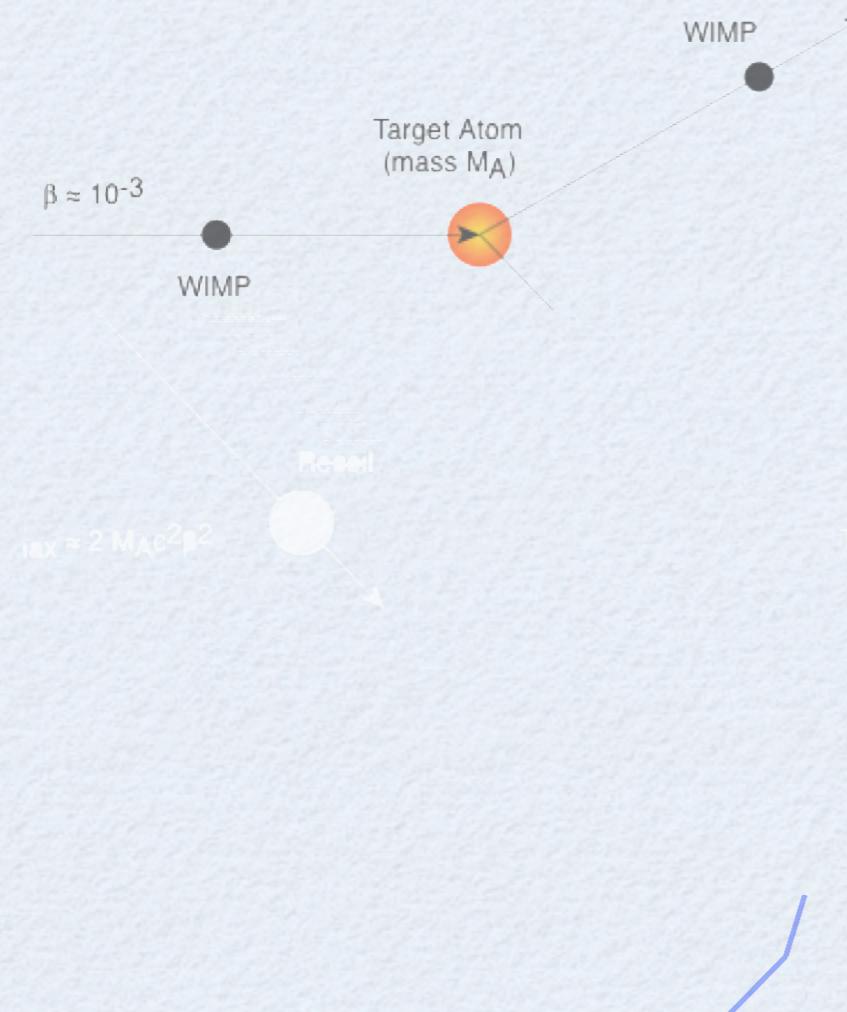


14th Geant4
Users and Collaboration Workshop
(Users Session)

Complete Monte Carlo simulation
of the optical response of the
WArP detector for Dark Matter search



F. Di Pompeo
(INFN-Laboratori Nazionali del Gran Sasso)

Laboratori Nazionali del Sud - INFN
Catania (Italy), October 15 - 17, 2009

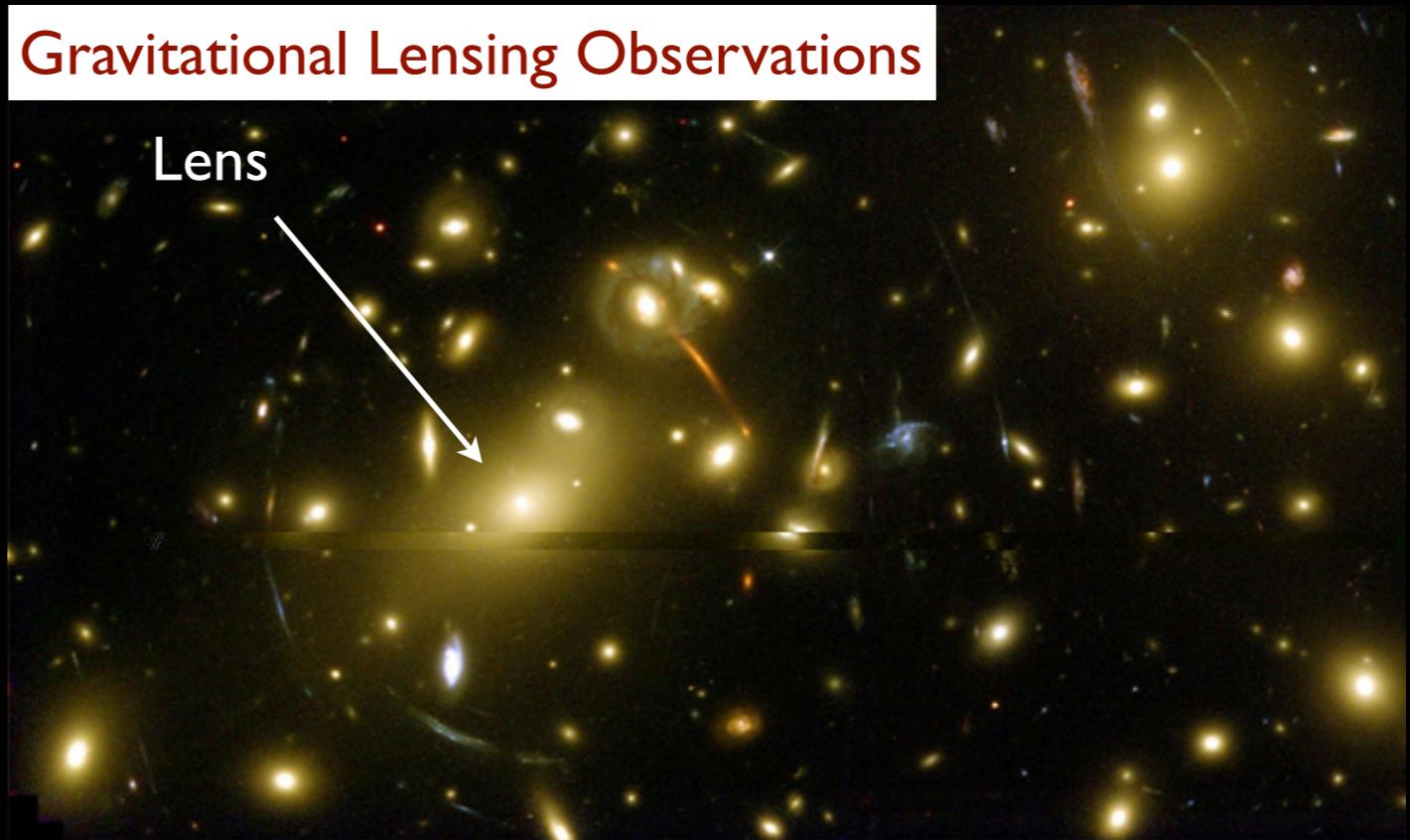
Dark Matter Indirect Evidences



Dynamic of galaxies in Clusters



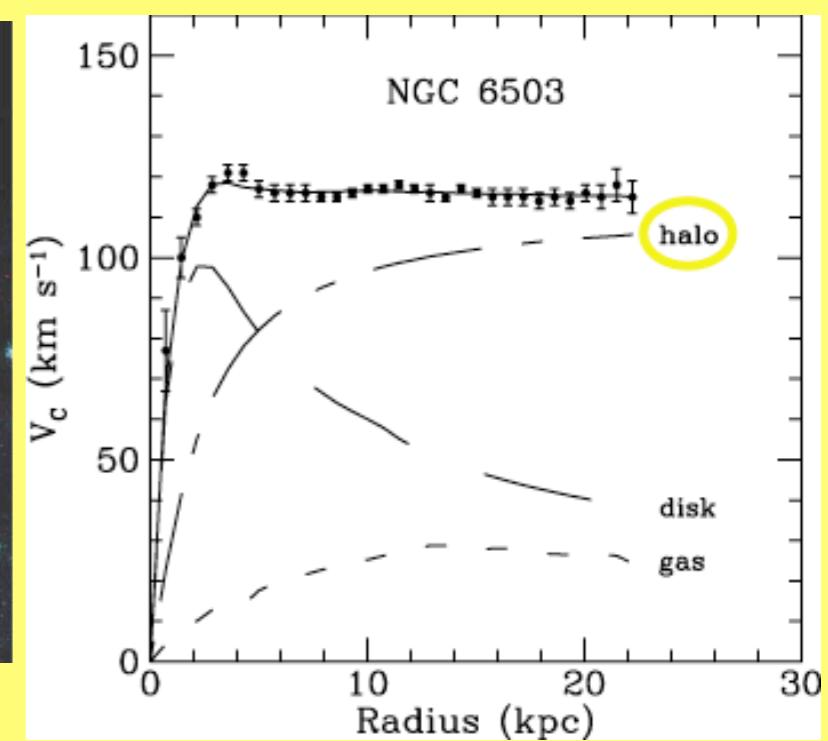
Merger Cluster



Gravitational Lensing Observations



Galaxy Rotation Curves



WIMP Direct Detection

Weakly
Interacting
Massive
Particle

1. They are expected to weakly interact with nuclei of ordinary matter.



expected rate below 0.1 event per kg of target per day

2. Ionizing particle is the recoiling nucleus



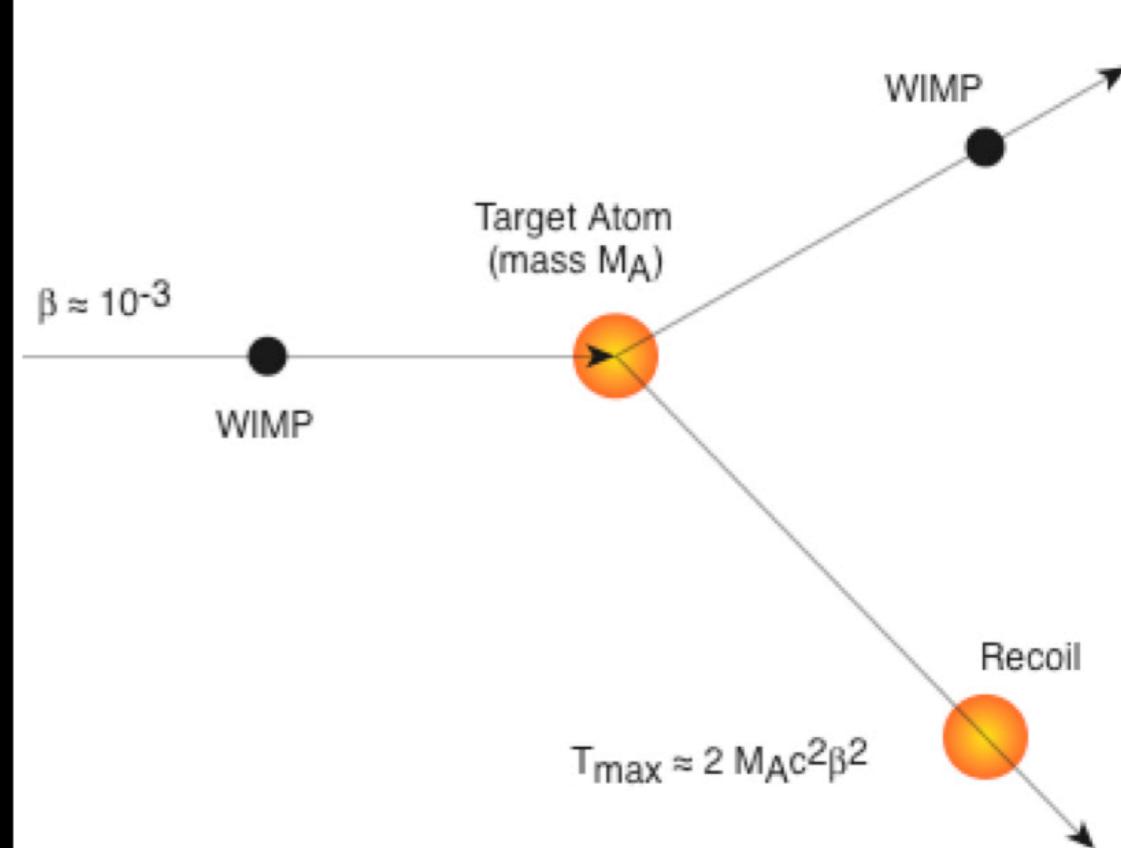
recoil energy very low (10-100 keV)

3. Energy range of interest is dominated by natural radioactivity



Need of very low radioactivity material and particle discrimination techniques to reject background events

Irreducible Background: Argon recoils induced by Neutron elastic scatterings



Dark Matter & Liquid Argon

Why noble liquids?

- High scintillation Yield
- Simultaneous measurement of scintillation and ionization (particle discrimination)
- Potentiality to be extended to multi-ton volumes

Scintillator	Nal(Tl)	Liquid Argon	Liquid Xenon
Photon Yield [ph/MeV]	4.3×10^4	4.0×10^4	4.2×10^4
Fast Decay Time [ns]	-	6	2.2
Slow Decay Time [ns]	250	1200-1500	27

Why liquid Argon?

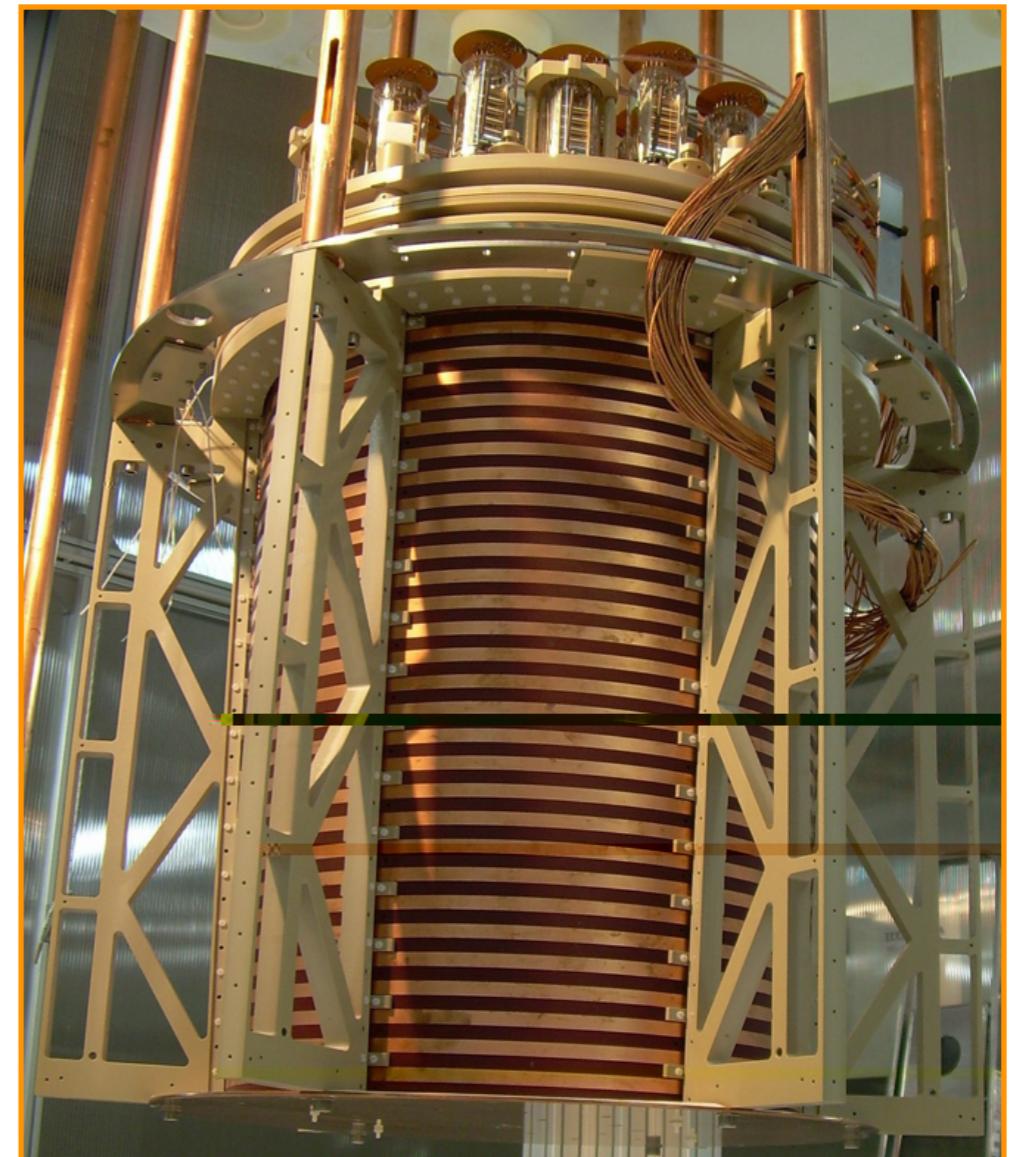
- Scintillation decay times very different ($\tau_f \approx 6$ ns , $\tau_s \approx 1200-1500$ ns) 
- Argon Technology fully operational
- Easily available (1% of atmosphere)  low cost

WArP detectors

2.3
liters



100
liters



Working since 2005 @ LNGS
R&D + WIMP run

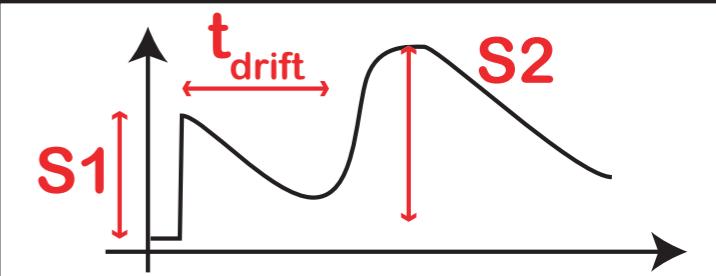
Test Run in single phase @ LNGS
(sept09)

WArP Detection

- Each interaction in liquid argon induce both atomic excitation and ionization
- By means of electric fields free ionization electrons are drifted, extracted and accelerated in the gas phase to produce electroluminescence
- prompt (S_1) & electroluminescence (S_2) signals are detected through the same set of PMTs.

Technique

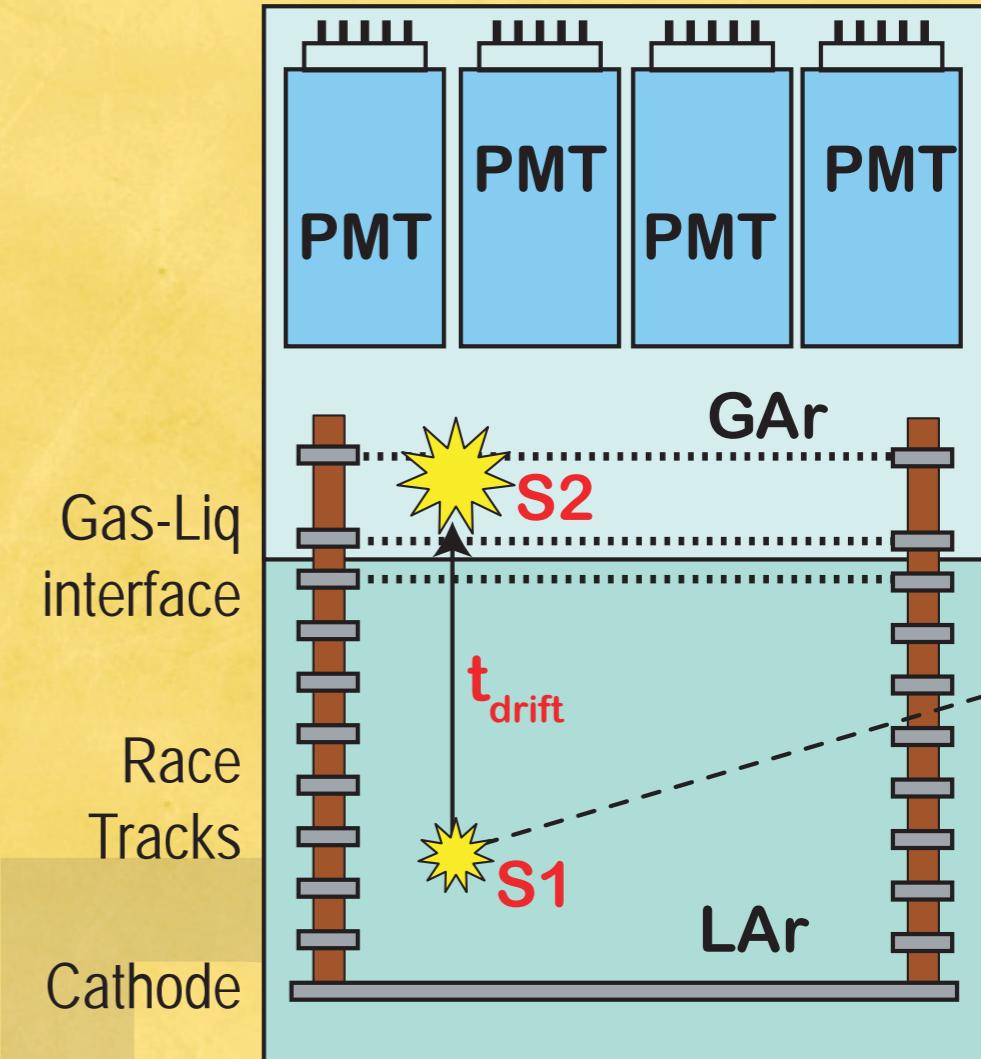
Integrated light signal



Gas-Liq
interface

Race
Tracks

Cathode

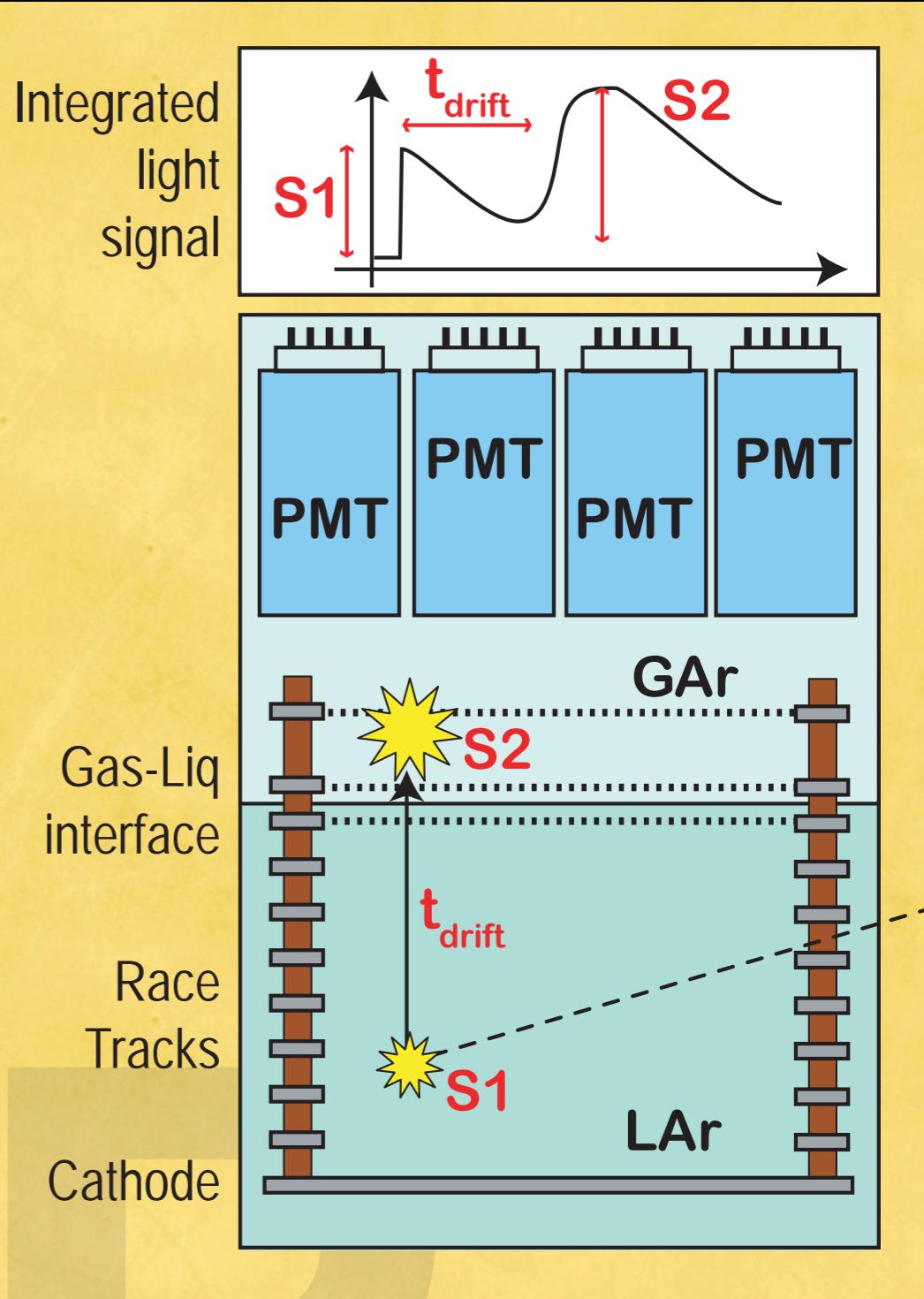


WArP Detection Technique

For a given energy deposition, the nature of the primary recoiling particle affects both:

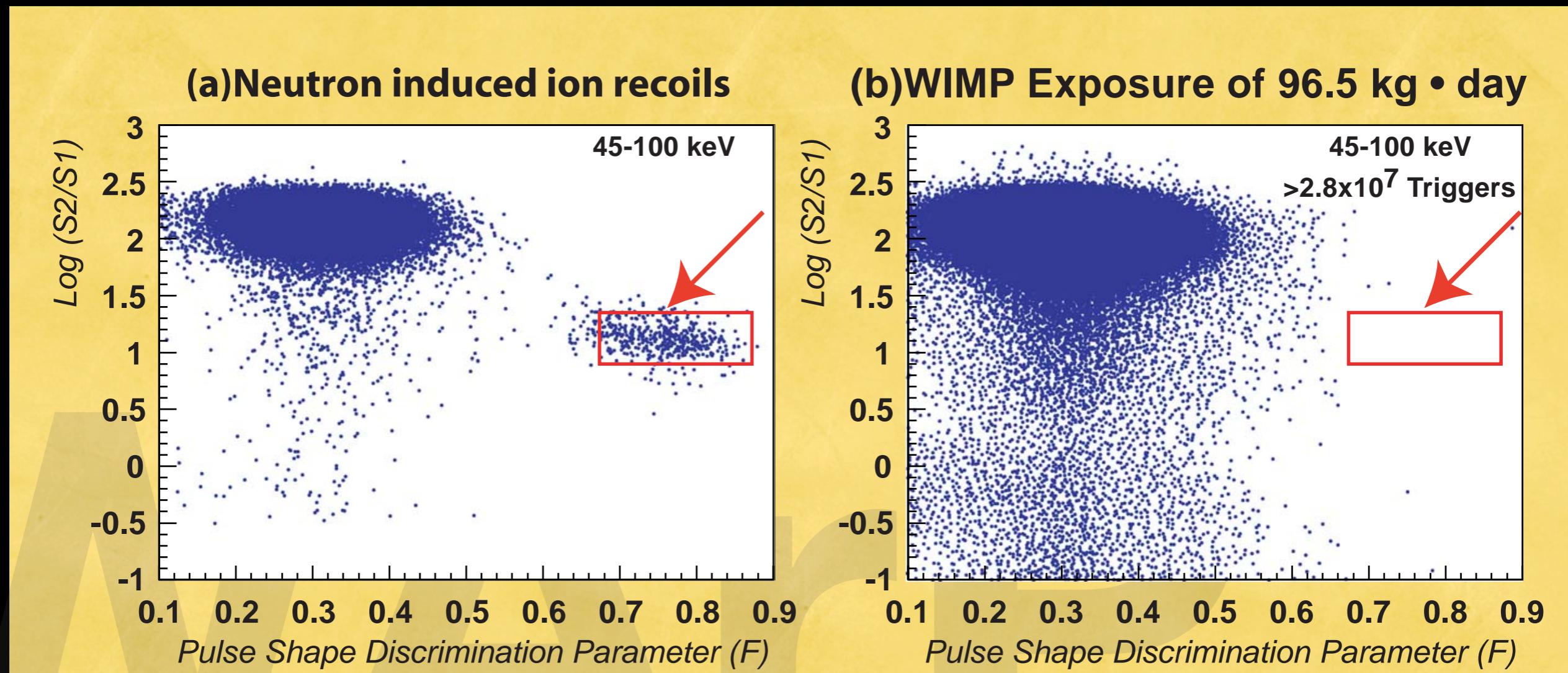
- The ratio between fast/low light intensity and hence the shape of the S1 signal
- The amount of ionization surviving recombination and hence the ratio between S1 and S2 signals

Double discrimination technique:
unique feature of Argon due to the wide difference between the two decay times.



