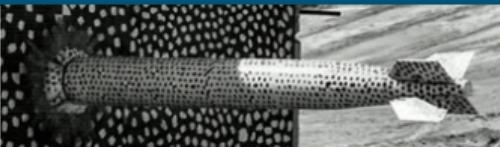




Sandia
National
Laboratories

The ALEGRA Finite Element Code and Trilinos



Presented by:

Tim Fuller



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Overview

What is ALEGRA?

Applications

Trilinos usage in ALEGRA

Conclusion



What is ALEGRA?



What is ALEGRA?

ALEGRA began development circa 1996 to solve the transient equations of Lagrangian solid dynamics (balance laws)

- conservation of mass

$$\frac{d}{dt} (\rho \, dv) = \frac{d}{dt} (dm) = 0$$

- conservation of momentum

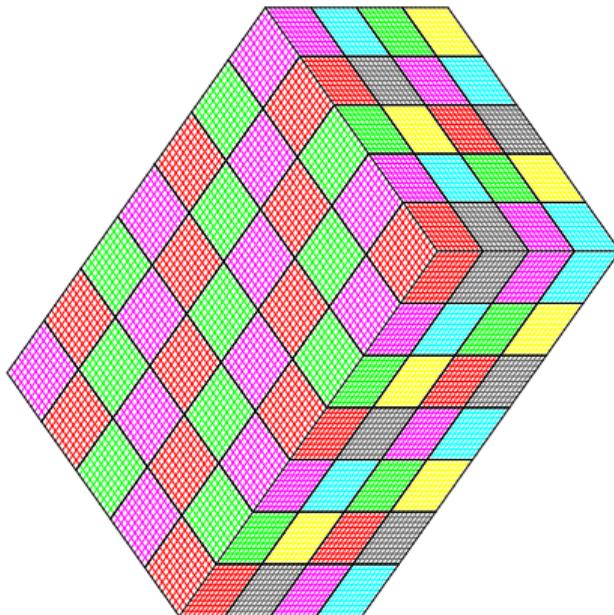
$$\rho \frac{d}{dt} \dot{\boldsymbol{U}} = \rho \ddot{\boldsymbol{U}} = \boldsymbol{\nabla} \cdot \boldsymbol{\sigma} + \boldsymbol{f}$$

- conservation of energy

$$\rho \frac{d}{dt} e = \rho s = \boldsymbol{\sigma} : \boldsymbol{D} - \boldsymbol{\nabla} \cdot \boldsymbol{q}$$

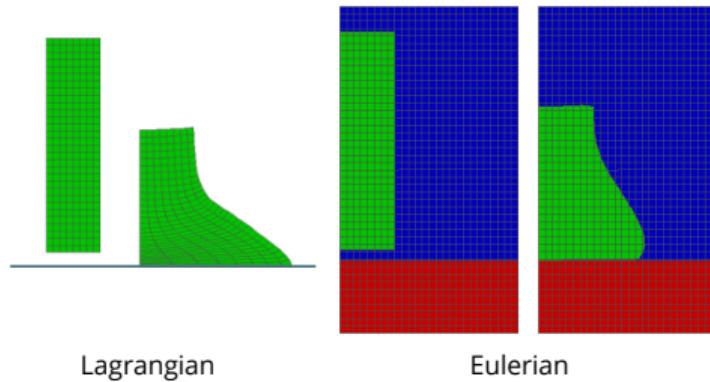
What is ALEGRA?

... using a finite element discretization ...



What is ALEGRA?

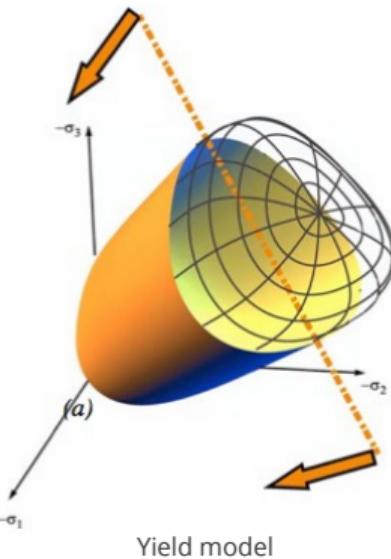
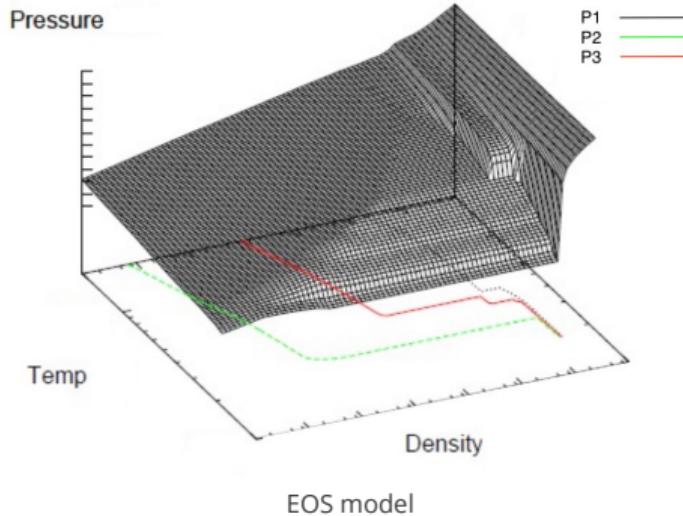
... with Lagrangian, Eulerian, Arbitrary Lagrangian-Eulerian (ALE), or XFEM capability ...



