

# Principles

*Notes on Statistical Mechanics for Molecular Simulation*

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# Preface

Oxford Graduate Texts

Tuckerman, M. E. (2010). Statistical mechanics: Theory and molecular simulation.

Oxford: Oxford University Press.

As a freshman in college physics, there must be errors.

For normal examples, I won't number the equations.

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# Chapter 1 Classical Mechanics

## § 1.1 Math recap

Take the derivative with respect to vectors

## § 1.2 Newton's laws of motion

Review basic laws.

Newton's second law:

$$\mathbf{F} = m \frac{d^2 \mathbf{r}}{dt^2} = m \ddot{\mathbf{r}} \quad (1.1)$$

And Newton's third law:

$$\mathbf{F}_{AB} = -\mathbf{F}_{BA} \quad (1.2)$$

The definition of work:

$$W_{AB} = \int_A^B \mathbf{F} \cdot d\mathbf{l} \quad (1.3)$$

In statistical mechanics, a particle  $i$  will experience a force  $\mathbf{F}_i(\mathbf{r}_1, \mathbf{r}_2, \dots, \mathbf{r}_N, \dot{\mathbf{r}}_i)$ , which is determined by the positions of all other particles and its velocity.

## § 1.3 Phase space: visualization

We use position and momentum in the phase space because

$$\mathbf{F} = \frac{d\mathbf{p}}{dt} \quad (1.4)$$

Phase space:  $3N$  position variables and  $3N$  momentum variables constitute the microscopic state of the system at time  $t$ , if the dimensionality is 3. These variables forms a so-called *phase space vector*  $x$ . And by solving the Newton's second law, we obtain a *trajectory*:

$$x_t = (\mathbf{r}_1(t), \dots, \mathbf{r}_N(t), \mathbf{p}_1(t), \dots, \mathbf{p}_N(t)) \quad (1.5)$$

We can visualize the trajectory for a single one-dimensional particle. Some examples are (not explained here):

- a free particle with momentum  $p$ .
- a harmonic oscillator, with an elliptical phase space.
- a particle crossing a "hill" potential

Visualizing a many-particle system is hard. Maybe we can consider a particular cut or surface representing a set of variables of interest, known as a *Poincaré section*.

## § 1.4 Lagrangian formulation of classical mechanics

### 1.4.1 Conservative forces

Conservative forces, means **energy conservation**. Defined as the negative gradient of  $U$ , the *potential energy function*

$$\mathbf{F} = -\nabla U(\mathbf{r}_1, \dots, \mathbf{r}_N) \quad (1.6)$$

where  $\nabla = \frac{\partial}{\partial \mathbf{r}}$ . Or for each particle  $i$

$$\mathbf{F}_i = -\frac{\partial U}{\partial \mathbf{r}_i} \quad (1.7)$$

An important property is that, the work done by conservative forces only depends on the difference of  $U$  of the start and end state, i.e. independent of the path taken.

$$W_{AB} = U_B - U_A$$

The kinetic energy of the system is given by

$$K(\dot{\mathbf{r}}_1, \dot{\mathbf{r}}_2, \dots, \dot{\mathbf{r}}_N) = \frac{1}{2} \sum_{i=1}^N m_i \dot{\mathbf{r}}_i^2 \quad (1.8)$$

### 1.4.2 Formulation

Here we introduce the *Lagrangian* of a system

$$\mathcal{L} = K(\dot{\mathbf{r}}_1, \dots, \dot{\mathbf{r}}_N) - U(\mathbf{r}_1, \dots, \mathbf{r}_N) \quad (1.9)$$

which is the difference between the kinetic energy and potential energy. In contrast, the total energy is the sum of the two:

$$E = K(\dot{\mathbf{r}}_1, \dots, \dot{\mathbf{r}}_N) + U(\mathbf{r}_1, \dots, \mathbf{r}_N) \quad (1.10)$$

#### The Euler-Lagrangian equation

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\mathbf{r}}_i} = \frac{\partial \mathcal{L}}{\partial \mathbf{r}_i} \quad (1.11)$$

which can be verified with the substitution of eqn. 1.7 and 1.8 (poco).

