ECOLE POLYTECHNIQUE FÉDÉRALE DE LAUSANNE

MASTER THESIS

Evaluation of optical aberrations using Phase Diversity

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in the

Astrophysics laboratory Basic Sciences

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Declaration of Authorship

I, Jordan VOIRIN, declare that this thesis titled, "Evaluation of optical aberrations using Phase Diversity" and the work presented in it are my own. I confirm that:

- This work was done wholly or mainly while in candidature for a research degree at this University.
- Where any part of this thesis has previously been submitted for a degree or any other qualification at this University or any other institution, this has been clearly stated.
- Where I have consulted the published work of others, this is always clearly attributed.
- Where I have quoted from the work of others, the source is always given. With the exception of such quotations, this thesis is entirely my own work.
- I have acknowledged all main sources of help.
- Where the thesis is based on work done by myself jointly with others, I have made clear exactly what was done by others and what I have contributed myself.

Signed:			
Date:			

"Thanks to my solid academic training, today I can write hundreds of words on virtually any topic without possessing a shred of information, which is how I got a good job in journalism."

Dave Barry

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Abstract

Physics Basic Sciences

Master in Applied Physics

Evaluation of optical aberrations using Phase Diversity

by Jordan VOIRIN

The Thesis Abstract is written here (and usually kept to just this page). The page is kept centered vertically so can expand into the blank space above the title too...

Acknowledgements

The acknowledgments and the people to thank go here, don't forget to include your project advisor. . .

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List of Abbreviations

WFS WaveFront Sensor

Physical Constants

Speed of Light $c_0 = 2.99792458 \times 10^8 \,\mathrm{m \, s^{-1}}$ (exact)

xxi

List of Symbols

a distance

P power $W(J s^{-1})$

 ω angular frequency rad

xxiii

For/Dedicated to/To my...

Chapter 1

Introduction

???

Chapter 2

Phase Diversity Experiment

2.1 Theoretical Background

2.2 Experimental Setup

The design of the experiment was already done by Bouxin (2017). The system is built according to her plans and specificationsS.

The experiment is mounted on a pressurized legs optical table. The assembly contains six main components: a light source, an entrance pupil, an imaging system, a converging lens to focus the beam on the camera, a camera and a wavefront sensor.

2.2.1 Light source

The final application of the phase diversity will be to characterize the optical aberrations induced by the imperfect optical path to a scientific detector of a telescope. For this reason, the light source has to simulate a distant

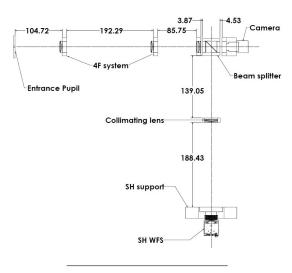


FIGURE 2.1: Experimental setup schema with the relevant distances, (Bouxin, 2017).

TABLE 2.1: Optical Components

#	Components	Model	Reference
1	Pigtailed laser diode	Thorlabs, LPS-635-FC	A.1
2	Converging lens, f = 11 mm	Thorlabs, A220TM-A	A.2
3	Pinhole, 10 μm	Thorlabs, P10S	A.3
4	Converging lens, f = 200 mm	Thorlabs, AL100200	A.4
5	3.2 mm Hole milled in metal sheet		
6	Converging lens, f = 100 mm	Thorlabs, AC254-100-A	A.5
7	Converging lens, f = 80 mm		
8	Camera CMOS	Ximea, MQ013MG-E2	A.6
9	Converging lens, f = 100 mm		
10	Shack-Hartman WFS	Thorlabs, WFS150-5C	A.7

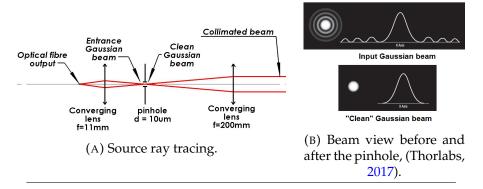


FIGURE 2.3: Source schema and pinhole effect on the beam.

star aberration-free wavefront. A distant star wavefront is considered planar since the object distance, z, is far greater than the telescope size, r, see Fig. 2.2. The source of our experiment must then be characterized by planar wavefront.

In order to obtain such a planar wavefront at the entrance pupil, the light source consist of a "pigtailed laser diode", a f=11mm converging lens, a pinhole and a f=200 mm converging lens, see Table 2.1. The pigtailed laser diode emits a Gaussian beam centred at 637.5 nm slightly diverging. The converging lens concentrates the beam at the center of the $10\mu m$ pinhole to filter the noise. The second converging lens collimates the beam, obtaining a collimated beam with a planar wavefront, see Fig. 2.3a and 2.3b.

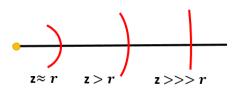


FIGURE 2.2: Wavefront curvature for different source's distances, *z. r* represents the characteristic size of the arc of interest.

2.2.2 Entrance pupil

The entrance pupil of our optical system is a circular aperture of 3.2 mm diameter placed after the collimating lens of the light source. It is milled in a metal plate and centred in his support, to avoid positioning with a XY table. The diameter is chosen in available material to fit in the different detector's surfaces.

