

Bangladesh Olympiad on Astronomy and Astrophysics

FORMULA BOOK

Instructions

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1 Useful Constant

Table 1: Physical and orbital characteristics of selected bodies in the Solar System

Property	Sun ①	Mercury ♀	Venus ♀	Earth ⊕	Moon €	Mars 💍	Jupier 2I	Saturn わ	Uranus 🗆	Neptune Ψ
Mass m / Kg	1.989×10^{30}	3.302×10^{23}	4.868×10^{24}	5.972×10^{24}	7.348×10^{22}	6.419×10^{23}	1.899×10^{27}	5.685×10^{26}	8.681×10^25	1.024×10^{26}
Radius R / m	6.963×10^8	2.439×10^6	6.051×10^6	6.370×10^6	1.738×10^6	3.396×10^6	7.149×10^7	6.027×10^7	2.556×10^7	2.476×10^7
Orbital Semi- Major axis, a / m	-	5.791×10^{10}	1.82×10^{11}	1.496×10^{11}	3.843×10^{8}	2.279×10^{11}	7.785×10^{11}	1.433×10^{12}	2.877×10^{12}	4.503×10^{12}
Orbital period T	-	87.97 days	224.70 days	365.24 days	27.322 days (sidereal) 29.531 days (synodic)	683.97 days	11.86 years	29.46 years	84.32 years	164.79 years
Orbital Eccentricity ε	-	0.205	0.0067	0.0167	0.0549	0.0933	0.0488	0.0557	0.0444	0.0112

Table 2: Commonly used fundamental constants and unit denitions

Units and Physical Quantities	Universal Constants
Average Solar Temperature, $T_{\odot}=5778~K$ Atomic mass unit, $u=1.660539\times 10^{-27}~kg$	Planck's Constant $h = 6.62606957 \times 10^{-34} m^2 \ kg \ s^{-1}$ Reduced Planck's Constant $\frac{h}{2\pi} = \frac{h}{2\pi}$ Gravitational Constant $G = 6.67384 \times 10^{-11} N \ m^2 \ kg^{-2}$ Speed of Light $c = 2.99792458 \times 10^8 \ m \ s^{-1}$ Boltzmann's Constant $k_B = 1.3806488 \times 10^{-23} \ J \ K^{-1}$ Coulomb Constant $k_e = \frac{1}{4\pi\epsilon_0} = 8.98755179 \times 10^9 \ N \ m^2 \ C^{-2}$ Stefan-Boltzmann Constant $\sigma = 5.67 \times 10^{-8} \ W \ m^2 \ K^{-4}$ Electronic charge $q_e = 1.602 \times 10^{-19} \ C$ Fine structure constant $\alpha = \frac{k_e(q_e)^2}{\hbar c} \approx \frac{1}{137}$ Wien's Displacement Constant $b = 2.89776829 \times 10^{-3} \ m \ K$ Hubble Constant $H_0 = 67.80 \pm 0.77 \ km \ s^{-1} \ Mpc^{-1}$ (as of 03/13)



2 Useful Formula

Table 3: Mathematical Formulae

Description	Formula		
Arc length on a circle is proportional to circular angle in radians	$s = r\theta$ (Give the circumference when $\theta = 2\pi$)		
Law of sines	$\frac{\sin A}{a} = \frac{\sin B}{b} = \frac{\sin C}{c} = 2R \text{ (on a plane)}$ $\frac{\sin A}{\sin a} = \frac{\sin B}{\sin b} = \frac{\sin C}{\sin c} \text{ (on a sphere)}$		
Law of cosines	$c^2 = a^2 + b^2 - 2ab \cos C \text{ (on a plane)}$ $\cos c = \cos a \cos b + \sin a \sin b \cos C$ (on a sphere)		
Small-angle approximations ($x\ll 1$, x in radians)	$sin x \approx x$ $cos x \approx 1 - \frac{x^2}{2}$ $tan x \approx x$		
First-order binomial expansion	$(1+x)^y \approx 1 + xy$		
Dot product of vectors	$\overrightarrow{a} \cdot \overrightarrow{b} = a_x b_x + a_y b_y + a_z b_z = \overrightarrow{a} \overrightarrow{b} \cos\theta$ $(\overrightarrow{a} + \overrightarrow{b}) \cdot \overrightarrow{c} = \overrightarrow{a} \cdot \overrightarrow{c} + \overrightarrow{b} \cdot \overrightarrow{c}$ $ \overrightarrow{a} = \sqrt{\overrightarrow{a} \cdot \overrightarrow{a}}$		
Cross product of vectors	$ \overrightarrow{a} \times \overrightarrow{b} = \overrightarrow{a} \overrightarrow{b} \sin\theta$ $\overrightarrow{a} \times \overrightarrow{b} = -\overrightarrow{b} \times \overrightarrow{a}$ $(\overrightarrow{a} \times \overrightarrow{b}) \cdot \overrightarrow{c} = \overrightarrow{a} \cdot (\overrightarrow{b} \times \overrightarrow{c})$ $= (\overrightarrow{b} \times \overrightarrow{c}) \cdot \overrightarrow{a}$ $(\overrightarrow{a} + \overrightarrow{b}) \times \overrightarrow{c} = \overrightarrow{a} \times \overrightarrow{c} + \overrightarrow{b} \times \overrightarrow{c}$ $\widehat{x} \times \widehat{y} = \widehat{z}$		



