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Let's step back a moment and summarize the formalism for the  $T=0$  calculations.

- While we arrived at the Feynman rules for the energy density at  $T=0$  by way of the path integral for the partition function in imaginary or Euclidean time  $\tau$ , we could also start formally in Minkowski space (e.g. real time) in the canonical ensemble (which means in practice that we don't use a chemical potential).
- We won't rederive the results here, but simply summarize the formulas that lead to the Feynman rules we've given.
- We will find it useful to switch between the formalisms, depending on the problem under consideration.

Instead of the partition function with sources, we have the closely related generating functional,

$$Z[\eta, \eta^\dagger] = \int \mathcal{D}(\psi, \bar{\psi}) e^{i \int d^4x [\mathcal{L}(x) + \eta^\dagger(x) \psi(x) + \bar{\psi}(x) \eta(x)]}$$

where the integration  $\int d^4x = \int d^3x \int dt$  is over real time and the Lagrangian (density)  $\mathcal{L}(x)$  is (for the delta-function potential)

$$\mathcal{L} = \bar{\psi}_\alpha(x) \left( i \frac{\partial}{\partial t} + \frac{\nabla^2}{2m} \right) \psi_\alpha(x) - \frac{\lambda}{2} \bar{\psi}_\alpha(x) \bar{\psi}_\beta(x) \psi_\beta(x) \psi_\alpha(x)$$

• So this is essentially a continuation from  $\tau$  to  $t$ .

Check:  $-\int d\tau \rightarrow -i \int dt$ ,  $\frac{\partial}{\partial \tau} \rightarrow -i \frac{\partial}{\partial t}$  so we cancel the minus signs and have  $+i \int d^4x \bar{\psi} \left( i \frac{\partial}{\partial t} \right) \psi$ .

Similarly  $-\int d\tau \bar{\psi} \left( -\frac{\nabla^2}{2m} \right) \psi \rightarrow +i \int d\tau \bar{\psi} \frac{\nabla^2}{2m} \psi$

• If we want to include a chemical potential  $\mu$ , we take

$$i \frac{\partial}{\partial t} \rightarrow i \frac{\partial}{\partial t} + \mu$$

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The development of a perturbation expansion follows closely that of the partition function: remove the interaction term from the Lagrangian using

$$\boxed{\psi \rightarrow \frac{i\delta}{\delta\eta}} \quad \left[ \text{since } \frac{i\delta}{\delta\eta(x)} e^{i\int d^4y \psi(y)\eta(y)} = e^{i\int d^4y \psi(y)\eta(y)} \psi(x) \right]$$

$$\boxed{\bar{\psi} \rightarrow -i\frac{\delta}{\delta\bar{\eta}}} \quad \left[ \text{since } -i\frac{\delta}{\delta\bar{\eta}(x)} e^{i\int d^4y \bar{\psi}(y)\eta(y)} = e^{i\int d^4y \bar{\psi}(y)\eta(y)} \bar{\psi}(x) \right]$$

and complete the square in the path integral using

$$\boxed{\bar{\psi} G^{-1} \psi + \bar{\psi} \eta + \eta^\dagger \psi = (\bar{\psi} + \eta^\dagger G) G^{-1} (\psi + G \eta) - \eta^\dagger G \eta}$$

(where all of the integrations are suppressed for simplicity).

The result is

$$\boxed{Z[\eta, \eta^\dagger] = Z_0 e^{-i\lambda \int d^4x \left( \frac{i\delta}{\delta\eta(x)} \frac{i\delta}{\delta\bar{\eta}(x)} \frac{-i\delta}{\delta\eta(x)} \frac{-i\delta}{\delta\bar{\eta}(x)} \right)} e^{-i \int d^4x_1 d^4x_2 \eta^\dagger(x_1) G(x_1, x_2) \eta(x_2)}}$$

where  $x \equiv (t, \vec{x})$  and

$$iG_{\alpha\beta}^0(x_1, x_2) = \sum_{\alpha\beta} \frac{1}{2} \sum_{\vec{k}} e^{i\vec{k} \cdot (\vec{x}_1 - \vec{x}_2)} e^{-i\omega_k(t_1 - t_2)} \\ \times [\theta(t_1 - t_2 - \eta) \theta(|\vec{k}| - k_f) - \theta(t_2 - t_1 + \eta) \theta(k_f - |\vec{k}|)] \\ \rightarrow \sum_{\alpha\beta} \int \frac{d^3k}{(2\pi)^3} e^{i\vec{k} \cdot (\vec{x}_1 - \vec{x}_2)} e^{-i\omega_k(t_1 - t_2)} [\theta(t_1 - t_2 - \eta) \theta(|\vec{k}| - k_f) - \theta(t_2 - t_1 + \eta) \theta(k_f - |\vec{k}|)]$$