This equation helps generate the equations of motion. Take the example of the one-dimensional harmonic oscillator, where

$$\begin{aligned} U(x) &= \frac{1}{2}kx^2 \\ K(\dot{x}) &= \frac{1}{2}m\dot{x}^2 \\ \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{x}} &= \frac{d(m\dot{x})}{dt} = m\ddot{x} \\ \frac{\partial \mathcal{L}}{\partial x} &= -kx \end{aligned}$$

It also helps to verify conservation of the energy.

$$\begin{aligned} \frac{dE}{dt} &= \sum_{i=1}^N m\dot{\mathbf{r}}_i \ddot{\mathbf{r}}_i + \sum_{i=1}^N \frac{\partial U}{\partial \mathbf{r}_i} \frac{d\mathbf{r}_i}{dt} \\ &= \sum_{i=1}^N m\dot{\mathbf{r}}_i \ddot{\mathbf{r}}_i - \sum_{i=1}^N \mathbf{F}_i \dot{\mathbf{r}}_i \\ &= \sum_{i=1}^N m\dot{\mathbf{r}}_i \ddot{\mathbf{r}}_i - m\ddot{\mathbf{r}}_i \dot{\mathbf{r}}_i \\ &= 0 \end{aligned}$$

### 1.4.3 Generalized coordinates

The power of Lagrangian formulation lies in the fact that the equations in an arbitrary coordinate system can be derived easily in order to address a particular problem. A set of  $3N$  *generalized coordinates* are related to the original Cartesian coordinate via

$$q_\alpha = f_\alpha(\mathbf{r}_1, \dots, \mathbf{r}_N) \quad \alpha = 1, 2, \dots, 3N \quad (1.12)$$

It is assumed that the transformation has a unique inverse

$$\mathbf{r}_i = \mathbf{g}_i(q_1, \dots, q_{3N}) \quad i = 1, 2, \dots, N \quad (1.13)$$

Thus, through the chain rule, we have

$$\dot{\mathbf{r}}_i = \sum_{\alpha=1}^{3N} \frac{\partial \mathbf{r}_i}{\partial q_\alpha} \dot{q}_\alpha \quad (1.14)$$

The kinetic energy can then be written as<sup>(1)</sup>

$$\begin{aligned}
 K(\mathbf{q}, \dot{\mathbf{q}}) &= \frac{1}{2} \sum_{i=1}^N m_i \sum_{\alpha=1}^{3N} \frac{\partial \mathbf{r}_i}{\partial q_\alpha} \dot{q}_\alpha \sum_{\beta=1}^{3N} \frac{\partial \mathbf{r}_i}{\partial q_\beta} \dot{q}_\beta \\
 &= \frac{1}{2} \sum_{\alpha=1}^{3N} \sum_{\beta=1}^{3N} \sum_{i=1}^N m_i \frac{\partial \mathbf{r}_i}{\partial q_\alpha} \frac{\partial \mathbf{r}_i}{\partial q_\beta} \dot{q}_\alpha \dot{q}_\beta \\
 &= \frac{1}{2} \sum_{\alpha=1}^{3N} \sum_{\beta=1}^{3N} G_{\alpha\beta} \dot{q}_\alpha \dot{q}_\beta
 \end{aligned} \tag{1.15}$$

where

$$G_{\alpha\beta}(\mathbf{q}) = \sum_{i=1}^N m_i \frac{\partial \mathbf{r}_i}{\partial q_\alpha} \cdot \frac{\partial \mathbf{r}_i}{\partial q_\beta} \tag{1.16}$$

is the function of  $\mathbf{q}$  (so is  $U(\mathbf{q})$ ).  $\alpha$  and  $\beta$  as indices, these elements forms a matrix  $G$ , the *mass metric matrix*. Then the Lagrangian is expressed as a function of  $\mathbf{q}$  and  $\dot{\mathbf{q}}$ . Adopting the Euler-Lagrangian equation, considering  $q_\gamma$  as  $q_\alpha$ :

$$\frac{d}{dt} \left( \sum_{\beta=1}^{3N} G_{\gamma\beta} \dot{q}_\beta \right) = \sum_{\alpha=1}^{3N} \sum_{\beta=1}^{3N} \frac{\partial G_{\alpha\beta}}{\partial q_\gamma} \dot{q}_\alpha \dot{q}_\beta - \frac{\partial U}{\partial q_\gamma} \tag{1.17}$$

### 1.4.4 Example: two-particle system

A two particle system subject to a potential  $U$  which only depends on the distance between them. We can write:

$$\mathcal{L} = \frac{1}{2} m_1 \dot{\mathbf{r}}_1^2 + \frac{1}{2} m_2 \dot{\mathbf{r}}_2^2 - U(|\mathbf{r}_1 - \mathbf{r}_2|) \tag{1.18}$$