ALEGRA = **ALE** General Research Application

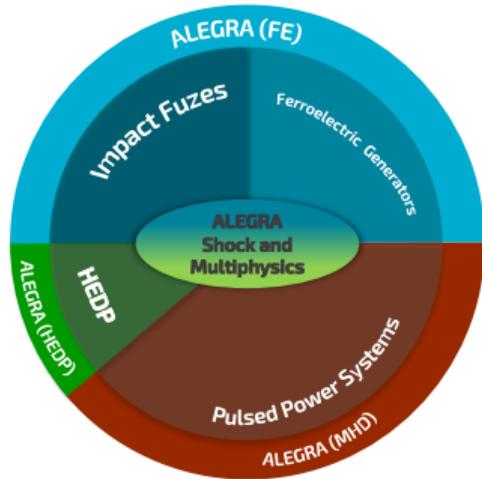
What is ALEGRA?

... using material data to close the system ...



What is ALEGRA?

... and operator splitting to add multiphysics.



Who uses ALEGRA?

Analysts developing models for

- Solid mechanics
- Shock hydrodynamics
- Electromagnetics
- Radiation hydrodynamics

Physicists and engineers interested in

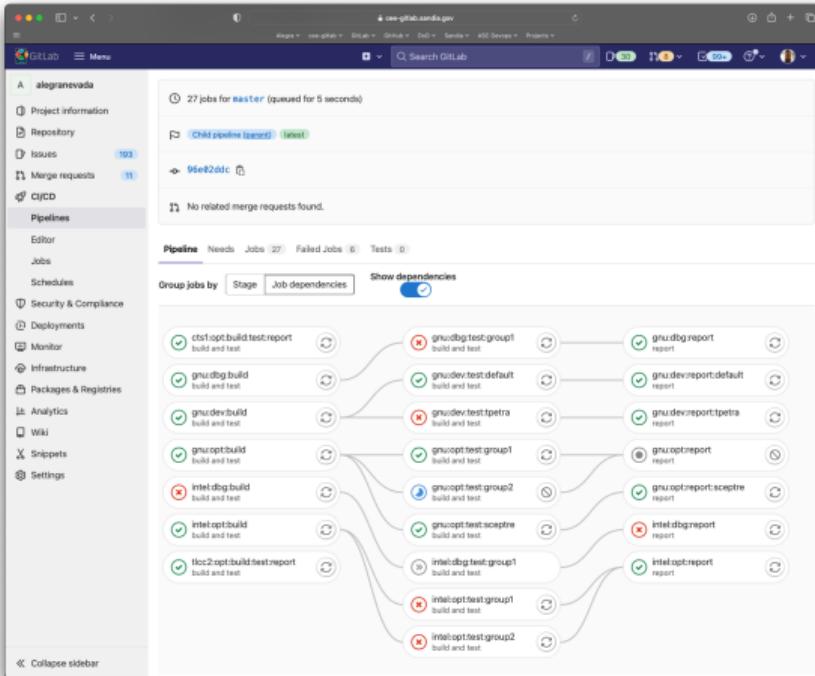
- Pulsed power
- Terminal ballistics
- Radiation effects

Civil servants and contractors in DOE and DoD physics/engineering R&D institutions

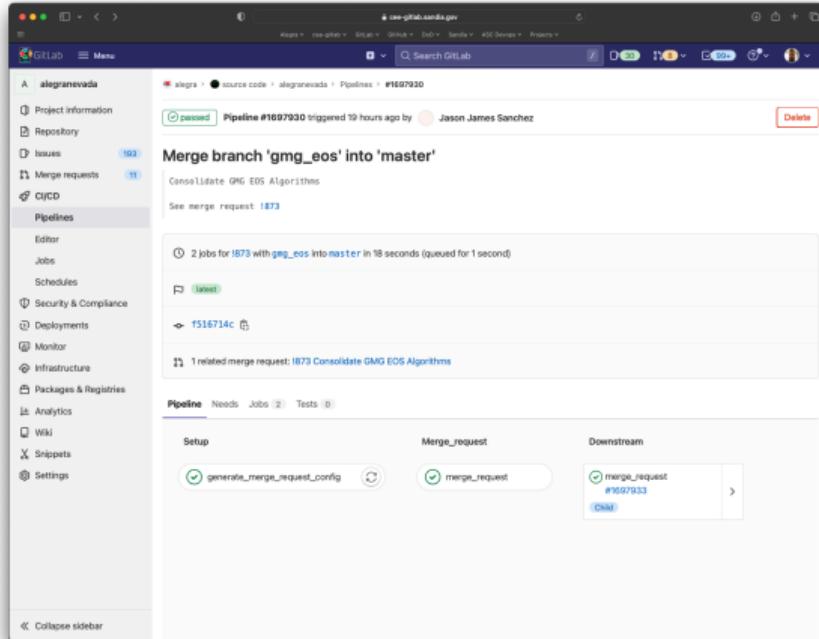
ALEGRA devops

- Git repository served on GitLab
- Spack for dependency management
- vvtest for testing
- CMake build system
- GitLab runners for CI/CD
- Toolset to glue it all together
- Toolset is python library used by developers and users to build, run, and test ALEGRA