We can see where the Feynman rules originate:

- $-i\lambda$  for the vertex (the  $\frac{1}{2}$  gets cancelled) and  $\int d^4x$  for each vertex from the term we pulled out.
- same symmetry factor and spin sum
- $\left( -i\frac{\delta}{\delta\eta(x)} \right) \left( \frac{i\delta}{\delta\bar{\eta}(x)} \right) \left[ -i \int d^4x_3 d^4x_4 \eta^\dagger(x_3) G_{\alpha\beta}^0(x_3, x_4) \eta(x_4) \right] = iG_{\alpha\beta}^0(x_1, x_2)$  for each line.

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There are two different direct derivations (as opposed to going through the finite temperature formalism) of perturbation theory at  $T=0$  for fixed particle number.

- The first is detailed in Negele and Orland chp. 3 and the 2nd in Fetter and Walecka section 6.
- Here we'll simply motivate the Negele/Orland construction.

First introduce the (normalized)  $N$ -particle eigenfunctions & the one-body Hamiltonian  $H_0$  and the full Hamiltonian

$$\hat{H} = \hat{H}_0 + \hat{V}$$

$$\hat{H}_0 |\Phi_n^N\rangle = E_0^{(n)} |\Phi_n^N\rangle \quad \text{with} \quad \langle \Phi_n^N | \Phi_n^N \rangle = 1$$

and

$$(\hat{H}_0 + \hat{V}) |\Psi_n^N\rangle = E_n^{(n)} |\Psi_n^N\rangle \quad \text{with} \quad \langle \Psi_n^N | \Psi_n^N \rangle = 1$$

Completeness of the full eigenstates says

$$1 = \sum_{n,N} |\Psi_n^N\rangle \langle \Psi_n^N|$$

(we have to sum over the excited states labeled by  $n$  for each particle number  $N$ ).

The basic idea is that a ratio like

$$\frac{\langle \Phi_0^N | e^{-i(\hat{H}_0 + \hat{V})T_0} | \Phi_0^N \rangle}{\langle \Phi_0^N | e^{-i\hat{H}_0 T_0} | \Phi_0^N \rangle} = \sum_n \frac{\langle \Phi_0^N | e^{-i\hat{H}_0 T_0} | \Psi_n^N \rangle \langle \Psi_n^N | \Phi_0^N \rangle}{e^{-E_0^{(0)} T_0} \langle \Phi_0^N | \Phi_0^N \rangle}$$

$\nearrow$   
 no sum  
 over  $N$  if  $\hat{H}_0$   
 commutes with  $N$

$$= \sum_n |\langle \Phi_0^N | \Psi_n^N \rangle|^2 e^{-i(E_n^{(n)} - E_0^{(0)})T_0}$$

goes to  $e^{2\text{Re} \langle \Phi_0^N | \Psi_0^N \rangle - i(E_0^{(0)} - E_0^{(0)})T_0}$  if  $\text{Im} T_0 \rightarrow -\infty$ .  
 That is, if we take  $\frac{1}{i_0} \ln(\text{ratio}) \rightarrow E - E_0$  (dropping the (0) superscript),  
 which is what we want.

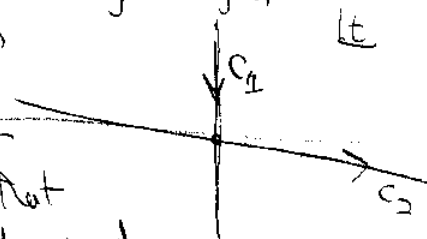
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Note the close similarity to  $\Omega - \Omega_0 = -\frac{1}{\beta} \ln \frac{Z}{Z_0}$ .

• Important assumptions:  $|\Phi_0\rangle$  is not orthogonal to  $|\Psi_0\rangle$  and that  $|\Psi_0\rangle$  is non-degenerate. If these are not fulfilled, we have to adjust our formalism.

• We can accomplish the condition  $\text{Im } T_0 \rightarrow -\infty$  by taking a contour in the complex time  $t$  plane on which we evaluate the matrix element with a path integral. In imaginary time, we had a contour that ran vertically ( $C_1$ ). Now we take one that runs for  $-T_0/2$  to  $+T_0/2$  with a slight downward slope:  $t = (1 - i\tilde{\eta})\tilde{T}$  where  $\tilde{T}$  is real and  $\tilde{\eta}$  is real and arbitrarily small.



