## EP20BTECH11015-Assignment-5

February 20, 2023

## 0.1 EP20BTECH11015 - ASSIGNMENT 3

```
[]: import numpy as np
import matplotlib.pyplot as plt
import scipy.stats as st
from scipy.optimize import curve_fit
from scipy import optimize as opt
import sklearn as skl
import pandas as pd
```

1.

Download the asteroid dataset from http://astrostatistics.psu.edu/datasets/asteroid\_dens.dat. Apply the Shapiro-Wilk test to both the asteroid density values and the natural logarithm of the density values.

From the p values, which of these is closer to a Gaussian distribution?

Verify this by plotting histograms of both density and its logarithm and overlaying the best-fit normal distribution (Look up stats.norm.fit) (25 points)

```
[]: aster = pd.read_csv('asteroid_dens.dat', sep=' ')
shapiro_val = st.shapiro(aster['Dens'])
shapiro_ln_val = st.shapiro(np.log(aster['Dens']))

print('p-value of Shapiro test for density: ', shapiro_val.pvalue)
print('p-value of Shapiro test for log density: ', shapiro_ln_val.pvalue)
#print('Normal Fit',*st.norm.fit(aster['Dens'], loc=shapiro_val.statistic))
```

```
p-value of Shapiro test for density: 0.051220282912254333 p-value of Shapiro test for log density: 0.5660613775253296
```

From the p-values, we can see that the natural logarithm of the density values is closer to a Gaussian distribution.

2. Download the Hipparcos star catalog from http://iith.ac.in/~shantanud/HIP\_star.dat. Detailed explanation of the columns in this dataset can be found in http://astrostatistics.psu.edu/datasets/HIP\_star.html under "Dataset".

Calculate using two-sample t-test whether the color (B- V) of the Hyades stars differs from the non-Hyades ones.

The Hyades stars have Right Ascension between  $50 \circ$  and  $100 \circ$ , declinations between 0 and  $25 \circ$ , proper motion in RA between 90 and 130 mas/year, proper motion in DEC between -60 and -10 mas/year.

Any other star which does not satisfy any of the above conditions is considered a non-Hyades star. (25 points)

Mean of B-V for stars in Haydes cluster: 0.6208924731182797 Mean of B-V for stars not in Haydes cluster: 0.7546207159177456 Means are close but not equal

## NULL HYPOTHESIS: The color of the Hyades stars is same as the non-Hyades stars.

```
p-value: 0.00019212855310889943
p-value is smaller than 0.05, so we reject the null hypothesis
I.e The colors of haydes and non-haydes stars are DIFFERENT
3.
```

The T90 distribution for Beppo-Sax T90 data can be found at http://www.iith.ac.in/~shantanud/beppoSax.txt.

Apply GMM to log10 of T90 data and find the optimum number of components using AIC and BIC by plotting BIC as a function of number of componts (20 points)

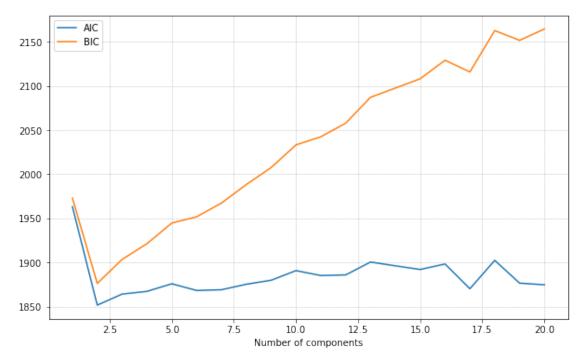
(Hint: Look at the source code for astroML figure 6.6)

```
[]: beppo = np.loadtxt('beppoSax.txt').reshape(-1, 1)
    log10beppo = np.log10(beppo)

gmm_fit = [skl.mixture.GaussianMixture(i).fit(log10beppo) for i in range(1, 21)]
    aic = [i.aic(log10beppo) for i in gmm_fit]
    bic = [i.bic(log10beppo) for i in gmm_fit]

#Plotting AIC and BIC vs number of components

plt.figure(figsize=(10, 6))
    plt.plot(range(1, 21), aic, label='AIC')
    plt.plot(range(1, 21), bic, label='BIC')
    plt.legend()
    plt.grid(True, ls=':', which='both')
    plt.xlabel('Number of components')
    plt.show()
```



From the graph, we see that the optimum number of components is 2.