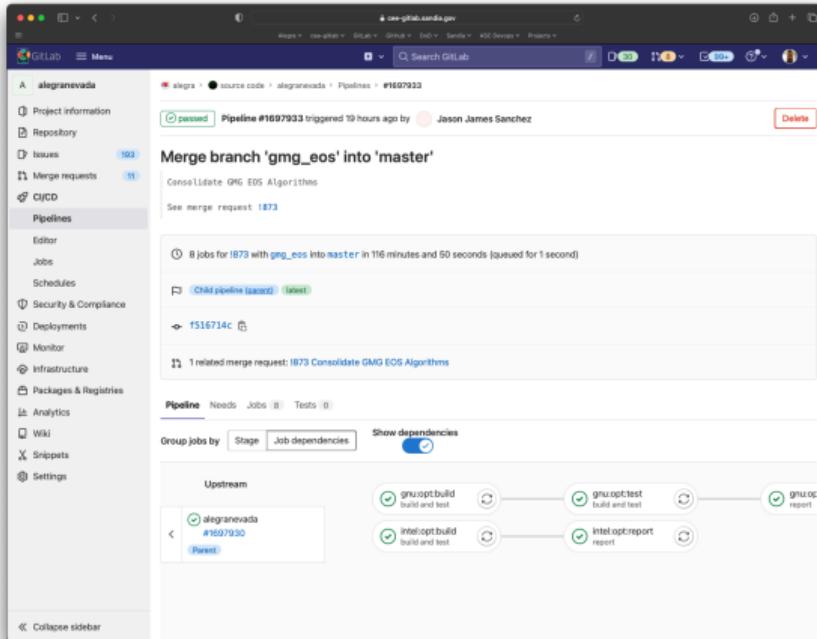
ALEGRA Devops



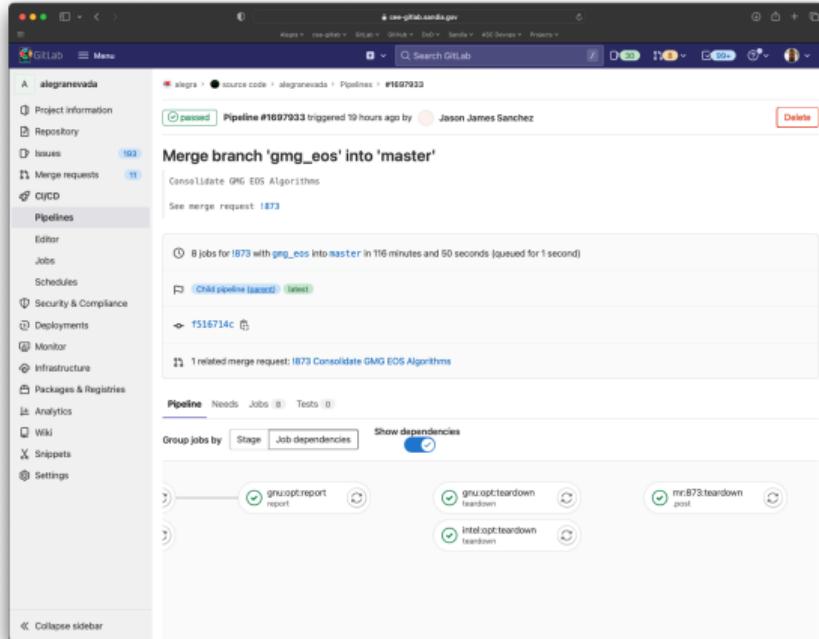
ALEGRA Devops



ALEGRA Devops



ALEGRA Devops

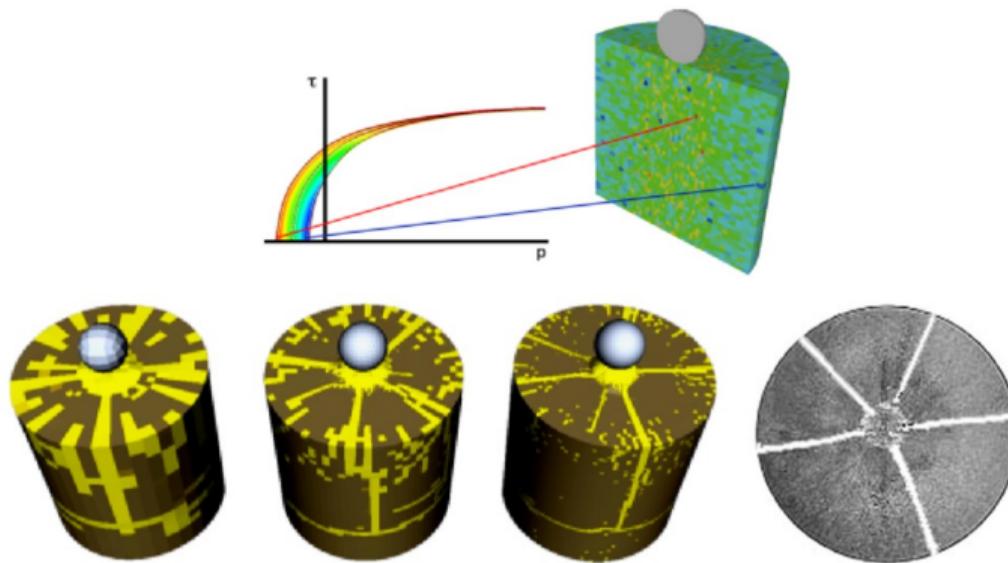




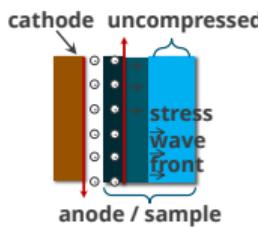
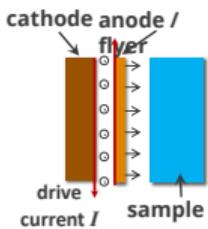
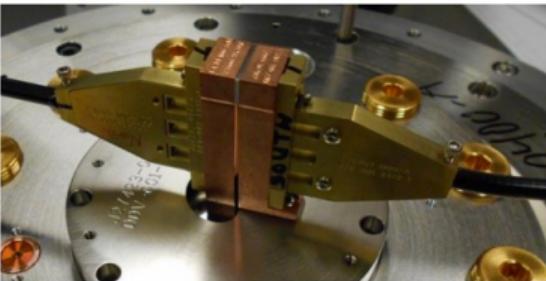
Applications



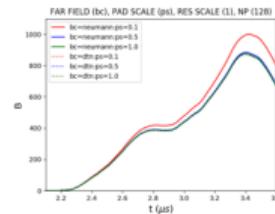