- We just need to remember that slope occasionally when it acts to make integrands converge that would otherwise simply oscillate.
- We take  $T_0 \rightarrow \infty$  after we cancel the overall  $T_0$  factor.
- We can break up the interval into little pieces of size  $\epsilon$  and remark our path integral construction.
- The basic substitutions are  $\tau \rightarrow i\epsilon$ ,  $\int_0^{\beta} d\tau \rightarrow \int_0^{\beta/2} d\tau$ ,  $(-1)^n \rightarrow (-i)^n$ , as implemented in the Feynman rules given earlier.

• Since the replica proof goes through as before, the log of the ratio is given by the sum of totally connected ("linked") diagrams:

$$E - E_0 = \lim_{b \rightarrow \infty} \frac{1}{b} \sum (\text{all linked diagrams})$$

• The 2nd direct derivation involves introducing an adiabatic (slow) switching on of the potential:  $H_\epsilon(t) = H_0 + e^{-\epsilon|t|} \hat{V}$  and applying the so-called Gell-Mann and Low Theorem. See Fetter and Walecka for details.

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• So by now we have diagrammatic methods to calculate the ground state energy or finite temperature thermodynamic functions in perturbation theory.

• What else can we learn about?

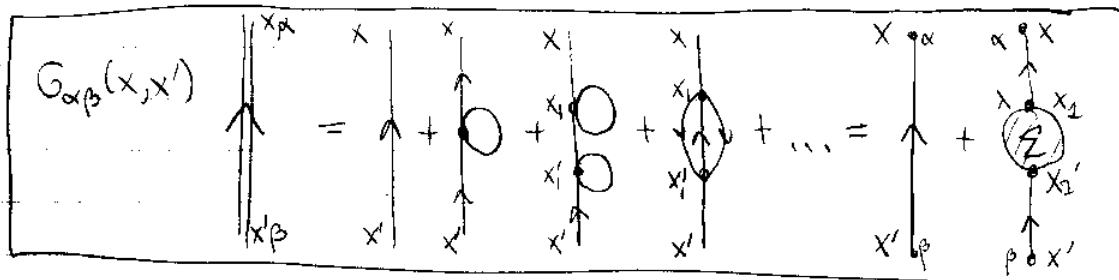
• How do we make non-perturbative approximations?

• To address both of these questions (in part), let's examine further the one-body Green's function

$$G_{\alpha\beta}(x, x') = \frac{\int \mathcal{D}(\psi) \psi_{\alpha}(x) \psi_{\beta}^{\dagger}(x') e^{i \int d^4x \mathcal{L}}}{\int \mathcal{D}(\psi) e^{i \int d^4x \mathcal{L}}} = \langle \mathbb{I}_0 | T[\psi_{\alpha}(x) \psi_{\beta}^{\dagger}(x')] | \mathbb{I}_0 \rangle$$

for which we have a diagrammatic expansion,

• Every diagram in the expansion is either  $G_{\alpha\beta}^0(x, x')$  or starts with one  $G^0$  at  $x'$  and ends with another at  $x$ . So



is the general structure, where the "self-energy"  $\Sigma(x, x')$  stands for all of the possible (corrected) diagram insertions.

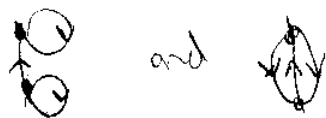
• In equation form, this is an integral equation:

$$G_{\alpha\beta}(x, x') = G_{\alpha\beta}^0(x, x') + \int d^4x_1 \int d^4x_2 G_{\alpha\lambda}^0(x, x_2) \Sigma_{\lambda\mu}(x_2, x_1) G_{\mu\beta}^0(x_1, x')$$

which serves to define the self energy.

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But we can go further. Compare two 2nd order contributions to  $\Sigma$ :






⇒ each part of the left diagram looks like a self-energy piece, because there is a single line joining them.

⇒ call it 1PR "one-particle reducible"

• The right diagram is 1PI "one-particle irreducible" because it does not fall into two pieces when a single line is cut (unlike the left diagram).

• Check: are the following 1PR or 1PI?



• The diagrams in  $\Sigma$  that are 1PI are called the "proper" self-energy and designated  $\Sigma^*$  =  +  +  + ...