Table 4: Classical Astrophysics

Description	Formula
Kinetic Energy	$E_{kin} = \frac{1}{2}mv^2$
Newton's Universal Law of Gravitation	$\overrightarrow{F} = -\frac{Gm_1m_2}{r^2}\widehat{r}$
Gravitational Potential Energy	$E_{pot} = -\frac{Gm_1m_2}{r}$
Gravitational binding energy of a uniform sphere	$U = -\frac{3}{5} \frac{GM^2}{R}$
Roche Limit for a small, rigid body of density ρ_2 approaching a larger body of density ρ_1 and radius R	$d_{Roche} = 1.26R \times (\frac{\rho_1}{\rho_2})^{-\frac{1}{3}}$
Angular Velocity ω and angular momentum l	$v = r\omega; \ \omega = 2\pi f = \frac{2\pi}{T}; \ l = I\omega = mr^2\omega$ (for orbiting bodies)
Centripetal acceleration and force	$a_c = \omega^2 r = \frac{v^2}{r}; \ F_c = ma_c$
Kepler's 3rd Law	$T^2 = \frac{4\pi^2}{G(m_1 + m_2)}a^3$
Hydrostatic Equilibrium	$\frac{dP}{dR} = -\rho_r \frac{GM_r}{R^2}$
Quantisation of energy-momentum	$E = hf = \hbar\omega; \ p = \frac{h}{\lambda} = \hbar k$
Planck's Law for intensity per unit frequency	$I_f = \frac{2\pi h f^3}{c^2} \frac{1}{e^{\frac{hf}{kt}} - 1}$
Stefan-Boltzmann Law	$L = 4\pi R^2 \sigma T^4$
Wien's Displacement Law	$\lambda_{max} = \frac{b}{T}$



Table 5: Relativistic Expressions

Description	Formula
Lorentz Factor	$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$
Velocity Addition	$u' = \frac{u+v}{1+\frac{uv}{c^2}}$
Relativistic Doppler Effect	$f_{observed} = f_{source} \sqrt{\frac{c+v}{c-v}}$
Relativistic Redshift	$z = \sqrt{\frac{c+v}{c-v}} - 1 \approx \frac{v}{c}$
Schwarzschild Radius	$r_s = \frac{2GM}{c^2}$
Redshift	$z = \frac{\lambda_{observed} - \lambda_{emitted}}{\lambda_{emitted}}$

Table 6: Practical Astronomy

Description	Formula
Keplerian orbital ellipse as a function of angular deviation from periapsis	$r = \frac{a(1 - \epsilon^2)}{1 + \epsilon \cos \phi}$
Orbital Eccentricity in terms of other parameters	$\epsilon = \frac{a - r_{periapsis}}{a} = \frac{r_{apoapsis} - a}{a} = \frac{r_a - r_p}{r_a + r_p}$
Rayleigh resolution criterion with aperture diameter D	$sin\Delta\phi_{min} = 1.220 \frac{\lambda}{D}$
Rocket Equation	$\Delta v = v_{exh} \log_e \frac{m_i}{m_j}$



Table 7: Distance Determination and Some Empirical Results

Description	Formula
Pogson's Law	$m_1 - m_2 = -2.5 \log(\frac{B_1}{B_2})$
Absolute Bolometric Magnitude	$M_{bol} = -2.5 \log \frac{L}{L_{\odot}} + 4.7554$
Distance modulus: difference between apparent and absolute magnitude	$m - M = 5\log\frac{d}{10pc}$
Relationship between Luminosity and Absolute Magnitude	$\frac{L_1}{L_2} = 10^{\frac{M_2 - M_1}{2.5}}$
Determining distance d in parsecs using an ob-served parallax \mathcal{P} in arcseconds	$D\approx\frac{1}{p}$
Color Indices	$B - V = M_B - M_V = m_B - m_V$ = 2.5 log($\frac{f(V)}{f(B)}$)
Period-Luminosity relationship for Cepheid variable stars, with period P in days	$M = -2.76 \log_{10} P - 1.4$
Absolute magnitude of RR Lyrae stars	$M \sim 0.75$
Absolute magnitude of Type Ia supernovae (at peak)	$M \sim -19.3$
Tully-Fisher Relation	$L \propto V^4$
Mass-Luminosity Relation for Main Sequence stars	$L \propto M^{3.5}$
Hubble's Law	$v = H_0 d$