CLASS

the Cosmological Linear Anisotropy Solving System¹



Julien Lesgourgues, Deanna C. Hooper TTK, RWTH Aachen University

Kavli Institute for Cosmology, Cambridge, 11-13.09.2018

¹ code developed by Julien Lesgourgues & Thomas Tram plus many others

Lecture 3: coding with class

- 1 overall structure of class
- 2 dynamical indexing rules
- input parameters
- 4 error management rules
- adding features

In CLASS, what is a module?

- a file include/xxx.h containing some declarations
- a file source/xxx.c containing some functions
- each module is a associated with a structure xx, containing all what other modules need to know, and nothing else
- some fields in this structure are filled in the input.c module (input parameters relevant for this module)
- all other fields are filled by a function xxx_init(...)
- "executing a module" ≡ calling xxx_init(...)



In include/background.h: localise struct background
In source/background.c: localise background_init()

List of structures associated to modules:

module structure ab. * main content

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau)),$ and harmonic spectra are given by $C^{ij}_l=4\pi\int\frac{dk}{k}\mathcal{P}(k)\Delta^i_l(k)\Delta^j_l(k).$

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters

In a flat universe, line-of-sight integrals read $\Delta_l^i(k) = \int d\tau S^i(k,\tau) j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C_I^{ij} = 4\pi \int \frac{dk}{k} \mathcal{P}(k) \Delta_I^i(k) \Delta_I^j(k)$.

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters background quantities as funct. of $ au$
background.c	background	ba	pba	

In a flat universe, line-of-sight integrals read $\Delta_l^i(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C_I^{ij} = 4\pi \int \frac{dk}{k} \mathcal{P}(k) \Delta_I^i(k) \Delta_I^j(k)$.

List of structures associated to modules:

module	structure	ab.	*	main content
input.c background.c thermodynamics.c	precision background thermodynamics	pr ba th	ppr pba pth	all precision parameters background quantities as funct. of τ thermo. quantities as funct. of z

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C^{ij}_l=4\pi\int \frac{dk}{k}\mathcal{P}(k)\Delta^j_l(k)\Delta^j_l(k)$.

◆ロ → ◆昼 → ◆ き → り へ ○

List of structures associated to modules:

module	structure	ab.	*	main content
input.c background.c thermodynamics.c	precision background thermodynamics	pr ba th	ppr pba pth	all precision parameters background quantities as funct. of τ thermo, quantities as funct, of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$

In a flat universe, line-of-sight integrals read $\Delta_l^i(k) = \int d\tau S^i(k,\tau) j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C_I^{ij} = 4\pi \int \frac{dk}{k} \mathcal{P}(k) \Delta_I^i(k) \Delta_I^j(k)$.

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters
background.c	background	ba	pba	background quantities as funct. of $ au$
thermodynamics.c	thermodynamics	th	pth	thermo. quantities as funct. of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$
primordial.c	primordial	pm	ppm	primordial spectra $\mathcal{P}(k)$

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C^{ij}_l=4\pi\int \frac{dk}{k}\mathcal{P}(k)\Delta^j_l(k)\Delta^j_l(k)$.

◆ロ → ◆団 → ◆ 豆 → ◆ 豆 ・ 夕 へ ○

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters
background.c	background	ba	pba	background quantities as funct. of $ au$
thermodynamics.c	thermodynamics	th	pth	thermo. quantities as funct. of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$
primordial.c	primordial	pm	ppm	primordial spectra $\mathcal{P}(k)$
nonlinear.c	nonlinear	nl	pnl	non-linear corrections $lpha_{ m NL}(k, au)$

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C^{ij}_l=4\pi\int \frac{dk}{k}\mathcal{P}(k)\Delta^j_l(k)\Delta^j_l(k)$.

→□ → ◆□ → ◆ = → ◆ = → ◆ < ○</p>

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters
background.c	background	ba	pba	background quantities as funct. of $ au$
thermodynamics.c	thermodynamics	th	pth	thermo. quantities as funct. of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$
primordial.c	primordial	pm	ppm	primordial spectra $\mathcal{P}(k)$
nonlinear.c	nonlinear	nl	pnl	non-linear corrections $lpha_{ m NL}(k, au)$
transfer.c	transfers	tr	ptr	harmonic transfer functions $\Delta^i_l(k)$

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C^{ij}_l=4\pi\int \frac{dk}{k}\mathcal{P}(k)\Delta^i_l(k)\Delta^j_l(k)$.