Detection Technique: results

- ✓ Double discrimination power estimated using a 2.3 liters detector inducing argon recoils by means of neutron source (*Astropart. Phys.* 28 (2008), 495).
- ✓ Fig. (a): Argon recoils populate the red box region
- ✓ Fig. (b): After removing neutron source we acquired ≈ 30 million of background events and no event fall in the argon recoil region



Need Of Optical Simulation

performance of WArP collection system

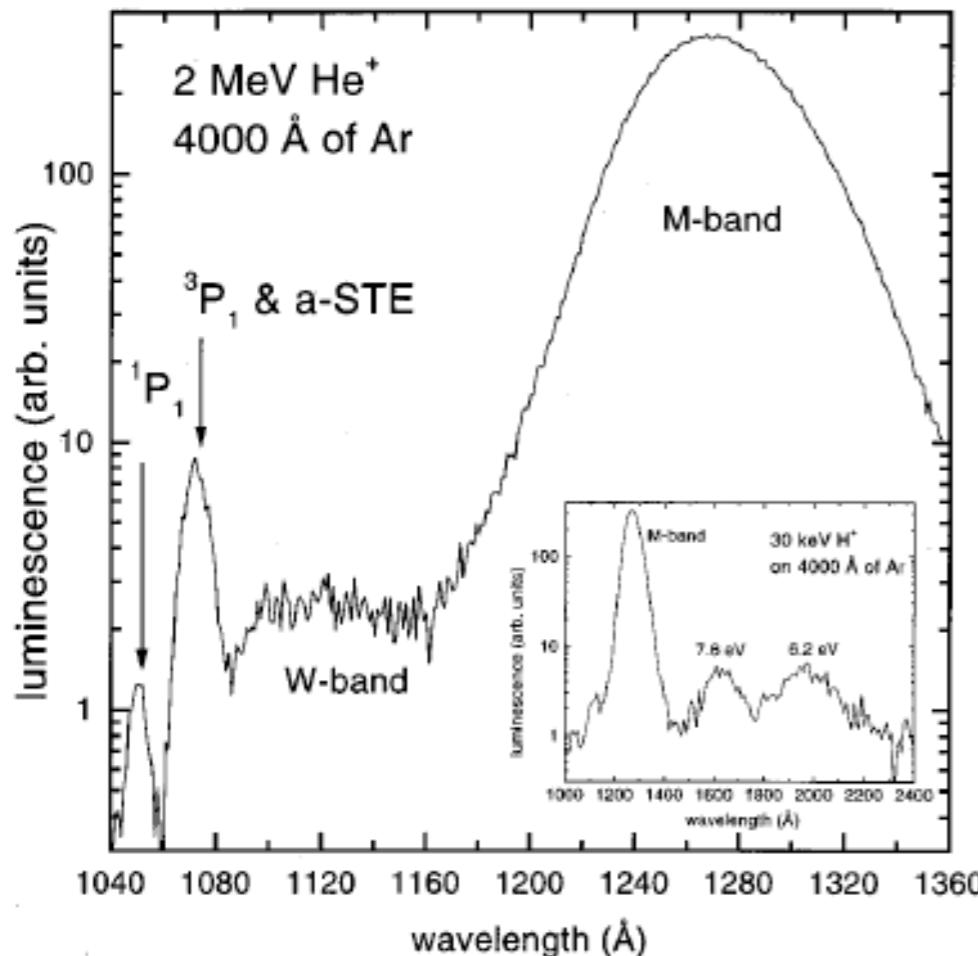
- Detection Efficiencies as a function of event Energy and Positions
- detected photons distribution on PMTs

produce Simulated Events like Raw Data

- Debugging of the WArP event reconstruction software by means of event-shape simulation (Preliminary.)
- performance of the pulse shape Discrimination (to be done)

WArP Detector Optical Response Simulated using GEANT4

LAr Scintillation Spectrum



Ar scintillation light is emitted with a narrow spectrum peaked @ 128 nm

UV optical properties to be implemented:

- refractive index of LAr and Gar
- TPB fluorescence efficiency

LAr Spectrum Parametrization

```
G4double Optic_Dete::ArScintillationSpectrum(const G4double kk)
{
    G4double waveL;
    waveL = exp(-0.5*((kk-128.0)/(2.929))*((kk-128.0)/(2.929)));
    return waveL;
}
```

Refractive Index Implementation

```
// Calculates the dielectric constant of LAr from the Bideau-Sellmeier formula.  
// See : A. Bideau-Mehu et al., "Measurement of refractive indices of Ne, Ar,  
// Kr and Xe ...", J. Quant. Spectrosc. Radiat. Transfer, Vol. 25 (1981), 395
```

```
G4double Optic_Dete::LArEpsilon(const G4double lambda)  
{  
    G4double epsilon;  
    if (lambda < (lowLambda * nanometer)) return 1.0e4; // lambda MUST be > 110.0 nm  
    epsilon = lambda / micrometer; // switch to micrometers  
    epsilon = 1.0 / (epsilon * epsilon); // 1 / (lambda)^2  
    epsilon = 1.2055e-2 * ( 0.2075 / (91.012 - epsilon) +  
        0.0415 / (87.892 - epsilon) +  
        4.3330 / (214.02 - epsilon) );  
    epsilon *= (8./12.); // Bideau-Sellmeier -> Clausius-Mossotti  
    epsilon *= (LArRho / GArRho); // density correction (Ar gas -> LAr liquid)  
    if (epsilon < 0.0 || epsilon > 0.99999) return 4.0e6;  
    epsilon = (1.0 + 2.0 * epsilon) / (1.0 - epsilon); // solve Clausius-Mossotti  
    return epsilon;  
}
```

```
// Calculates the refractive index of LAr  
  
G4double Optic_Dete::LArRefIndex(const G4double lambda)  
{  
    return ( sqrt(LArEpsilon(lambda)) ); // square root of dielectric constant  
}
```

```
G4MaterialPropertiesTable* LArMPT = new G4MaterialPropertiesTable();  
LArMPT->AddProperty("RINDEX", LAr_PPCK, LAr_RIND, NUMENTRIES);
```

Previous Work
from ICARUS simulation

J. Quant. Spectrosc. Radiat. Transfer,
Vol. 25 (1981), 395

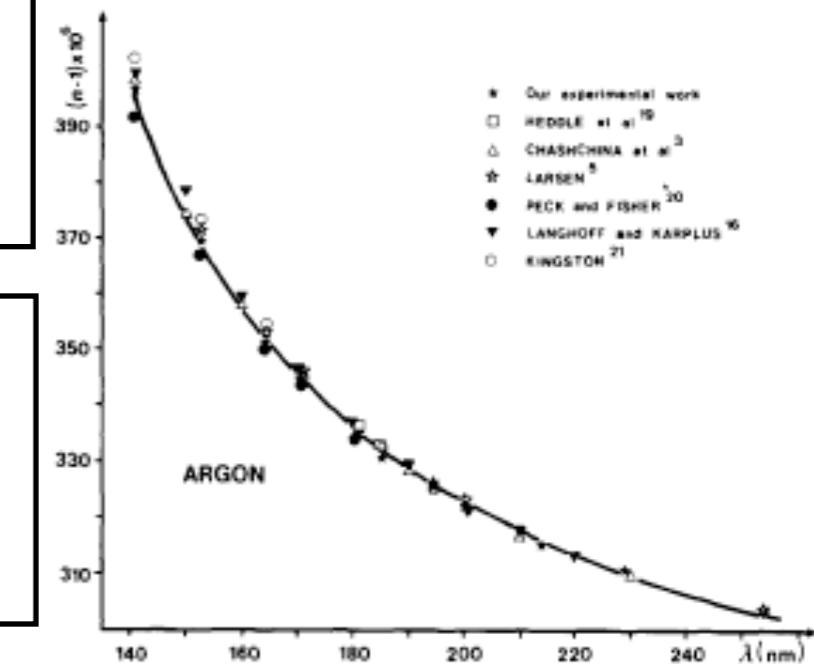
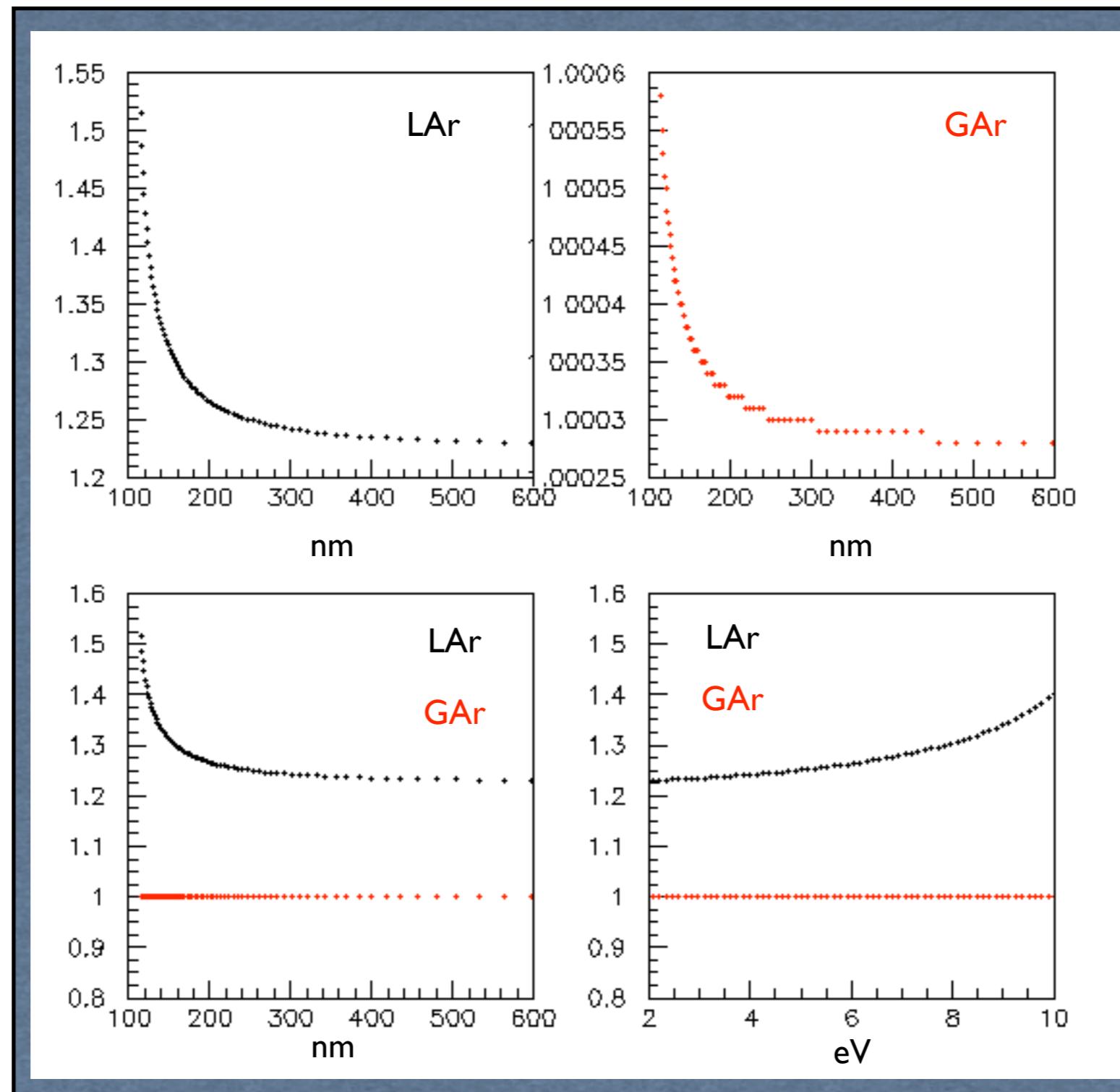


Fig. 2. Argon refractivity; the full line represents the Sellmeier formula.

Refractive Index Implementation

- GAr Refractive Index is practically constant in the range (100-600 nm)
- LAr RINDEX variation of the order of 25%



LAr Scintillation implementation

```
LArMPT->AddProperty("FASTCOMPONENT",LAr_SCPP,LAr_SCIN,num);
```

```
LArMPT->AddProperty("SLOWCOMPONENT",LAr_SCPP,LAr_SCIN,num);
```

```
LArMPT->AddProperty("FASTCOMPONENT",Scnt_spe,Scnt_pp,scinum);
```

```
LArMPT->AddProperty("SLOWCOMPONENT",Scnt_spe,Scnt_pp,scinum);
```

```
LArMPT->AddConstProperty("FASTTIMECONSTANT", 7.*ns);
```

```
LArMPT->AddConstProperty("SLOWTIMECONSTANT",1400.*ns);
```

I

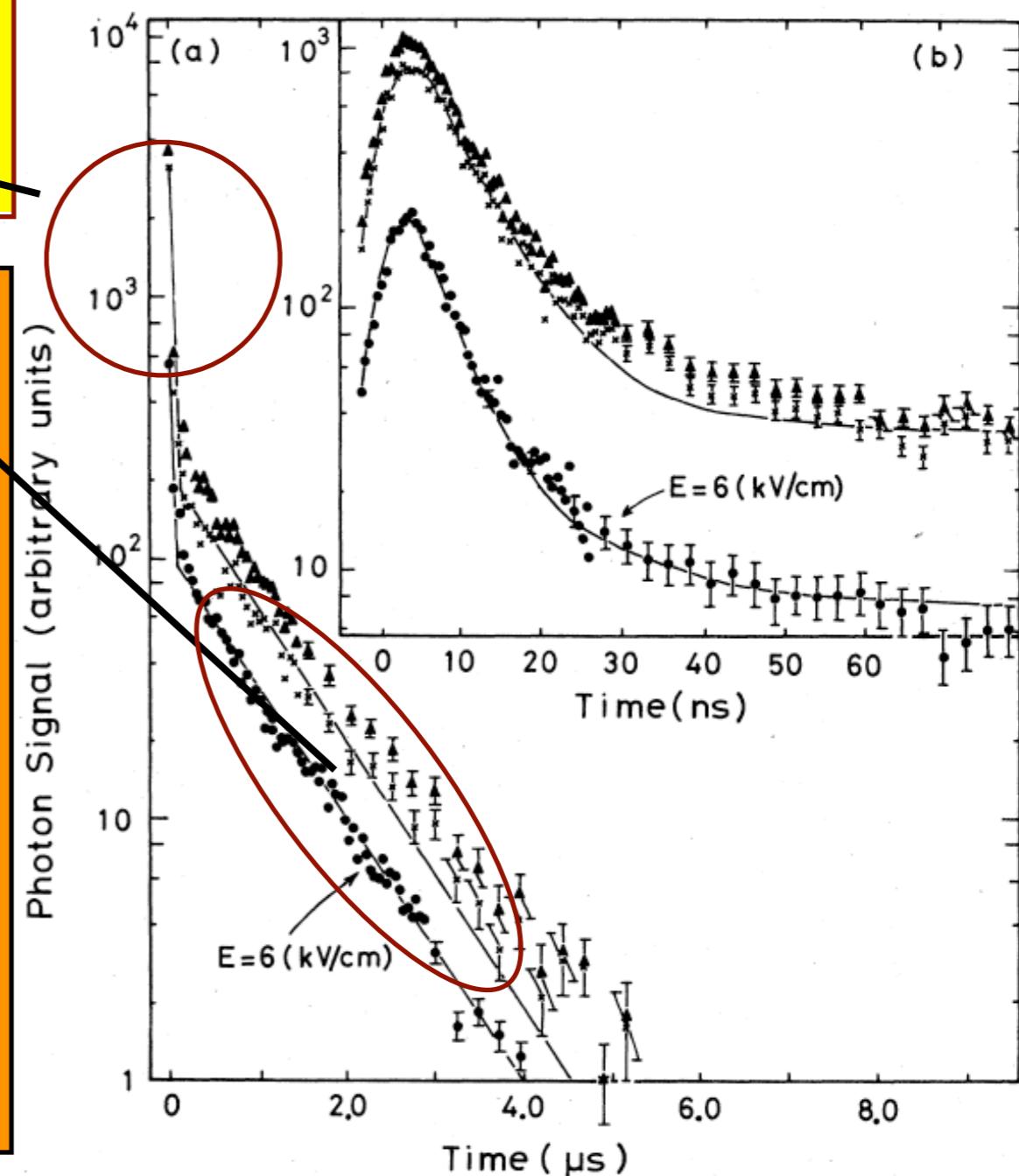
Fast and Slow wavelength spectra
can be separately consider

The presence of fast and slow
components is associated with Ar^{2*}
molecular transition to fundamental
dissociative state from a singlet and a
triplet excited molecular states.

Practically the same scint. spectrum

2

LAr decay time behavior



LAr Scintillation implementation

```
LArMPT->AddProperty("FASTCOMPONENT",LAr_SCPP,LAr_SCIN,num);  
LArMPT->AddProperty("SLOWCOMPONENT",LAr_SCPP,LAr_SCIN,num);  
  
LArMPT->AddProperty("FASTCOMPONENT",Scnt_spe,Scnt_pp,scinum);  
LArMPT->AddProperty("SLOWCOMPONENT",Scnt_spe,Scnt_pp,scinum);  
LArMPT->AddConstProperty("FASTTIMECONSTANT", 7.*ns);  
LArMPT->AddConstProperty("SLOWTIMECONSTANT",1400.*ns);  
LArMPT->AddConstProperty("YIELDRATIO",0.75);
```

Fast/Total ratio for nuclear recoils

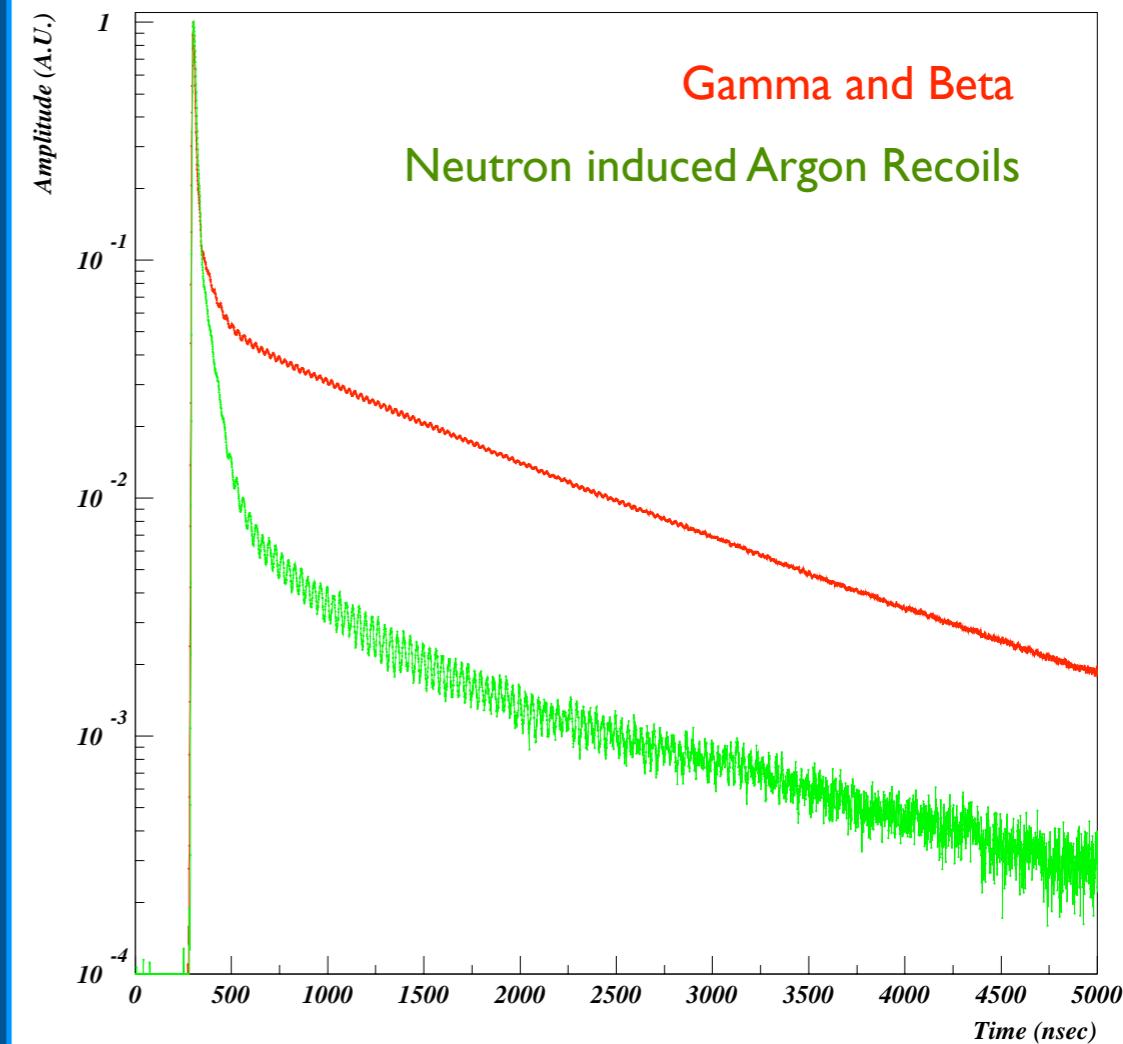
for other particles can bee sets in Physics.cc
(SetScintillationExcitationRatio)

experimental results

1. decay constants are not functions of ionization density
2. for high ionization density the fast components intensity increases at expense of the slow one

$$F_{\text{rec.}} = 0.75 \quad F_{e^-} = 0.23$$

Effect of ionization density on time dependence of luminescence

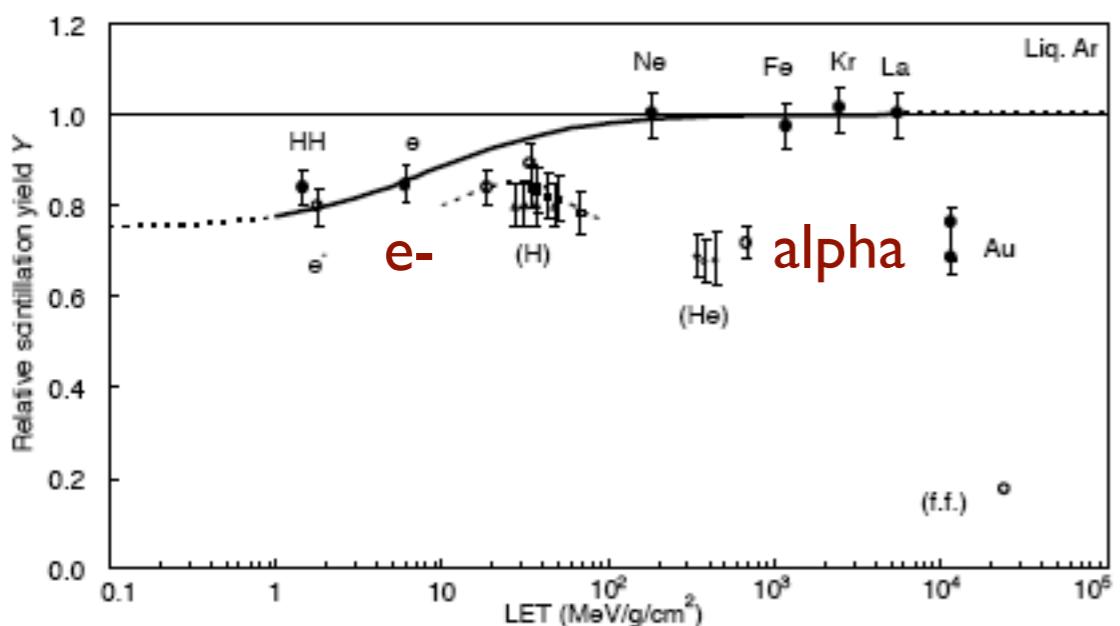


WArP DATA

LAr Scintillation implementation

```
LArMPT->AddProperty("FASTCOMPONENT",LAr_SCPP,LAr_SCIN,num);  
LArMPT->AddProperty("SLOWCOMPONENT",LAr_SCPP,LAr_SCIN,num);  
  
LArMPT->AddProperty("FASTCOMPONENT",Scnt_spe,Scnt_pp,scinum);  
LArMPT->AddProperty("SLOWCOMPONENT",Scnt_spe,Scnt_pp,scinum);  
  
LArMPT->AddConstProperty("FASTTIMECONSTANT", 7.*ns);  
LArMPT->AddConstProperty("SLOWTIMECONSTANT",1400.*ns);  
LArMPT->AddConstProperty("YIELDRATIO",0.75);  
  
G4double scint_yield=1.0/(19.5*eV);  
LArMPT->AddConstProperty("SCINTILLATIONYIELD",scint_yield);
```

ionization density effect on Scintillation Yield



Doke et al, J. Appl. Phys.
41 (2002) 1538

- Scintillation Yield (mean energy to produce a UV photon) dependent on the nature of the impinging particles



- for flat top response particles the mean energy to produce a photon is 19.5 eV
- for other particles can be sets in Physics.cc (ScintillationYieldFactor)

- Scintillation Yield dependent also on:
 - Field configuration
 - Quencher impurities

@ zero E field

$Y_{e^-} = 0.8 Y$ $Y_{\text{alpha}} = 0.7 Y$ $Y_{\text{recoils}} = 0.2 - 0.4$

LAr Scintillation implementation

Scintillation for electron in Physics.cc

```
theScintProcess4Electron = new G4Scintillation("Scintillation");
theScintProcess4Electron->SetScintillationYieldFactor(0.8);
theScintProcess4Electron->SetTrackSecondariesFirst(true);
theScintilProcessElectron->SetScintillationExcitationRatio(0.23);
theScintProcess4Electron->SetVerboseLevel(OpVerbLevel);
```