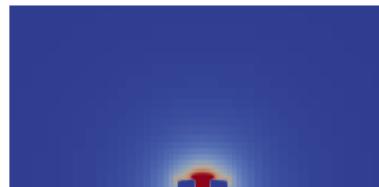
Validating theories of brittle damage



Dynamic material properties (DMP)



Compute Accurate Magnetic Field Drive History in the Gap at the Material Surface



Courtesy Allen Robinson

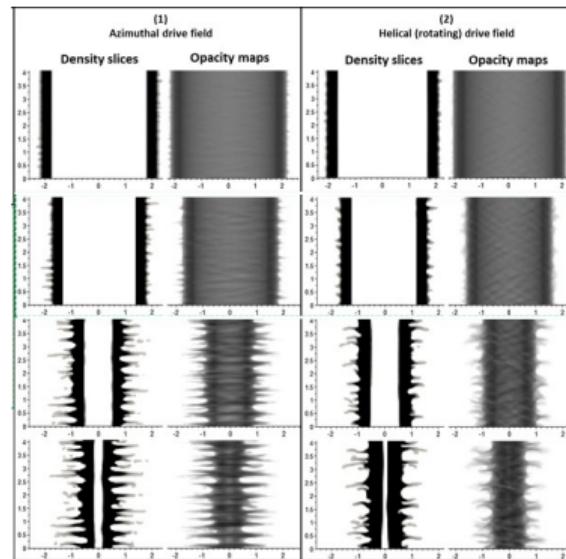
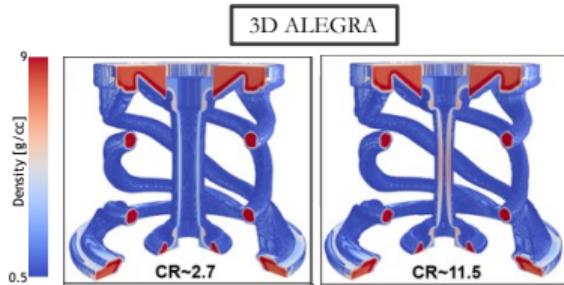
MRTI mitigation via screw pinch mechanism

Azimuthal drive field implosion

vs.