And knowing that (not important)

$$\begin{aligned}
 |\mathbf{r}_1 - \mathbf{r}_2| &= \sqrt{(x_1 - x_2)^2 + (y_1 - y_2)^2 + (z_1 - z_2)^2} \\
 \frac{\partial |\mathbf{r}_1 - \mathbf{r}_2|}{\partial x_1} &= \frac{2x_1}{2\sqrt{(x_1 - x_2)^2 + (y_1 - y_2)^2 + (z_1 - z_2)^2}} = \frac{x_1}{|\mathbf{r}_1 - \mathbf{r}_2|} \\
 \frac{\partial U(|\mathbf{r}_1 - \mathbf{r}_2|)}{\partial \mathbf{r}_1} &= \frac{dU(|\mathbf{r}_1 - \mathbf{r}_2|)}{d(|\mathbf{r}_1 - \mathbf{r}_2|)} \frac{\partial |\mathbf{r}_1 - \mathbf{r}_2|}{\partial \mathbf{r}_1} = \frac{dU(|\mathbf{r}_1 - \mathbf{r}_2|)}{d(|\mathbf{r}_1 - \mathbf{r}_2|)} \frac{\mathbf{r}_1}{|\mathbf{r}_1 - \mathbf{r}_2|} \\
 \frac{\partial U(|\mathbf{r}_1 - \mathbf{r}_2|)}{\partial \mathbf{r}_2} &= - \frac{dU(|\mathbf{r}_1 - \mathbf{r}_2|)}{d(|\mathbf{r}_1 - \mathbf{r}_2|)} \frac{\mathbf{r}_2}{|\mathbf{r}_1 - \mathbf{r}_2|}
 \end{aligned}$$

<sup>(1)</sup>separate everything containing  $q$  ( $i$ ) and the  $\dot{q}$  velocities, by summing over  $i$  first. Imagina a 3D cube, the z-axis is  $m_i$ , while on the x-y plane sits pairs of  $q$  terms...



,using eqn 1.11, we can get

$$\begin{aligned} m_1 \ddot{\mathbf{r}}_1 &= - \frac{dU(|\mathbf{r}_1 - \mathbf{r}_2|)}{d(|\mathbf{r}_1 - \mathbf{r}_2|)} \frac{\mathbf{r}_1}{|\mathbf{r}_1 - \mathbf{r}_2|} \\ m_2 \ddot{\mathbf{r}}_2 &= \frac{dU(|\mathbf{r}_1 - \mathbf{r}_2|)}{d(|\mathbf{r}_1 - \mathbf{r}_2|)} \frac{\mathbf{r}_1}{|\mathbf{r}_1 - \mathbf{r}_2|} \end{aligned}$$

We now want to introduce a more natural set of general coordinates: *center of mass* and *relative position*:

$$\begin{aligned} \mathbf{R} &= \frac{m_1 \mathbf{r}_1 + m_2 \mathbf{r}_2}{m_1 + m_2} \\ \mathbf{r} &= \mathbf{r}_1 - \mathbf{r}_2 \end{aligned}$$

Let  $M = m_1 + m_2$ . The inverse of this transformation is

$$\begin{aligned} \mathbf{r}_1 &= \mathbf{R} + \frac{m_2}{M} \mathbf{r} \\ \mathbf{r}_2 &= \mathbf{R} - \frac{m_1}{M} \mathbf{r} \end{aligned}$$

Substituted into eqn 1.18, the Lagrangian becomes

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} m_1 \left( \dot{\mathbf{R}} + \frac{m_2}{M} \dot{\mathbf{r}} \right)^2 + \frac{1}{2} m_2 \left( \dot{\mathbf{R}} - \frac{m_1}{M} \dot{\mathbf{r}} \right)^2 - U(|\mathbf{r}|) \\ &= \frac{1}{2} M \dot{\mathbf{R}}^2 + \frac{1}{2} \mu \dot{\mathbf{r}}^2 - U(|\mathbf{r}|) \end{aligned}$$

where *reduced mass*  $\mu = \frac{m_1 m_2}{M}$ .

Since the energy does not change as the center of mass moves,  $\frac{\partial \mathcal{L}}{\partial \mathbf{R}} = 0$ . Then the equation of motion will be

$$\begin{aligned} M \ddot{\mathbf{R}} &= 0 \\ \mu \ddot{\mathbf{r}} &= - \frac{dU}{d|\mathbf{r}|} \frac{\mathbf{r}}{|\mathbf{r}|} \end{aligned}$$

We can also transform  $\mathbf{r}$  further into spherical coordinates to obtain the one-dimensional equation of motion.

