# Space-Time DPG: Designing a Method for Parallel CFD

#### Truman Ellis

Leszek Demkowicz, Nathan Roberts, Jesse Chan, Robert Moser





#### **Motivation**

# TEXAS AT AUSTIN

#### DPG Summary

#### Overview of Features

- Robust for singularly perturbed problems
- Stable in the preasymptotic regime
- Designed for adaptive mesh refinement

#### DPG is a minimum residual method:

$$u_h = \underset{w_h \in U_h}{\arg \min} \frac{1}{2} \|Bw_h - l\|_{V'}^2$$

$$\downarrow \qquad \qquad \downarrow \qquad \qquad \downarrow$$

$$b(u_h, R_V^{-1}B\delta u_h) = l(R_V^{-1}B\delta u_h) \quad \forall \delta u_h \in U_h$$

where  $v_{\delta u_h} := R_V^{-1} B \delta u_h$  are the **optimal test functions**.

#### **Heat Equation**



Simplest Nontrivial Space-Time Problem

Equation is elliptic in space, but hyperbolic in time.

$$\frac{\partial u}{\partial t} - \epsilon \Delta u = f$$

This is really just a composite of Fourier's law and conservation of energy.

$$\boldsymbol{\sigma} - \boldsymbol{\epsilon} \nabla \boldsymbol{u} = 0$$

$$\frac{\partial u}{\partial t} - \nabla \cdot \boldsymbol{\sigma} = f$$

We can rewrite this in terms of a space-time divergence.

$$\frac{1}{\epsilon}\boldsymbol{\sigma} - \nabla u = 0$$

$$\nabla_{xt} \cdot \begin{pmatrix} -\boldsymbol{\sigma} \\ u \end{pmatrix} = f$$

### **Heat Equation**



**DPG** Formulation

Multiply by test function and integrate by parts over space-time element K.

$$\begin{pmatrix} \frac{1}{\epsilon} \boldsymbol{\sigma}, \boldsymbol{\tau} \end{pmatrix} + (u, \nabla \cdot \boldsymbol{\tau}) - \langle \hat{u}, \boldsymbol{\tau} \cdot \boldsymbol{n}_{x} \rangle = 0$$

$$- \left( \begin{pmatrix} -\boldsymbol{\sigma} \\ u \end{pmatrix}, \nabla_{xt} v \right) + \langle \hat{t}, v \rangle = f$$

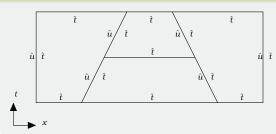
#### where

$$\hat{u} := \operatorname{tr}(u)$$

$$\hat{t} := \operatorname{tr}(-\boldsymbol{\sigma}) \cdot \boldsymbol{n}_{x} + \operatorname{tr}(u) \cdot n_{t}$$

- Trace  $\hat{u}$  defined on spatial boundaries
- Flux  $\hat{t}$  defined on all boundaries

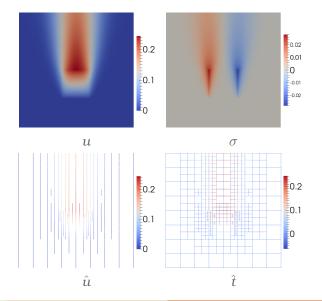
### Support of Trace Variables



# Heat equation

#### Pulsed Source Problem





# TEXAS AT AUSTIN

Strong Form

The compressible Navier-Stokes equations are

$$\frac{\partial}{\partial t} \left[ \begin{array}{c} \rho \\ \rho \mathbf{u} \\ \rho e_0 \end{array} \right] + \nabla \cdot \left[ \begin{array}{c} \rho \mathbf{u} \\ \rho \mathbf{u} \otimes \mathbf{u} + p \mathbf{I} - \mathbb{D} \\ \rho \mathbf{u} e_0 + \mathbf{u} p + \mathbf{q} - \mathbf{u} \cdot \mathbb{D} \end{array} \right] = \left[ \begin{array}{c} f_c \\ \mathbf{f}_m \\ f_e \end{array} \right] \,,$$

where

$$\mathbb{D} = 2\mu \mathbf{S}^* = 2\mu \left[ \frac{1}{2} \left( \nabla \mathbf{u} + (\nabla \mathbf{u})^T \right) - \frac{1}{3} \nabla \cdot \mathbf{u} \mathbf{I} \right],$$
$$\mathbf{q} = -C_p \frac{\mu}{Pr} \nabla T,$$

and (assuming an ideal gas EOS)

$$p = \rho RT.$$



First Order Space-Time Form

Writing this in space-time in terms of  $\rho$ ,  $\boldsymbol{u}$ , T,  $\mathbb{D}$ , and  $\boldsymbol{q}$ :

$$\mathbb{D} - \mu \left( \nabla \mathbf{u} + (\nabla \mathbf{u})^T \right) + \frac{2\mu}{3} \nabla \cdot \mathbf{u} \mathbf{I} = 0$$