Helical (rotating) drive field produced
by Z design return current structure

$$\Xi = \frac{B_z}{B_\phi} \sim 0.5$$

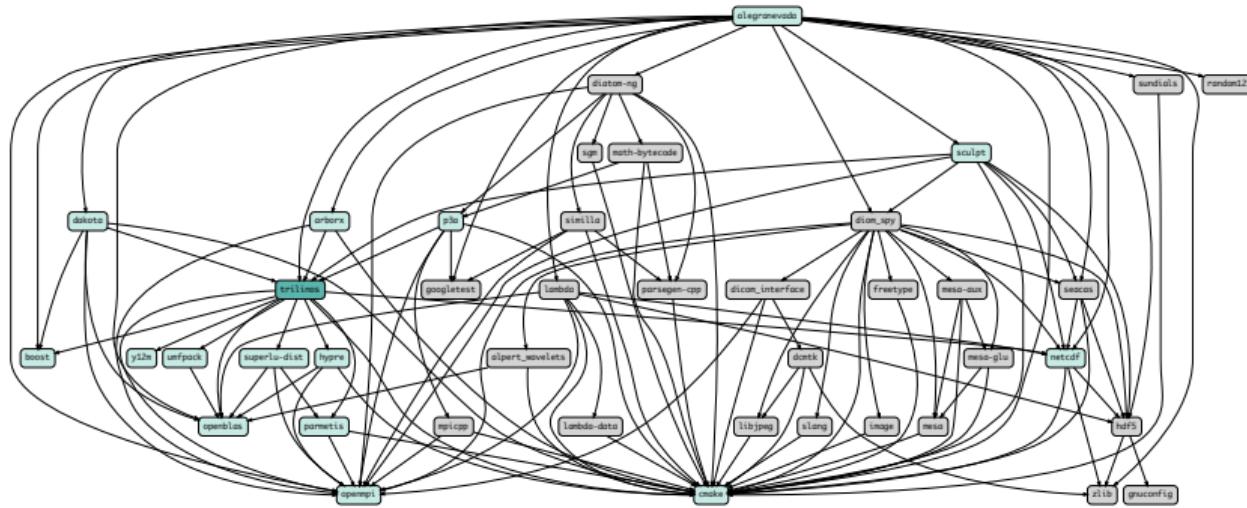




Trilinos usage in ALEGRA

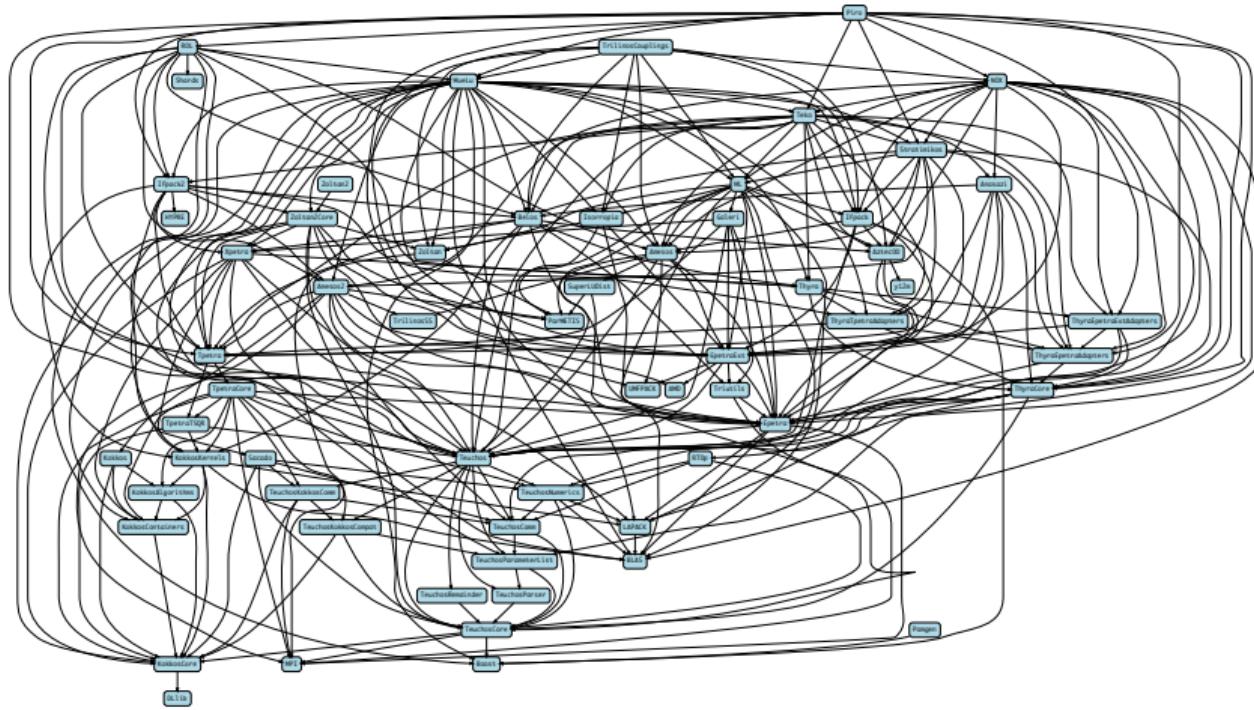


ALEGRA depends on many third-party libraries



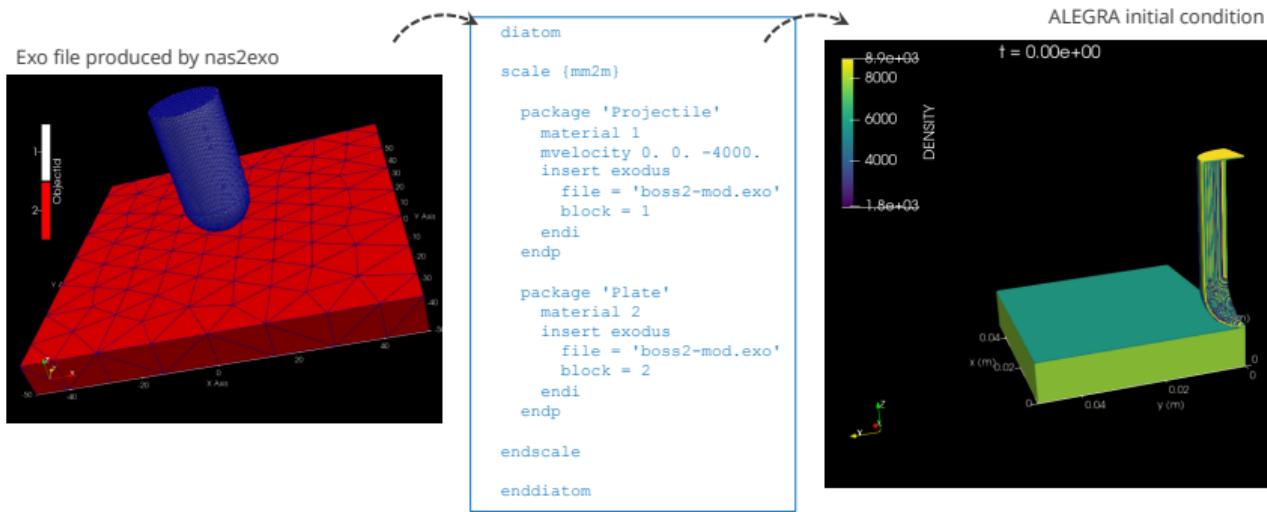
Trilinos plays a central role in ALEGRA by providing data structures and solvers required by key physics capabilities.

