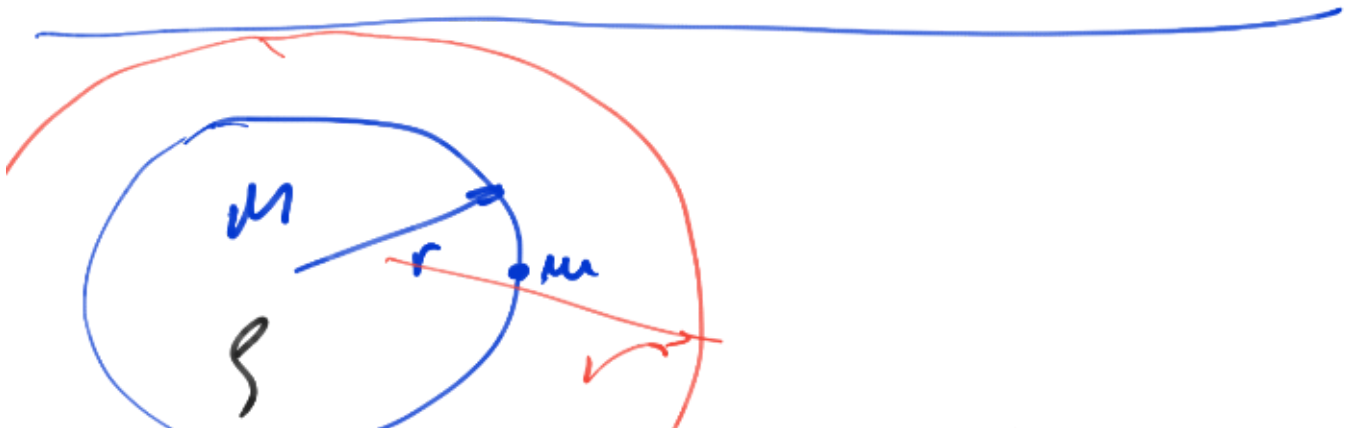


dist. Via Láctea y Andromeda
 $\sim 2 \text{ M.e.l.} \sim 0,6 \text{ Mpc.}$

$$v = \frac{1}{t_0} \cdot d = 67,3 \frac{\text{km/s}}{\text{Mpc}} \cdot 0,6 \text{ Mpc}$$

$$v = 40,38 \frac{\text{km}}{\text{s}}$$



$$\cancel{\frac{1}{2} m v^2} = \cancel{\frac{G M m}{r}} \Rightarrow v_e = \sqrt{\frac{2 G M}{r}}$$

$$v = H_0 r$$

$$V = \frac{4}{3} \pi r^3 \Rightarrow \rho = \frac{M}{V} = \frac{M}{\frac{4}{3} \pi r^3}$$

$$\Rightarrow M = \rho \cdot \frac{4}{3} \pi r^3$$

$$v_e = \sqrt{\frac{2 G \rho \frac{4}{3} \pi r^3}{r}}$$

$$v_e = \sqrt{\frac{8}{3} \pi G \rho r^2} \quad \left. \vphantom{\sqrt{\frac{8}{3} \pi G \rho r^2}}} \right\} \rho_c = \text{Densidad Crítica.}$$

$$v = H_0 r$$

$$\cancel{H_0^2 r^2} = \frac{8}{3} \pi G \cancel{\rho_c r^2}$$

$$\rho_c = \frac{H_0^2}{\frac{8}{3} \pi G} \Rightarrow \boxed{\rho_c = \frac{3 H_0^2}{8 \pi G}}$$

$$\rho_c = 6 \text{ protons/m}^3$$

$$\Omega_i = \frac{\rho_i}{\rho_c} \quad \text{y} \quad \Omega_{\text{tot}} = \sum_i \Omega_i$$

Abierto $\rightarrow \Omega_{\text{tot}} < 1 \Rightarrow \rho < \rho_c$ - Expansión perpetua