## § 1.5 Legendre transforms

## § 1.6 Hamiltonian formulation

Recall our mass metric matrix  $G$

$$G_{\alpha\beta}(\mathbf{q}) = \sum_{i=1}^N m_i \frac{\partial \mathbf{r}_i}{\partial q_\alpha} \frac{\partial \mathbf{r}_i}{\partial q_\beta} \quad (1.19)$$

## **§ 1.7 Summary of chapter**

# Chapter 2 内容示例

## § 2.1 用户环境示例

### GRE 备考指南

<https://qyxf.site/latest/GRE备考指南-v2.0.pdf>

### 军事理论教程

<https://qyxf.site/latest/军事理论教程.pdf>

### 分析力学笔记

<https://qyxf.site/latest/分析力学笔记-v1.0.pdf>

### 大学物理题解

<https://qyxf.site/latest/大物题解（上）.pdf>

### 实变函数习题解答

<https://qyxf.site/latest/实变函数习题解答.pdf>

### 计算方法撷英

<https://qyxf.site/latest/计算方法撷英-v1.1.pdf>

### 计算机程序设计指南

<https://qyxf.site/latest/计算机设计程序指南.pdf>

### Equation 1 *Euler-Lagrangian equation*

$$\frac{d^2}{dt^2} \frac{\partial \mathcal{L}}{\partial \dot{\mathbf{r}}_i} =$$

**定义 1** 极限就是超越自我。

**定理 1** 任何极限都可以直接观察得出。

**引理 2** 以上内容，纯属扯淡。

**推论 3** 这是一个推论。

注记

好好学习，天天向上。

警告

今天你学习了吗？

§ 2.2 列表样式

- 这是第一层
- 这也是第一层
  - 这是第二层
    - 这是第三层
- 1. 这是第一层
- 2. 这也是第一层
  - 1. 这是第二层
    - (1) 这是第三层

§ 2.3 正文示例

**微分学**（*differential calculus*）是微积分的一部分，是通过导数和微分来研究曲线斜率、加速度、最大值和最小值的一门学科，也是探讨特定数量变化速率的学科。微分学是微积分的两个主要分支之一，另一个分支则是**积分学**，探讨曲线下的面积。

Tab. 2.1 常用导数

原函数	导函数	原函数	导函数
$C$	0	$\ln x$	$\frac{1}{x}$
$x^\mu$	$\mu x^{\mu-1}$	$\sin x$	$\cos x$
$e^x$	$e^x$	$\cos x$	$-\sin x$

……几乎所有量化的学科中都有微分的应用。例如在物理学中，运动物体其位移对时间的导数即为其速度，速度对时间的导数就是加速度、物体动量对时间的导数即为物体所受的力，重新整理后可以得到牛顿第二运动定律  $F = ma$ 。化学反

应的化学反应速率也是导数。在运筹学中，会透过导数决定在运输或是设计上最有效率的做法。

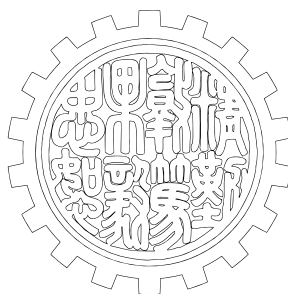


Fig. 2.1 V2 版本的封面图片

导数常用来找函数的极值。含有微分项的方程式称为**微分方程**，是自然现象描述的基础。微分以及其广义概念出现在许多数学领域中，例如复分析、泛函分析、微分几何、测度及抽象代数<sup>(1)</sup>。

## § 2.4 引导命令示例

**练习 1** 试用配方法求解方程：

$$ax^2 + bx + c = 0 \quad (2.1)$$

**解** 首先，方程左右两侧同除以  $a$ ，得到

$$x^2 + \frac{b}{a}x + \frac{c}{a} = 0$$

根据一次项来配方，按公式  $(x + A)^2 = x^2 + 2Ax + A^2$  配出常数项：

$$x^2 + \frac{b}{a}x + \left(\frac{b}{2a}\right)^2 + \frac{c}{a} - \left(\frac{b}{2a}\right)^2 = 0$$

配方并移项得到

$$\left(x + \frac{b}{2a}\right)^2 = \frac{b^2}{4a^2} - \frac{c}{a}$$

方程左右开方，得

$$x + \frac{b}{2a} = \pm \sqrt{\frac{b^2}{4a^2} - \frac{c}{a}}$$

<sup>(1)</sup>以上内容摘自维基百科中文词条 — 微分学：<https://zh.wikipedia.org/wiki/微分学>。

从而得到方程 (2.1) 之解为

$$x = -\frac{b}{2a} \pm \sqrt{\frac{b^2}{4a^2} - \frac{c}{a}} \quad (2.2)$$

该式即为一元二次方程的通用求根公式。

分析 在这一问题中，需要注意以下几点 [1, 2]:

- .....
- .....
- .....

# References

- [1] KNUTH D E. The  $\text{T}_{\text{E}}\text{X}$ book [M]. Addison-Wesley: Reading, 1986.
- [2] 刘海洋.  $\text{L}_{\text{A}}\text{T}_{\text{E}}\text{X}$  入门 [M]. 人民邮电出版社: 北京, 2013.