→ Defining a new Scintillation Process

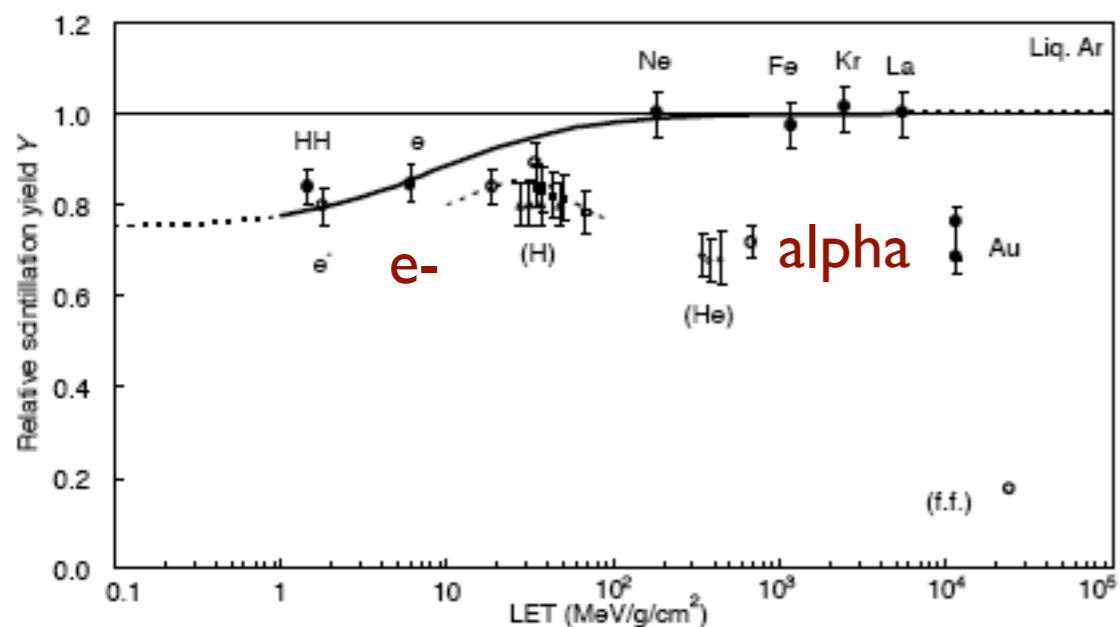
→ Yield Rescaling

→ Fast/Total ratio for electron

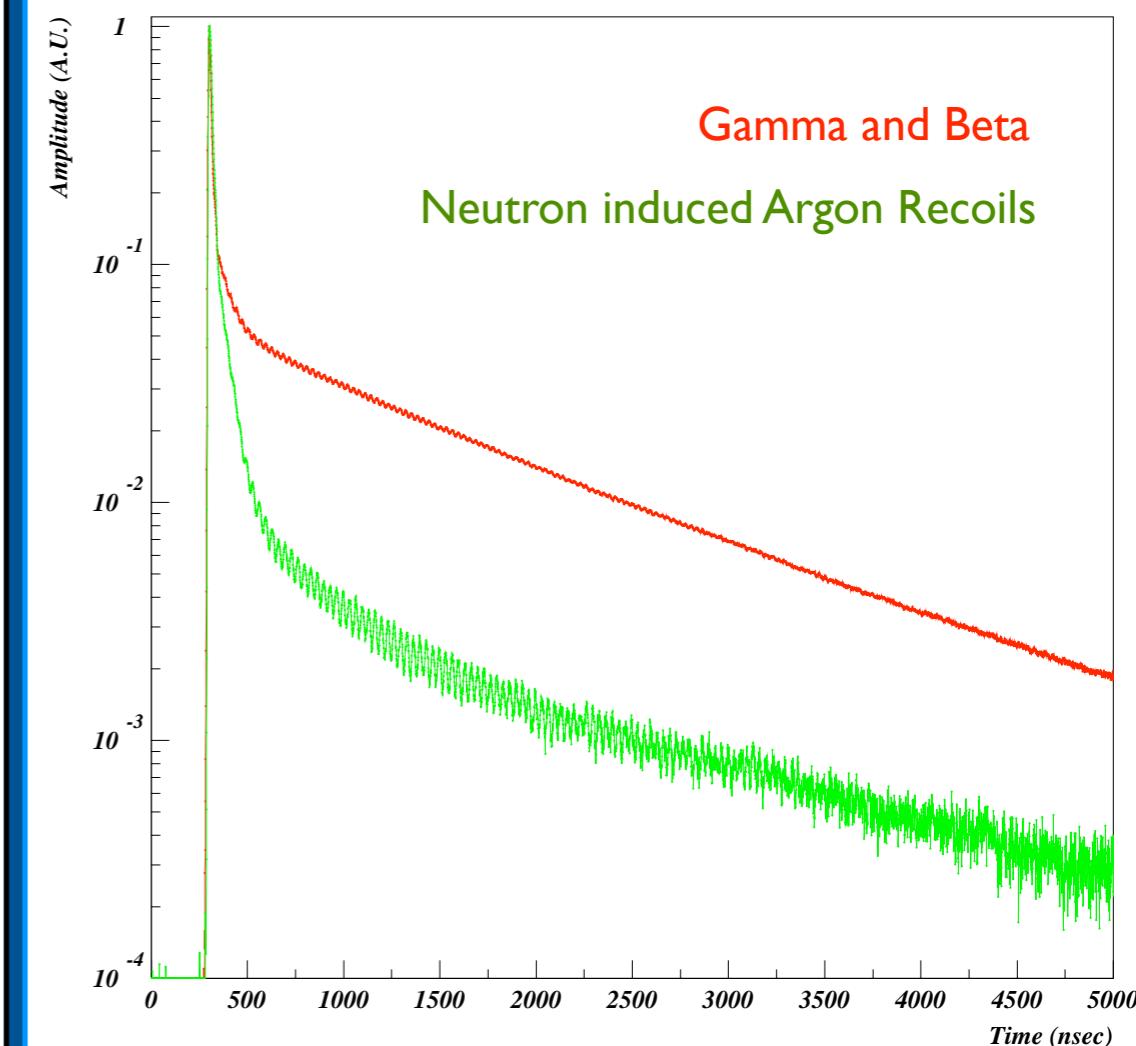
ionization density effect on Scintillation Yield

$$Y_{e^-} = 0.8 Y \quad Y_{\text{alpha}} = 0.7 Y \quad Y_{\text{recoils}} = 0.2 - 0.4$$

@ zero E field



Doke et al, J. Appl. Phys.
41 (2002) 1538



LAr Scintillation implementation

```
LArMPT->AddProperty("FASTCOMPONENT",LAr_SCPP,LAr_SCIN,num);  
LArMPT->AddProperty("SLOWCOMPONENT",LAr_SCPP,LAr_SCIN,num);  
  
LArMPT->AddProperty("FASTCOMPONENT",Scnt_spe,Scnt_pp,scinum);  
LArMPT->AddProperty("SLOWCOMPONENT",Scnt_spe,Scnt_pp,scinum);  
  
LArMPT->AddConstProperty("FASTTIMECONSTANT", 7.*ns);  
LArMPT->AddConstProperty("SLOWTIMECONSTANT",1400.*ns);  
  
LArMPT->AddConstProperty("YIELDRATIO",0.75);  
  
LArMPT->AddConstProperty("SCINTILLATIONYIELD",scint_yield);  
G4double fano = 0.11;// Doke et al, NIM 134 (1976)353  
LArMPT->AddConstProperty("RESOLUTIONSCALE",fano);
```

“[...]

The actual number of emitted photons during a step fluctuates around the mean number of photons with a width given by

*ResolutionScale*sqrt(MeanNumberOfPhotons)
[...]"*

statistical yield fluctuation can be broadened or narrower (impurities, fano factor)



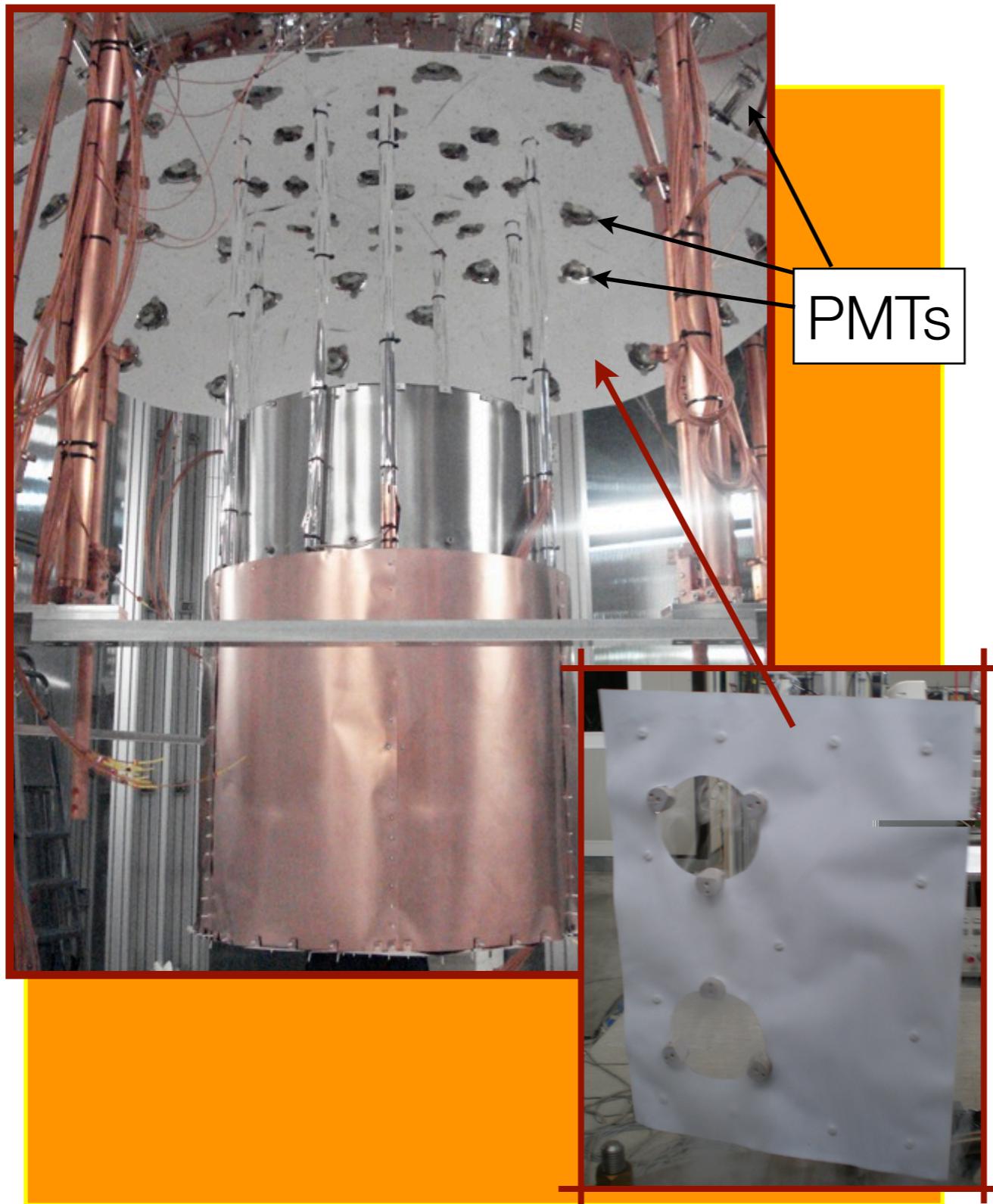
LAr Fano factor = 0.11

(Doke et al, NIM 134 (1976)353)



RESOLUTIONSCALE according to the presence of the fano reduction

light collection system



Two kind of layers
in visible range

- ✓ TPB evaporated on VM2000*: an high reflective layer in order to improve light collection efficiency
- ✓ TPB-Polystyrene mixture: an high transparency layer for coating the PMT windows in order to shift direct XUV light and transmit optical photons

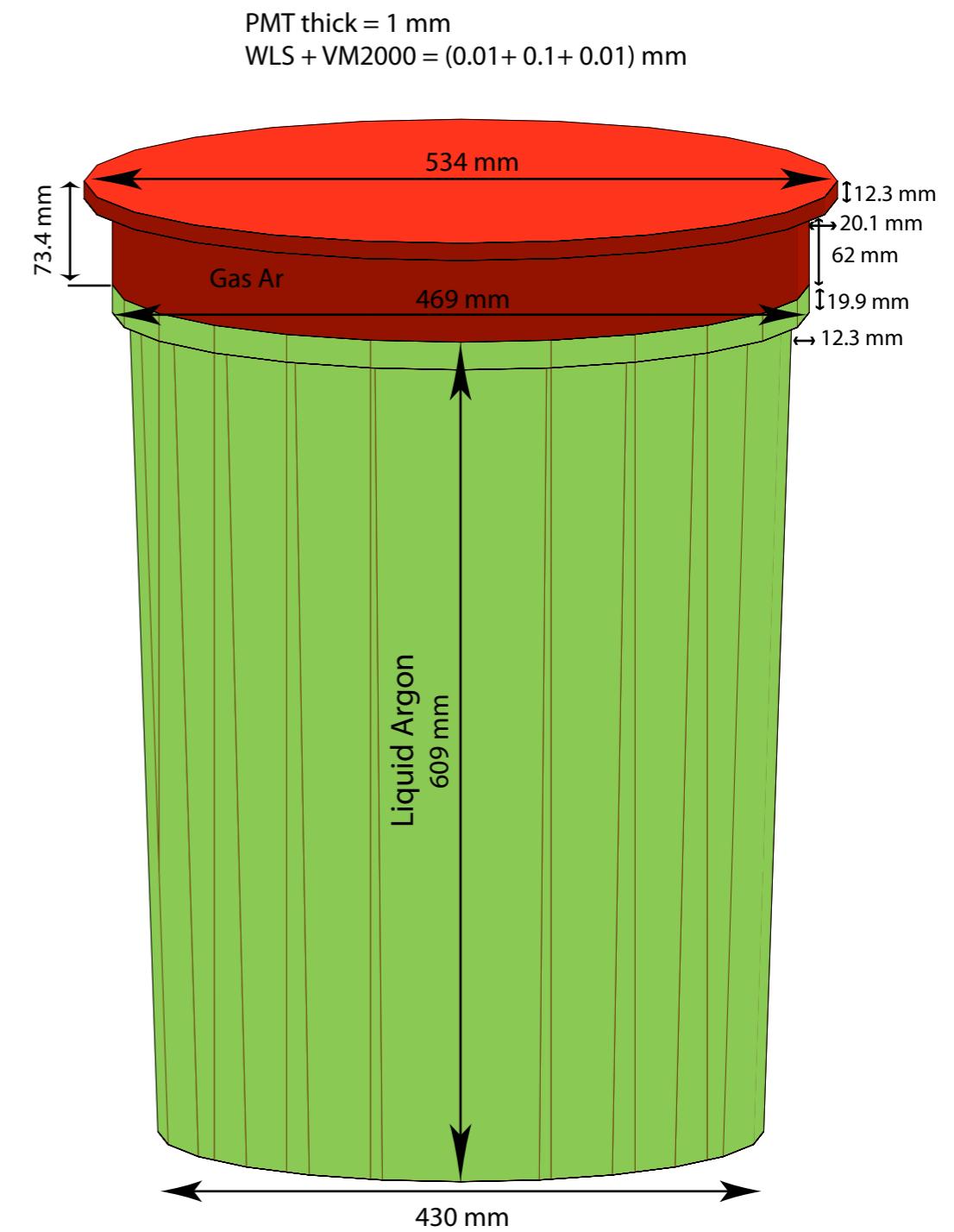
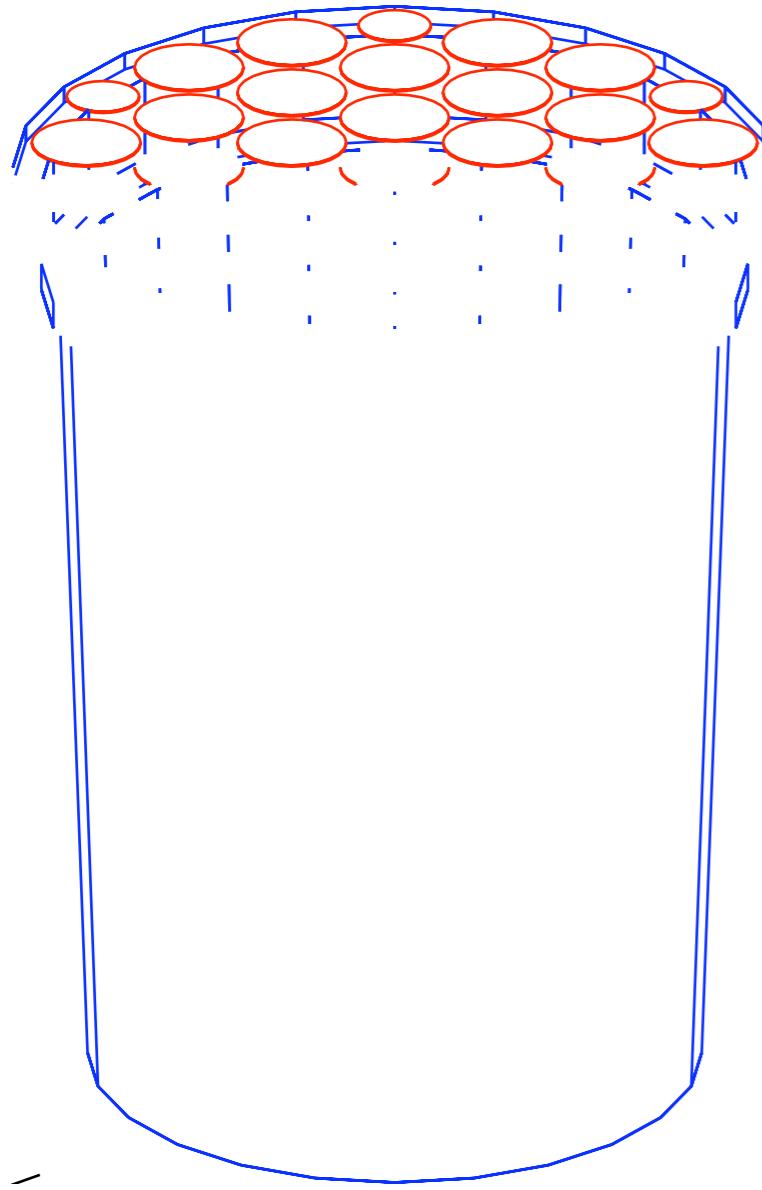
*dielectric mirror

In order to characterize the optical properties of the light collection system a series of measurements have been performed in collaboration with ENEA

Detector Geometry Simulation

Detailed Simulation of the inner detector:

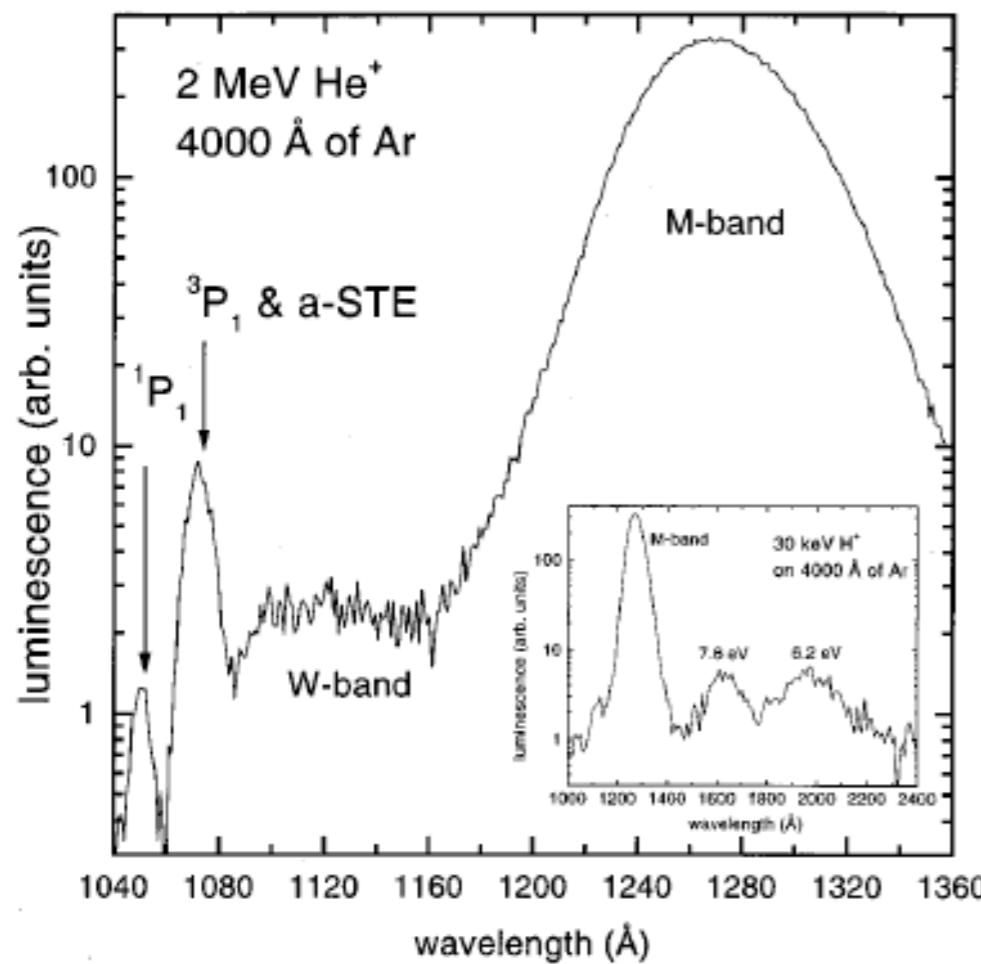
- Liquid & Gas Argon Volumes
- TPB evaporated Mirror foils
- TPB-Polystyrene coated PMTs Glass



WA_P WaveLength Shifting

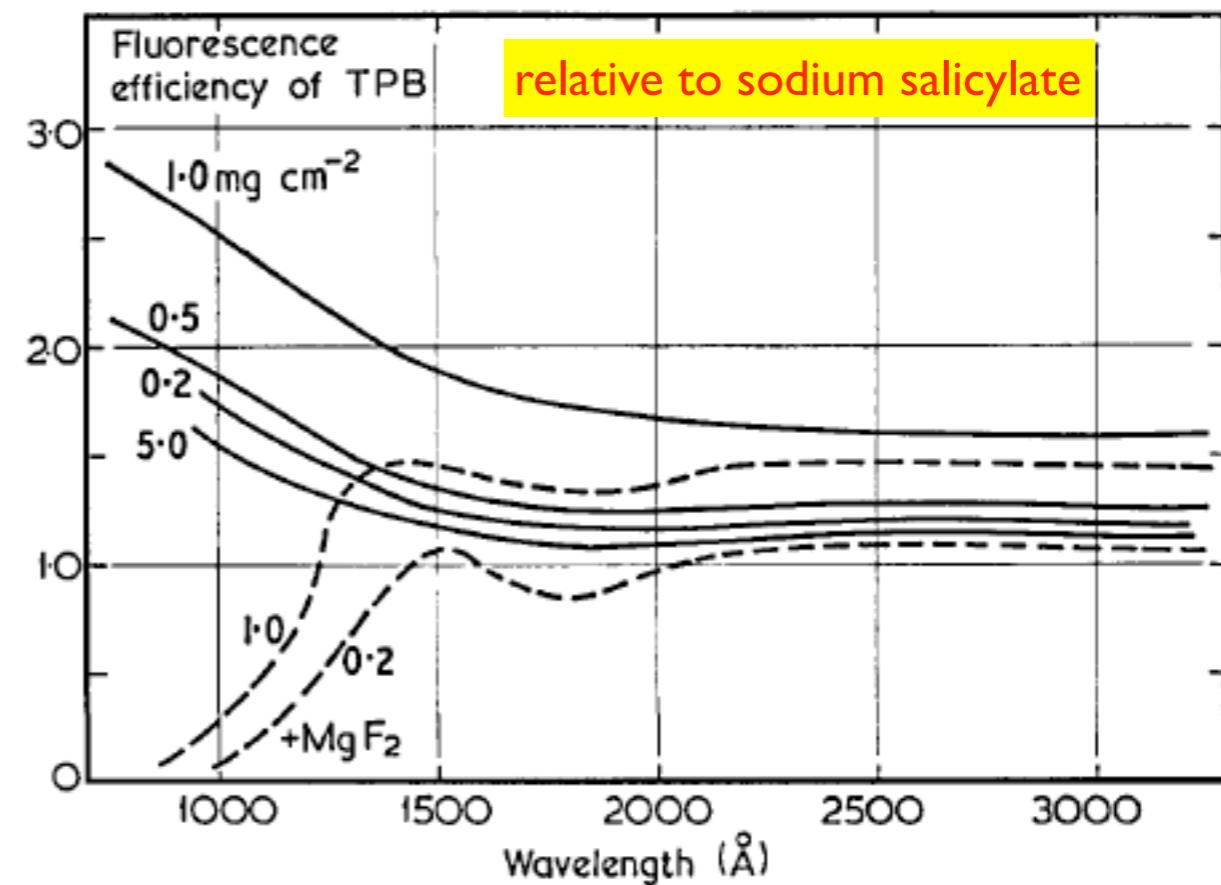
Discrimination power is very sensitive to the number of collected photons

Argon scintillation light is emitted with a narrow spectrum peaked @ 128 nm



PMTs efficiency is very low in the XUV wavelength region (glass transmittance)

Wavelength shift is needed. Solution we choose is Tetra-Phenyl-Butadiene (TPB)



TPB: high XUV conversion efficiency @ $\lambda = 128$ nm probably higher than 1 (depending on TPB thickness)

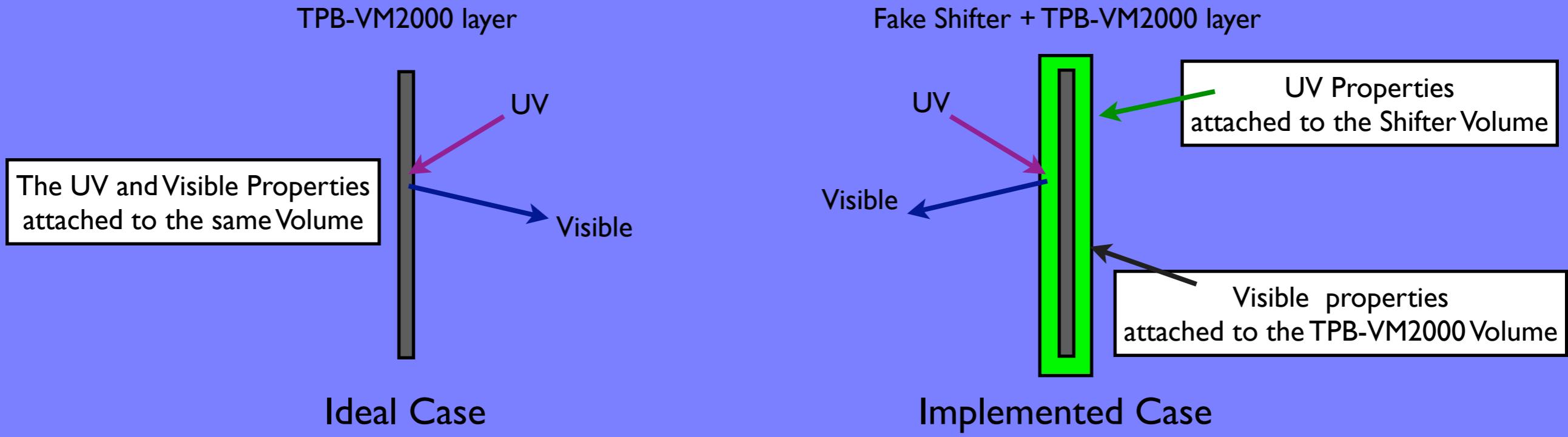
Detector Geometry Simulation

A unique layer with wavelength shifting and Optical Boundary Process properties (T, R,..) but:

- Wavelength Shifting is attached to a Material;
- Optical Boundary Processes implemented using the concept of Surface.

Implementation Strategy

WLS properties attached to a thin volume surrounding the layers with Optical Boundary Properties



In similar cases can be useful an implementation of WLS attached to a Surface!