• Diagrammatically,  $\Sigma$  and  $\Sigma^*$  are related by

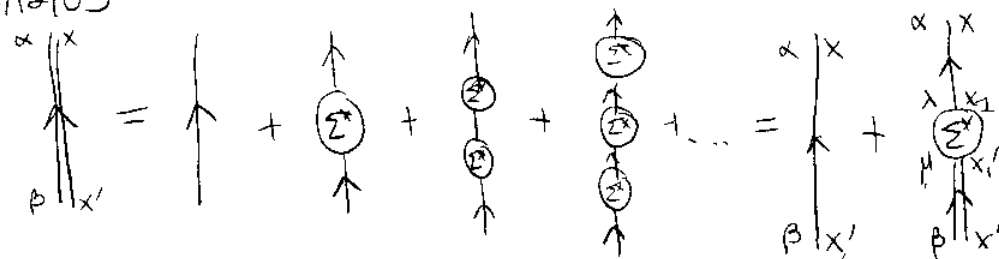
$$\Sigma = \Sigma^* + \begin{array}{c} x_1 \\ \circlearrowleft \Sigma^* \\ x_2 \\ \uparrow \\ x_3 \\ \circlearrowleft \Sigma^* \end{array} + \begin{array}{c} \circlearrowleft \Sigma^* \\ \uparrow \\ \circlearrowleft \Sigma^* \\ \uparrow \\ \circlearrowleft \Sigma^* \end{array} + \dots$$

or, in equations (suppressing spin indices)

$$\Sigma(x_1, x_1') = \Sigma^*(x_1, x_1') + \int d^4x_2 d^4x_3' \Sigma^*(x_1, x_2) G^0(x_2, x_3') \Sigma^*(x_3', x_1') + \dots$$

• Now we can insert this equation back into our original equation for  $G$  to derive another integral equation:

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$$\Rightarrow G_{\alpha\beta}(x, x') = G_{\alpha\beta}^0(x, x') + \int d^4x_1 d^4x'_1 G_{\alpha\lambda}^0(x, x_1) \Sigma^*(x_1, x'_1)_{\lambda\mu} G_{\mu\beta}(x'_1, x')$$

This is "Dyson's Equation" for the propagator ("2-point function").

- If the system of interest is translationally invariant, it almost always is best to introduce Fourier transforms in three-momentum and frequency  $\Rightarrow$  four-momentum

$$\Sigma^*(x, x')_{\alpha\beta} = \Sigma^*(x-x')_{\alpha\beta} = \int \frac{d^4k}{(2\pi)^4} e^{ik \cdot (x-x')} \Sigma^*(k)_{\alpha\beta}$$

$$G_{\alpha\beta}(x, x') = \int \frac{d^4k}{(2\pi)^4} e^{ik \cdot (x-x')} G_{\alpha\beta}(k)$$

$$\begin{aligned} G_{\alpha\beta}^0(x, x') &= \int \frac{d^4k}{(2\pi)^4} e^{ik \cdot (x-x')} G_{\alpha\beta}^0(k) \\ &= \int \frac{d^4k}{(2\pi)^4} e^{ik \cdot (x-x')} \delta_{\alpha\beta} \left[ \frac{\theta(|\vec{k}| - k_F)}{w - w_F + i\eta} + \frac{\theta(k_F - |\vec{k}|)}{w - w_F - i\eta} \right] \end{aligned}$$

- Where  $k \cdot x \equiv \vec{k} \cdot \vec{x} - \omega t$ ,  $d^4k \equiv d^3k d\omega$
- We've substituted for  $G_{\alpha\beta}^0(k) \Rightarrow$  note the pole structure.

- To transform the equation for  $G_{\alpha\beta}$ , integrate over  $(x-x')$ :

$$\int d^4(x-x') e^{ik \cdot (x-x')} \left[ G_{\alpha\beta} = G_{\alpha\beta}^0 + G_{\alpha\lambda}^0 \Sigma^* G_{\mu\beta} \right]$$

$$\Rightarrow G_{\alpha\beta}(k) = G_{\alpha\beta}^0(k) + \int d^4(x-y) d^4x_1 d^4x'_1 \int \frac{d^4k_1}{(2\pi)^4} \frac{d^4k_2}{(2\pi)^4} \frac{d^4k_3}{(2\pi)^4} e^{-ik \cdot (x-x')} e^{ik_1 \cdot (x-x_1)} e^{ik_2 \cdot (x_1-x'_1)} e^{ik_3 \cdot (x'_1-x')} \\ \times G_{\alpha\lambda}^0(k_1) \Sigma^*(k_2)_{\lambda\mu} G_{\mu\beta}(k_3)$$