ALEGRA depends on many Trilinos packages



CAD insertion through SEACAS

The SEACAS `nas2exo` tool converts Nastran bulk data files to Exodus files that can be imported into an ALEGRA simulation



Epetra to Tpetra migration in MHD physics

Requirements

- Transition Epetra/AztecOO/ML to Tpetra/Belos/MueLu
- Enable 64 bit global indices to enable 2^{32} (or more) unknowns
- Ensure same results using Tpetra stack as Epetra stack

Strategy

- Support both Epetra and Tpetra via runtime switch
- Work with Trilinos developers to develop missing functionality in Tpetra stack
- Develop IO capabilities to compare results between Epetra/Tpetra stacks
- Nightly regression testing of both stacks

Lessons learned

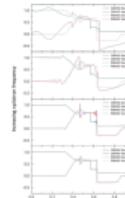
- Epetra does a lot of work under the hood that users of Tpetra must do themselves
- `Tpetra_FECrsGraph`, `Tpetra_FECrsMatrix`, and `Tpetra_FEMultiVector` had a non-intuitive interface (fixed)
- Belos and MueLu lack(ed) implementations for several solvers and preconditioners implemented in AztecOO/ML that users of ALEGRA depend on

3D GMHD(XMHD) ALEGRA

1) Original implementation of GMHD equations (Hall term) was found to contain a near null space but included non-physical oscillatory modes complicating scalable linear solves.

$$\int \left(\frac{\epsilon}{\Delta t} + \underbrace{\chi_\beta \sigma^{n+1} \Pi_P}_{\text{when Hall term dominates, } \approx \frac{1}{3} \text{ of spectrum is nearly 0}} \right) \mathbf{E}^{n+1} \cdot \Psi d\Omega + \int \frac{\Delta t}{\mu} \operatorname{curl} \mathbf{E}^{n+1} \cdot \operatorname{curl} \Psi d\Omega$$

solve for \mathbf{E}^{n+1}



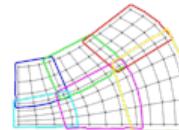
2) Reformulated into 2x2 system which avoids element projections in Ohm's law to produce a matrix more amenable to solver technology.

Block edge element matrix

$$\begin{bmatrix} \mathbb{A}_{11} & \mathbb{A}_{12} \\ \mathbb{A}_{21} & \mathbb{A}_{22} \end{bmatrix} \begin{bmatrix} \mathbf{E}^{n+1} \\ \mathbf{J}^{n+1} \end{bmatrix} = \begin{bmatrix} \mathbf{b}_1 \\ \mathbf{b}_2 \end{bmatrix}, \quad \begin{aligned} \mathbb{A}_{11}(\hat{\mathbf{E}}_i, \hat{\mathbf{E}}_j) &= \frac{\epsilon}{\Delta t} \int_{\Omega} \hat{\mathbf{E}}_i \cdot \hat{\mathbf{E}}_j d\Omega + \frac{\Delta t}{\mu} \int_{\Omega} \operatorname{curl} \hat{\mathbf{E}}_i \cdot \operatorname{curl} \hat{\mathbf{E}}_j d\Omega & \mathbb{A}_{12}(\hat{\mathbf{E}}_i, \hat{\mathbf{J}}_j) &= \int_{\Omega} \hat{\mathbf{E}}_i \cdot \hat{\mathbf{J}}_j d\Omega \\ \mathbb{A}_{21}(\hat{\mathbf{J}}_i, \hat{\mathbf{E}}_j) &= - \int_{\Omega} \sigma^{n+1} (\hat{\mathbf{J}}_i \cdot \hat{\mathbf{E}}_j) d\Omega & \mathbb{A}_{22}(\hat{\mathbf{J}}_i, \hat{\mathbf{J}}_j) &= \int_{\Omega} \left(\left(1 + \frac{\tau}{\Delta t} \right) (\hat{\mathbf{J}}_i \cdot \hat{\mathbf{J}}_j) - \frac{e\tau}{m_e} [\hat{\mathbf{J}}_i \cdot (\mathbf{B}^n \times \hat{\mathbf{J}}_j)] \right) d\Omega \end{aligned}$$

3) Designed patch preconditioners to resolve near null space errors and complement existing algebraic multigrid preconditioners.

Tuminaro Led Late Start LDRD – “Composing preconditioners for multiphysics PDE systems with applications to Generalized MHD”
SAND2022-12164R



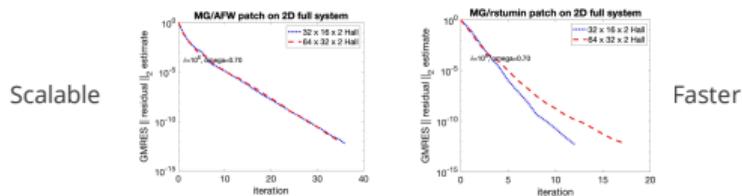
3D GMHD(XMHD) ALEGRA

ALEGRA GMHD Solver Strategy going Forward

- 1) Devise and implement a production 3D algebraic multigrid software strategy based on MatLab based geometric multigrid analysis of sample Alegra matrices.