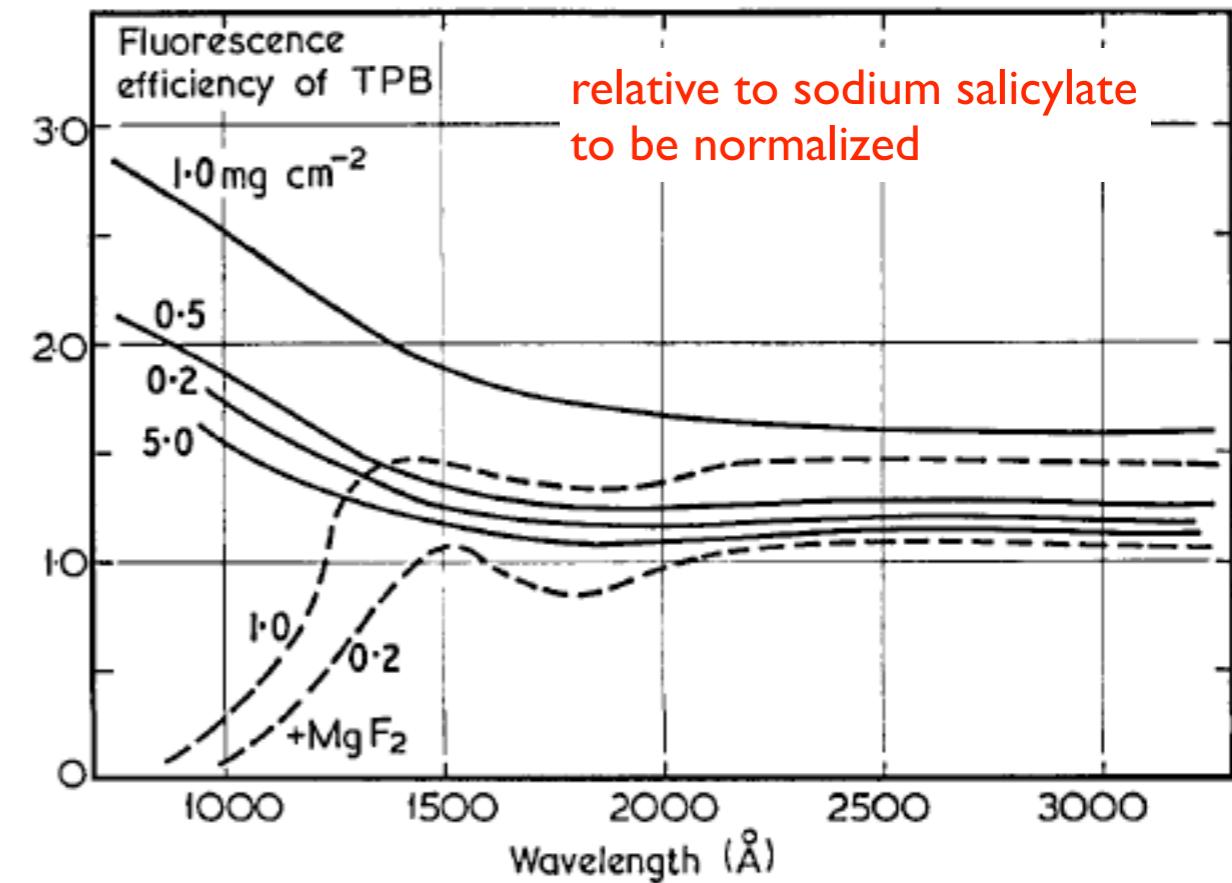
Wavelength Shifter Implementation

- Photon Absorption (as function of λ)
- Photon Emission Spectrum (as function of λ)
- Time delay from Absorption to Reemission

```
//-----Liquid argon Shifter
```

```
density = 1.4*g/cm3;  
  
G4Material* WSLAr = new G4Material  
  (name="WLSLiquidArgon", density, ncomponents=1);  
  
WSLAr->AddElement(Argon, 1.0);
```

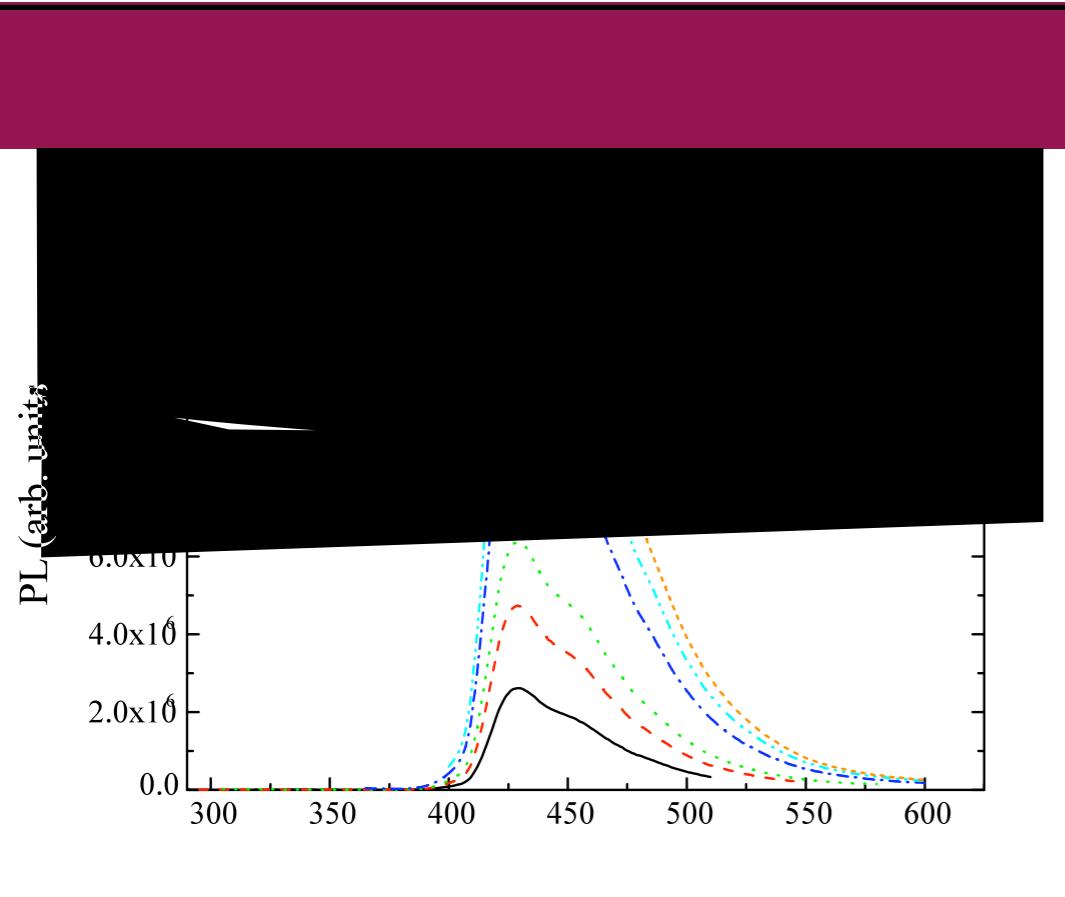
```
//-----
```



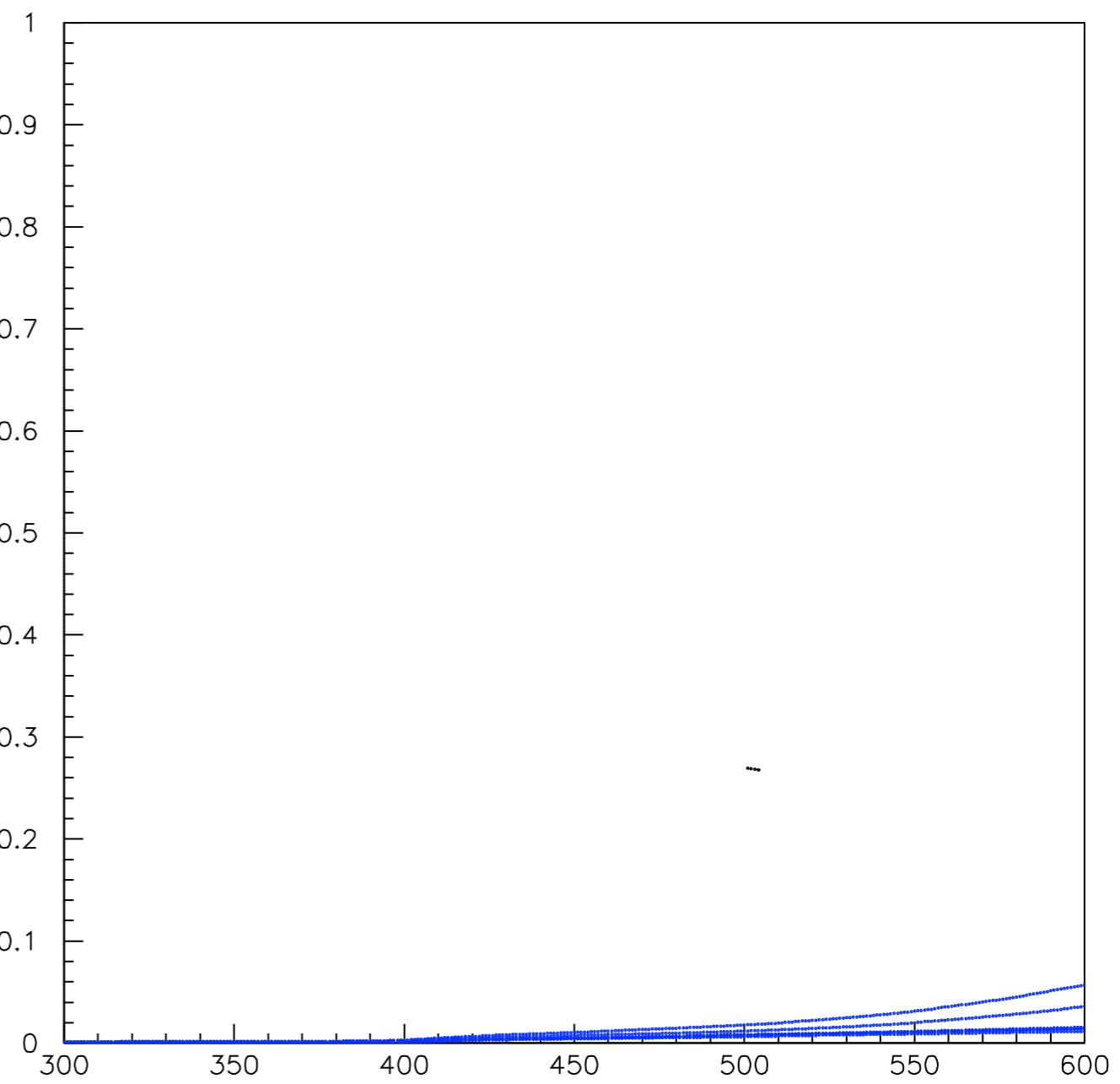
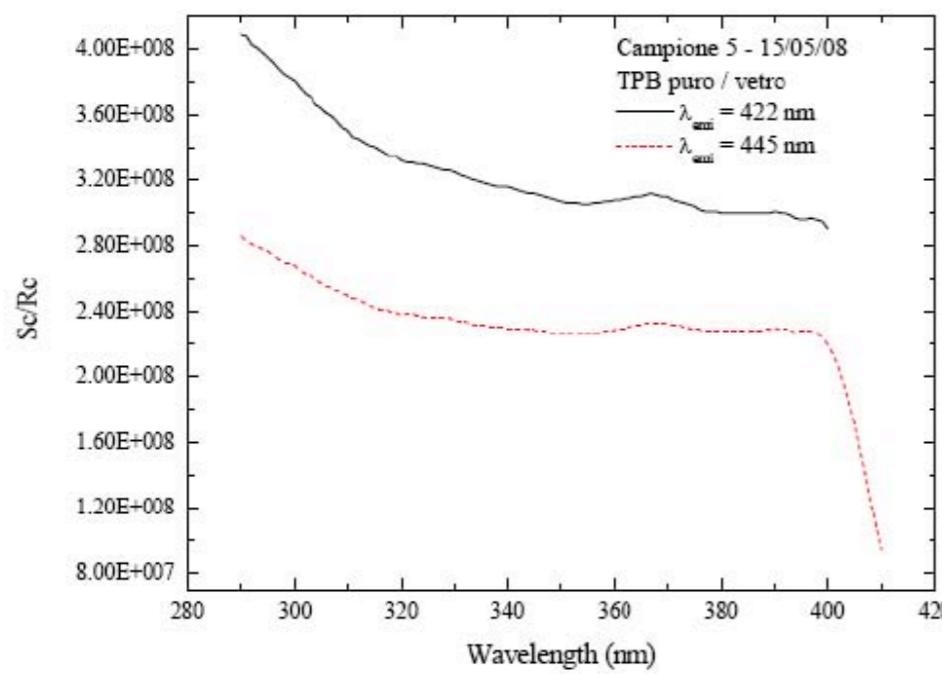
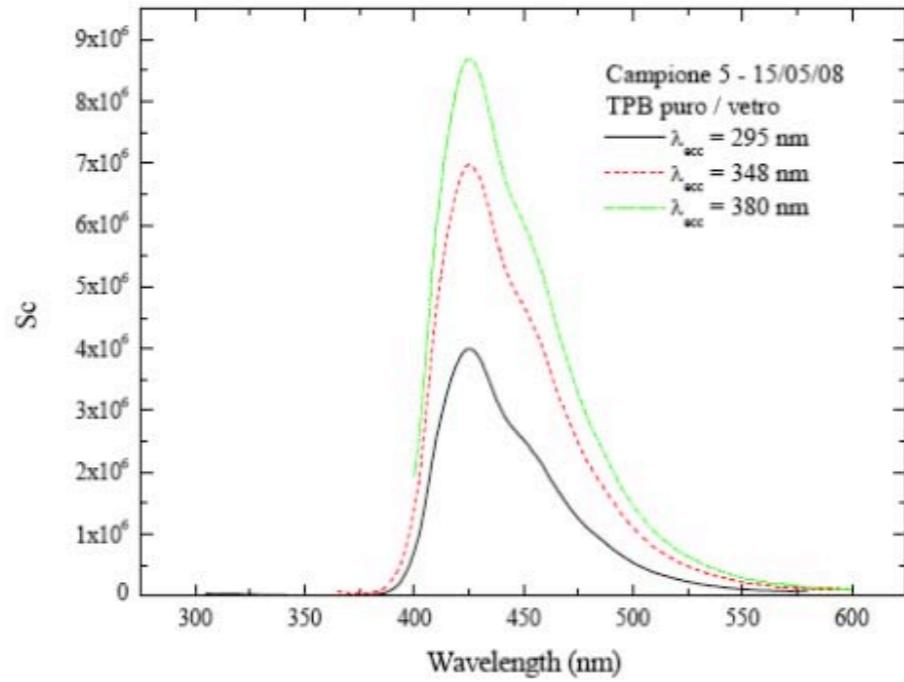
```
const G4int qua=4;  
G4double ene_abso_WLS[qua]={10.0*eV, 9.0*eV, 8.99*eV, 2.*eV};  
G4double abso_WLS[qua]={0.001*nm, 0.001*nm, 100000*m, 100000*m}//100% shift efficiency @ 128 nm
```

```
G4MaterialPropertiesTable* WLSMPT = new G4MaterialPropertiesTable();  
WLSMPT -> AddProperty("WLSABSLLENGTH", ene_abso_WLS, abso_WLS, qua);  
WLSMPT -> AddProperty("WLSCOMPONENT", WLSEnergy, WLS_Emission, NUMWLS);  
WLSMPT -> AddConstProperty("WLSTIMECONSTANT", 0.01*ns);
```

Wavelength Shifter Implementation



Evaporated TPB on glass

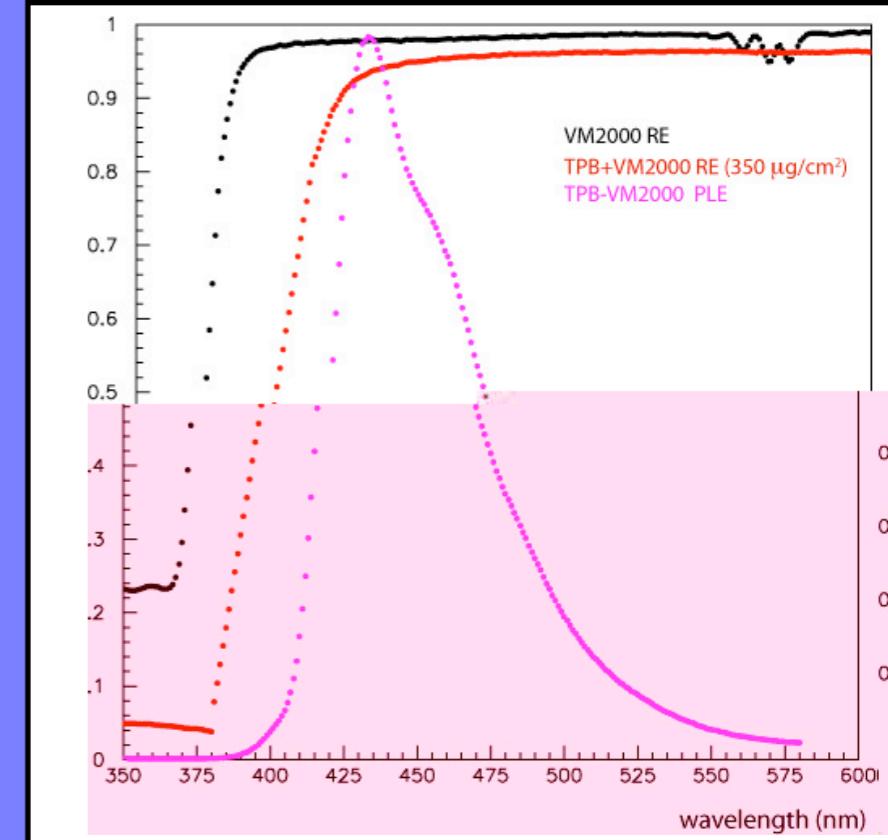


Boundary Processes Implementation

Evaporated TPB on VM2000

1. A very diffusive substrate (TPB) on a Mirror
2. Layer RINDEX not known
3. Also TPB thickness not known
4. VM2000 has almost null T: fully characterized by R

- SkinSurface
- dielectric-dielectric
- groundfrontpainted (only Lambertian Reflection)

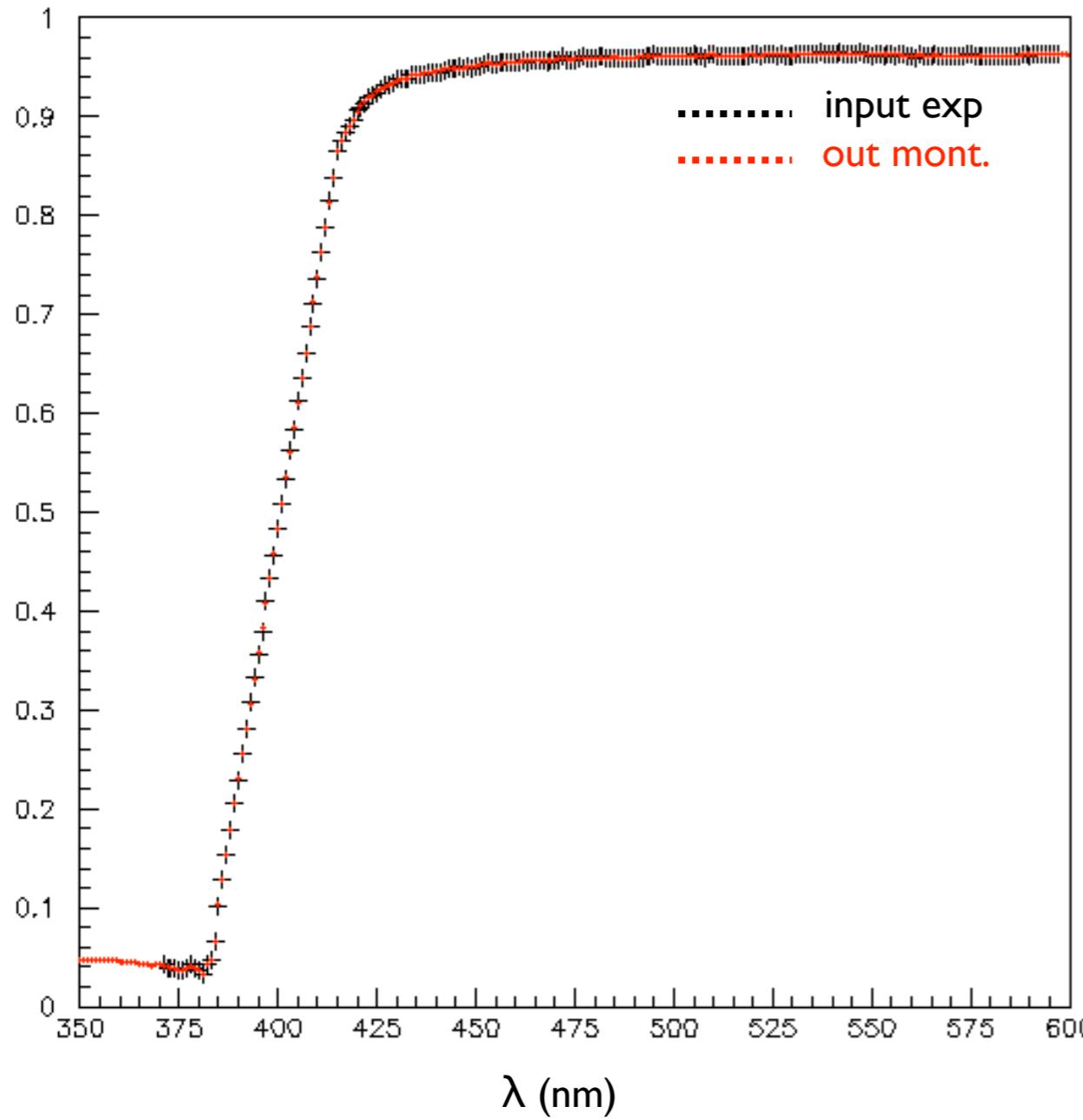


```
G4OpticalSurface* OpSurface = new G4OpticalSurface("OpS1");
G4LogicalSkinSurface* Surface = new G4LogicalSkinSurface("OpS1", VM2000_log, OpSurface);
OpSurface -> SetType(dielectric_dielectric);
OpSurface -> SetFinish(groundfrontpainted);
G4MaterialPropertiesTable *OpSurfaceProperty = new G4MaterialPropertiesTable();
OpSurfaceProperty -> AddProperty("REFLECTIVITY", pp, reflectivity, NUM);
```

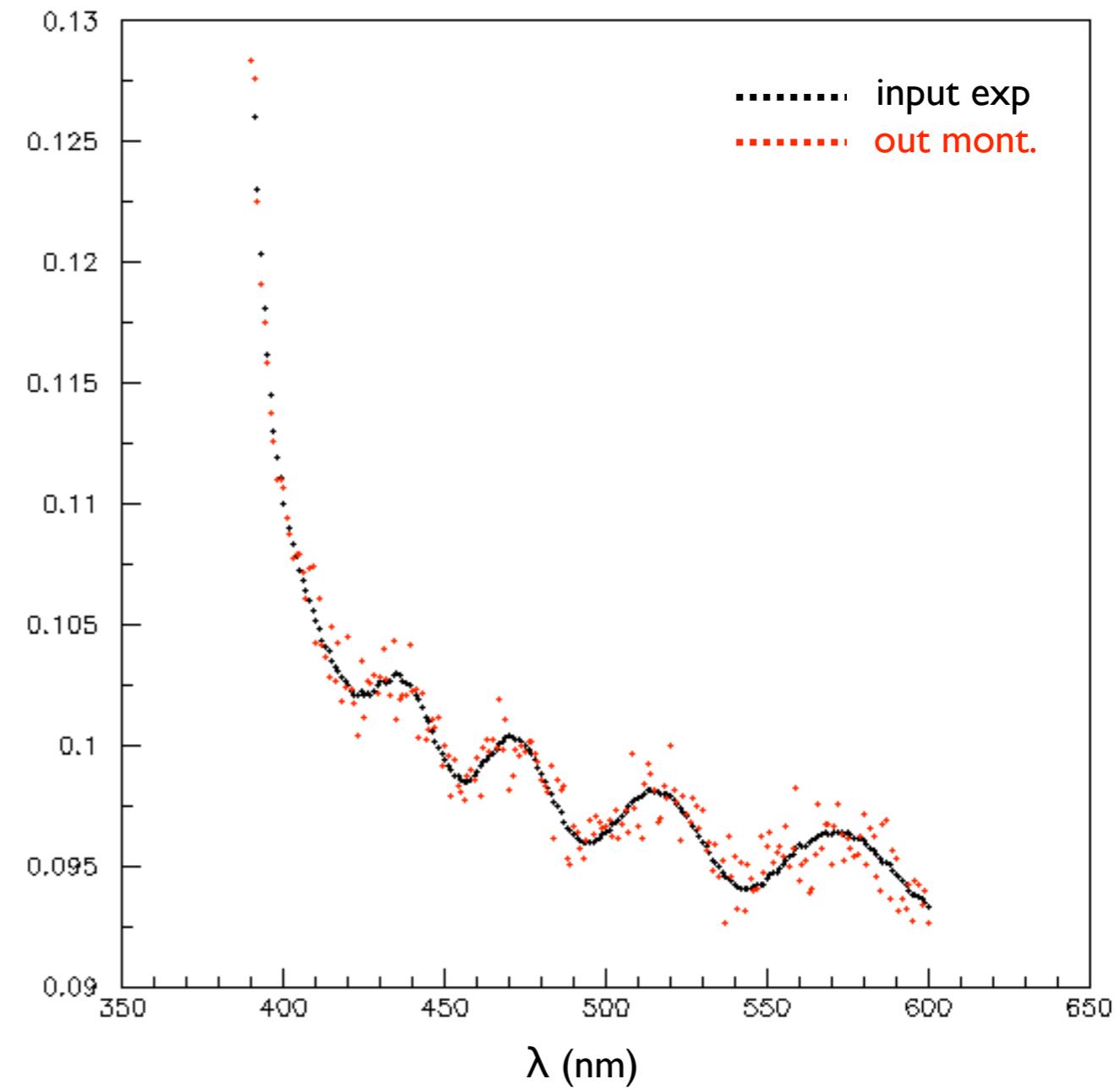
OpProperties Debugging

100000 Primary optical photon incidents with random angle on the surfaces

TPB-VM2000 Reflectance



TPB-Polystyrene Reflectance

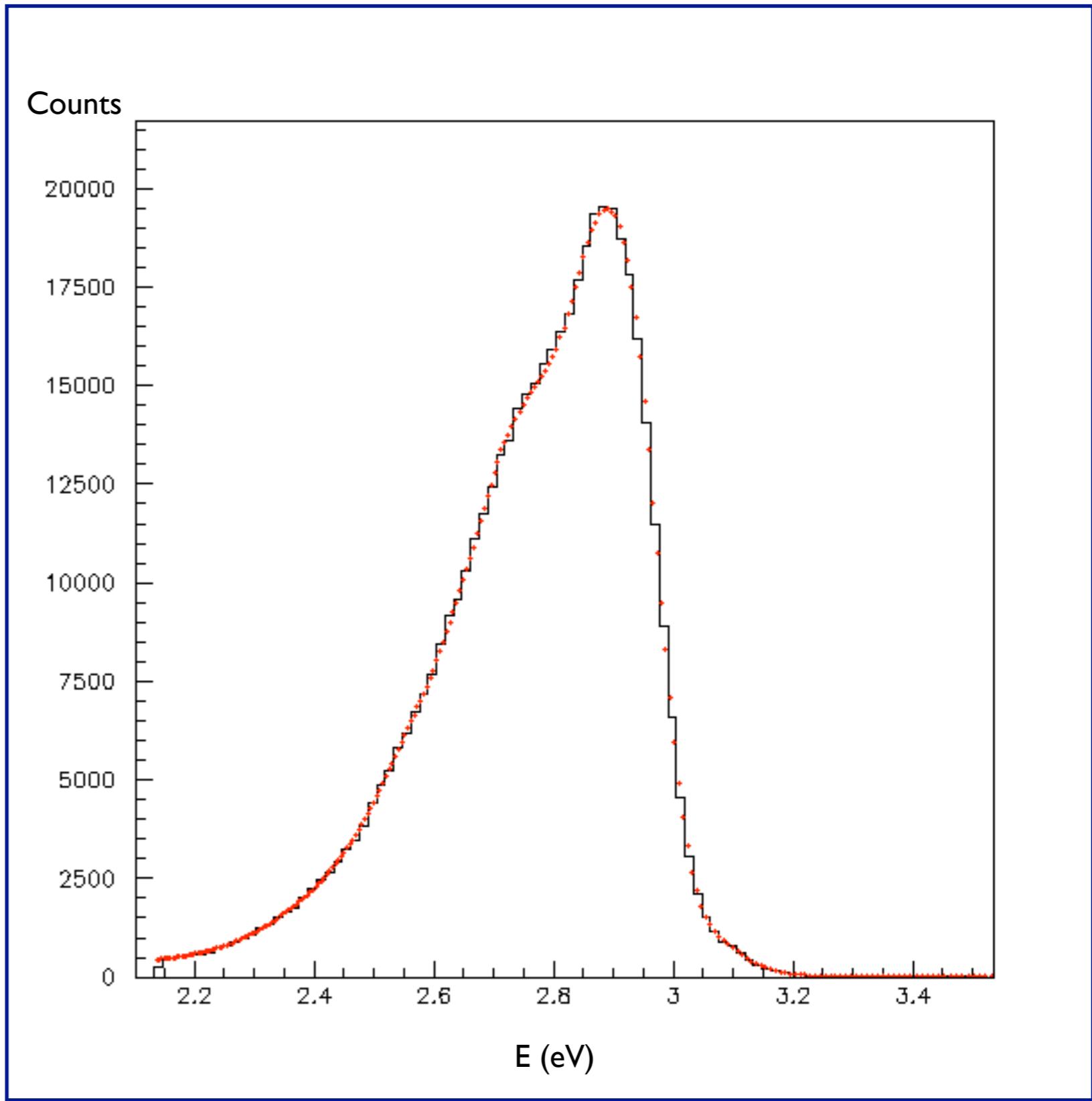


OpProperties Debugging

Fluorescence Spectrum

vs

Energy Spectrum of the
optical photons detected by
the PMTs



Resolution Study

1. Monoenergetic e^- events (20,100,250,500) keV
2. Uniformly distributed in the LAr Volume
3. Fitted with the sum of 4 gaussian with:
 - $\sigma^2 = a + b \text{ (phe)} + (c \text{ phe})^2$
 - free μ_i (also with constraints on μ_i ratios)