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Now carry out the integrals:

$$x_2 \text{ integral} \Rightarrow (2\pi)^4 \delta^4(k_1 - k_2)$$

$$k_1 \text{ integral} \Rightarrow \text{sets } k_1 = k_2$$

$$x_1' \text{ integral} \Rightarrow (2\pi)^4 \delta^4(k_2 - k_3)$$

$$k_2 \text{ integral} \Rightarrow \text{sets } k_2 = k_3$$

$$(x-x') \text{ integral} \Rightarrow \int d^4(x-x') e^{ik(x-x')} e^{ik_3(x-x')} \rightarrow (2\pi)^4 \delta^4(k-k_3)$$

$$k_3 \text{ integral} \Rightarrow \text{sets } k_3 = k$$

When the dust settles, we get

$$G_{\alpha\beta}(k) = G_{\alpha\beta}^0(k) + G_{\alpha\lambda}^0(k) \Sigma^*(k) \lambda_{\lambda\mu} G_{\mu\beta}(k)$$

 $\Rightarrow$  an algebraic (matrix) equationIf, as in the case of the  $\lambda \delta^3(\vec{x}-\vec{x}')$  interaction,  $G, G^0$  and  $\Sigma^*$  are diagonal in the spin indices, then

$$\Rightarrow G(k) \delta_{\alpha\beta} = G^0(k) \delta_{\alpha\beta} + G^0(k) \Sigma^*(k) G(k) \delta_{\alpha\lambda} \delta_{\lambda\mu} \delta_{\mu\beta}$$

$$\Rightarrow G(k) = G^0(k) + G^0(k) \Sigma^*(k) G(k)$$

so we can just solve it!

$$[1 - G^0(k) \Sigma^*(k)] G(k) = G^0(k) \Rightarrow G(k) = \frac{G^0(k)}{1 - G^0(k) \Sigma^*(k)} = \frac{1}{[G^0(k)]^{-1} - \Sigma^*(k)}$$

$$\text{Since } [G^0(k)]^{-1} = [G^0(\vec{k}, \omega)]^{-1} = \omega - \omega_{\vec{k}} = \omega - \frac{E_{\vec{k}}^0}{\hbar} \quad (\text{now put } \hbar=1 \text{ again!})$$

$$\Rightarrow G_{\alpha\beta}(k) = G_{\alpha\beta}(\vec{k}, \omega) = \frac{1}{\omega - E_{\vec{k}}^0 - \Sigma^*(\vec{k}, \omega)} \delta_{\alpha\beta}$$

So instead of poles at  $\omega = E_{\vec{k}}^0$ , they are at the solutions to  $\omega_{\text{pole}} = E_{\vec{k}}^0 + \Sigma^*(\vec{k}, \omega_{\text{pole}})$ . What does this mean?



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What good is Dyson's equation?

- If we approximate  $\Sigma^*(\mathbf{k}, \omega)$ , eg. to some order in the perturbation expansion in  $\lambda$ , then we get an infinite order approximation to  $G$ . (That is, all powers of  $\lambda$  contribute.)
- For example, we could take  $\Sigma^* \doteq \Sigma_1^* = (1 - \frac{1}{g})\lambda g$  and the  $G$  would have poles at  $E_{\mathbf{k}}^{\text{el}} = E_{\mathbf{k}}^0 + \hbar \Sigma_1^*$ .

• If we have an approximation to  $\Sigma^*$  and  $G$ , how do we get a new approximation to the energy?

• If we think of the diagrams

$$G = \text{---} + \text{---} \bigcirc \text{---} + \dots \equiv \text{---} \bigcirc \text{---}$$

$$\Sigma^* = \text{---} \bigcirc \text{---} + \text{---} \bigcirc \bigcirc \text{---} + \text{---} \bigcirc \bigcirc \bigcirc \text{---} + \dots \equiv \text{---} \bigcirc \text{---}$$

it looks like we can generate the diagrams in the  $E - E_0$  expansion by using close the  $G$  ends on the free ends of  $\Sigma^*$

$$\Rightarrow \text{---} \bigcirc \Sigma^* \text{---} = \text{---} \bigcirc \bigcirc \text{---} + \text{---} \bigcirc \bigcirc \bigcirc \text{---} + \text{---} \bigcirc \bigcirc \bigcirc \bigcirc \text{---} + \dots$$

• This does, in fact, generate the relevant diagrams. But that's not enough  $\Rightarrow$  we need to get the factors in front correct, and  $G\Sigma^*$  doesn't do that.