∢ロト ∢部 ▶ ∢ 差 ▶ ∢ 差 ▶ ○ 差 ○ 釣 Q C

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters
background.c	background	ba	pba	background quantities as funct. of $ au$
thermodynamics.c	thermodynamics	th	pth	thermo. quantities as funct. of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$
primordial.c	primordial	pm	ppm	primordial spectra $\mathcal{P}(k)$
nonlinear.c	nonlinear	nl	pnl	non-linear corrections $lpha_{ m NL}(k, au)$
transfer.c	transfers	tr	ptr	harmonic transfer functions $\Delta_l^i(k)$
spectra.c	spectra	sp	psp	linear and/or non-linear $P(k,z)$, C_ℓ 's

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau)),$ and harmonic spectra are given by $C^{ij}_l=4\pi\int\frac{dk}{k}\mathcal{P}(k)\Delta^j_l(k)\Delta^j_l(k).$

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters
background.c	background	ba	pba	background quantities as funct. of $ au$
thermodynamics.c	thermodynamics	th	pth	thermo. quantities as funct. of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$
primordial.c	primordial	pm	ppm	primordial spectra $\mathcal{P}(k)$
nonlinear.c	nonlinear	nl	pnl	non-linear corrections $lpha_{ m NL}(k, au)$
transfer.c	transfers	tr	ptr	harmonic transfer functions $\Delta_I^i(k)$
spectra.c	spectra	sp	psp	linear and/or non-linear $P(k,z)$, C_{ℓ} 's
lensing.c	lensing	le	ple	lensed CMB C_ℓ 's
_	_		_	

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau)),$ and harmonic spectra are given by $C^{ij}_l=4\pi\int\frac{dk}{k}\mathcal{P}(k)\Delta^j_l(k)\Delta^j_l(k).$

◆ロト ◆団ト ◆草ト ◆草ト 草 りゅ(

List of structures associated to modules:

module	structure	ab.	*	main content
input.c	precision	pr	ppr	all precision parameters
background.c	background	ba	pba	background quantities as funct. of $ au$
thermodynamics.c	thermodynamics	th	pth	thermo. quantities as funct. of z
perturbations.c	perturbs	pt	ppt	source functions $S^i(k,t)$
primordial.c	primordial	pm	ppm	primordial spectra $\mathcal{P}(k)$
nonlinear.c	nonlinear	nl	pnl	non-linear corrections $lpha_{ m NL}(k, au)$
transfer.c	transfers	tr	ptr	harmonic transfer functions $\Delta_I^i(k)$
spectra.c	spectra	sp	psp	linear and/or non-linear $P(k,z)$, C_{ℓ} 's
lensing.c	lensing	le	ple	lensed CMB C_ℓ 's
output.c	output	op	pop	description of output format

In a flat universe, line-of-sight integrals read $\Delta^i_l(k)=\int d\tau S^i(k,\tau)j_l(k(\tau_0-\tau))$, and harmonic spectra are given by $C^{ij}_l=4\pi\int \frac{dk}{k}\mathcal{P}(k)\Delta^j_l(k)\Delta^j_l(k)$.

Each module contains:

- a function xxx_init(...) filling the structure xx
- a function xxx_free(...) freeing all the memory allocated to this structure
- some functions xxx_external_1(...), ..., xxx_external_n(...) that can be called from other modules (e.g. to read correctly or interpolate the content of the structure xx)
- some functions xxx_internal_1(...), ..., xxx_internal_m(...) that are called only inside the module, within xxx_init(...)

Each module contains:

- a function xxx_init(...) filling the structure xx
- a function xxx_free(...) freeing all the memory allocated to this structure
- some functions xxx_external_1(...), ..., xxx_external_n(...) that can be called from other modules (e.g. to read correctly or interpolate the content of the structure xx)
- some functions xxx_internal_1(...), ..., xxx_internal_m(...) that are called only inside the module, within xxx_init(...)

Following order always respected in xxx.c:

```
xxx_external_1(...)
...
xxx_external_n(...)
xxx_init(...)
xxx_free(...)
xxx_internal_1(...)
...
xxx_internal_m(...)
```

Each module contains:

- a function xxx_init(...) filling the structure xx
- a function xxx_free(...) freeing all the memory allocated to this structure
- some functions xxx_external_1(...), ..., xxx_external_n(...) that can be called from other modules (e.g. to read correctly or interpolate the content of the structure xx)
- some functions xxx_internal_1(...), ..., xxx_internal_m(...) that are called only inside the module, within xxx_init(...)