Fit results

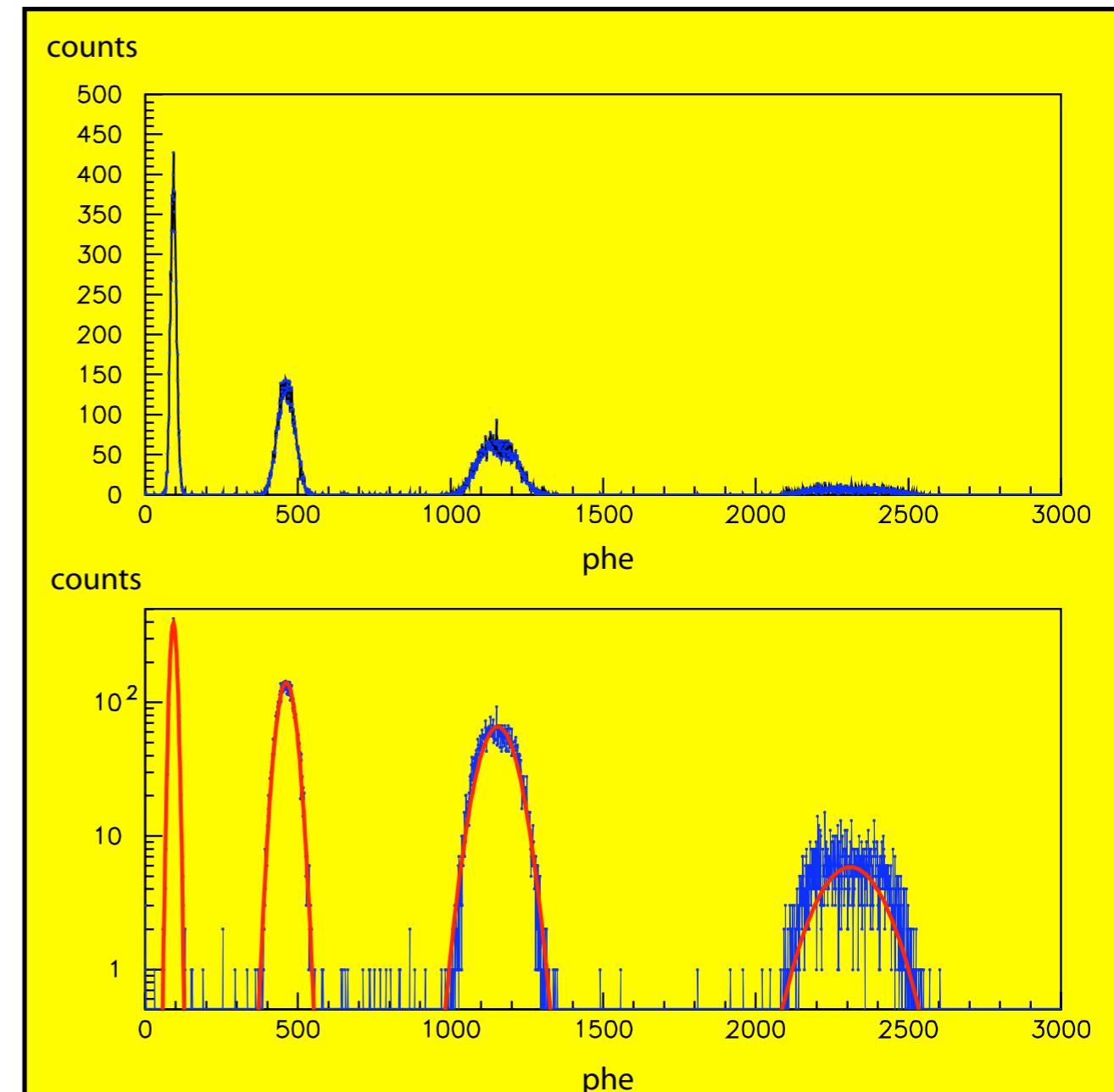
$\mu_1 = 92.8 (0.1)$
 $\mu_2 = 461.9 (0.3)$ According to the
 $\mu_3 = 1154.1 (0.6)$ primary E ratios
 $\mu_4 = 2306 (4)$

$a = 1 (5)$
 $b = 0.84 (0.05)$
 $c = 0.039 (0.001)$

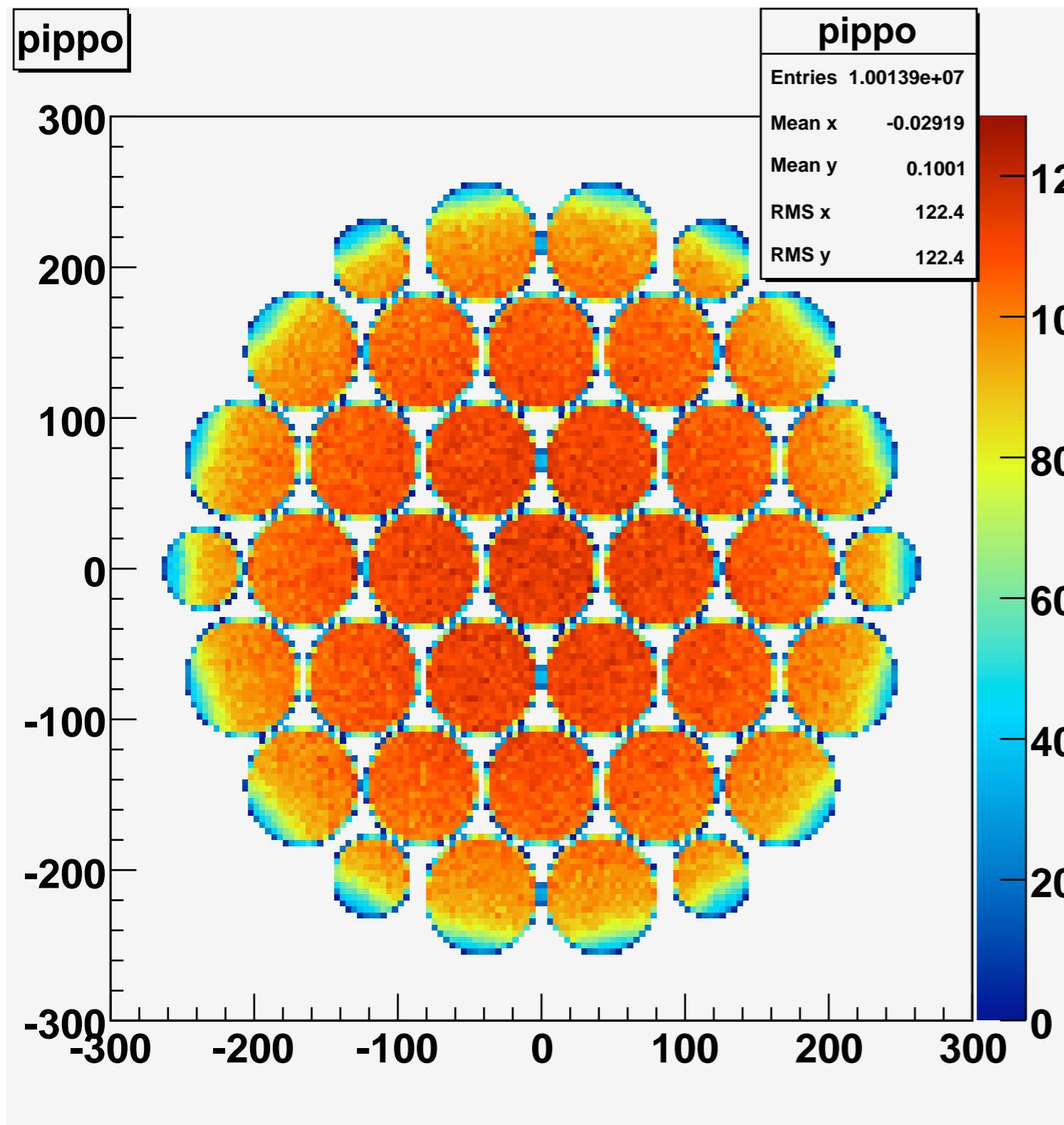
b according with expected binomial fluctuations

$$1 - \text{det_eff} \approx 0.89$$

Montecarlo Light Yield = 4.6 phe/keV



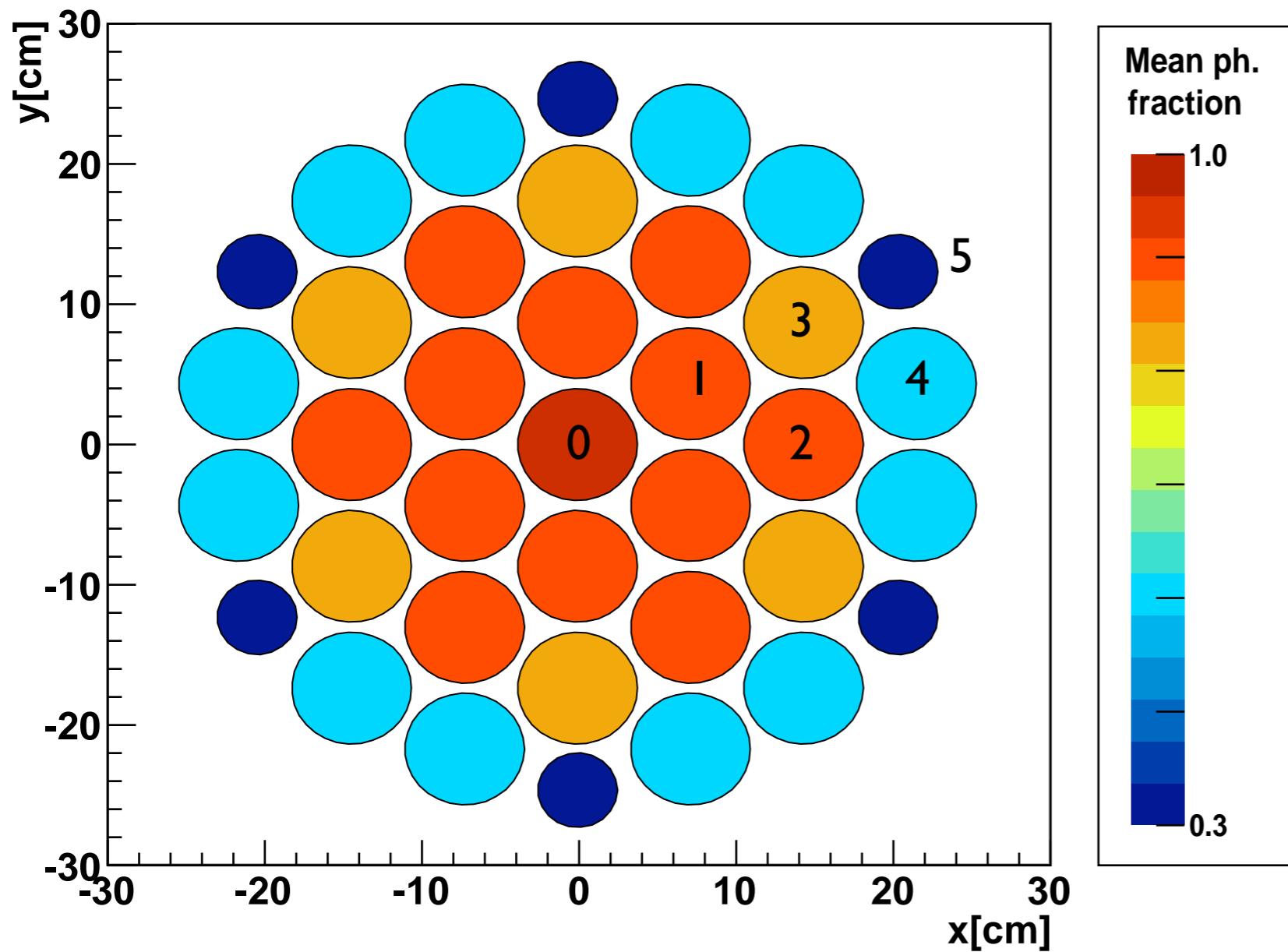
X-Y distribution on PMTs



- Primary Generated uniformly in the LAr volume
- There is a visible different collection efficiency, decreasing with R

distribution of hits on PMTs

- The number of photons detected by each phototube has been considered
- effective sensible surfaces for 3inch and 2inch ($r=24$ mm $r=34$ mm)
- The considered quantity is the relative efficiency, normalized to the most efficient PMT

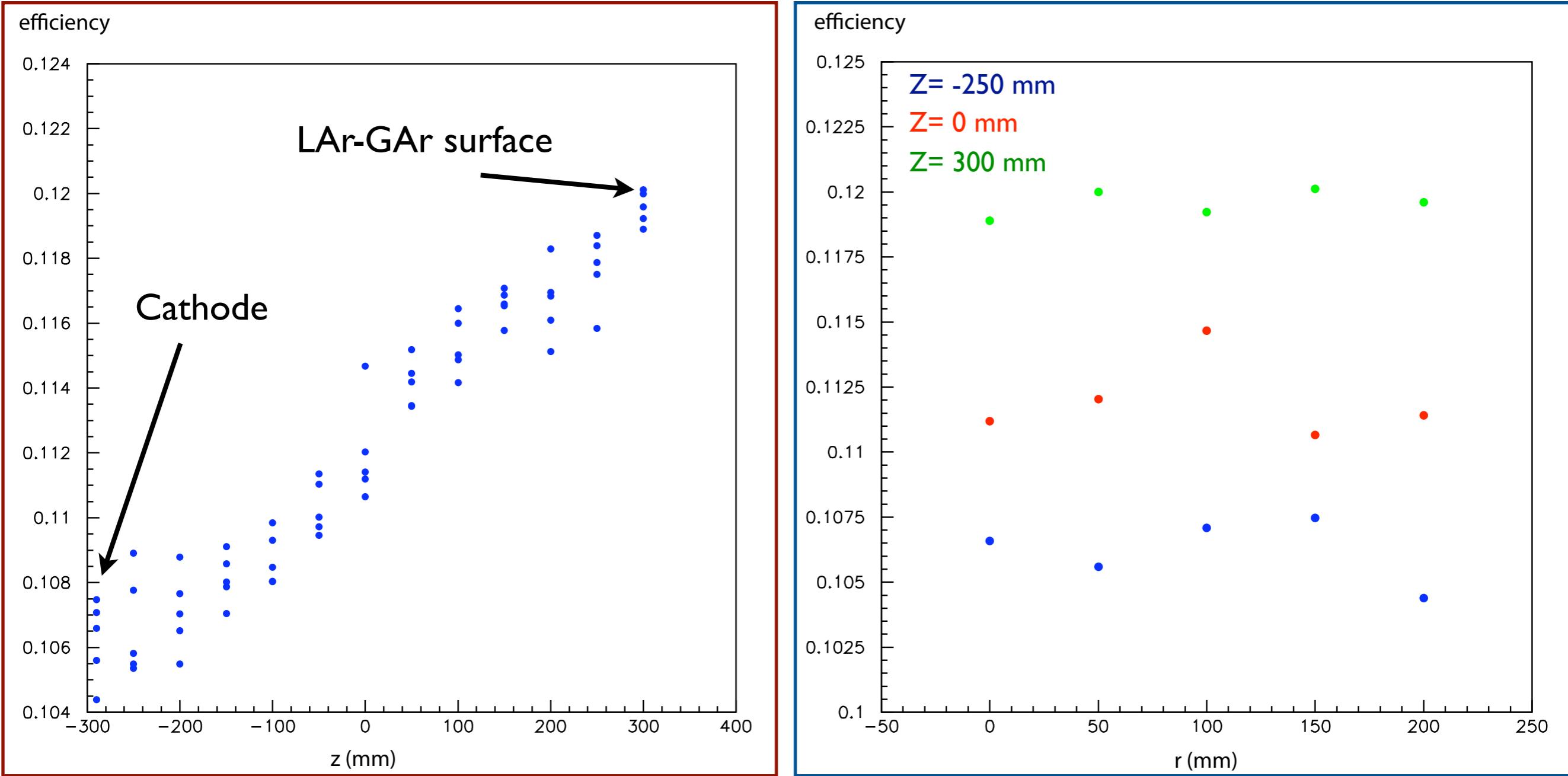


- the Most efficient is the PMT in the center
- Clear dependence on R
- 2inch PMTs have of course also smaller photocathode surface

<u>Efficiency ring values:</u>
1. 0.99
2. 0.96
3. 0.94
4. 0.82
5. 0.33

Efficiency Vs Z,R

- bunches of 100.000 UV primary photons generated in a grid of 5 cm in Z & R
- Detection Efficiency defined as the ration between $N_{\text{detected}}/N_{\text{Primary}}$



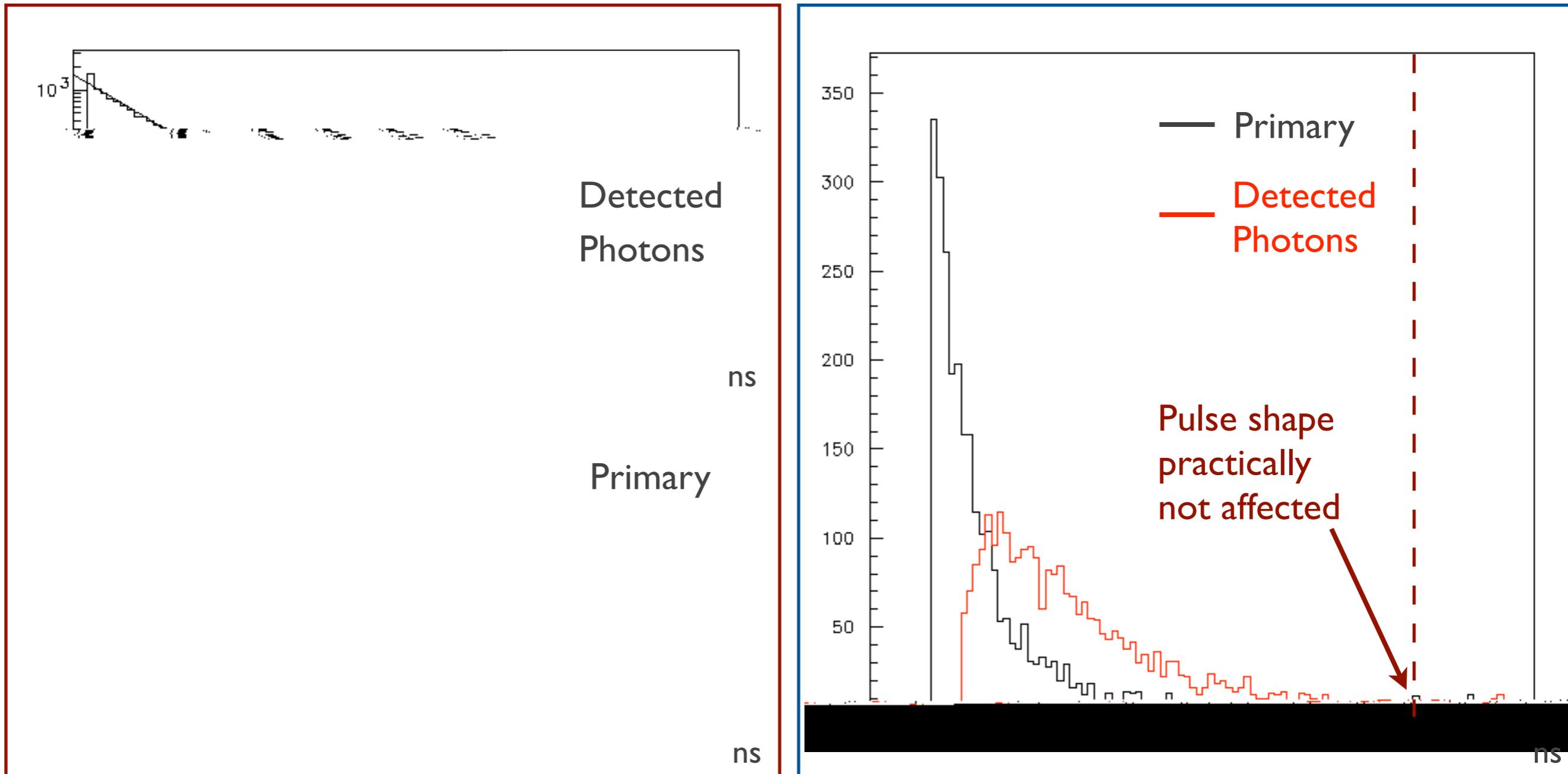
1. Efficiency of the order of 10%

2. Increasing with Z (max variation 10 %)

3. no R dependence

Time Distribution

- 100.000 UV primary photons uniformly in the LAr Volume
- primary time fixed at zero (left) and according LAr scintillation decay times (right)

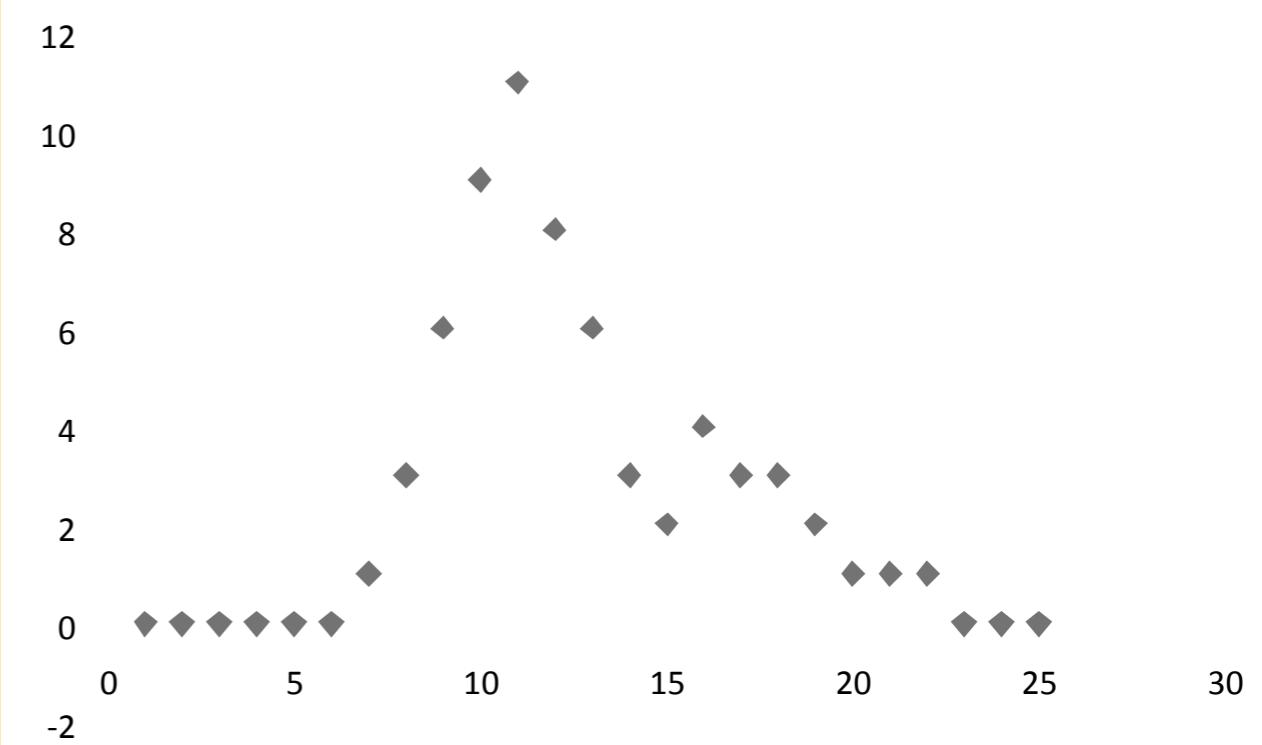


Detector time response is approximately exponential with “decay time” of 10.5 ns

Signal Simulation

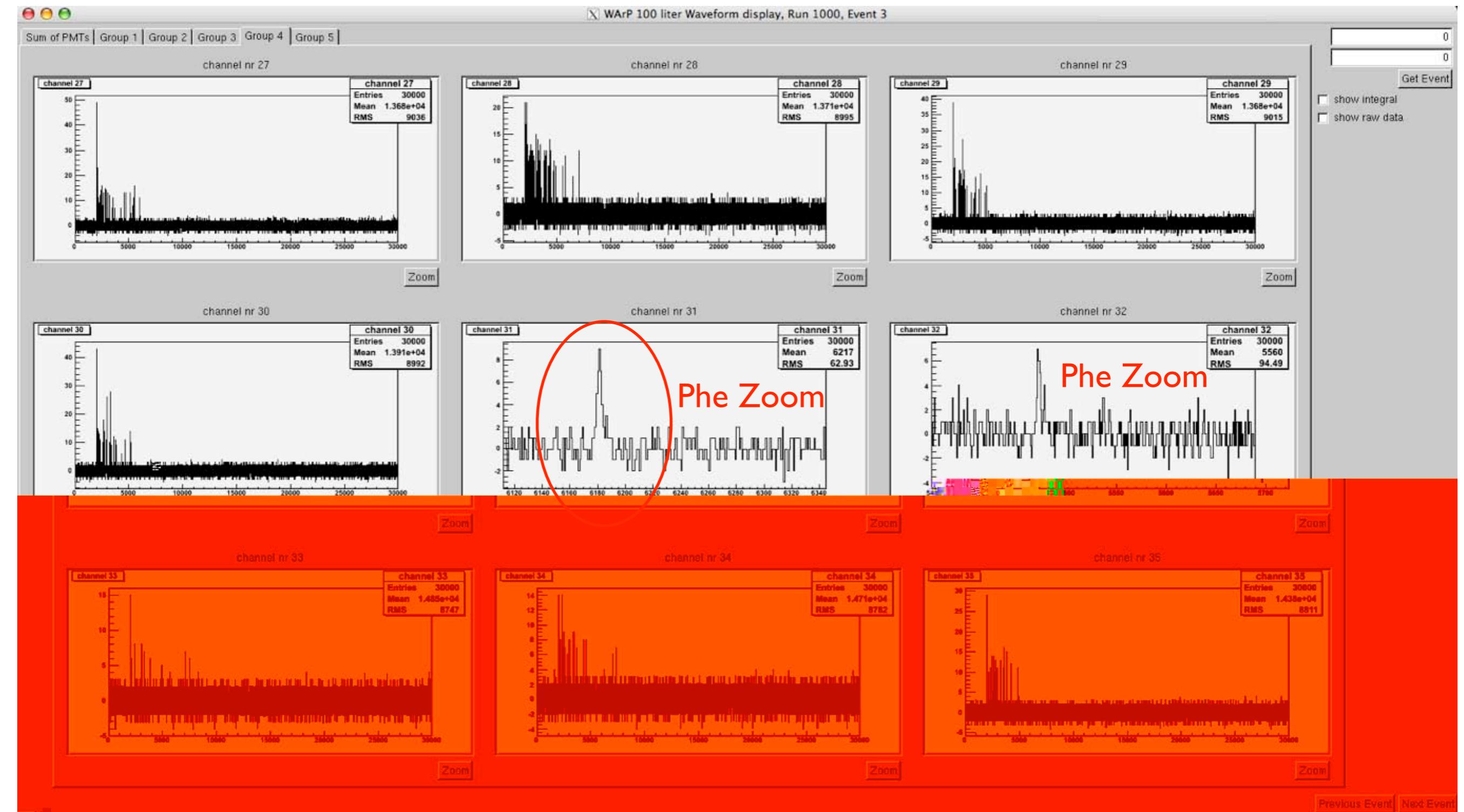
A dedicated C++ code has been made to simulate the electronics response:

- SER
 - SER shape obtained looking at real single photonelectrons shape in the acquired data
 - Convolution of the single hits
 - SER amplitude values and its spread for each phototube according to experimental values