• However, if we introduce a dimensionless parameter  $\alpha$  that ranges from 0 to 1, and multiply  $\hat{V}$  by  $\alpha$ :

$$\hat{A}(\alpha) \equiv \hat{A}_0 + \alpha \hat{A}_1 = \hat{A}_0 + \alpha \hat{V}$$

Then  $\alpha$  interpolates from  $\hat{A}_0$  to  $\hat{A}$ . Claim:

$$E - E_0 = -\frac{i}{2} \int_0^1 \frac{d\alpha}{\alpha} \int d\mathbf{x} d\mathbf{x}' dt' \Sigma^{*\alpha}(\mathbf{x}t; \mathbf{x}'t') G^\alpha(\mathbf{x}'t'; \mathbf{x}t)$$

$$= -\frac{i}{2} \mathcal{N} \int_0^1 \frac{d\alpha}{\alpha} \int \frac{d^4 k}{(2\pi)^4} e^{i\omega\eta} \Sigma_{00}^{*\alpha}(\mathbf{k}, \omega) G_{00}^\alpha(\mathbf{k}, \omega)$$

$\times$  includes  $\Sigma$  and  $\alpha$

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The notation  $\Sigma^\alpha$  and  $G^\alpha$  simply means the proper self-energy and Green's function for a particular value of  $\alpha$ .

- This is actually trivial in perturbation theory since we just have  $\lambda \rightarrow \alpha\lambda$ . So if we are working at  $O(\lambda^3)$ , all contributions have  $\alpha^3$  [so  $\Sigma^\alpha$  could be  $O(\alpha)$  while  $G^\alpha$  is  $O(\alpha^2)$  and all other combinations].
- In this case, the  $\alpha$  integral is trivial.

To show how the formula for  $E - E_0$  comes about, we will go back to the field operators in the Heisenberg picture, which appear in the definition of  $G$ :

$$iG_{\alpha p}(\vec{x}t, \vec{x}'t') = \langle \Psi_0^N | T[\hat{\psi}_{\alpha H}(\vec{x}t) \hat{\psi}_{\alpha H}^\dagger(\vec{x}'t')] | \Psi_0^N \rangle$$

where the subscript  $H$  is a reminder of the Heisenberg picture time dependence

$$\hat{O}_H(t) = e^{i\hat{H}t} \hat{O}_S e^{-i\hat{H}t} \Rightarrow i \frac{d}{dt} \hat{O}_H(t) = [\hat{O}_H(t), \hat{H}]$$

and  $|\Psi_0^N\rangle$  is the (normalized) fully interacting ground state.

Find the Heisenberg equation of motion for  $\hat{\psi}_{\alpha H}(\vec{x}t)$

$$\begin{aligned} i \frac{d}{dt} \hat{\psi}_{\alpha H}(\vec{x}t) &= [\hat{\psi}_{\alpha H}(\vec{x}t), \hat{H}] \\ &= [\hat{\psi}_{\alpha H}(\vec{x}t), \int d^3z \hat{\psi}_{\alpha'}^\dagger(\vec{z}) T(\vec{z}) \hat{\psi}_{\alpha'}(\vec{z}) \\ &\quad + \frac{1}{2} \int d^3z d^3z' \hat{\psi}_{\alpha'}^\dagger(\vec{z}) \hat{\psi}_{\alpha'}^\dagger(\vec{z}') V(\vec{z}, \vec{z}') \hat{\psi}_{\alpha''}(\vec{z}) \hat{\psi}_{\alpha''}(\vec{z}')] \\ &\stackrel{V \rightarrow \lambda \delta^3(\vec{z}, \vec{z}')}{=} [\hat{\psi}_{\alpha H}(\vec{x}t), \int d^3z \hat{\psi}_{\alpha'}^\dagger(\vec{z}) T(\vec{z}) \hat{\psi}_{\alpha'}(\vec{z}) + \frac{\lambda}{2} \int d^3z \hat{\psi}_{\alpha'}^\dagger(\vec{z}) \hat{\psi}_{\alpha'}^\dagger(\vec{z}) \hat{\psi}_{\alpha''}(\vec{z}) \hat{\psi}_{\alpha''}(\vec{z})] \end{aligned}$$

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We can carry out the commutators using

$$\{\hat{\psi}_\alpha(\vec{x}), \hat{\psi}_\beta^\dagger(\vec{x}')\} = \{\hat{\psi}_\alpha(\vec{x}), \hat{\psi}_\beta(\vec{x}')\} = 0$$

$$\{\hat{\psi}_\alpha(\vec{x}), \hat{\psi}_\beta^\dagger(\vec{x}')\} = \delta_{\alpha\beta} \delta^3(\vec{x} - \vec{x}')$$