$\Omega_{\text{tot}} = 1 \Rightarrow \rho = \rho_c$ - Asintótico

$\Omega_{\text{tot}} > 1 \Rightarrow \rho > \rho_c$ - Colapso
→ Cerrado

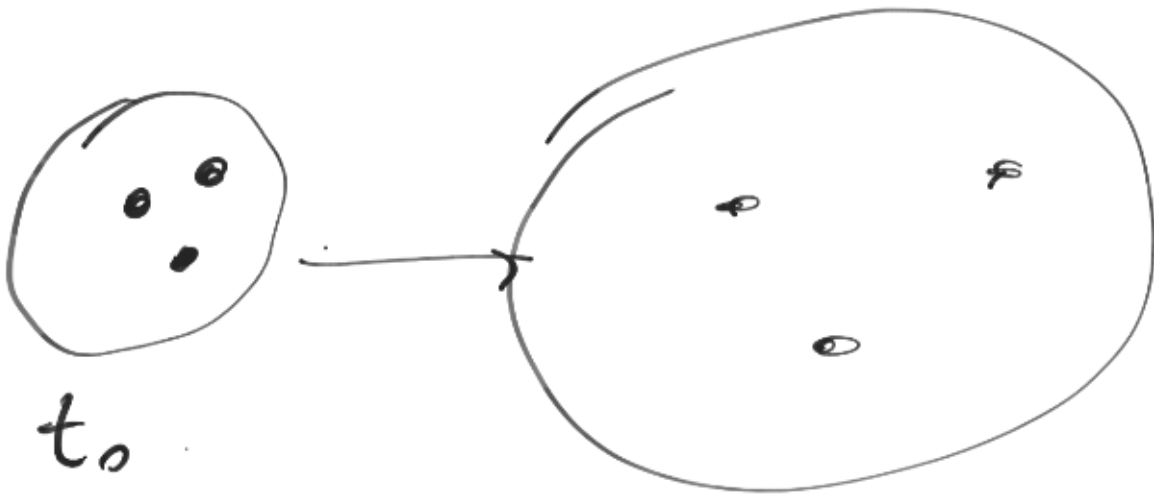
Plano.

Λ CDM

$$\Omega_{\text{tot}} = 1,001 \pm 0,001$$

$$\left. \begin{array}{l} \Omega_K \\ \Omega_m \\ \Omega_\gamma \\ \Omega_\nu \end{array} \right\} \Sigma = 1$$

i j



t_0



$E \propto \frac{1}{\lambda}$

$$g_m \sim 1/r^3 \sim \frac{1}{t^3}$$

$$g_g \sim 1/r^4 \sim \frac{1}{t^4}$$

$$\Omega_{\text{tot}} = 1.$$

$$\Omega_{\text{tot}} = \Omega_{\text{Cold}} + \Omega_m + \Omega_r + \Omega_{\text{disc.}}$$

$10^{-5} + 0,05 + 0,002 +$



Energia
Oscura

$$\Omega_H = 0,259 \quad \Omega_\Lambda = 0,68$$

Dark
Matter

Dark
Energy

Modelo Estándar Cosmológico

$$8\pi G^{\mu\nu} = R^{\mu\nu}$$



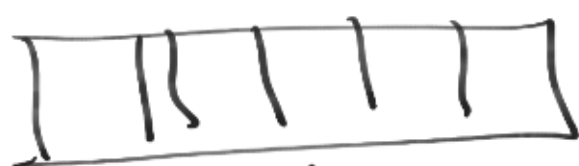
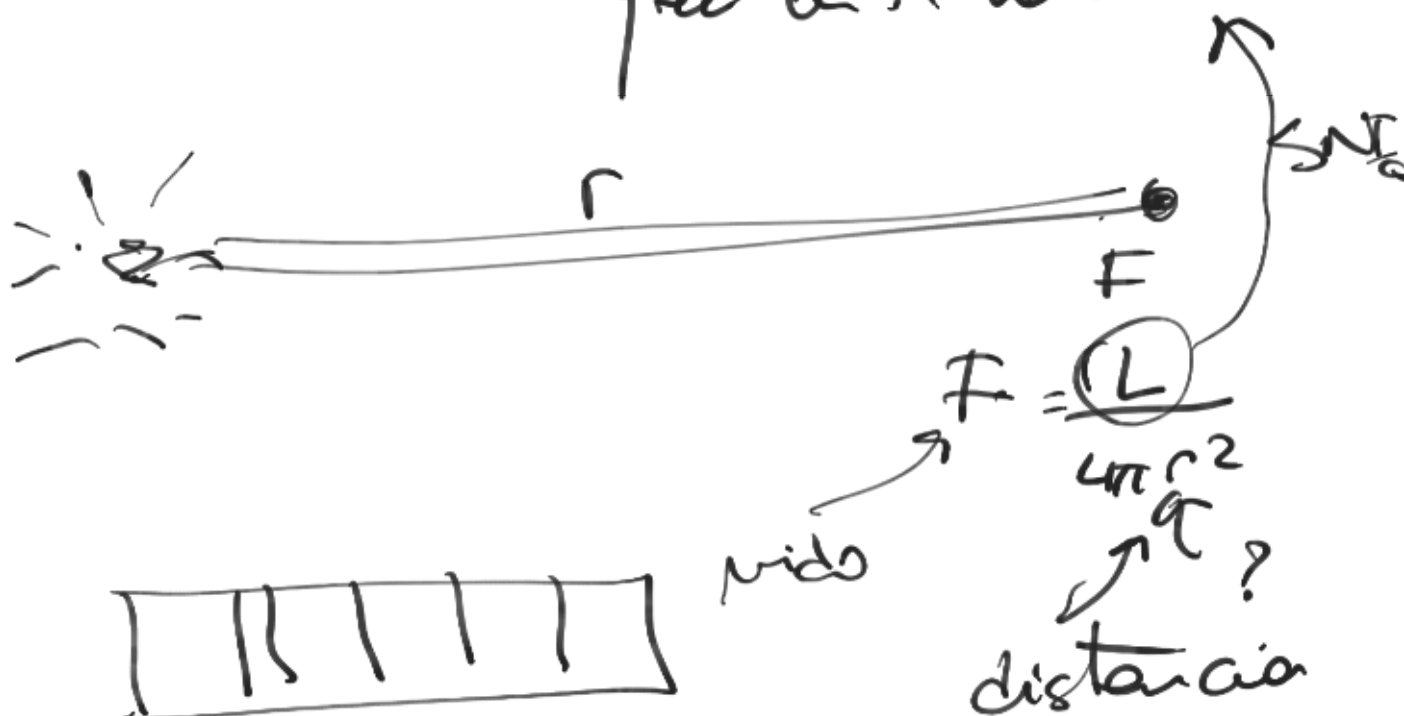
F.R.W \rightarrow Universo en
expansión.