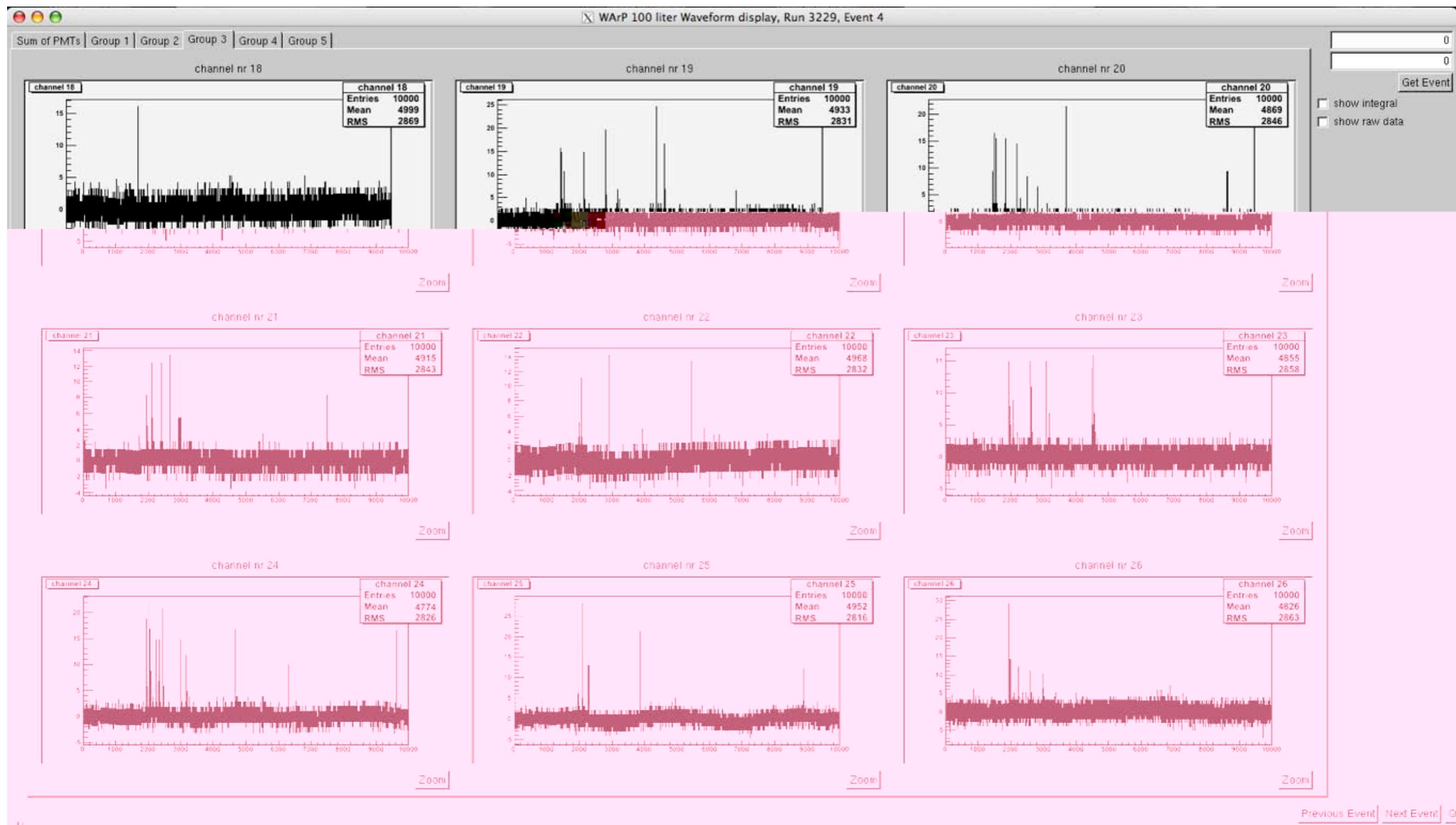


- noise
 - noise simulated using simple gaussian fluctuations around the baseline
 - Noise from exp. data: to be implemented

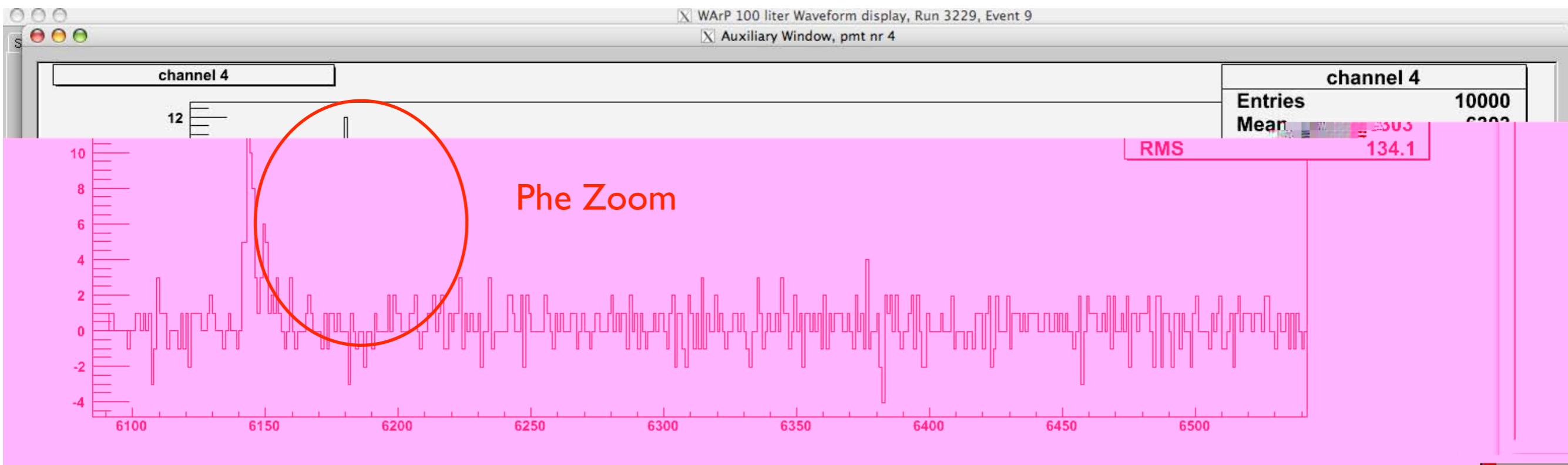
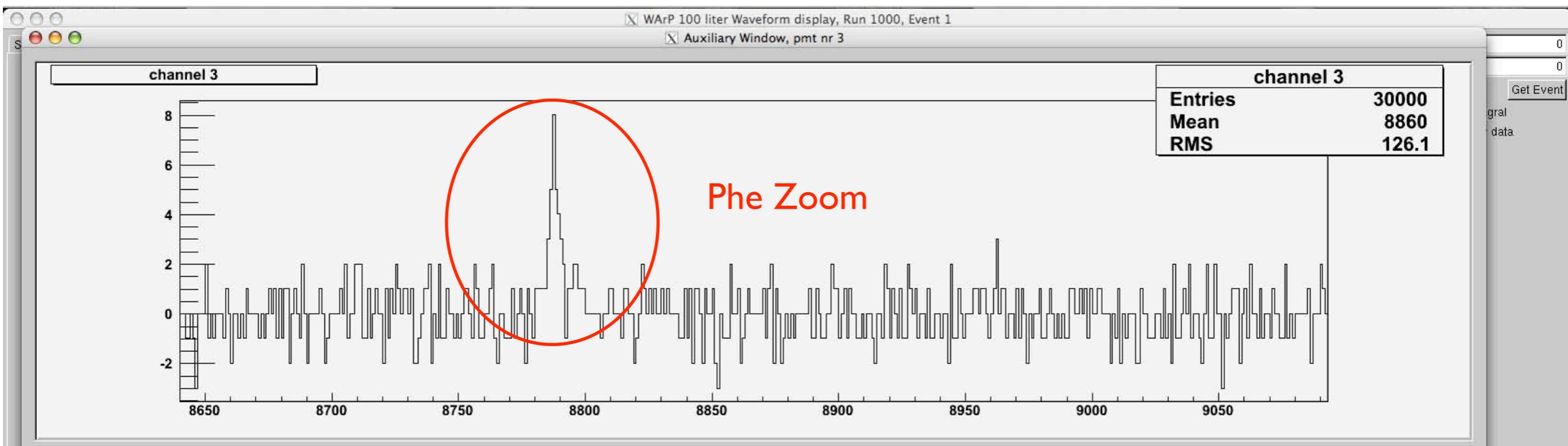
Signal Simulation



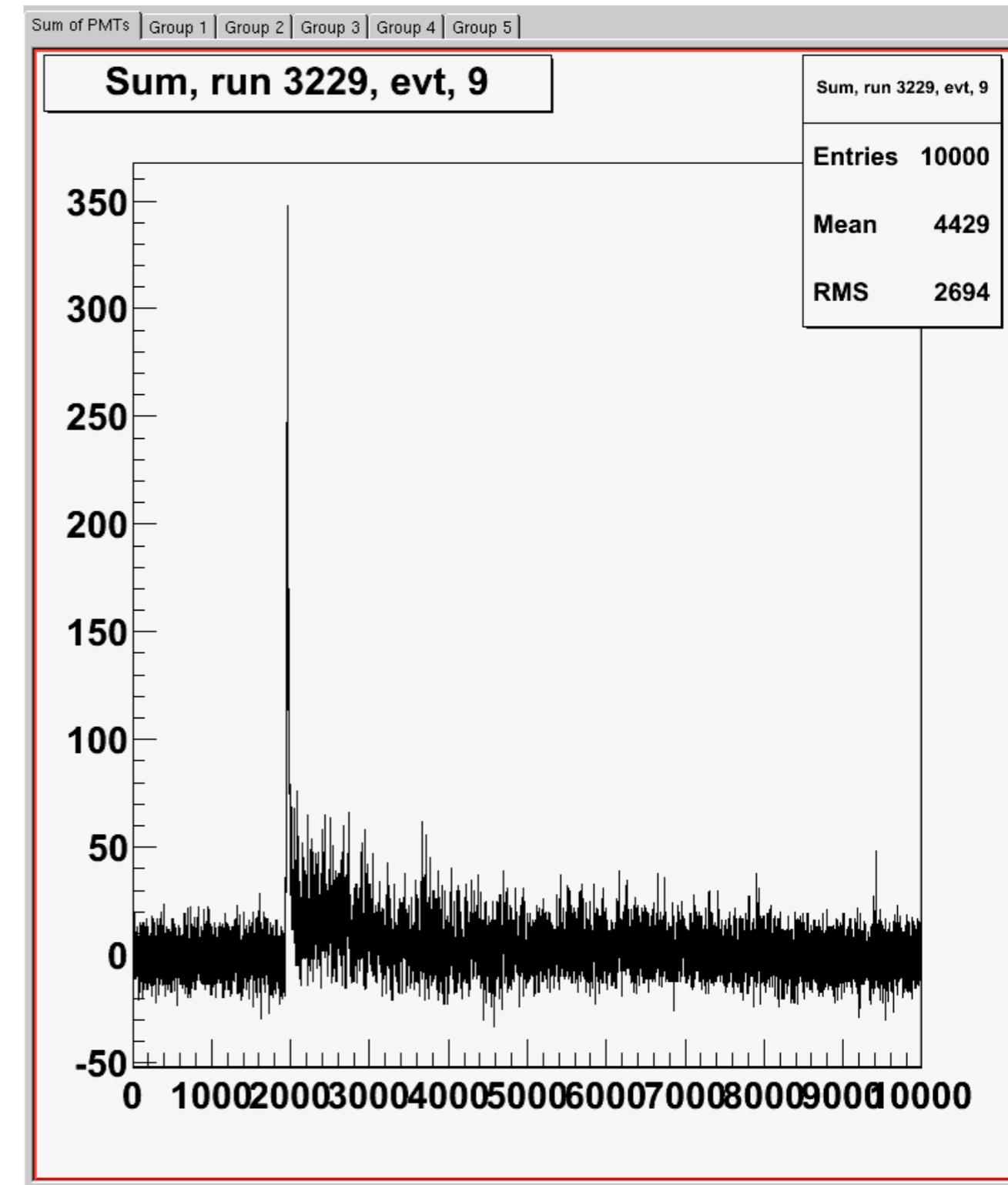
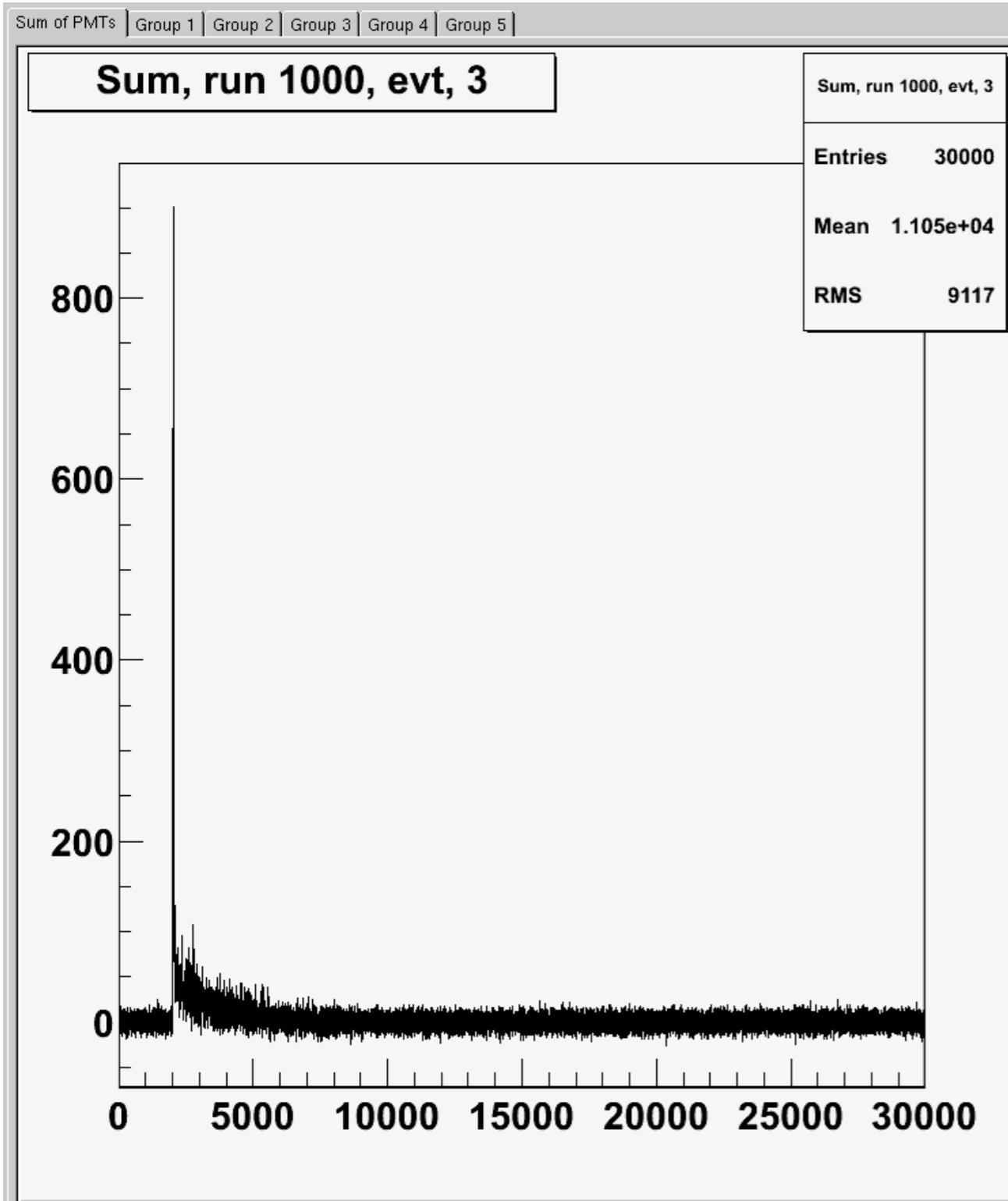
Real Events



Signal Simulation

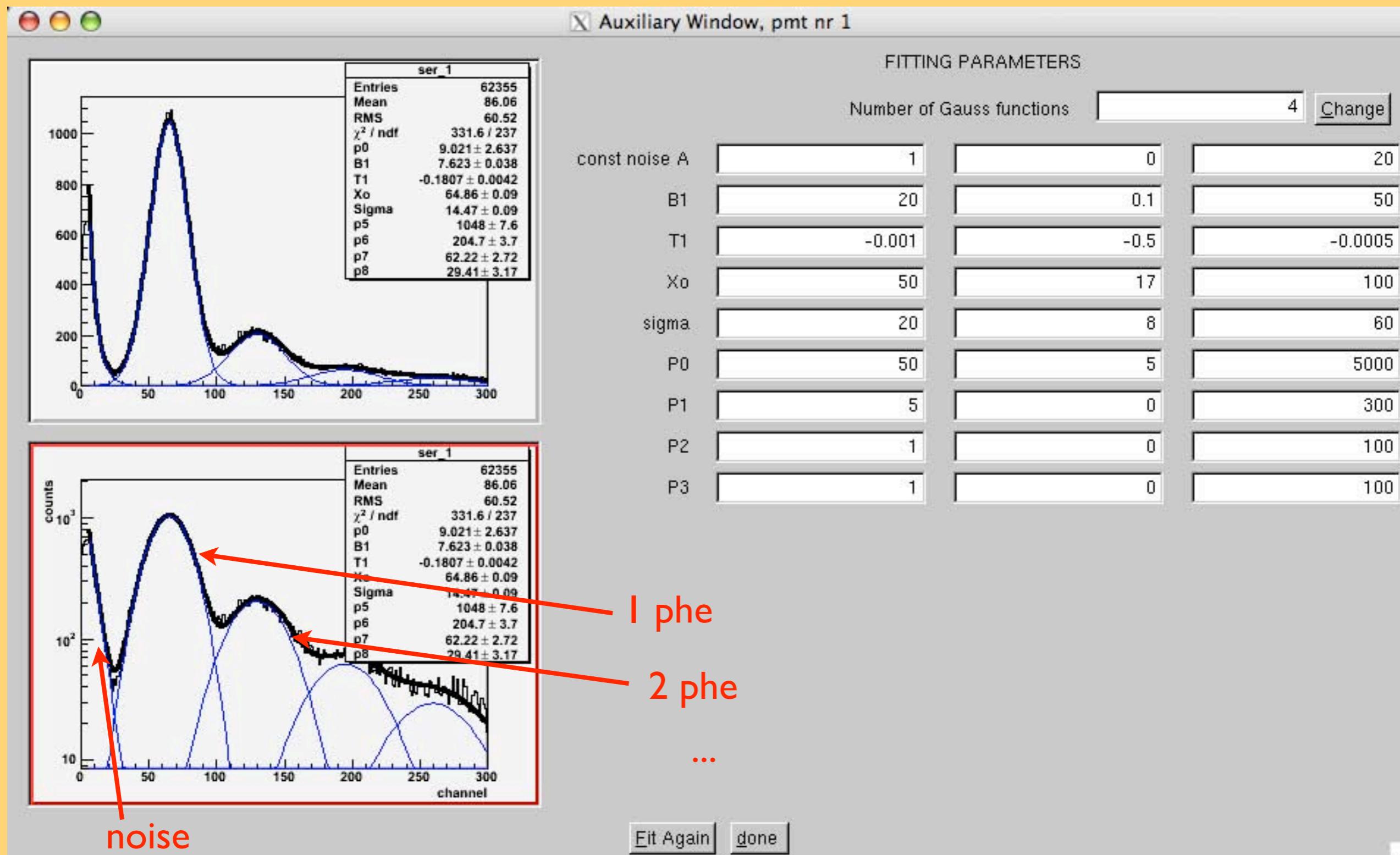


Simulation Vs Exp. Data

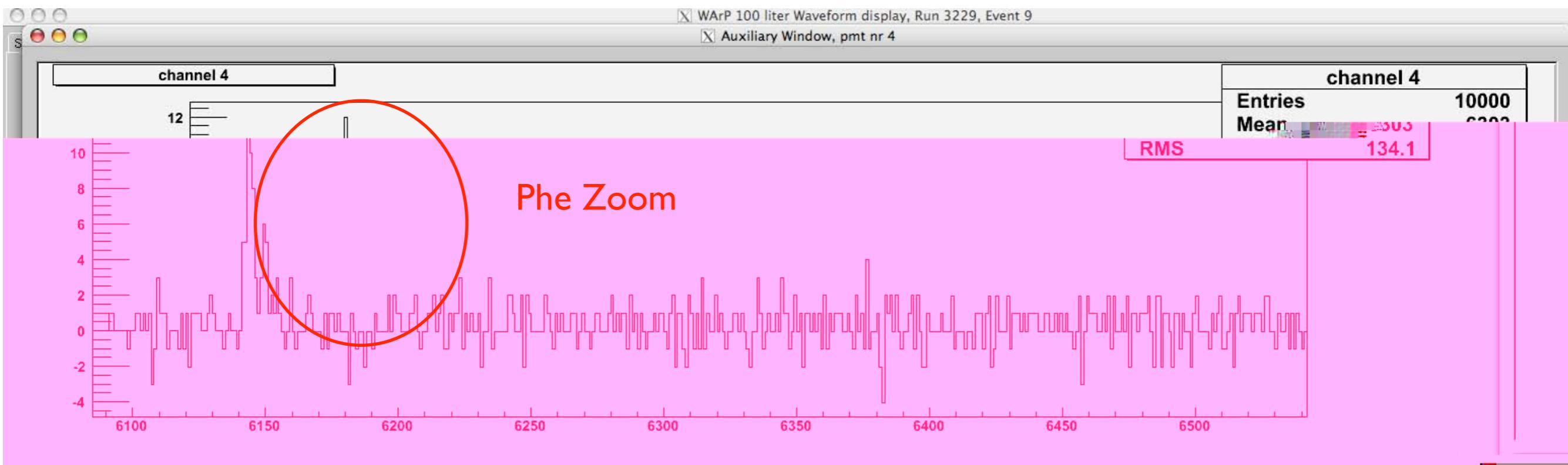
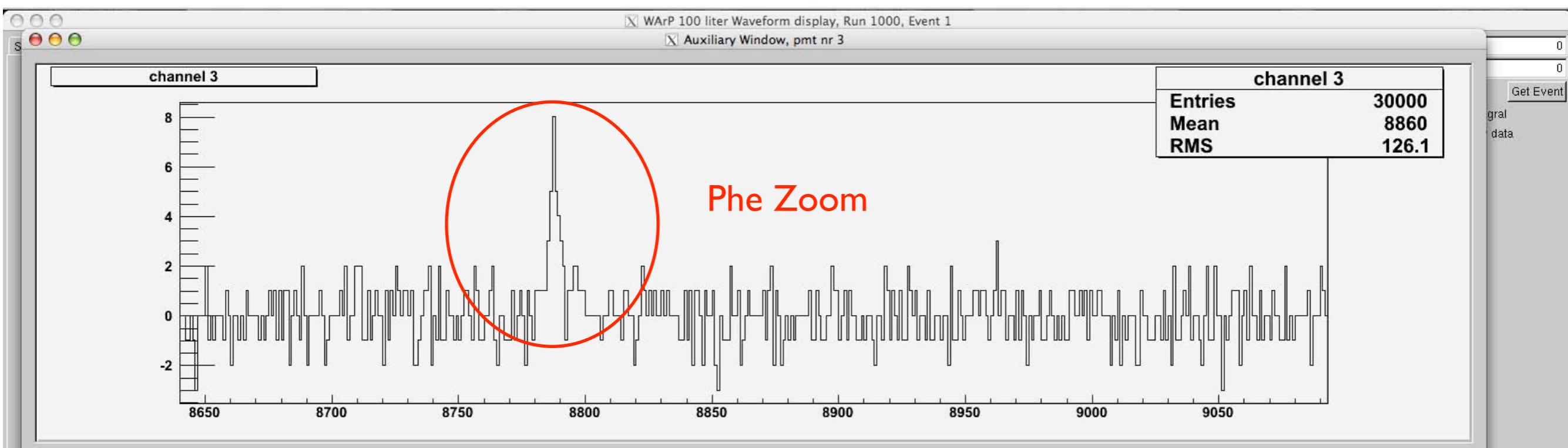


SER fit debugging

SER fitting algorithm search for isolated peaks on the tail of the acquired signal



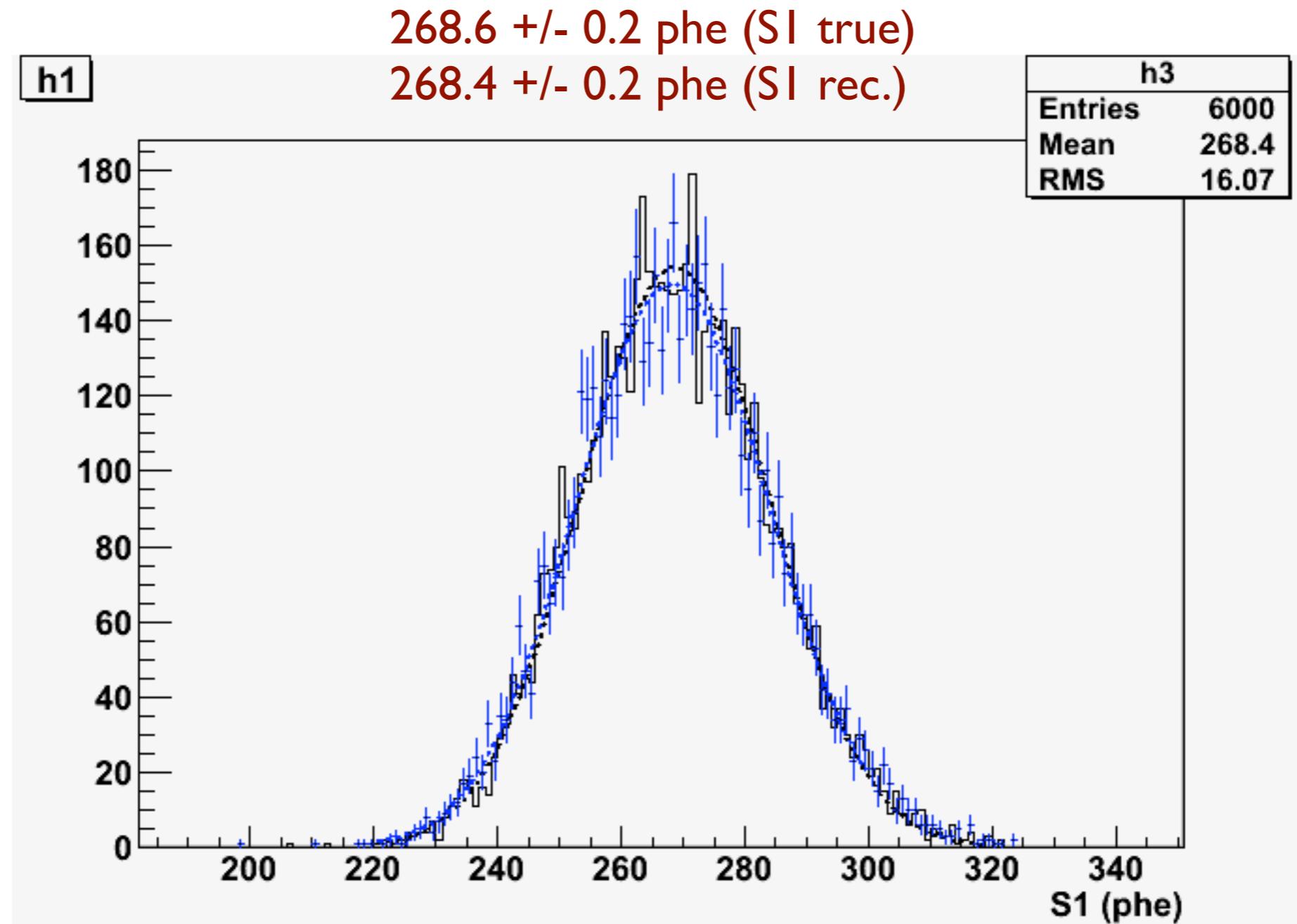
Simulated Vs Real



Software Efficiency

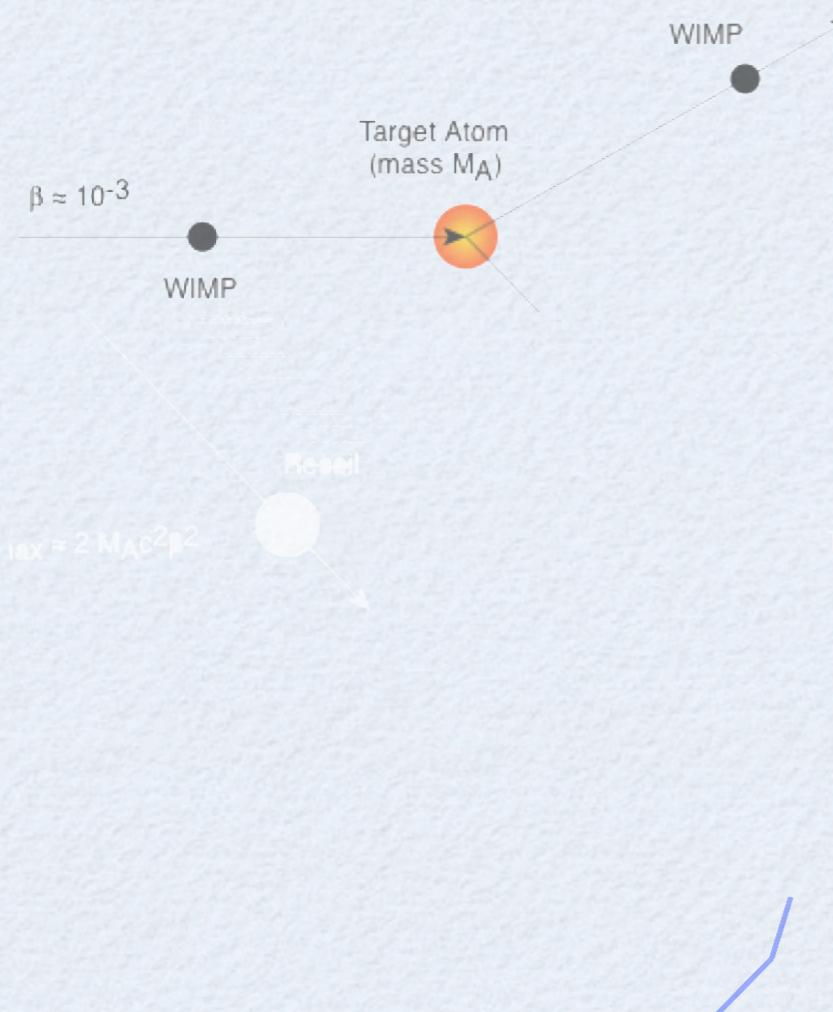
Reconstruction Software efficiency has been estimated:

- monoenergetic primary ($e^- E=50$ keV)
- number of detected photons VS reconstructed SI



Conclusions

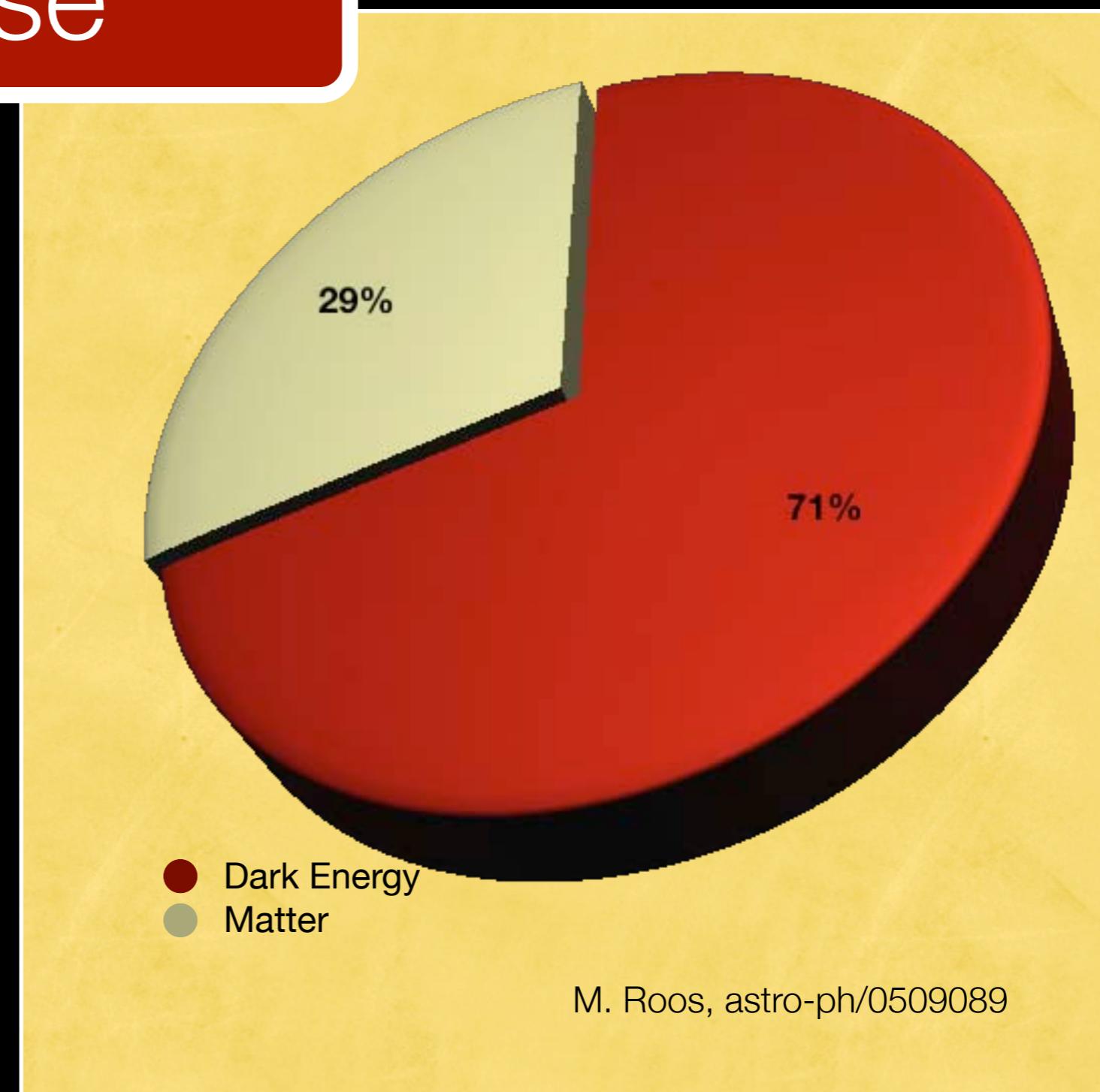
- A GEANT4 simulation of the optical response of the WArP inner Detector has been implemented
- visible & NUV Optical Properties Measurements done at ENEA labs
- UV measurements to be done in Nov09-Dic09
- Study of detector efficiencies done
- Electronics response simulated with a dedicated C++ code to obtain events like raw data
- Analysis Code debugging (preliminary)
- Preliminary results show that the reconstruction algorithm works well but
 - to be implemented Experimental noise
 - a more detailed analysis will be done



Backup

Universe

- CMB
- Supernovae Ia
- Gravitational lensing



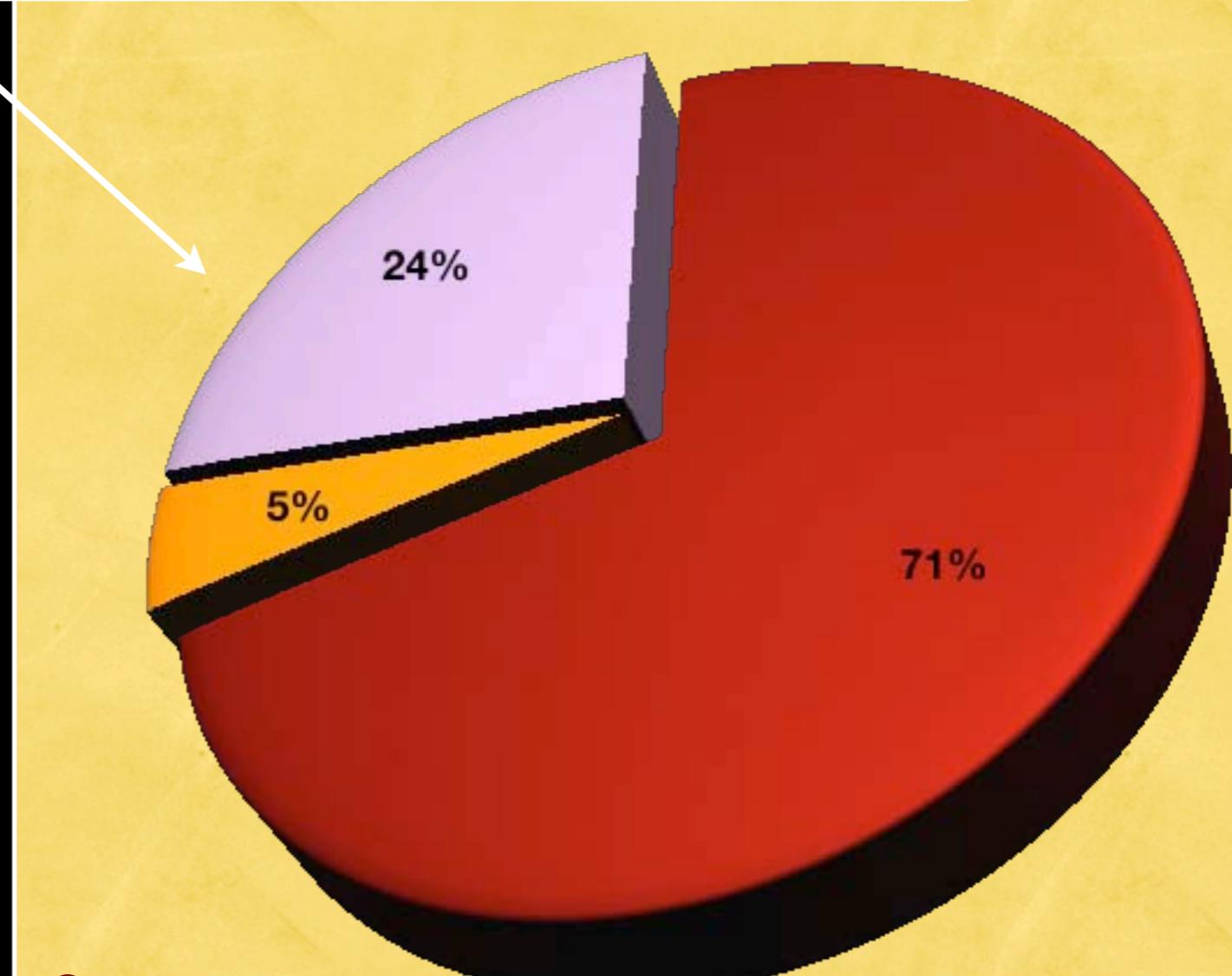
What is Dark Matter made of?

- Non-Baryonic
- Neutral
- Relic
- Cold

SUSY extension of standard model provides a Good candidate

WIMP

Dark Matter



● Dark Energy
● Baryons
● Dark Matter

M. Roos, astro-ph/0509089

Substrate Characterizations

In order to optimize the light collection system dedicated measurements have been done

Measured samples can be divided as follows:

- TPB evaporated over glass
- TPB evaporated over VM2000
- TPB-Polystyrene mixture over glass
- VM2000

Measurements done*:

- Photoluminescence (PL)
- Photoluminescence Excitation (PLE)
- Direct Transmittance
- Hemispheric Transmittance & Reflectance

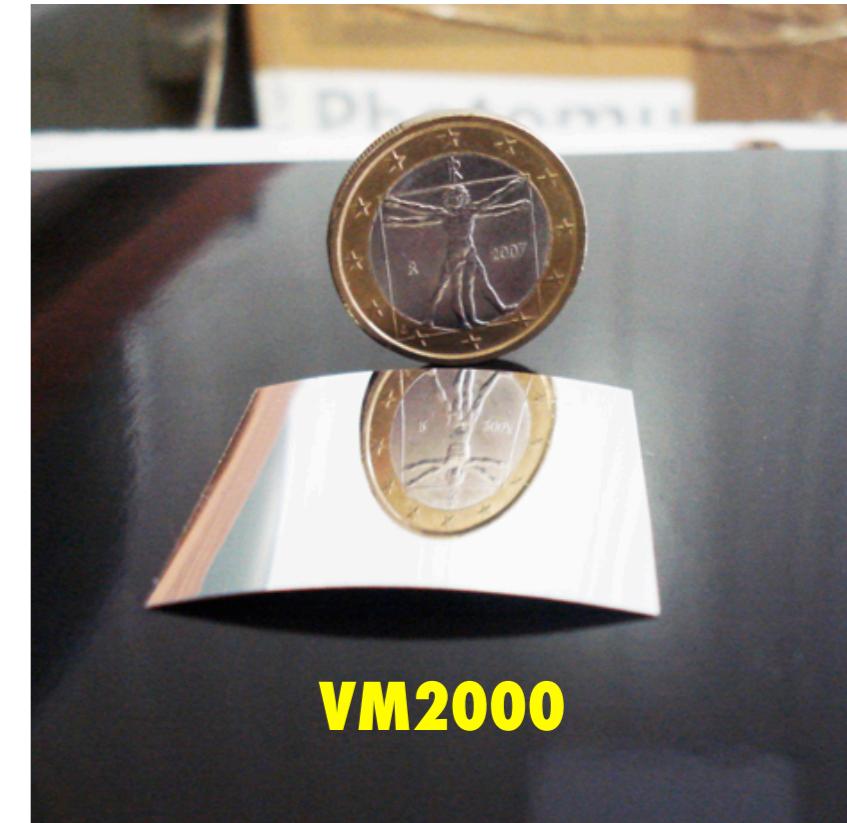
* in collaboration with ENEA:

R.M. Montereali and M.A. Vincenti

Dip. Tecnologie Fisiche e Nuovi Materiali, FIM-FISACC, ENEA C.R. Frascati

E. Nichelatti

Dip. Tecnologie Fisiche e Nuovi Materiali, FIM-FISOTT, ENEA C.R. Casaccia

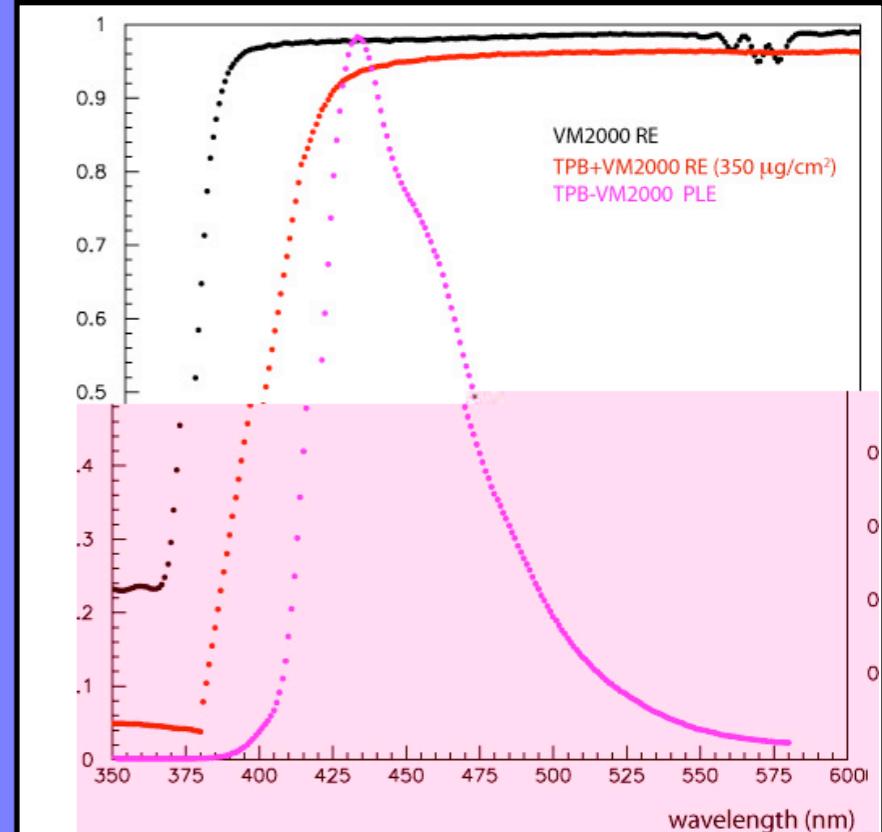


Boundary Processes Implementation

Evaporated TPB on VM2000

- A much diffusive substrate (TPB) on a Mirror (slide...)
- Layer RINDEX not known
- VM200 has almost null T: fully characterized by R

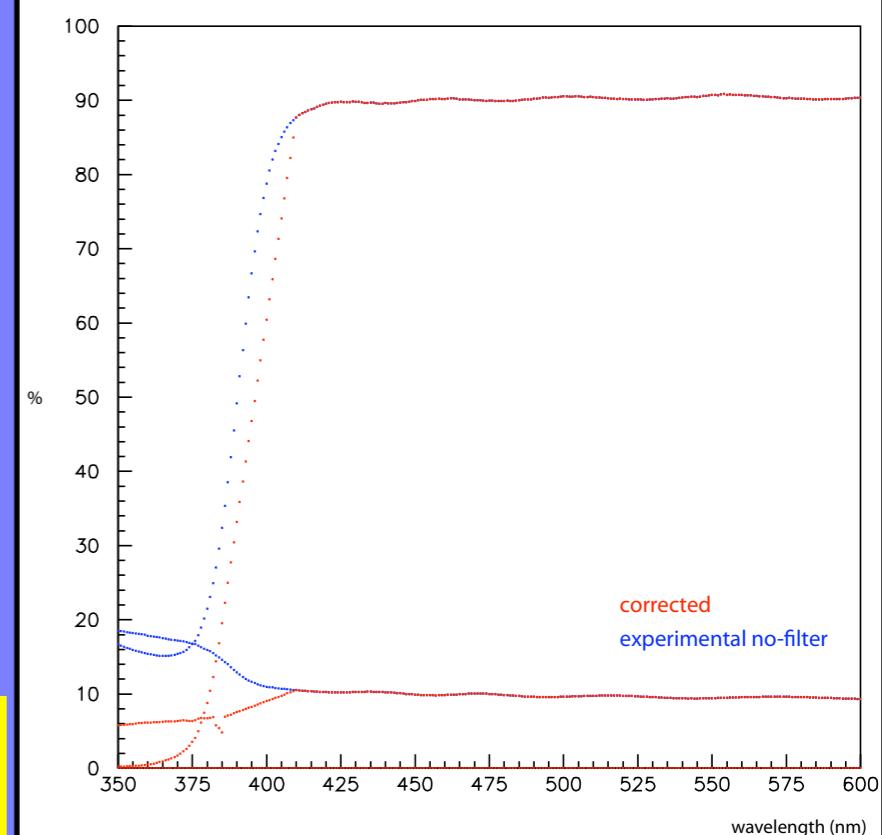
- A much diffusive substrate (TPB) on a Mirror (slide...)
- Layer RINDEX not known
- VM200 has almost null T: fully characterized by R



TPB Polystyrene Mixture

- Glass coated with TPB-Polystyrene mixture
 - Practically zero Absorption
-
- Used only the Reflectance
 - The absorbed light is collected according to the PMT Q.E. (EFFICIENCY=Q.E.)

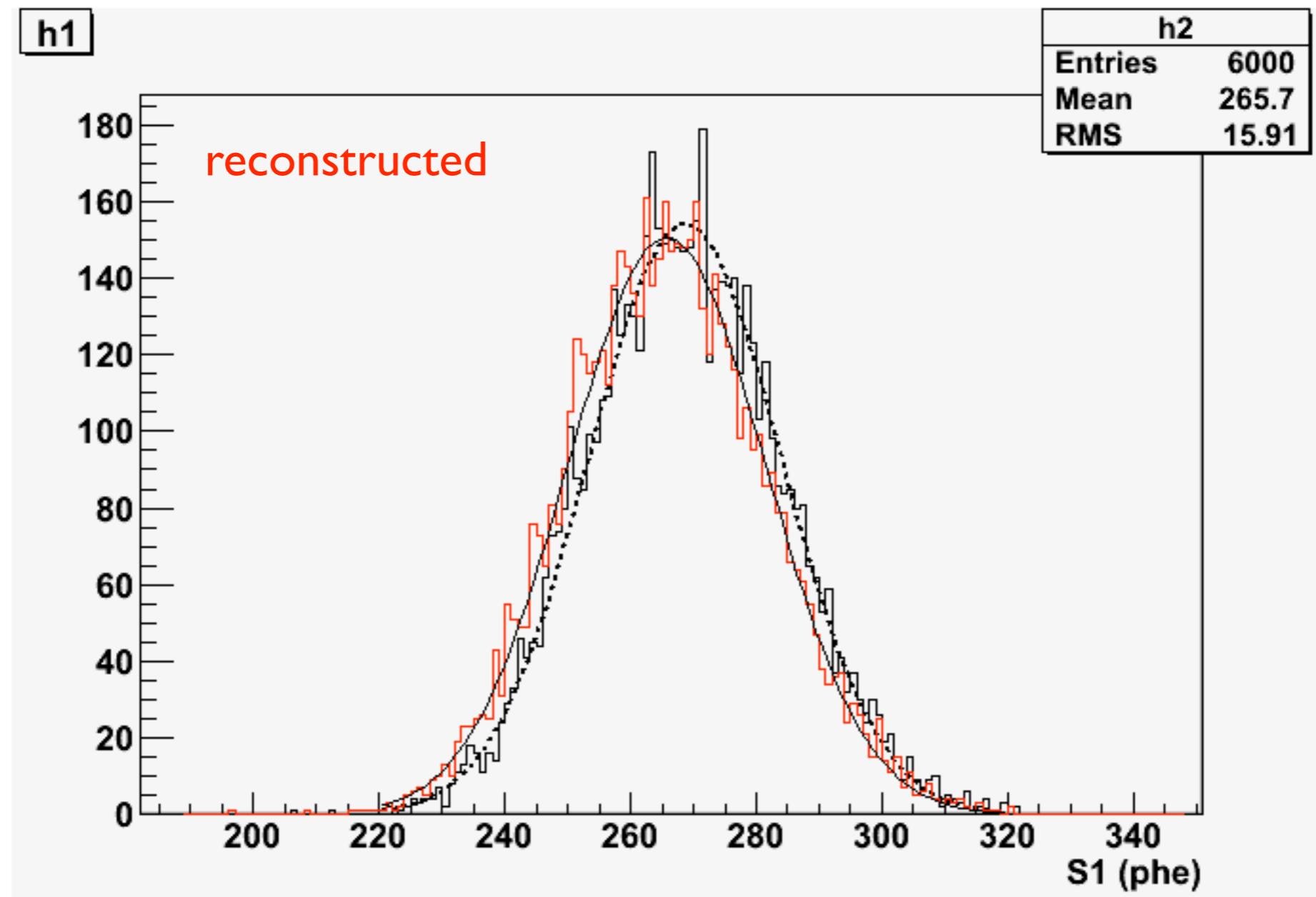
```
OpPMTProperty -> AddProperty("REFLECTIVITY", pp, refPMT, NUM);  
OpPMTProperty -> AddProperty("EFFICIENCY", pp, effiPMT, NUM);
```



Reconstruction Software Efficiency

Rec. Software efficiency has been estimated:

- monoenergetic primary ($e^- E=50$ keV)
- number of detected photons VS reconstructed SI

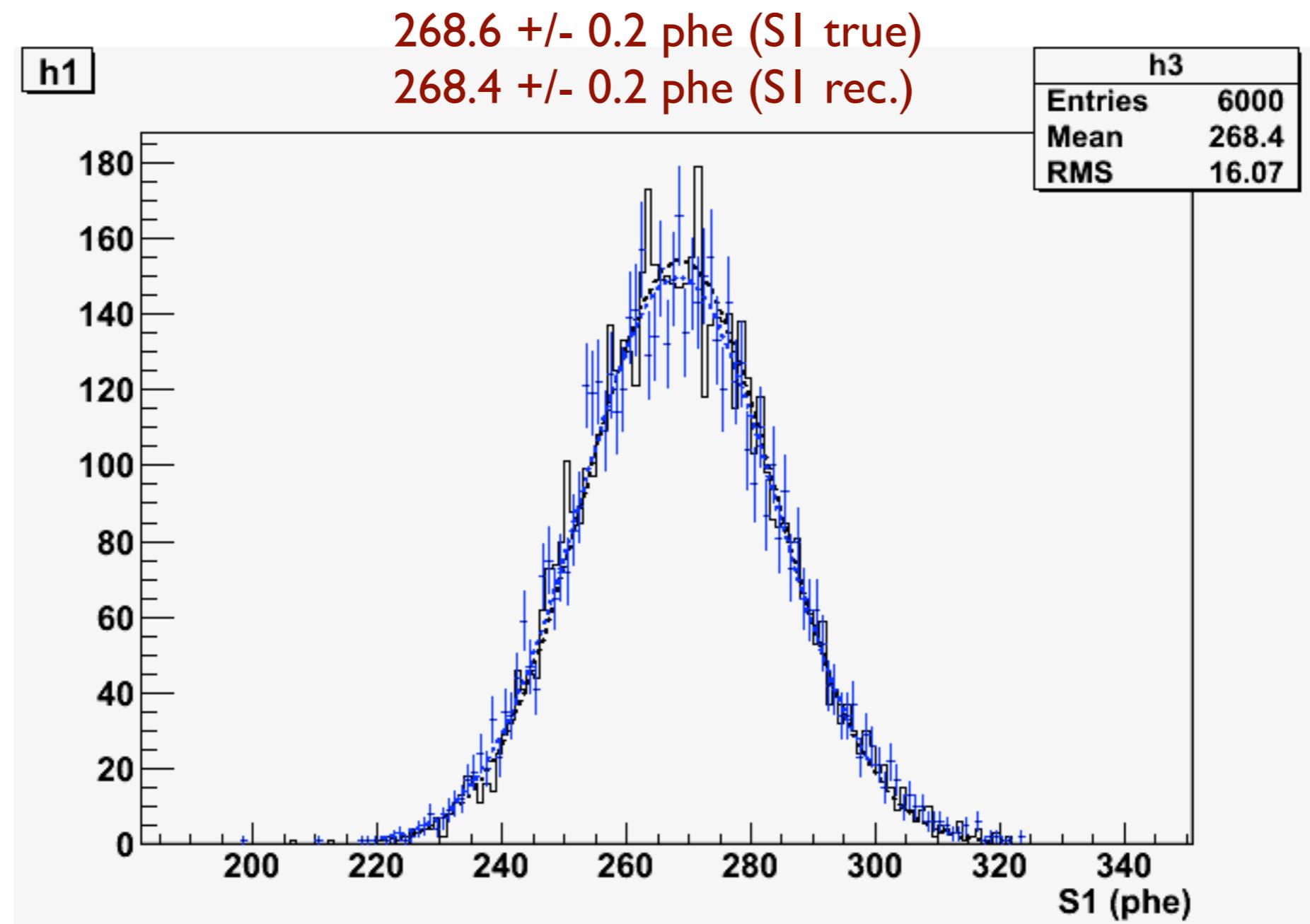


systematic underestimation of the order of 1%

Software Efficiency

It is due to the fact that the SI integral is estimated in a time window of 7 μ s

- the mean fraction of light lost is almost 1%
- applying the expected correction factor ($1 - 0.75 \cdot \exp(-7/1.6)$) the two spectra are well superimposed



Liquid Argon scintillation light emission

An interaction in argon produces Atomic

✓ excitation
✓ ionization

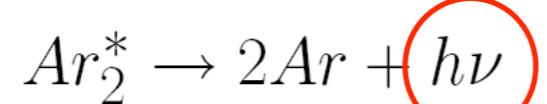


emission of 128 nm luminescence through 2 processes

self trapped exciton luminescence

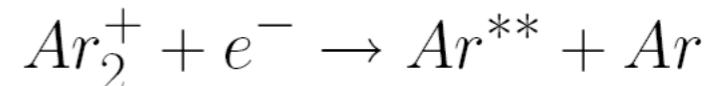
1)

$Ar^* :$



2)

$Ar^+ :$



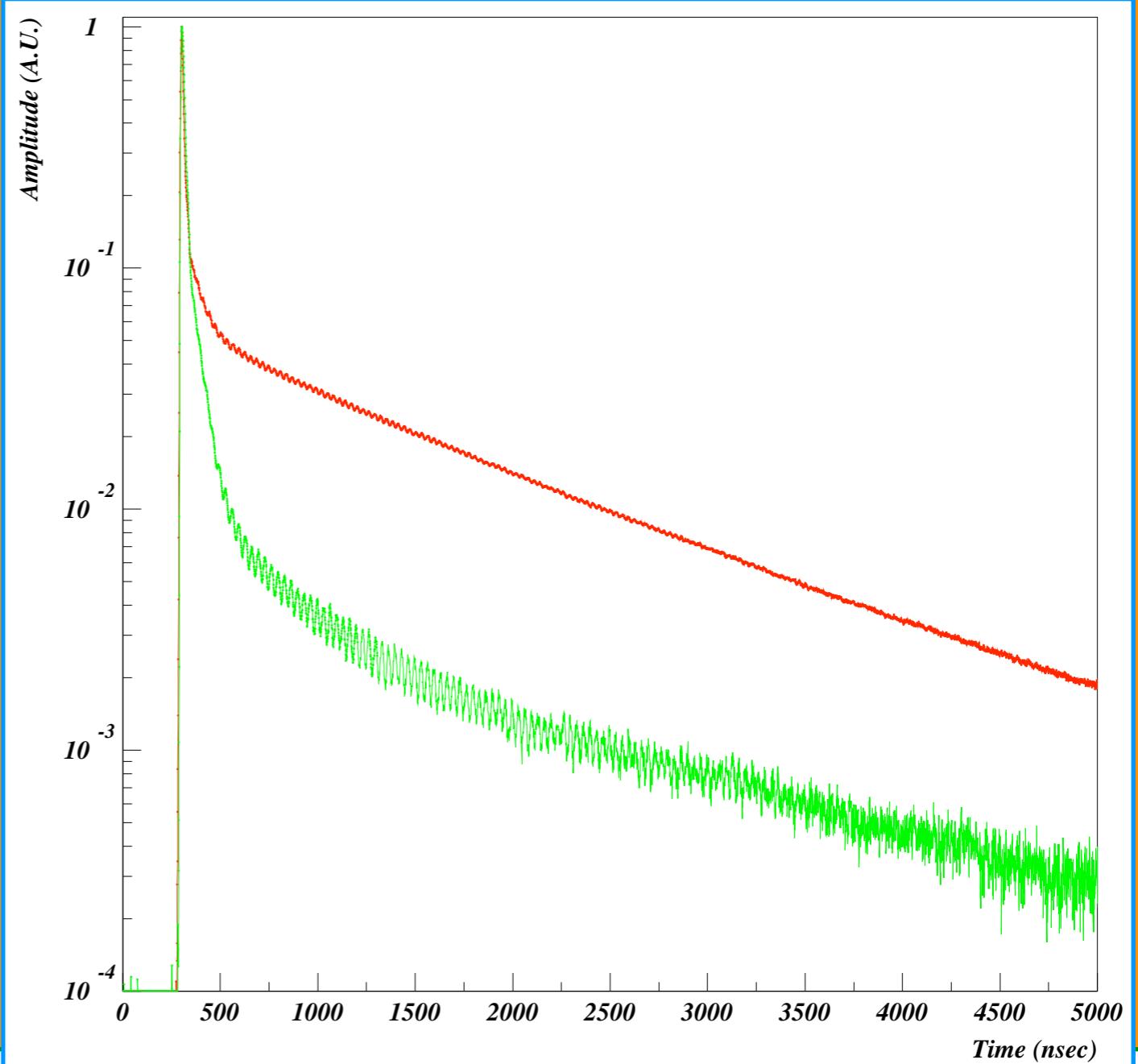
both processes

- ✓ ending up with the same radiative reaction
- ✓ inducing the emission of a 128 nm UV photon



Effect of ionization density on time dependence of luminescence

1. decay constants are not functions of ionization density
2. for high ionization density the fast components intensity increases at expense of the slow one



Evaporation Procedure



- Internal Diameter 63 cm
- Volume 186 lt
- Evacuation time \approx 30 min.
- heater filament: tantalum foil
- graphite furnace temperature monitored by a thermocouple

Evaporation procedure is done @ LNGS by members of the WArP collaboration

Experimental Setup

HeT and HeR measures

- Spectrophotometer UV-Vis-NIR Perkin-Elmer mod. Lambda 19 equipped with a integrating sphere (150 mm diameter).
- Wavelength experimental window: 200-600 nm (reduced to 250-600 nm).
- Filter: bandpass filter (250-390 nm), thickness 2.1 mm.