2.2.3 Pupil imaging system

The phase diversity technique requires PSFs images as input, which means that the beam as to be focused onto the detector surface. To analyse the aberration in the pupil plane, one needs to focus an image of the beam passing through the entrance pupil. The simplest assembly to achieve this goal is the 4F system, which consist of two converging lenses of focal 100 mm. The two lenses are separated by 200 mm, see Fig. 2.1. This places the image of the entrance pupil 100 mm after the second converging lens.

2.2.4 Acquisitions

The image of the entrance pupil, obtained with the 4F system, is focused onto a CMOS Ximea camera by a f = 80 mm converging lens to acquire the PSFs for the phase diversity wavefront retrieval. The camera has surface composed by 1280×1024

2.3. Results 5

pixels of 5.3 μ m, see Appendix A.6. It is mounted on sliding support in order to be able to acquire in/out-of-focus images. A beam splitter is placed in the converging beam to separate it in two. The second beam is collimated and a Shack-Hartman WFS is placed on the entrance pupil image plane, to check the results of the phase diversity wavefront retrieval. The Shack-Hartman WFS has a 39 X 31 lenslet grid and a CCD with a resolution of 1280x1024 pixels of 4.65 μ m, see Appendix A.7.

2.3 Results

This section presents the results of the phase diversity experiment, with the introduction of different sources of aberration.

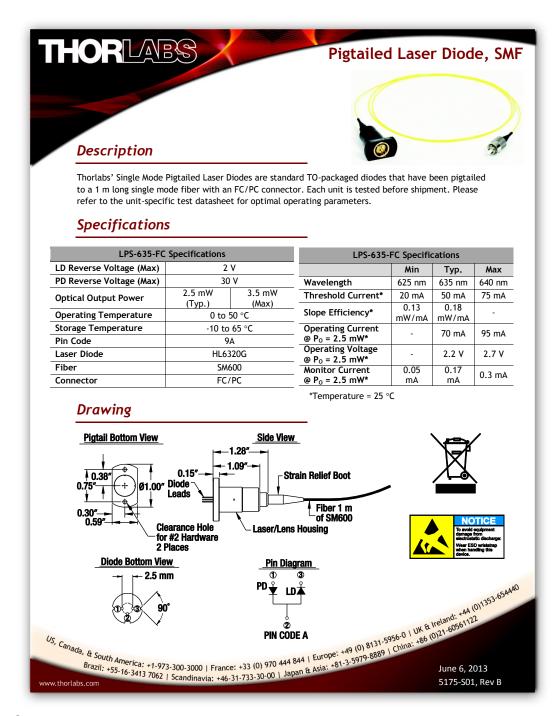
2.3.1 Parallel plane plate

The first source of aberration studied in this work is a tilted parallel plane plate which is used as a calibrated source of astigmatism.

