>dependent</u> distribution</b> <b>values, distribution numbers</b>

## SI (source information) Card

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FORM: SIn option entries

blank

or

- H entries are histogram bin boundaries (default)
- L entries are discrete values
- A entries are points where probability density distribution is defined
- S entries are distribution numbers

## SI Card Examples

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**SDEF ERG=D1**

**SI1 H .01 .1 1.0 3.0 14.0 \$ bins**

**SDEF POS=D1**

**SI1 L 0. 0. 0. 10. 0. 0. \$ xyz values**

**SDEF ERG=D1**

**SI1 S 3 4 5 \$ other distributions #**

**⋮**

## SP and SB Cards

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SP (source probability) Card

SB (source bias) Card

FORM:            SPn    **option**    entries

                 OR    SPn    F    a    b

                 SBn    same form as SPn

blank or

- D**        entries are bin probabilities (default)
- C**        entries are cumulative bin probabilities
- V**        entries are probability proportional to volume
  
- F**        built-in function
- a, b**     parameters for function

## SP Card Examples

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SDEF ERG=D1

SI1 .01 0.1 1.0 3.0 14.0

SP1 0 2. 1. 1. 6. (bin probabilities)

*OR*

SP1 C 0 .2 .3 .4 1.0 (cumulative)

SDEF POS=0 1 0 RAD=D2 CEL=D1

SI2 H 0 3

SP2 -21 2

SI1 L 30 40 (proportional to volume)

SP1 V (special case for CEL)

(see input file "svol1")

## Cartesian Volume

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<b>point:</b>	<b>SDEF</b>	<b>X=5</b>	<b>Y=3</b>	<b>Z=6</b>	
<b>line:</b>	<b>SDEF</b>	<b>X=D1</b>	<b>Y=3</b>	<b>Z=6</b>	
	<b>SI1 H</b>	<b>-50</b>	<b>50</b>		
	<b>SP1</b>	<b>0</b>	<b>1</b>		
<b>area:</b>	<b>SDEF</b>	<b>X=D1</b>	<b>Y=D2</b>	<b>Z=6</b>	
	<b>SI2 H</b>	<b>0</b>	<b>10</b>		
	<b>SP2</b>	<b>0</b>	<b>1</b>		
<b>volume:</b>	<b>SDEF</b>	<b>X=D1</b>	<b>Y=D2</b>	<b>Z=D3</b>	<b>CEL=5</b>

Cell Rejection:

if x,y,z are not in cell 5,  
then reject and sample x,y,z again

## SP and SB General Example

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SDEF	ERG=D1	POS=x	y	z		
SI1	H	$E_1$	$E_2$	...	$E_k$	
SP1	D	0	$P_2$	...	$P_k$	\$ bin probabilities
SB1	D	0	$B_2$	...	$B_k$	\$ bin probabilities

**Biasing some energy bins at the expense of others.  
(think variance reduction)**

## EXAMPLE PROBLEMS

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Problem source2

Photon Sources in Lead Shield  
(Volumetric Source)



## **Pb Shield -- SDEF Volumetric Source**

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1) Copy file shield to source2

Visual Editor Source plots

2) Change SDEF card to be a spatial distribution in xyz to surround bottom sphere.

Without SDEF Cell=40

3) Run. Look at starting cell locations in Table 110

4) Add “cell=40” to sdef card for cell acceptance – source particles will only be in bottom sphere

With SDEF Cell=40

5) Run. Look at starting cell locations in Table 110

**HINT: See slide #23**

## EXAMPLE PROBLEMS

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Problem source3

$^{60}\text{Co}$  Photon Sources in Lead Shield

(Multiple Discrete Energies)

(OPTIONAL)



## Pb Shield – SDEF $^{60}\text{Co}$ Source

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- 1) Copy file source2 to source3
- 2) Make the source a  $^{60}\text{Co}$  source:
  - $^{60}\text{Co}$  emits 2 photons per decay, 1.173 and 1.332 MeV
  - Make the distribution in energy equally probable
- 3) Run the problem
- 4) Look at the energies in Table 110



## More about SDEF Distributions

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- So far, each distribution has been independent and fully defined by the user.
- SDEF Distributions can also be built in functions.
- SDEF Distribution can also depend on previous distributions on the SDEF card.