Following order always respected in xxx.c:

```
xxx_external_1(...)
...
xxx_external_n(...)
xxx_init(...)
xxx_free(...)
xxx_internal_1(...)
...
xxx_internal_m(...)
```

Remark: a module in the CLASS code is very similar to a "class" in C++. We enjoy the structure of C++ with the speed and readability of C.

Following order always respected in xxx.c:

```
xxx_external_1(...)
...
xxx_external_n(...)
xxx_init(...)
xxx_free(...)
xxx_internal_1(...)
...
xxx_internal_m(...)
```



Count number of external and internal functions in source/background.c:
Search "int background." starting from top

The main() function of CLASS located in main/class.c could only contain:

```
int main() {
 input_init_..(..,ppr,pba,pth,ppt,ptr,ppm,psp,pnl,ple,pop);
 background_init(ppr,pba);
 thermodynamics_init(ppr,pba,pth);
 perturb_init(ppr,pba,pth,ppt);
 primordial_init(ppr,ppt,ppm);
 nonlinear_init(ppr,pba,pth,ppt,ppm,pnl);
 transfer_init(ppr,pba,pth,ppt,pnl,ptr);
 spectra_init(ppr,pba,ppt,ppm,pnl,ptr,psp);
 lensing_init(ppr,ppt,psp,pnl,ple);
 output_init(pba,pth,ppt,ppm,ptr,psp,pnl,ple,pop)
 /* all calculations done, free the structures */
 lensing_free(ple);
 spectra_free(psp);
 transfer_free(ptr);
 nonlinear_free(pnl);
 primordial_free(ppm);
 perturb_free(ppt);
 thermodynamics_free(pth);
 background_free(pba);
```

• Indexing is very generic in CLASS, same rules apply everywhere.

- Indexing is very generic in CLASS, same rules apply everywhere.
- Example: we want to define the indices of a vector of background quantities (stored in the background table).

- Indexing is very generic in CLASS, same rules apply everywhere.
- Example: we want to define the indices of a vector of background quantities (stored in the background table).
- We choose an abreviation of 2 letters for these indices, _bg_.

- Indexing is very generic in CLASS, same rules apply everywhere.
- Example: we want to define the indices of a vector of background quantities (stored in the background table).
- We choose an abreviation of 2 letters for these indices, _bg_.
- Then we declare all possible indices index_bg_<black="blabla"> in include/background.h (more precisely, inside the structure background, because these indices are necessary for manipulating the background table).

- Indexing is very generic in CLASS, same rules apply everywhere.
- Example: we want to define the indices of a vector of background quantities (stored in the background table).
- We choose an abreviation of 2 letters for these indices, _bg_.
- Then we declare all possible indices index_bg_<black="index-bg_declaration">black=round.h
 (more precisely, inside the structure background, because these indices are necessary for manipulating the background table).
- We also declare flags saying whether these indices need to be defined or not.

In include/background.h:

```
struct background {
    /** input parameters with assigned in the input module*
    double OmegaO_cdm;
    /** flags and indices **/
    int has_cdm; // can take values _TRUE_ or _FALSE_
    . . . .
    int index_bg_rho_cdm;
    . . .
    int bg_size;
    /** interpolation table **/
    double * background_table;
```

In source/background.c, the function background_indices() called at the beginning of background_init() assigns numerical value to indices, that the user will never need to know (quantities always written symbolically as y[pba->index_bg_rho_cdm])

```
int background_indices(pba,...) {
    /* initialize all flags */
    if (pba->0mega0_cdm != 0.)
       pba->has_cdm = _TRUE_;
    /* initialize all indices */
    index_bg=0;
    class_define_index(pba->index_bg_rho_cdm,
                       pba->has_cdm,
                       index_bg,
                        1):
    class_define_index(pba->index_bg_rho_fld,
                       pba->has_fld,
                       index_bg,
                        1):
    pba->bg_size = index_bg;
```

This logic is followed everywhere for all groups of indices! Examples:

- in background.c: index_bg_... for all background variables
- in background.c: index_bi_... subset of backg. var. integrated over time
- in thermodynamics.c: index_th_... for all thermodynamics variables
- in perturbations.c: index_pt_... perturbation var. integrated over time
- in perturbations.c: index_mt_... metric perturbations
- in perturbations.c: index_md_... list of modes (scalar, vector, tensor)
- in perturbations.c: index_ic_... list of initial conditions (AD, CDI, NID...)
- in perturbations.c: index_tp_... list of type of required source (temperature, polarisation, matter fluctuation...)
- in perturbations.c: index_ap_... list of approximations that may be used
- etc. etc.