$$8\pi G^{\mu\nu} + \Lambda_g = R^{\mu\nu}$$

constante
cosmológica

Hubble expansion

SN Ia \rightarrow Sequence muy buena
de cal. de la edad
pero con la SN Ia



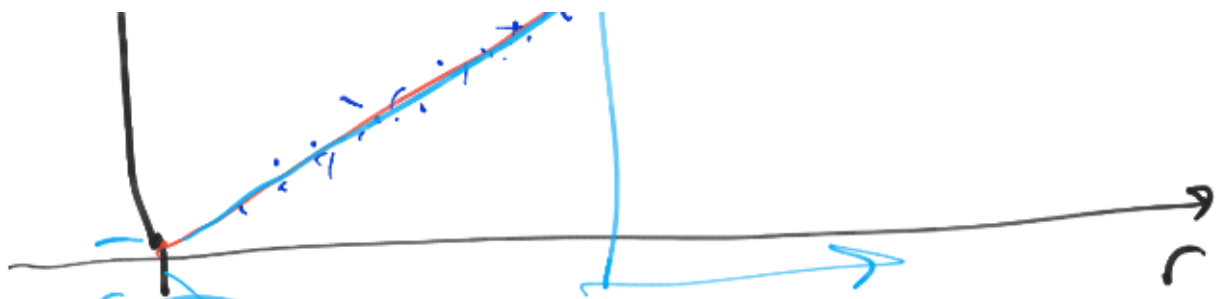
$z \sim H_0$

z espectro

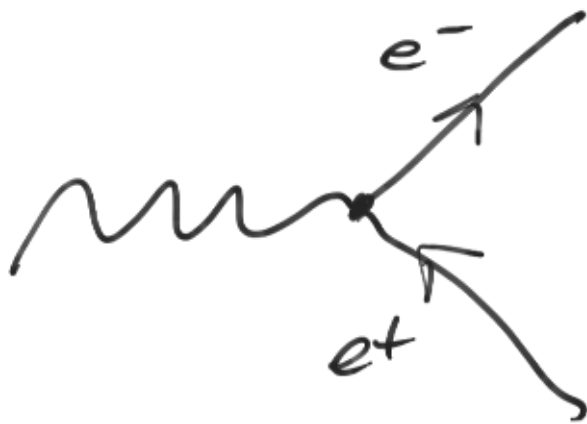
aceleración

H_0

A



Dark Energy \rightarrow Constant



$$E_\gamma \geq 2m_e c^2$$

$$E_\gamma \geq 1.022 \text{ MeV}$$

$$\langle E_\gamma \rangle < 1.022 \text{ MeV}$$

$$\phi(E) \propto e^{-E/kT}$$

Factor of Boltzmann



t \rightarrow E

\rightarrow



Plasma.

Nuclei samples ($H, D, {}^3He, {}^3H$
 $+ Be, Li, \dots$)

positrons, electrons, fission.

\rightarrow Combustion. etc. no labelation

completo, que aparece
Oscilación de fases, rad. de frenado

$\langle E \rangle \sim eV \rightarrow$ Atómos Simples.
 \downarrow
Neutrinos.

El Universo se hace transparente

$\langle E \rangle \sim eV \sim T \approx 4000K$

$378000 \leftrightarrow z = 1100$

$T \approx 3,6 K$

$T \approx 2,7 K$