## SDEF Built-in Functions

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Some distributions already built into MCNP, mostly used for energy distributions of neutron sources

Example:

**SDEF**      **ERG=D1**

**SP1**      **-3**      \$ Watt fission; default **a** and **b**

Example:

**SDEF**      **ERG=D1**

**SP1**      **-4**      \$ Gaussian fusion spectrum; default **a** and **b**

## Built-in Functions for Source Probability and Bias Specification

Source Variable	Function No. and Input Parameters	Description
ERG	-2 $a$	Maxwell fission spectrum
ERG	-3 $a b$	Watt fission spectrum
ERG	-4 $a b$	Gaussian fusion spectrum
ERG	-5 $a$	Evaporation spectrum
ERG	-6 $a b$	Muir velocity Gaussian fusion spectrum
ERG	-7 $a b$	Spare
DIR, RAD, or EXT	-21 $a$	Power law: $p(x)=c x ^a$
DIR or EXT	-31 $a$	Exponential: $p(\mu)=ce^{a\mu}$
TME, X, Y, Z	-41 $a b$	Gaussian distribution of time or position

**CAUTION:** Some defaults depend on which Source Variable.

## Dependent Source Distributions

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**Want to make the energy emitted a function of location?**

- 1) Use **FUNCTION** of preceding Source Variable on SDEF card.
  - Example: SDEF Y=D20 ERG = **F**Y = D45
- 2) Change its source information card (SI) to dependent source card (DS).
  - Example: DS45 L 0.14 1.88
- 3) Remove SP card for the dependant source, since the probability of something is now correlated to the preceding source variable.
- 4) Must match number of selections on SI and DS cards.

## DS (dependent source) Card

---

**FORM:**    **DSn**    **option**    **entries**

blank or

- H**    continuous distribution values (default)
- L**    discrete values
- S**    distribution numbers

**Other choices for **option** exist. See the MCNP manual for more detail.**



## EXAMPLE PROBLEMS

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### Example Problem source4

$^{99m}\text{Tc}$  and  $^{38}\text{S}$  point sources in shield

(Source variables are functions of other source variables)

## Dependant $\gamma$ sources in Pb Shield

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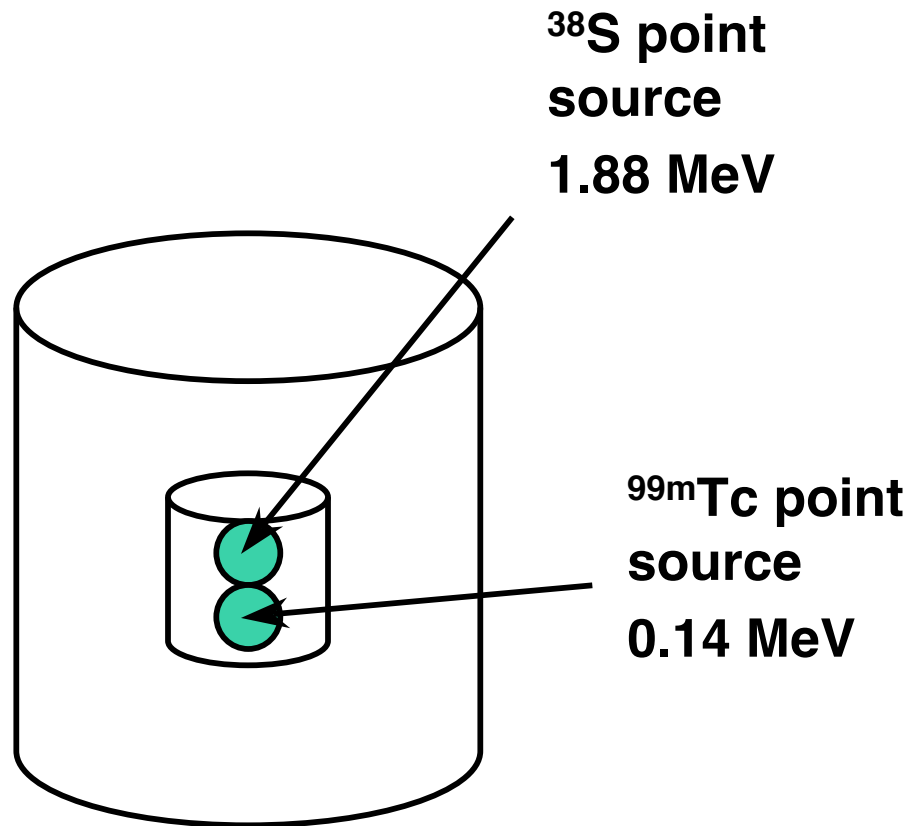
1) Copy **shield** to **source4**

2) Delete the SDEF, SI and SP cards.