PL and PLE measures

- Spectrophotometer UV-Vis-NIR Jobin Yvon Fluorolog-3 FL-1

DT measures

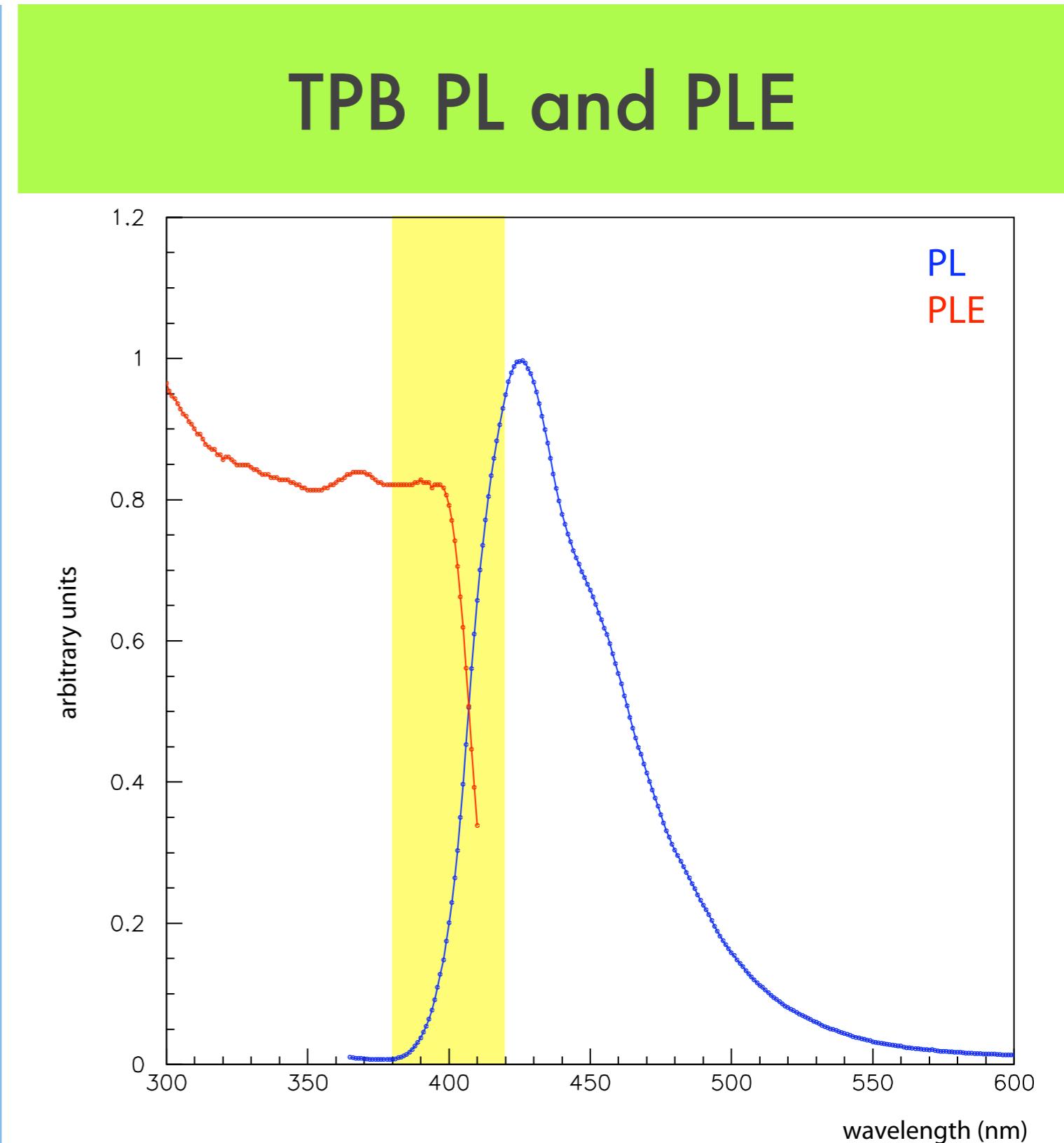
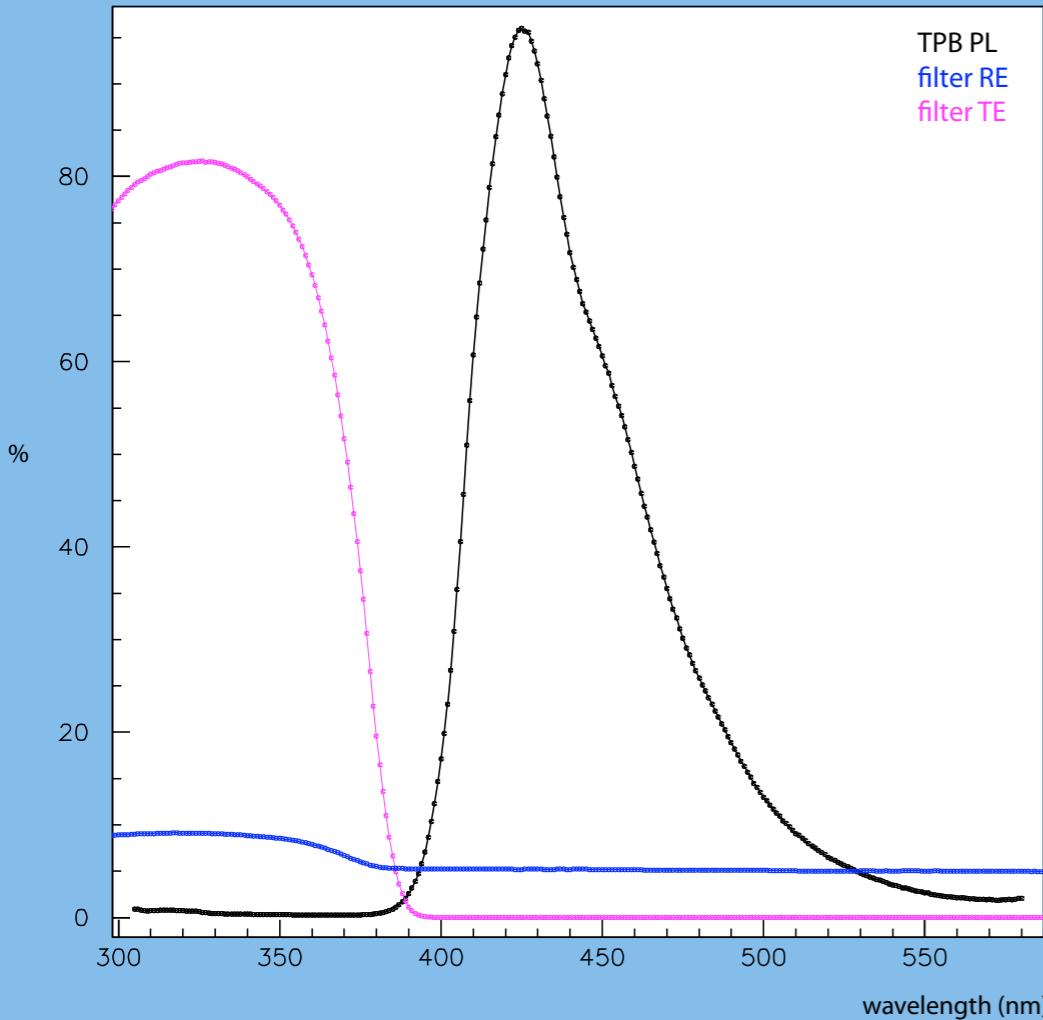
- Spectrophotometer UV-Vis-NIR Perkin-Elmer mod. Lambda 19
- Wavelength experimental window: 200-1400 nm

TPB PL and PLE

- TPB PLE is very large up to 420 nm
- 390-420 nm: TPB PL and PLE overlaps
- VM2000 exhibit a similar behavior

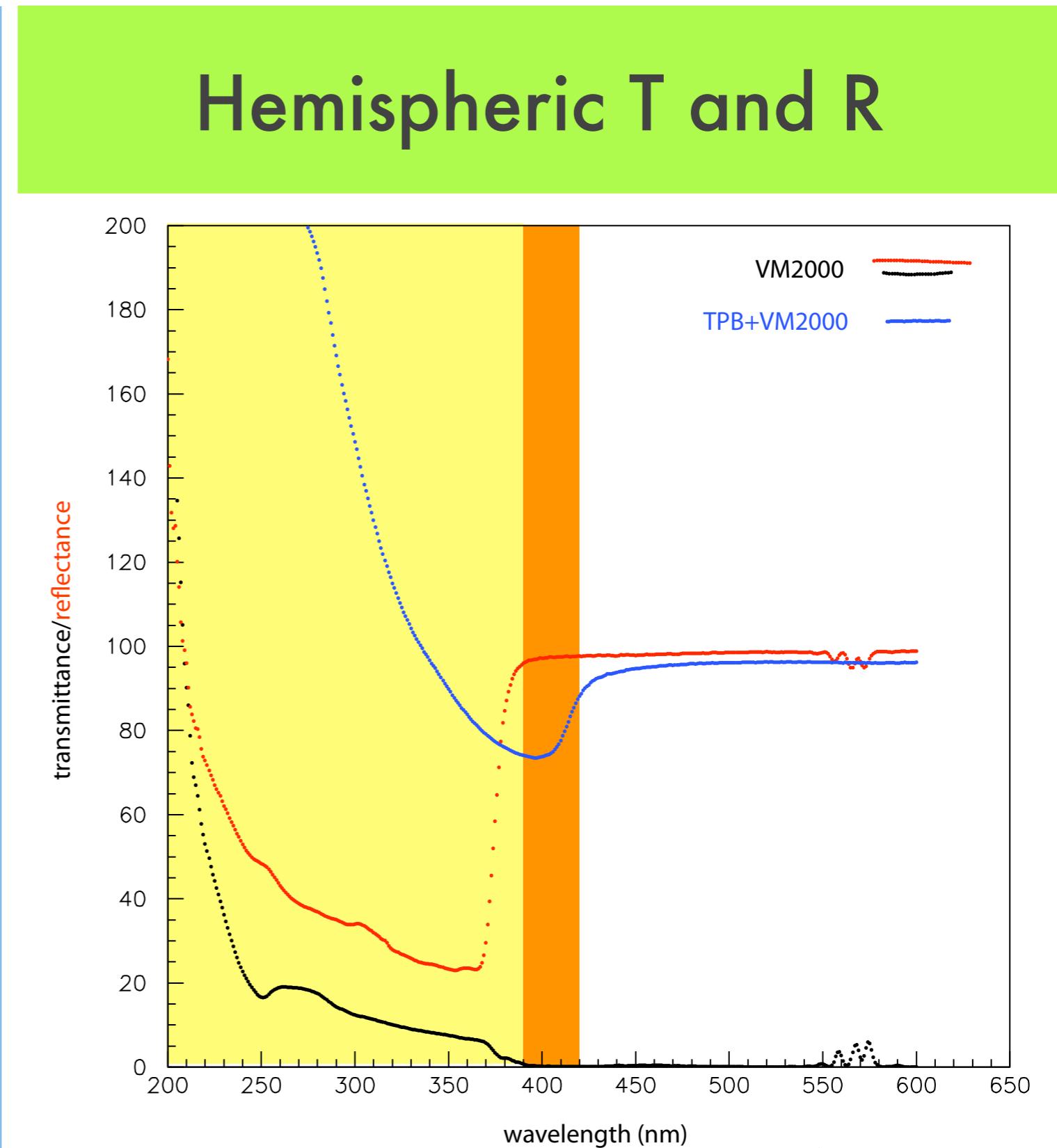
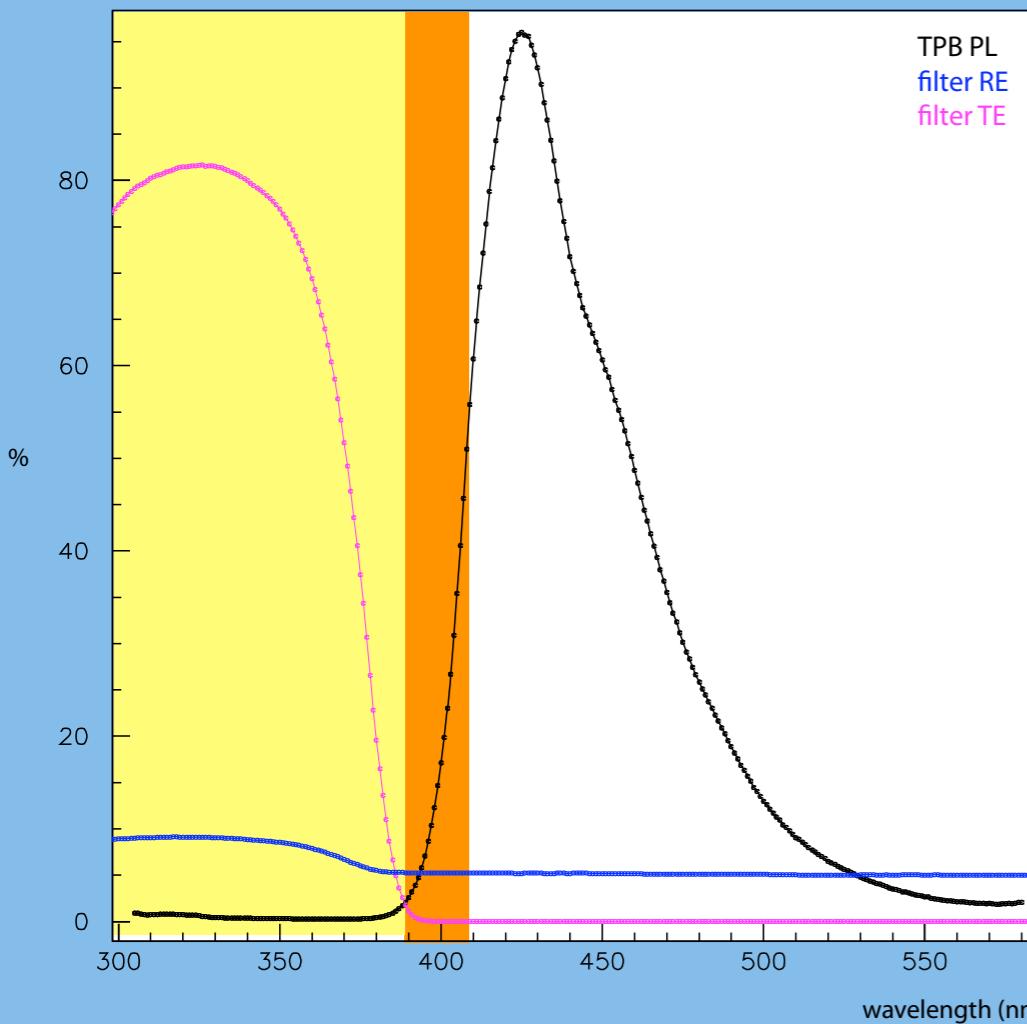


Need: filter to remove spurious PL



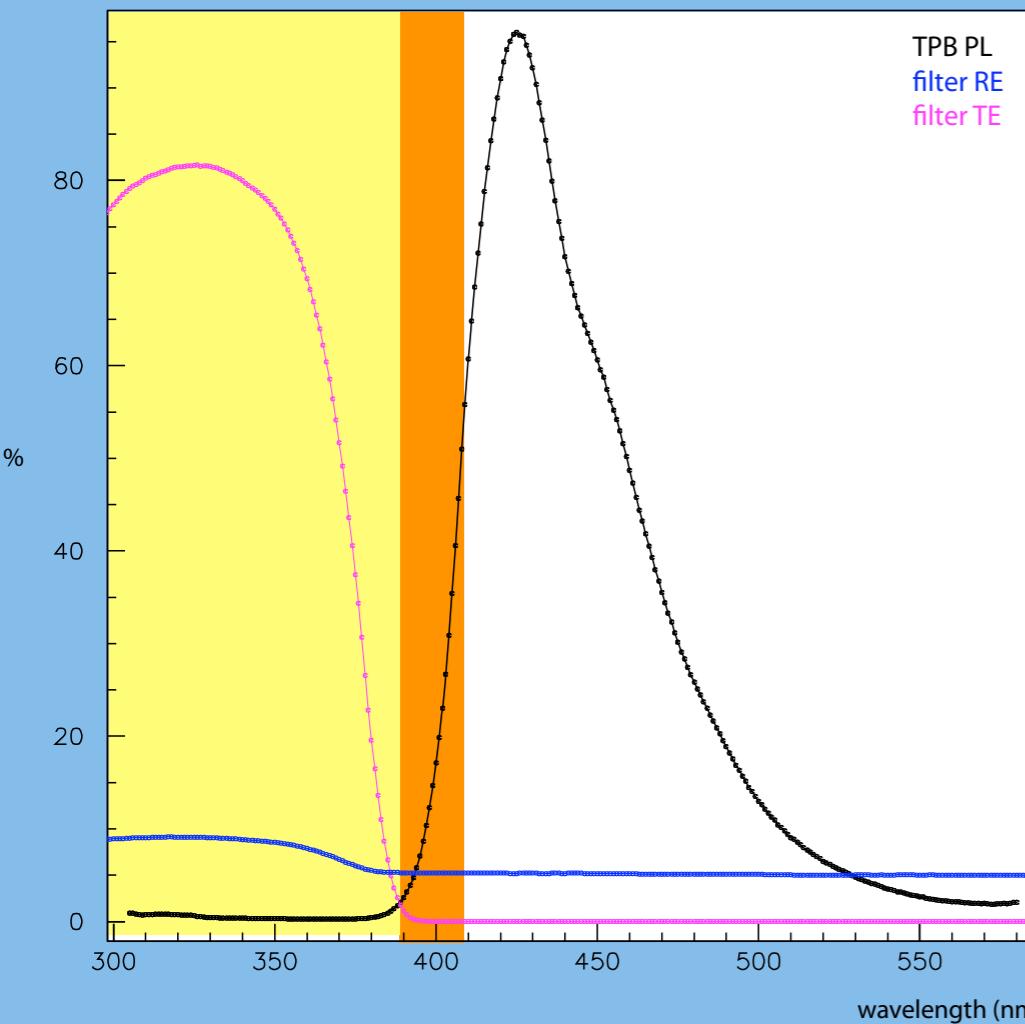
Evaporated TPB on VM2000

- PL affects strongly hemispheric properties measurements
- pass band filter (250-390 nm)
- Linear interpolation (390-420 nm)

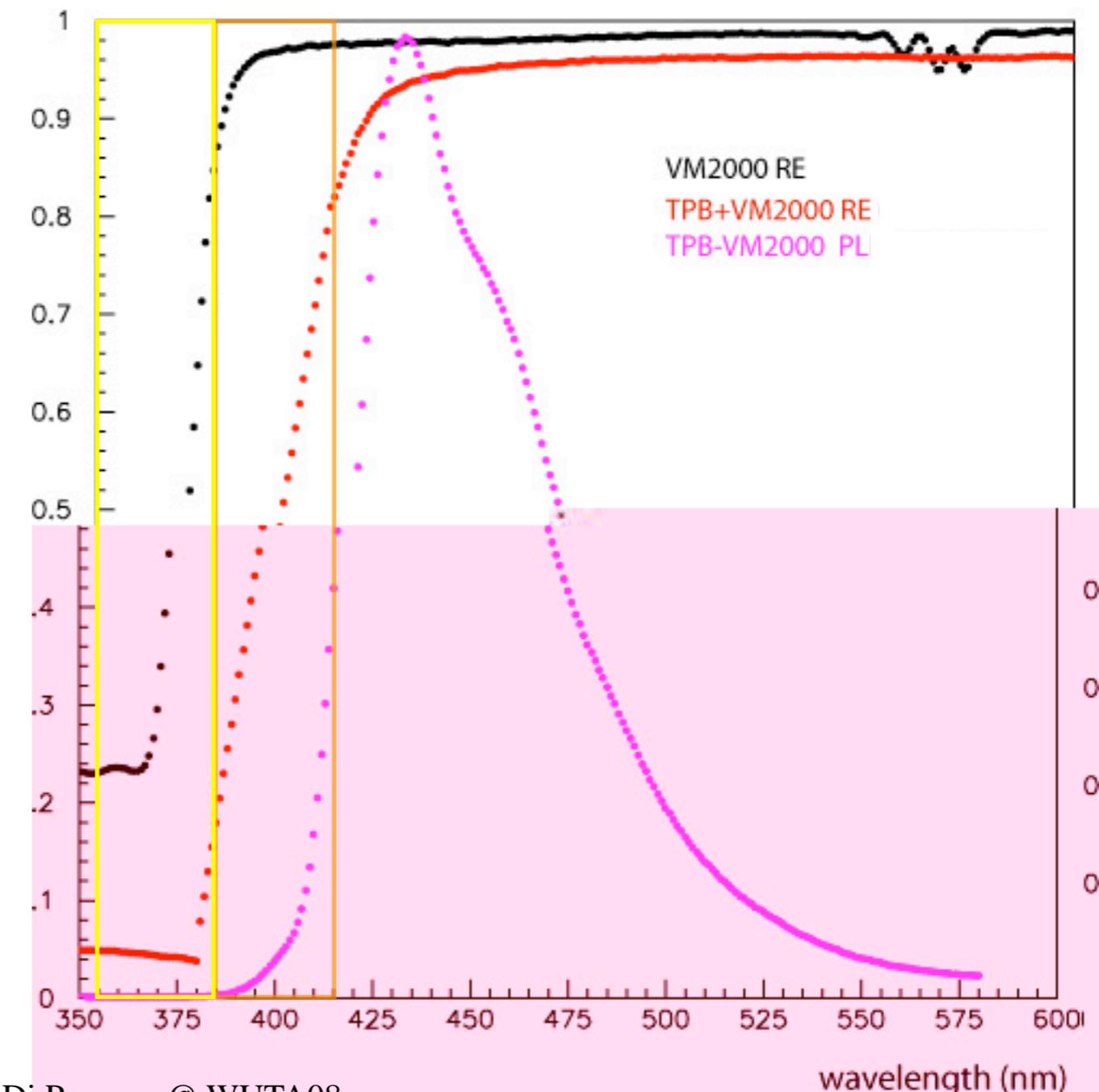


Evaporated TPB on VM2000

- PL affects strongly hemispheric properties measurements
- pass band filter (250-390 nm)
- Linear interpolation (390-420 nm)

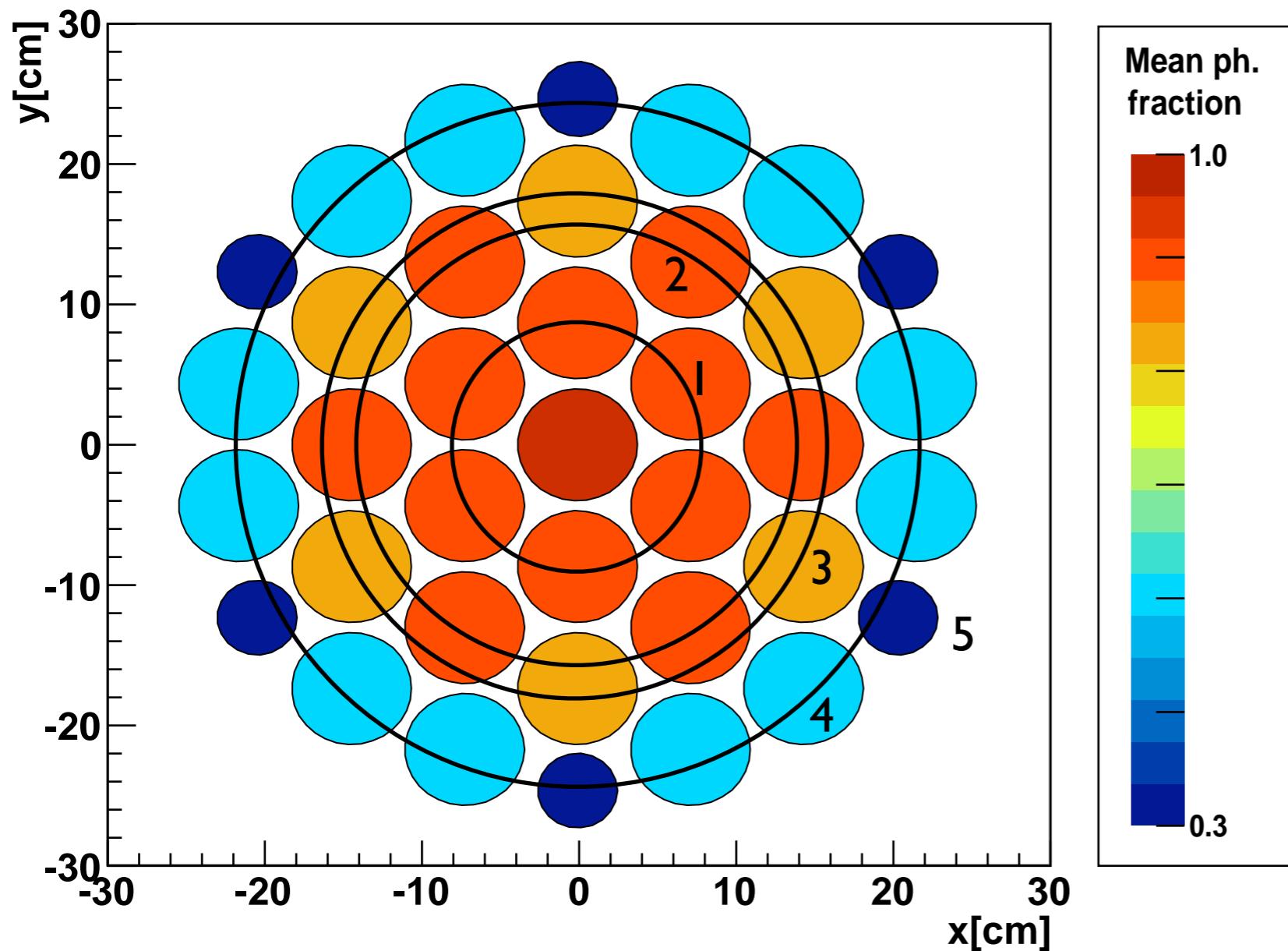


TPB-VM2000 sample



hits distribution on PMTs

- The number of photon detected by each phototube has been considered
- effective sensible surfaces for 3inch and 2inch ($r=?$ $r=?$)
- The considered quantity is the relative efficiency, normalized to the most efficient PMT



- Most efficient is the PMT in the center
- Clear dependence on R
- 2inch PMTs have also smaller photocatode surface

Efficiency ring values:

- 1.
- 2.
- 3.
- 4.
- 5.