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Aim:

To determine the dispersive power of frism.

Affaratus required:

Spectrometer, brism, brism clamp, mercury valour lamp, lens

Formula used:

The formula for obtaining dispersive power of prism is,

w = 4v - 42

Here,  $\mu_{\nu}$  is refractive index of prism corresponding to violet colour light Similarly,  $\mu_{\nu}$  and  $\mu_{\nu}$  are corresponding to red & yellow colour lights

Theory:

The repractive index of the material of the prism can be calculated by,

 $\sin \left[ A + \delta m \right]$ 

sin [A]

where, Sm is angle of minimum deviation.

A is the angle of friem.

Upon passage through the prism, the white light is separated into its com-- forent colours - VIBGYOR. The separation of visible light into its different colours is known as dispersion.



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## Observation table:

Line	Vernier	Refracted nay		Direct readings			Difference	Mean, sm	
aurice.		MSR			MSR	VSR	Total	Sm	
Violet	٧,	70	0	70	110	20	110.3334	40.3334	40.25835
	Va	249.5		249.9834	290	10	290.1667	40.1833	ritareals A
Yellow	Vı	70.5	18	70.8	110	20	110.3334	39.5334	39.53335°
70000	Vz	250.5		250.6334	290	10	290.1667	39.5333	
Red	Vi	71	23	71.3834	110	20	110.3334	38.95	38.9495°
	V <sub>2</sub>	251	13	251,2167	290	10	290.1667	38.949	ne de all

## Calculations:

Angle of min. deviation, Sv = 40.25835°

$$\delta_y = 39.53335^{\circ}$$

$$\mu_{v} = \frac{\sin\left[\frac{40.25835 + 59.98}{2}\right]^{\circ}}{\sin\left[\frac{59.98}{2}\right]^{\circ}} = \frac{\sin\left(50.119175^{\circ}\right)}{\sin\left(29.99^{\circ}\right)} = 1.5352237$$

$$\mu_{y} = \frac{\sin\left[\frac{39.53335 + 59.98}{2}\right]}{\sin\left[\frac{59.98}{2}\right]^{\circ}} = \frac{\sin\left(49.756675^{\circ}\right)}{\sin\left(29.99^{\circ}\right)} = 1.527$$

$$\mu_{9} = \frac{38.9495 + 59.98}{2}^{\circ} = \frac{\cancel{\text{sin}} (49.46475^{\circ})}{\cancel{\text{sin}} (59.98)^{\circ}} = 1.52047219$$

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	The formula used for obtaining dispersive fower of frism is, $\omega = \frac{K v^- K r}{K r}$
	μμ-1
_	where, $\mu_{\nu}$ is refractive index of friem corresponding to violet colour light. similarly, $\mu_{r}$ & $\mu_{y}$ are corresponding to red & yellow colour lights.
	Least count of spectrometer:
	One main scale division (N) = 0.5° = 30 minutes
	Number of divisions on vernier (V) = 30
	Least count (LC) = N _ 30' _ 1 minute  V 30
	V 30
	Procedure:
1.	Determine the least count of spectrometer.
2	Set the telescope by focusing on distant object.
3.	Determine angle of brism
4.	Rotate vernier table so that different components of light from collimator falling on one of the face of prism & emerges through other face.  The telescope is turned to view the repracted images of different coloured elite on other face.
	falling on one of the face of prism & emerges through other face
5.	The telescope is turned to view the repracted images of different coloured
	slits on other face.
6	The vernier table is slowly turned in such a direction that the images of
	Slits is move directed towards the directed ray.  It will be found that at a certain fosition, images are stationary for a
7	It will be found that at a certain fosition, images are stationary for a
	colour for some moment.
8.	Note the readings on main scale & vernier scale for violet, red & yellow
9.	Carefully remove the friem from friem table.
10.	Carefully remove the prism from prism table. Turn the telescope parallel to collimator & note direct ray readings.
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Dispersive forcer, 
$$\omega = \frac{\mu_v - \mu_h}{\mu_h - 1} = \frac{1.5352237 - 1.52047219}{1.527 - 1}$$

$$\omega = \frac{0.01475}{0.527} = 0.027988$$

Dispersive power of brism is 0.028.

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the telescope forallet to willimater & note hired may mading.

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Result:	
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Angle of the frism = 59.9 Dispossive power of frism	= 0.027988 \( \sigma 0.028
Precautions & Lources of	ever:
1 light coming from slit	should be narrow and bright.
2. Telescope must be focus	d.
3. Keadings of vernier sca	le should be taken carefully
4. Lero error of vernier &	ale must be taken into account. welled with a spirit level before flacing friem.
5. Poism Table should be I	welled with a spirit level before placing frum.
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