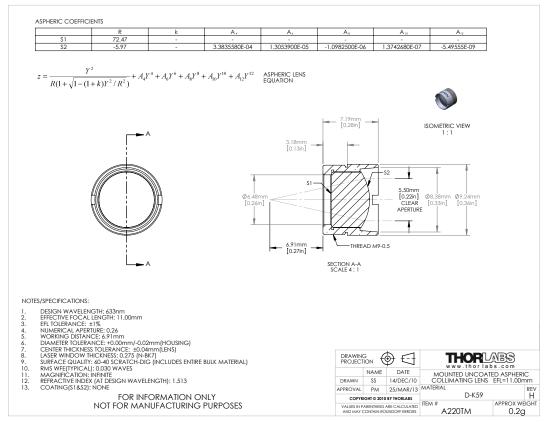
Appendix A

Optical Component Datasheets

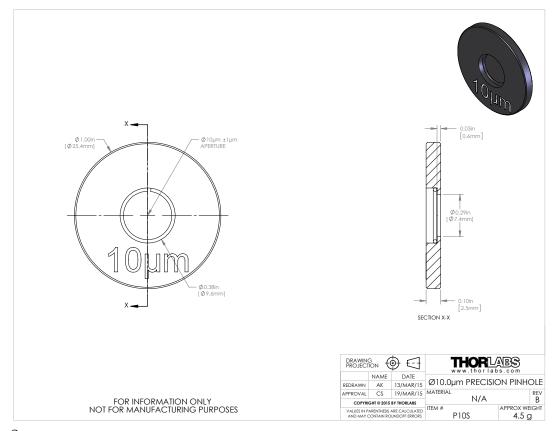
A.1 Pigtailed laser diode



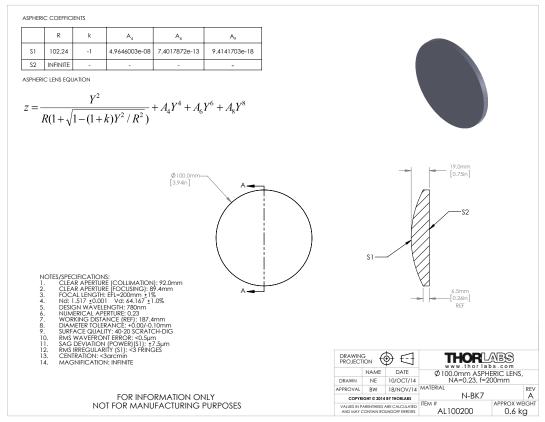
A.2 Converging lens A220TM-A, f = 11 mm



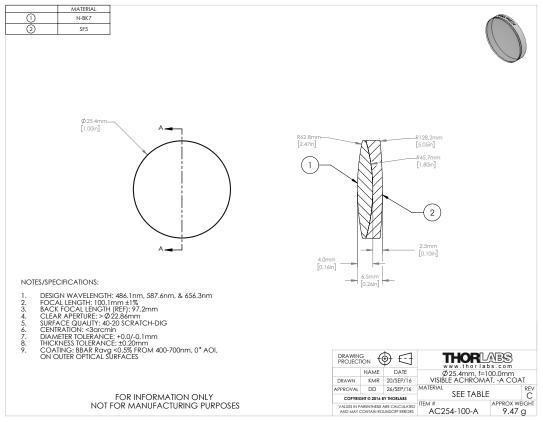
A.3 Pinhole 10 μ m



A.4 Converging lens AL100200, f = 200 mm



A.5 Converging lens AC254-100-A, f = 100 mm



A.6 Ximea Camera, MQ013MG-E2



Bits per pixel: 8, 10

Dynamic range: 60 dB

Frame rates: 60 fps

On-chip binning: 1x1, 2x2

Image data interface: USB 3.0

Data I/O: GPIO IN, OUT

Power requirements: 0.9 Watt

Lens mount: C or CS Mount

Weight: 26 grams

Dimensions WxHxD: 26 x 26 x 26 mm

Operating
environment:

Source: www.ximea.com/en/products/usb3-vision-cameras-xiq-line/mq013mg-e2

Customs tariff code: 8525.80 30 (EU) / 8525.80 40 (USA)

EAR99

ECCN:

Shack-Hartmann wavefront sensor, WFS150-5C

8 Appendix

8.1 Technical Data

8.1.1 WFS150/300

Item #	WF\$150-5C	WF\$150-7AR	WF\$300-14AR	
Microlenses				
Microlens Array	MLA150M-5C	MLA150M-7AR	MLA300M-14AR	
Substrate Material		Fused Silica (Quartz)	•	
Number of Active Lenslets		Software Selectable		
Max. Number of Lenslets	39 >	: 31	19 x 15	
Camera				
Sensor Type		CCD		
Resolution	max. 1280	x 1024 pixels, Software	Selectable	
Aperture Size		5.95 mm x 4.76 mm		
Pixel Size		4.65 µm x 4.65 µm		
Shutter		Global		
Exposure Range		79 µs - 65 ms		
Frame Rate		max. 15 Hz		
Image Digitization	8 bit			
Wavefront Measurement				
Wavefront Accuracy 1)	λ/15 rms @ 633 nm λ/50 rms @		λ/50 rms @ 633 nm	
Wavefront Sensitivity 2)	λ/50 rms @ 633 nm		λ/150 rms @ 633 nm	
Wavefront Dynamic Range 3)	> 100 \(\hat{Q} \) 633 nm		> 50 A @ 633 nm	
Local Wavefront Curvature 4)	> 7.4 mm > 10.0 mm		> 40.0 mm	
External Trigger Input				
Save Static Voltage level		0 to 30 V DC		
LOW Level	0.0 V to 2.0 V			
HIGH Level		5.0 V to 24 V		
Input current		> 10 mA		
Min Pulse Width	100 µs			
Min. Slew Rate	35 V / msec			
Common Specifications				
Optical Input	C-Mount			
Power Supply	<1.5 W, via USB			
Operating Temperature Range ⁵)	+5 to +35 °C			
Storage Temperature Range	-40 to 70 °C			
Warm-Up Time for Rated Accuracy	15 min			
Dimensions (W x H x D)	32.0 mm x 40.4 mm x 45.5 mm			
Weight	0.1 kg			

¹⁾ Absolute accuracy using internal reference. Measured for spherical wavefronts of known RoC.

All technical data are valid at 23 \pm 5°C and 45 \pm 15% rel. humidity (non condensing)

@ 2007 - 2015 Thorlabs GmbH

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Source: WFS Series Operation Manual, www.thorlabs.com

²⁾ Typical relative accuracy. Achievable after, and with respect to a user calibration, 10 image averages

³⁾ Over entire aperture of wavefront sensor
4) Radius of wavefront curvature over single lenslet aperture

Bibliography

Bouxin, A. (2017). "Phasor diversity to measure the static aberrations of an optical system". MA thesis. HEIG-VD.

Thorlabs (2017). *Principles of Spatial Filters*. Thorlabs. URL: https://www.thorlabs.com/newgrouppage9.cfm?objectgroup_id=1400.