$$\mathbf{q} + C_p \frac{\mu}{Pr} \nabla T = 0$$

$$\nabla_{xt} \cdot \begin{pmatrix} \rho \mathbf{u} \\ \rho \end{pmatrix} = f_c$$

$$\nabla_{xt} \cdot \begin{pmatrix} \rho \mathbf{u} \otimes \mathbf{u} + \rho R T \mathbf{I} - \mathbb{D} \\ \rho \mathbf{u} \end{pmatrix} = \mathbf{f}_m$$

$$\nabla_{xt} \cdot \begin{pmatrix} \rho \mathbf{u} \left( C_v T + \frac{1}{2} \mathbf{u} \cdot \mathbf{u} \right) + \mathbf{u} \rho R T + \mathbf{q} - \mathbf{u} \cdot \mathbb{D} \\ \rho \left( C_v T + \frac{1}{2} \mathbf{u} \cdot \mathbf{u} \right) \end{pmatrix} = f_e.$$



**DPG** Formulation

Multiplying by test functions and integrating by parts:

$$\begin{split} (\mathbb{D},\mathbb{S}) + (2\mu\mathbf{u},\nabla\cdot\mathbb{S}) - \left(\frac{2\mu}{3}\mathbf{u},\nabla\operatorname{tr}\mathbb{S}\right) - \langle 2\mu\hat{\mathbf{u}},\mathbb{S}\mathbf{n}_{x}\rangle + \left\langle\frac{2\mu}{3}\hat{\mathbf{u}},\mathbb{S}\mathbf{n}_{x}\right\rangle &= 0 \\ (\mathbf{q},\tau) - \left(C_{p}\frac{\mu}{Pr}T,\nabla\cdot\tau\right) + \left\langle C_{p}\frac{\mu}{Pr}\hat{T},\tau_{n}\right\rangle &= 0 \\ - \left(\left(\begin{array}{c}\rho\mathbf{u}\\\rho\end{array}\right),\nabla_{xt}v_{c}\right) + \langle\hat{t}_{c},v_{c}\rangle &= (f_{c},v_{c}) \\ - \left(\left(\begin{array}{c}\rho\mathbf{u}\otimes\mathbf{u} + \rho RT\mathbf{I} - \mathbb{D}\\\rho\mathbf{u}\end{array}\right),\nabla_{xt}\mathbf{v}_{m}\right) + \langle\hat{\mathbf{t}}_{m},\mathbf{v}_{m}\rangle &= (\mathbf{f}_{m},\mathbf{v}_{m}) \\ - \left(\left(\begin{array}{c}\rho\mathbf{u}\left(C_{v}T + \frac{1}{2}\mathbf{u}\cdot\mathbf{u}\right) + \mathbf{u}\rho RT + \mathbf{q} - \mathbf{u}\cdot\mathbb{D}\\\rho\left(C_{v}T + \frac{1}{2}\mathbf{u}\cdot\mathbf{u}\right)\end{array}\right),\nabla_{xt}v_{e}\right) + \langle\hat{t}_{e},v_{e}\rangle &= (f_{e},v_{e})\;, \end{split}$$

where  $\hat{u}$  and  $\hat{T}$  are spatial traces and  $\hat{t}_c$ ,  $\hat{t}_m$ , and  $\hat{t}_e$  are fluxes.



Flux and Trace Variables

Spatial traces and fluxes are defined as follows:

$$\hat{\mathbf{u}} = \operatorname{tr}(\mathbf{u}) 
\hat{T} = \operatorname{tr}(T) 
\hat{t}_c = \operatorname{tr}(\rho \mathbf{u}) \cdot \mathbf{n}_x + \operatorname{tr}(\rho) n_t 
\hat{t}_m = \operatorname{tr}(\rho \mathbf{u} \otimes \mathbf{u} + \rho R T \mathbf{I} - \mathbb{D}) \cdot \mathbf{n}_x + \operatorname{tr}(\rho \mathbf{u}) n_t 
\hat{t}_e = \operatorname{tr}\left(\rho \mathbf{u}\left(C_v T + \frac{1}{2}\mathbf{u} \cdot \mathbf{u}\right) + \mathbf{u}\rho R T + \mathbf{q} - \mathbf{u} \cdot \mathbb{D}\right) \cdot \mathbf{n}_x 
+ \operatorname{tr}\left(\rho\left(C_v T + \frac{1}{2}\mathbf{u} \cdot \mathbf{u}\right)\right) n_t.$$

#### Linearization

Fluxes, traces, and  ${\bf q}$  are linear in the above bilinear form, but we need to linearize in  $\rho$ ,  ${\bf u}$ , T, and  ${\mathbb D}$  (Jacobian and residual not shown here).

Sod Shock Tube