and

$$[A, \hat{B}\hat{C}] = \{A, B\}\hat{C} + B\{\hat{A}, \hat{C}\}$$

When the dust settles,

$$\begin{aligned} \frac{\partial}{\partial t} \hat{\psi}_{\alpha_H}(\vec{x}, t) &= T_x \hat{\psi}_{\alpha_H}(\vec{x}, t) + \int d^3z \hat{\psi}_{\alpha_H}^\dagger(\vec{z}, t) V(\vec{x}, \vec{z}') \hat{\psi}_{\beta'}(\vec{z}', t) \hat{\psi}_{\beta''}(\vec{z}'', t) \\ &\quad \text{with } V \Rightarrow \lambda \delta^3(\vec{x} - \vec{z}) \\ &= T_x \hat{\psi}_{\alpha_H}(\vec{x}, t) + \lambda \hat{\psi}_{\alpha_H}^\dagger(\vec{x}, t) \hat{\psi}_{\alpha_H}(\vec{x}, t) \hat{\psi}_{\alpha_H}(\vec{x}, t) \end{aligned}$$

So by taking a time derivative on the field, we generate the potential.  $\Rightarrow$  exploits that to find  $\langle \psi_0 | \hat{V} | \psi_0 \rangle$  and then  $\langle \psi_0 | \hat{H} | \psi_0 \rangle = E$  in terms of  $G$ .

The idea is that we can evaluate any one-body observable if we know  $G$ . In second quantization,

$$\hat{O} = \int d^3x d^3x' \hat{\psi}_\alpha^\dagger(\vec{x}) O(\vec{x}', \vec{x})_{\alpha\alpha'} \hat{\psi}_{\alpha'}(\vec{x})$$

$$\Rightarrow \langle \psi_0 | \hat{O} | \psi_0 \rangle \equiv \langle \hat{O} \rangle = -i \int d^3x d^3x' O(\vec{x}', \vec{x}) G_{\alpha\alpha'}(\vec{x}, t; \vec{x}', t^+)$$

where the  $t^+$  ensures the correct ordering of  $\hat{\psi}^\dagger$  and  $\hat{\psi}$ .

Note that we take a trace in the spin indices.

We can also accommodate time derivatives and local operators by taking a limiting procedure on the arguments of  $G_{\alpha\beta}(\vec{x}, t; \vec{x}', t')$ . (see below!)

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Examples:

• The kinetic energy

$$\langle T \rangle = -i \int d^3x \lim_{\vec{x}' \rightarrow \vec{x}} \left[ \frac{\hbar^2}{2m} \nabla^2 \text{tr} G(\vec{x}t, \vec{x}'t^+) \right]$$

where  $\text{tr} G \rightarrow G_{\alpha\alpha}$ 

• IF  $\hat{S} = \int d^3x \hat{f}(\vec{x})$  with  $\hat{f}(\vec{x}) = \hat{\psi}_{p\alpha}^\dagger(\vec{x}) \hat{g}_{p\alpha}(\vec{x}) \hat{\psi}_\alpha(\vec{x})$

the  $\langle \hat{f}(\vec{x}) \rangle = -i \lim_{t \rightarrow t^+} \lim_{\vec{x}' \rightarrow \vec{x}} \hat{g}_{p\alpha}(\vec{x}) G_{\alpha p}(\vec{x}t, \vec{x}'t^+)$

$$= -i \lim_{\vec{x}' \rightarrow \vec{x}} \text{tr} \hat{f}(\vec{x}) G_{\alpha p}(\vec{x}t, \vec{x}'t^+)$$

$\Rightarrow$  the density  $g(\vec{x})$  [most interesting for a non-uniform system]

$$\Rightarrow \left[ \hat{g}_{p\alpha}(\vec{x}) = \hat{\psi}_{p\alpha}^\dagger(\vec{x}) \hat{\psi}_p(\vec{x}) \delta_{\alpha p} \right] \Rightarrow g(\vec{x}) = -i \text{tr} G(\vec{x}t, \vec{x}t^+)$$

- We can check this result using  $G^0$  for a uniform system:

$$g = -\text{tr} \delta_{\alpha p} \int \frac{d^3k}{(2\pi)^3} e^{i\vec{k} \cdot (\vec{x} - \vec{x})} e^{-i\vec{k} \cdot (\vec{x} - \vec{x})} [-\theta(k_F - k)]$$

$$= g \int \frac{d^3k}{(2\pi)^3} \theta(k_F - k) = \frac{g k_F^3}{6\pi^2} \quad \checkmark$$