Two types of patch preconditioners have been investigated:

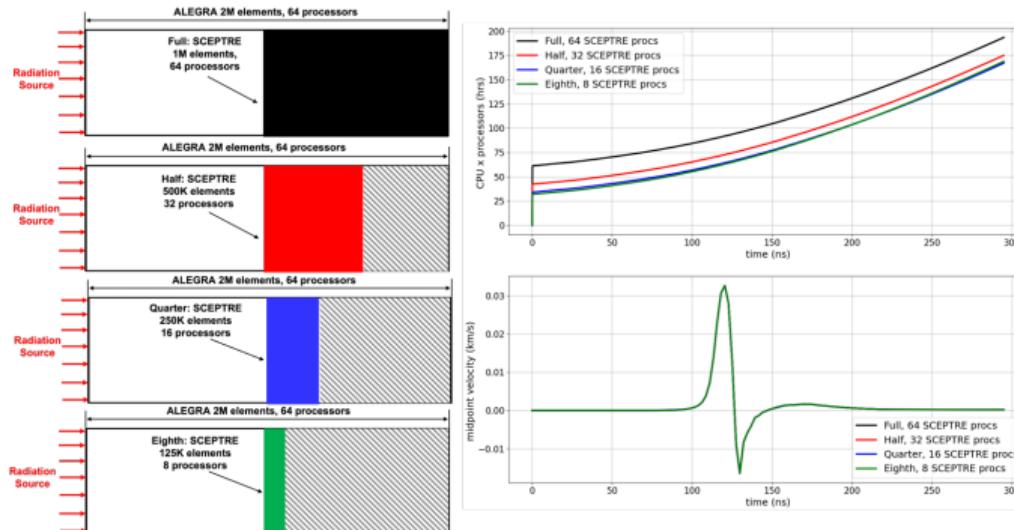
Arnold-Faulk-Winter (MG/AFW) and Adjacent Elements (MG/rstumin)



- 2) Trilinos impact: work should result in a better patch preconditioner infrastructure useful for dealing with multiple near null space preconditioning issues.
- 3) Will also investigate additional XMHD model extensions and the associated impact on discretization and solver strategy.

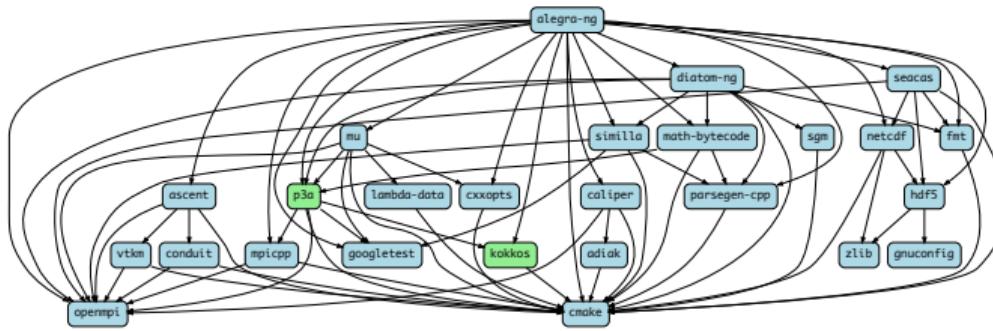
ALEGRA-SCEPTRE coupling through STK

- ALEGRA-SCEPTRE interface allows radiation transport physics subdomains, for which SCEPTRE is only applied to relevant parts of a problem
- Coupling achieved via STKSearch, STKTransfer, and MPI MPMD interfaces



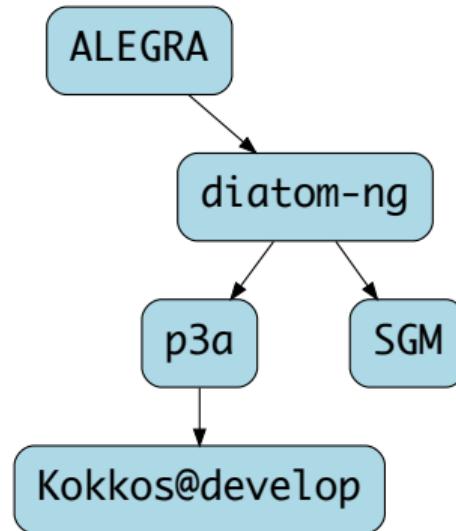
ALEGRA performance portability through Kokkos

The ALEGRA project has a performance portability effort to rewrite ALEGRA functionality using Kokkos and modern C++17