3) Create a new SDEF:

2 point sources, each in the middle of the two spheres.

- Make the  $y=2$  point emit 1.88 MeV photons ( $^{38}\text{S}$ ) and the  $y=0$  point emit 0.14 MeV photons ( $^{99\text{m}}\text{Tc}$ )
- Hint: Y should be a distribution with two discrete values.
- Hint: ERG is a dependant distribution of Y, and has two discrete values



## More Dependent Source Examples

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SDEF		POS=D1		ERG=FPOS=D2			
SI1	L	0.0	0.0	0.0	10.0	0.0	0.0
SP1			1			1	
DS2	S	3	4		\$ other distributions		
SI3	H	2.0	10.0	14.0	\$ below 2, 2-10, 10-14		
SP3	D	0	0.5	0.5	\$ prob: 0, 1/2, 1/2		
SP4	-5				\$ evaporative spectrum		

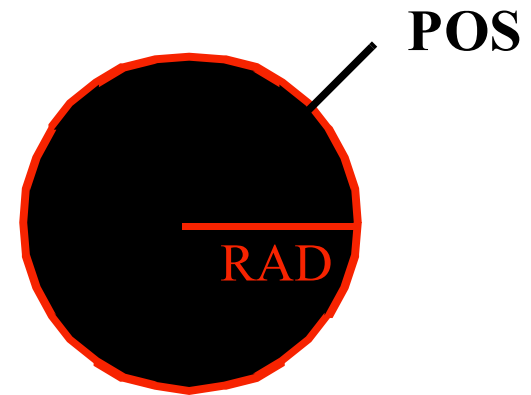
SDEF		RAD=D1		POS		FRAD		D4
SI1	S	2	3					
•								
•								
•								
DS4	L	0.0	0.0	0.0	10.0	0.0	0.0	



## Spherical Volume Sources

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uniform in volume:



```
SDEF RAD=D1 POS = 2 0 5
```

```
SI1 H 0 3
```

```
SP1 -21 2 $ density proportional to  $R^2$  (default)
```

can source bias with a function

```
SB1 -21 1.5 $ density proportional to  $R^{1.5}$ 
```

## Uniform Spherical Volume

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SP   -21   2

WHY  $R^2$  ???

Want  $p(r) \propto$  incremental volume

$$\frac{dV}{dr} = \frac{d(4/3 \pi r^3)}{dr} = 4\pi r^2$$

Prob of choosing radius  $r$

$$p(r) \propto r^2$$

## Spherical Volume Source Example

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Uniform in a spherical shell:

```
SDEF  RAD=D1      POS      0      0      0
SI1    2          3
SP1    -21        2      $ default
```

## Clipped Sphere and Shell Example

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```
10      100    -11.4    -1    imp:p=1
30      200    -1.12    +1    imp:p=1
```

```
1  PX  2.0
```

```
SDEF  RAD=D1      POS  FRAD  D5      CEL  FRAD  D6
SI1   S          2      4              $  dist #2 & #4
SP1   D          1      1
SI2   H          0              3      $  solid
SI4   H          2              3      $  shell
DS5   L    0    0    0      4    0    0
DS6   L          10              30      $  cell acceptance
```

(see file “clip1”)

## EXAMPLE PROBLEMS

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### Problem source5

$^{99m}\text{Tc}$  and  $^{38}\text{S}$  spherical sources in shield

(Source functions and spherical sources)

(OPTIONAL)

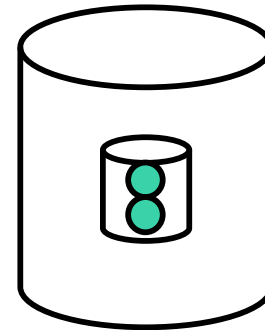




## E sources in Pb Shield – using Spherical Sources

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- 1) Copy source4 to source5
- 2) Make both sources into spherical sources homogeneously filling their geometric spheres. Use RAD



**HINT:** Must use POS, not X Y Z

- 3) Run the problem
- 4) Change the source into a thin shell covering the two spheres, like surface contamination. (**OPTIONAL**)

Visual Editor slices (XY plane @ origin),  
source particles colored by energy



## Source5 Comments

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### CAUTION:

If the source starting location is always going to be on an MCNP defined surface, must add **SUR** to **SDEF** card. Otherwise particles could get lost.

Alternatively, change particle starting locations (rad, pos, or x y z) to start them in a cell, not on a surface (see file source5b).

Example:

```
SDEF    RAD=1.0    SUR=1
```

or

```
SDEF    RAD=0.999
```



## Source5 Comments (cont.)

---

For previous example with two spheres, **SUR** must be a distribution.

- For better efficiency, use a dependent distribution
- **SUR** can NOT depend on **POS**
- **SDEF** changed so both **POS** and **SUR** depend on energy (see file source5c)



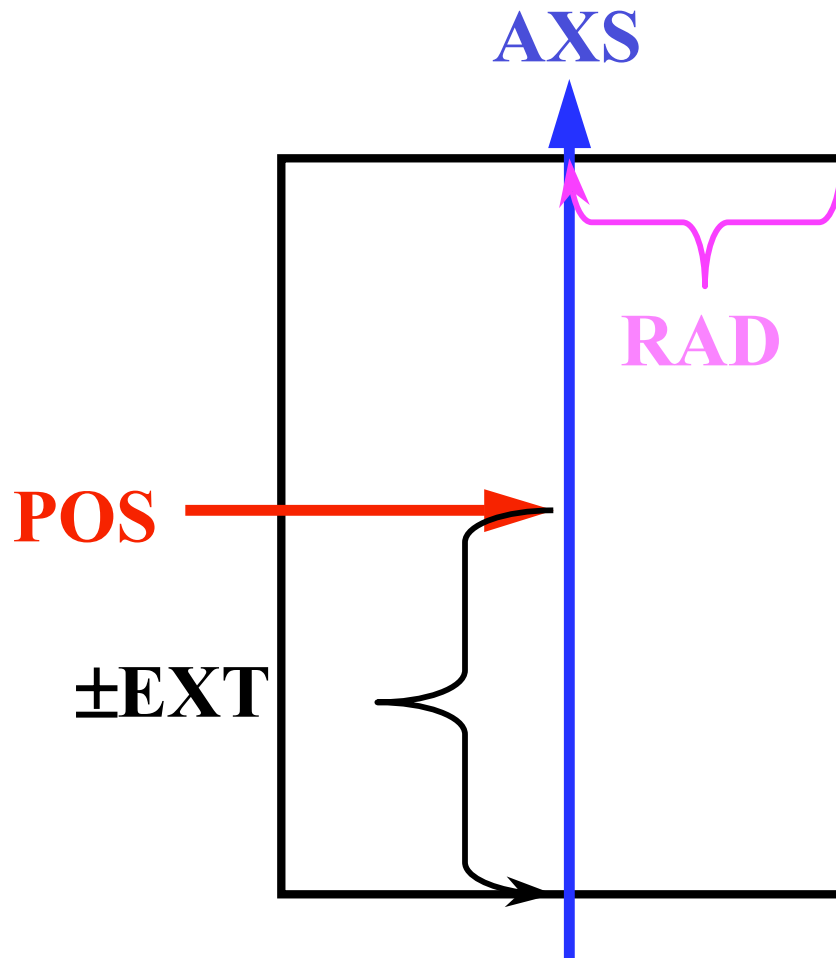
## Cylindrical Volume Sources

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- A cylindrical volume source is created by adding the AXS (Axis) vector variable to the SDEF card
- The RAD keyword now only applies orthogonally to the AXS vector
- The EXT (Extent) keyword applies along the AXS vector, and defines how far along the axis from POS will be sampled.

## Cylindrical Volume Sources

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**AXS** vector  $u \ v \ w$

**POS** vector  $x \ y \ z$

**RAD** distribution  $D_n$

**EXT** distribution  $D_n$

## Cylindrical Volume (cont)

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```
SDEF POS 0 0 0 RAD=D1 EXT=D2 AXS 0 1 0
SI1 H 0 5.6 $ inner and outer radii
SP1 -21 1 $ default density proportional to R
SI2 H -7 7 $ height
SP2 -21 0 $ default density constant with Y
```

**Add cell acceptance/rejection:**

```
SDEF POS 0 0 0 RAD=D1 EXT=D2 AXS 0 1 0 CEL 5
```

If the cell acceptance rate is too low, the problem is terminated for inefficiency. (EFF Default < 1 in 100)

Source Efficiency = 0.6149 in Cell 5

## Uniform Cylindrical Volume

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SP -21 1

WHY  $R^1$  ???