Check in your include/*.h files!

Input management in class

```
Terminal
                                                        Python wrapper
           file xxx.ini
input_init_from_argument(...)
             (parser)
                                                          .set(...)
struct file_content fc; (all parameter names/values stored as arrays of strings)
                               input_init(...)
                         input_read_parameters(...)
     (assign all default values + interprete input + update some parameters)
       relevant parameters only get stored in the structures of each module
```

For special parameters requiring a shooting method: repeated calls of input_read_parameters(...) from input_init(...) until shooting target is met.

Input management in class

For normal parameters (no shooting): example of CDM density:

```
/** - Omega_O_cdm (CDM) */
class_call(parser_read_double(pfc, "Omega_cdm",&param1,&
    flag1, errmsg),
           errmsg,
           errmsg);
class_call (parser_read_double(pfc, "omega_cdm", &param2, &
    flag2, errmsg),
           errmsg,
           errmsg);
class_test(((flag1 == _TRUE_) && (flag2 == _TRUE_)),
           errmsg,
           "In input file, you can only enter one of
               Omega_cdm or omega_cdm, choose one");
if (flag1 == _TRUE_)
  pba->0mega0_cdm = param1;
if (flag2 == _TRUE_)
  pba->0mega0_cdm = param2/pba->h/pba->h;
```

Input management in class

For shooting parameters, establish mapping between *target parameter*, *unknown* parameter and *level*. Currently:

target parameter	unknown parameter	level
$100 \times \theta_s$	h	thermodynamics
σ_8	A_s	spectra
$\Omega_{ m dcdm}$	$ ho_{ m dcdm}^{ m ini}$	background
***	•••	

... plus a few others (alternative parametrizations of decaying CDM, quintessence parameters).

If you need to add such parameters: see how it is done e.g. for 100*theta_s and replicate the structure!



Run with an input file containing only

 $omega_b = 0.07$

By following a few general rules, we get automatically some very informative error messages like:

```
Error in thermodynamics_init
=>thermodynamics_init(L:292) :error in
    thermodynamics_helium_from_bbn(ppr,pba,pth);
=>thermodynamics_helium_from_bbn(L:1031) :condition (omega_b
    > omegab[num_omegab-1]) is true; You have asked for an
    unrealistic high value omega_b = 7.e-02. The
    corrresponding value of the primordial helium fraction
    cannot be found in the interpolation table. If you
    really want this value, you should fix YHe to a given
    value rather than to BBN
```

We only wrote the piece starting with "You have asked...". All the rest was generated automatically by the code. This follows from following everywhere 5 rules.

Rule 1:

All functions are of type int, and return either _SUCCESS_ or _FAILURE_ (defined internally in include/common.h: #define _SUCCESS_ 0 , #define _FAILURE_ 1)

```
int function(input, &output) {
    ...
    if (something goes wrong) return _FAILURE_;
    ...
    return _SUCCESS_;
}
```

Rule 2:

```
All functions are called with the macro class_call(.,.,.) (all macros class_xxx(...) are defined in include/common.h):
```

Rule 3:

Each of the 9 main structures xx has a field called error_message. Any function in the module xxx.c is called xxx_something() and writes its error message in xx.error_message (if pxx is a pointer to xx, in pxx->error_message).

So if we are in perturb_init() and we call perturb_indices() we write:

But if we are in perturb_init() and we call background_at_tau() we write:

Rule 4:

Whenever an error could occur, we first write a test with the macro class_test(.,.,.):

```
class_test(condition, error_message, "Some text");
or
class_test(condition, error_message, "Some text and numbers
    %d %e",n,x);
```

Example:

In the text, no need to say in which function we are, or to write that the number of points is zero, or to put a \n, all this is done automatically.

Rule 5:

Always allocate memory with the macros class_alloc(), class_calloc(), class_realloc().