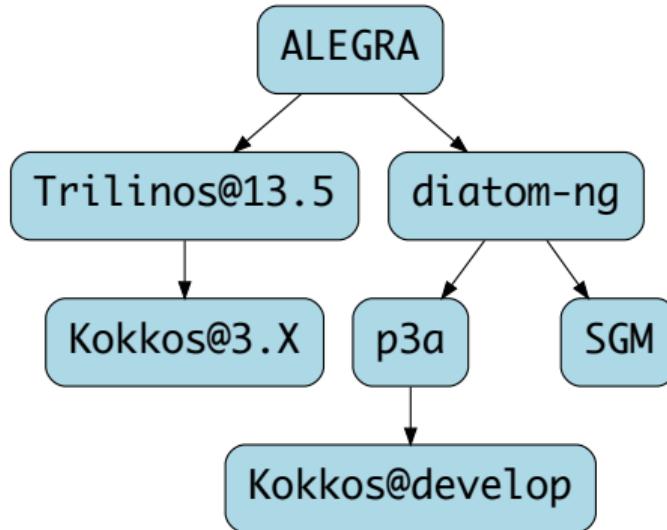


- Achieved across Intel CPUs, NVIDIA V100 GPUs, and NVIDIA A100 GPUs
- Developed SIMD types for guaranteed efficient outer loop vectorization on Intel and ARM CPUs
- SIMD types are being moved into Kokkos to benefit other Sandia applications
- Kokkos built standalone and for GPUs uses the built-in CMake support for CUDA, not the Kokkos compiler wrappers

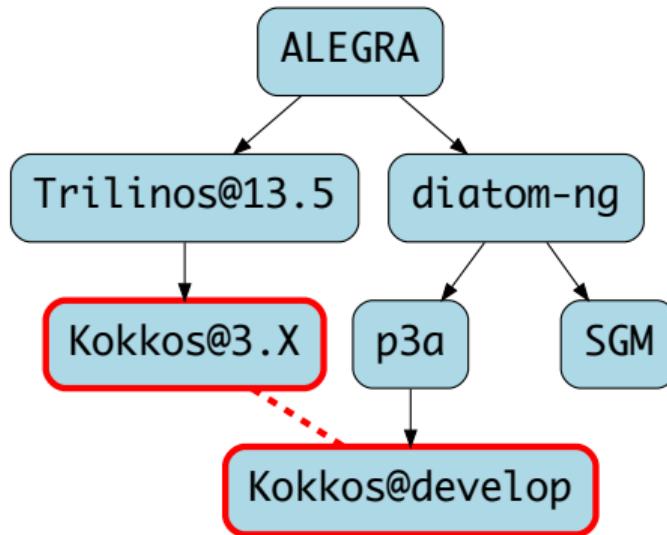
Next-gen geometry insertion



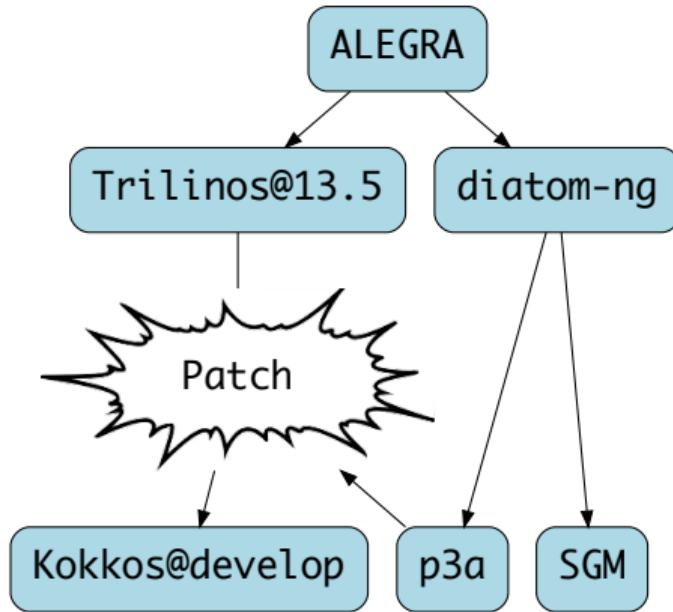
Next-gen geometry insertion



Next-gen geometry insertion

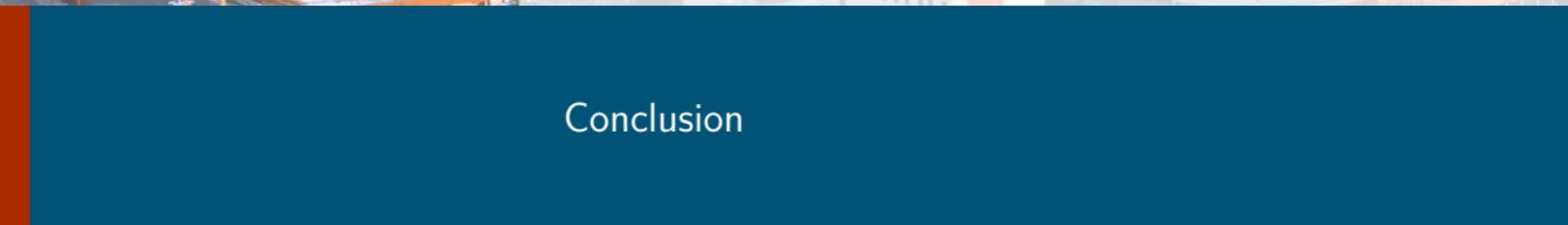


Next-gen geometry insertion





Conclusion



Conclusion

- ALEGRA depends on dozens of Trilinos packages (and sub-packages)
- Trilinos is a key component to ALEGRA delivering value to its customers
- ALEGRA uses both the Epetra and Tpetra stack and is currently migrating to fully Tpetra
- Infrequent Trilinos releases causes issues in a Spack eco system
- Inability to use intrinsic CMake CUDA support is an obstacle for ALEGRA's GPU strategy
- Inability to link external Kokkos causes significant devops issues