$p(r) \propto$  incremental volume

$$\frac{dV}{dr} = \frac{d(\pi r^2 h)}{dr} = 2\pi r h$$

Prob of choosing radius  $r$

$$p(r) \propto r^1$$

## Non-Isotropic Sources

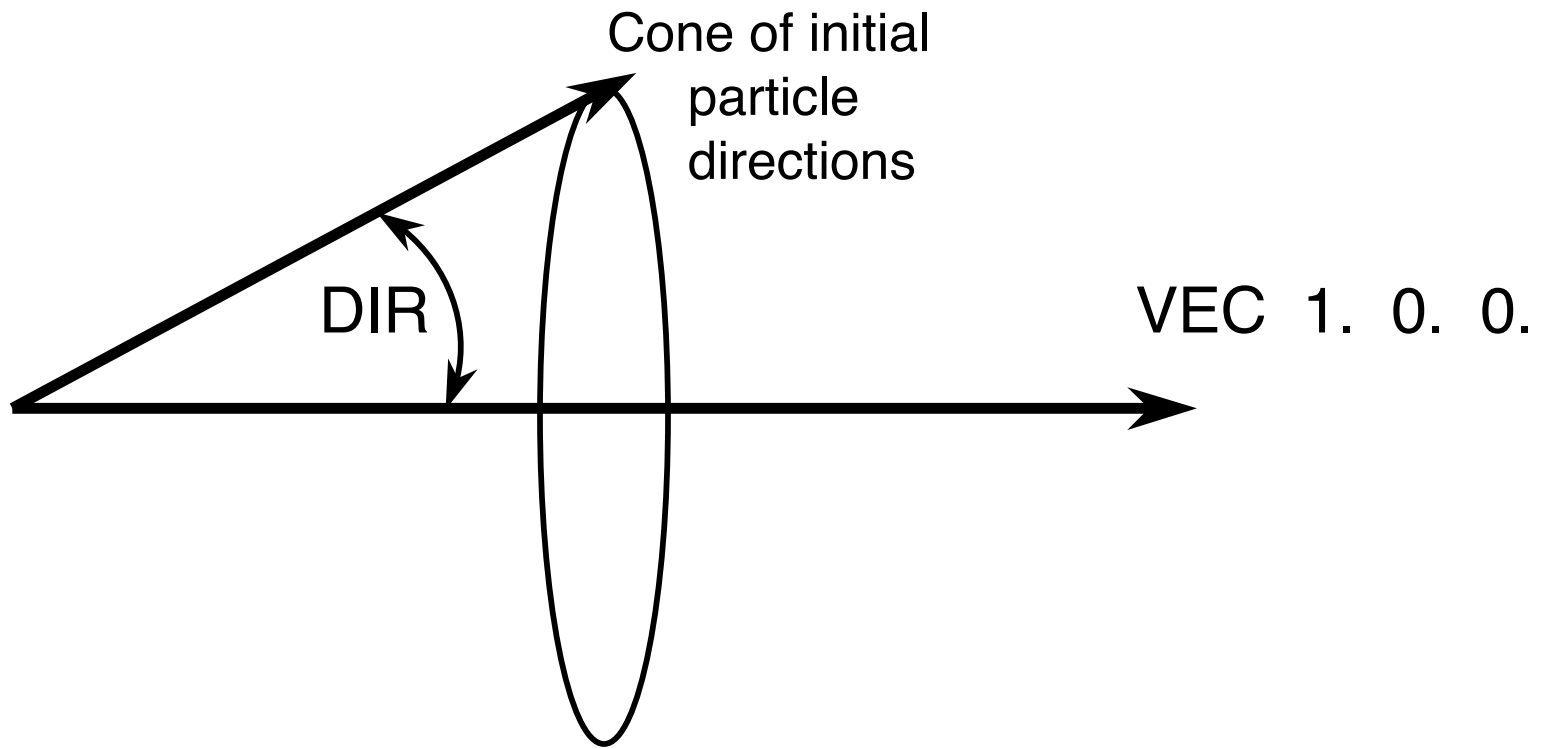
---

- A non-isotropic source is created by adding the VEC vector variable to the SDEF card, and the distribution from this vector with the variable DIR
- DIR is the cosine of the angle, 1=forward, along direction of VEC, -1 is opposite direction of VEC.
- The default for DIR is equiprobable in cosine, which results in an isotropic source.
- Setting DIR = value results in particles being emitted in a cone.
- Remember, particle starting location and direction are separate (independent) variables



## Non-Isotropic Sources

- A non-isotropic source is created by adding the VEC vector variable to the SDEF card, and the distribution from this vector with the variable DIR
- DIR is the cosine of the angle, 1.0=foreward, along direction of VEC.