Instead of

```
malloc(parray, N*sizeof(double));
```

use

```
class_alloc(parray, N*sizeof(double), pxx->error_message);
```

If allocation fails (N too big, null or negative), the function will automatically return a <code>_FAILURE_</code> and the code will return an appropriate error message:

```
Error running background_init
=>background_init(L:537):error in background_solve(ppr,pba);
=>background_solve(L:1303):could not allocate pvecback with
    size -8
```

Useful CLASS macros:

```
class_call(function, errmsg_input, errmsg_output);
class_call_parallel(...);
class_call_except(...,[line of code;line of code;...;]);

class_test(condition, errmsg_output, "message"[,args]);
class_test_parallel(...);
class_test_except(...,[line of code;line of code;...;]);
class_stop(errmsg_ouput, "message"[,args]);

class_alloc(pointer,size);
class_alloc_parallel(...);
class_realloc(...);
class_calloc(...);
```



You can see them in include/common.h files!

Few special cases:

in main/class.c there is no "higher level" so the 10 initialisation functions are called like e.g.:

```
int main(int argc, char **argv) {
    if (background_init(&pr,&ba) == _FAILURE_) {
    printf("\n\nError running background_init \n=>%s\n"
        ,ba.error_message);
   return _FAILURE_;
   }
```

Few special cases:

• in main/class.c there is no "higher level" so the 10 initialisation functions are called like e.g.:

```
int main(int argc, char **argv) {
   if (background_init(&pr,&ba) == _FAILURE_) {
     printf("\n\nError running background_init \n=>%s\n"
        ,ba.error_message);
   return _FAILURE_;
}
```

• the input module does not have an error message attached to its structure, and just uses the local variable errmsg. So inside this module, the calls read e.g.:

Few special cases:

in main/class.c there is no "higher level" so the 10 initialisation functions are called like e.g.:

```
int main(int argc, char **argv) {
    if (background_init(&pr,&ba) == _FAILURE_) {
    printf("\n\nError running background_init \n=>%s\n"
        ,ba.error_message);
    return _FAILURE_;
    }
```

 the input module does not have an error message attached to its structure, and just uses the local variable errmsg. So inside this module, the calls read e.g.:

```
class_call(background_ncdm_init(ppr,pba),
               pba->error_message,
               errmsg);
class_call(parser_read_file(...,errmsg),
               errmsg,
               errmsg);
```

 when calling external functions not in the 10 modules we must pass the error message as an argument:

```
class_call(array_interpolate(...,pba->error_message),
               pba->error_message,
               pba->error_message);
```

90 Q

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

the logic is always the same:

define an acronym easy to search in the C files (e.g. for early dark energy: earde is good, ede is bad because it is inside "redefine", "needed", etc.)

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

- define an acronym easy to search in the C files (e.g. for early dark energy: earde is good, ede is bad because it is inside "redefine", "needed", etc.)
- 2 think of the feature closest to yours, and find its acronym (e.g. for fluid: fld)

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

- define an acronym easy to search in the C files (e.g. for early dark energy: earde is good, ede is bad because it is inside "redefine", "needed", etc.)
- 2 think of the feature closest to yours, and find its acronym (e.g. for fluid: fld)
- grep for all occurences of fld in include/*.h and source/*.c (normally they are all within some "if (has_fld){ ...}" and you can search directly for occurences of has_fld)

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

- define an acronym easy to search in the C files (e.g. for early dark energy: earde is good, ede is bad because it is inside "redefine", "needed", etc.)
- 2 think of the feature closest to yours, and find its acronym (e.g. for fluid: fld)
- grep for all occurences of fld in include/*.h and source/*.c (normally they are all within some "if (has_fld){ ...}" and you can search directly for occurences of has_fld)
- 4 duplicate these occurences

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

- define an acronym easy to search in the C files (e.g. for early dark energy: earde is good, ede is bad because it is inside "redefine", "needed", etc.)
- 2 think of the feature closest to yours, and find its acronym (e.g. for fluid: fld)
- grep for all occurences of fld in include/*.h and source/*.c (normally they are all within some "if (has_fld){ ...}" and you can search directly for occurences of has_fld)
- duplicate these occurences
- 6 change fld into earde

If you want to implement:

- a new species
- a new approximation scheme to simplify some equations in some regime
- a new mathematical description of an existing species (switching on more precise corrections, etc.)
- a new observable or output (new source function, new transfer function, new spectrum...)

- define an acronym easy to search in the C files (e.g. for early dark energy: earde is good, ede is bad because it is inside "redefine", "needed", etc.)
- 2 think of the feature closest to yours, and find its acronym (e.g. for fluid: fld)
- grep for all occurences of fld in include/*.h and source/*.c (normally they are all within some "if (has_fld){ ...}" and you can search directly for occurences of has_fld)
- duplicate these occurences
- 6 change fld into earde
- 6 change some equations to describe the specific properties of your feature