- Note that we can calculate any n-body operator expectation value using the n-body Green's function:

$$G^{(n)}(x_1 t_1, \dots, x_n t_n; x'_1 t'_1, \dots, x'_n t'_n)$$

$$= (-i)^n \langle \Psi_0 | T [\hat{\psi}(x_1 t_1) \dots \hat{\psi}(x_n t_n) \hat{\psi}^\dagger(x'_n t'_n) \dots \hat{\psi}^\dagger(x'_1 t'_1)] | \Psi_0 \rangle$$

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So now we can use our result for  $i \frac{\partial}{\partial t} \hat{\psi}_{\alpha H}(\vec{x}, t)$  to write the expectation value of  $\hat{V}$  and  $\hat{H}$  in the ground state in terms of  $\hat{H}$ .

$$\Rightarrow \langle \hat{V} \rangle = \frac{1}{2} \int d^3x \int d^3x' \langle \psi_0 | \hat{\psi}_{\alpha}^{\dagger}(\vec{x}) \hat{\psi}_{\beta}^{\dagger}(\vec{x}') V(\vec{x}, \vec{x}') \hat{\psi}_{\beta}(\vec{x}') \hat{\psi}_{\alpha}(\vec{x}) | \psi_0 \rangle$$

$$= -\frac{1}{2} i \int d^3x \lim_{t \rightarrow t'} \lim_{\vec{x}' \rightarrow \vec{x}} \left[ i \frac{\partial}{\partial t} - T(\vec{x}) \right] G_{\alpha\alpha}(\vec{x}, t, \vec{x}', t')$$

and

$$E = \langle \hat{H} \rangle = \langle \hat{T} + \hat{V} \rangle$$

$$= -\frac{1}{2} i \int d^3x \lim_{t \rightarrow t'} \lim_{\vec{x}' \rightarrow \vec{x}} \left[ i \frac{\partial}{\partial t} + T(\vec{x}) \right] G_{\alpha\alpha}(\vec{x}, t, \vec{x}', t')$$

(recall that  $T(\vec{x}) = -\frac{\nabla^2}{2m}$  in our examples so far)

• We relate these expressions to the self energy using Dyson's equation:

$$G^{-1} = G_0^{-1} - \Sigma = i \frac{\partial}{\partial t} - T_x - \Sigma$$

so we can write  $(i \frac{\partial}{\partial t} - T_x) G = (G^{-1} + \Sigma) G = 1 + \Sigma G$

• The final ingredient is

$$\frac{\partial}{\partial \alpha} \langle \psi_0^N(\alpha) | \hat{H}(\alpha) | \psi_0^N(\alpha) \rangle = E(\alpha)$$

$$= \left( \frac{\partial}{\partial \alpha} \langle \psi_0^N(\alpha) | \right) \hat{H}(\alpha) | \psi_0^N(\alpha) \rangle + \langle \psi_0^N(\alpha) | \hat{H}(\alpha) \left( \frac{\partial}{\partial \alpha} | \psi_0^N(\alpha) \rangle \right)$$

$$+ \langle \psi_0^N(\alpha) | \frac{\partial \hat{H}(\alpha)}{\partial \alpha} | \psi_0^N(\alpha) \rangle$$

$$= E(\alpha) \frac{\partial}{\partial \alpha} \langle \psi_0^N(\alpha) | \psi_0^N(\alpha) \rangle + \langle \psi_0^N(\alpha) | \hat{V} | \psi_0^N(\alpha) \rangle$$

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So put it all together:

$$\begin{aligned}
 E - E_0 &= \int_0^1 d\alpha \frac{d}{d\alpha} E(\alpha) = \int_0^1 \frac{d\alpha}{\alpha} \langle \Psi_0^N(\alpha) | \alpha \hat{V} | \Psi_0^N(\alpha) \rangle \\
 &= -\frac{i}{2} \int_0^1 \frac{d\alpha}{\alpha} \int d^3x \left[ \left( i \frac{\partial}{\partial t} - T(\vec{x}) \right) G_{00}^\alpha(\vec{x}t, \vec{x}'t') \right]_{\vec{x}=\vec{x}'} \\
 &= -\frac{i}{2} \int_0^1 \frac{d\alpha}{\alpha} \int d^3x \left[ \int d^3x'' dt' \left( [G^\alpha]^\Gamma(\vec{x}t, \vec{x}'t') + \Sigma^\alpha(\vec{x}t, \vec{x}'t') \right) G^\alpha(\vec{x}'t', \vec{x}t') \right] \\
 &= -\frac{i}{2} \int_0^1 \frac{d\alpha}{\alpha} \int d^3x \int d^3x' dt' \Sigma_{00}^\alpha(\vec{x}t; \vec{x}'t') G_{00}^\alpha(\vec{x}t, \vec{x}'t')
 \end{aligned}$$

as desired.