## EXAMPLE PROBLEMS

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Problem    nbeam

Neutron beam source

(Cylindrical source, w/ direction)

(OPTIONAL)



## Example - Neutron Beam Source

- 1) Create a new file – nbeam
- 2) Make 1 cell – a large box
- 3) Create a disk neutron beam source:

Particle Starting Locations:

Disk at  $X=2.0$ , radius 2.5 cm

Disk facing along Y axis

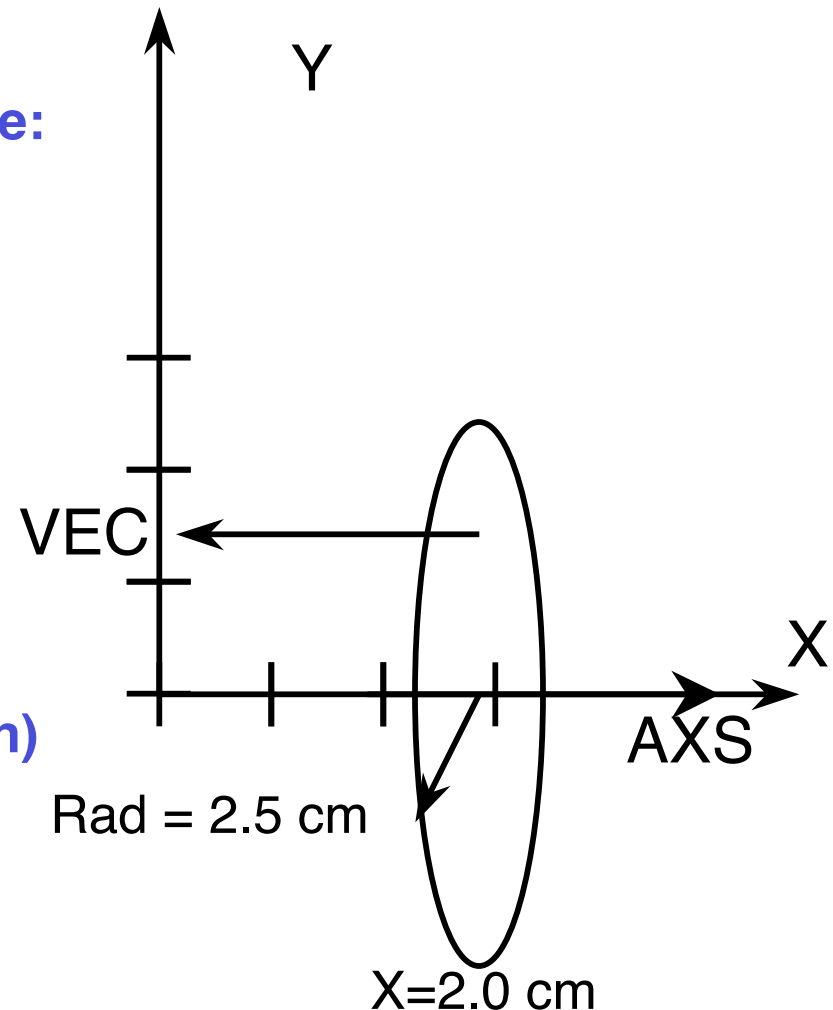
Particle starting Directions:

VEC along -X

$\mu = 1.0$

(start in VEC direction, no distribution)

- 4) Run Problem
- 5) Look at table 110



## problem nbeam

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Foreward Directed Neutron Disk Beam (Cylinder w/ Extent = 0.0)

100 0 -10 imp:n=1

102 0 10 imp:n=0

10 RPP -10.0 10.0 -10.0 10.0 -10.0 10.0

mode n

sdef pos 2. 0. 0. rad=d4 \$ Starting position, radius for disk  
axs 1. 0. 0. \$ What happens if no axs? - spherical source  
ext=0.0 \$ A disk source is a cylinder source w/ 0.0 extent  
vec -1. .0 0. \$ Starting direction for particles  
dir=1.0 \$ What happens if no dir? [isotropic source]  
erg=1.0 \$ Monoenergetic neutrons at 1 MeV  
si4 H 0.0 2.5 \$ Disk from zero to 2.5 cm  
sp4 -21 1 \$ Power Law, power=1 for disk (area) source  
nps 100  
print 10 110 170

