

$$\overrightarrow{\times}(t) = \overrightarrow{R}'(t) + \overrightarrow{\times}'(t)$$
 (0)

$$\overrightarrow{\nabla}(t) = \overrightarrow{R}(t) + \overrightarrow{\nabla}'(t) \tag{1}$$

$$\overrightarrow{\vartheta}(t) = \overrightarrow{R}(t) + \overrightarrow{\vartheta}'(t) \tag{2}$$

From (2), the accelerations experienced by the system o' are given by:

$$\vec{\vartheta}'(t) = \vec{\vartheta}(t) - \ddot{\vec{R}}(t) \tag{3}$$

Acceleration
"felt" by the
accelerometer
in 0'

Accoleration present in O (typically gravity)

Relative acceleration butween systems

Examples:

1. Galilean relativity principle, or why an accelerometer cannot directly measure speeds.

If  $\overrightarrow{R}(t)$  is not accelerated (for instance:  $\overrightarrow{R}(t) = \overrightarrow{R}_0 + \overrightarrow{V}_0 t$ ), then  $\overrightarrow{R}(t) = \overrightarrow{O}$  and, from (3),  $\overrightarrow{O}'(t) = \overrightarrow{O}'(t)$ . So both O and O' "experience" the same physics, and thus the speed  $\overrightarrow{V}_0$  is "invisible" for the accelerameter in O'.

2. Free-follers experience no occeleration.

Gravity is already present in 0, so  $\overrightarrow{\partial}(t) = \overrightarrow{g} = \begin{bmatrix} 0 \\ -9 \end{bmatrix}$ A free-folling object accelerates with  $\overrightarrow{g}$ , so  $\overrightarrow{R}(t) = \overrightarrow{g}$  too. From (3):  $\overrightarrow{\partial}'(t) = \overrightarrow{O}$ 

3. Centri fugal force

If O' moves in a circle around O, with

constant appeal, we have:

$$R(t) = \begin{bmatrix} R_0 \cos(\omega t) \\ R_0 \sin(\omega t) \end{bmatrix} \Rightarrow R(t) = \begin{bmatrix} -R_0 \omega^2 \cos(\omega t) \\ -R_0 \omega^2 \sin(\omega t) \end{bmatrix} = -\omega^2 R(t)$$

$$\overrightarrow{R}'(t) = \begin{bmatrix} R_0 \cos(\omega t) \\ R_0 \sin(\omega t) \end{bmatrix} \Rightarrow \overrightarrow{R}'(t) = \begin{bmatrix} -R_0 \omega^2 \cos(\omega t) \\ -R_0 \omega^2 \sin(\omega t) \end{bmatrix} = -\omega^2 \overrightarrow{R}'(t)$$

: (4)

(5)

From (2), the accelerations experienced by the system O Using (3) again:

(E) 
$$\overrightarrow{\partial}'(t) = \overrightarrow{\partial}'(t) + \omega^2 \overrightarrow{R}'(t)$$

The centrifugal term (1) (2)

de see that O'experiences" an extra acceleration.

present in O geoderation Using the relation between linear land plangular speed in a circular movement:

equation (4) takes the more familiar form:

$$\frac{\partial}{\partial t}(t) = \frac{\partial}{\partial t}(t) + \frac{v^2}{R_0} \frac{\overrightarrow{R}(t)}{R_0} = \frac{\partial}{\partial t}(t) + \frac{v^2}{R_0} \stackrel{\text{distinct of }}{R_0} \stackrel{\text{distinct of }$$

If only gravity and rotation apply: (8) not the 5 = (1) \$

$$\frac{1}{3}(t) = \begin{bmatrix} 0 \\ -9 \end{bmatrix} + \frac{\sqrt{2}}{R_0} \begin{bmatrix} \cos(\omega t) \\ \sin(\omega t) \end{bmatrix}$$

2. Free follers experience no occaleration. Gravity is already present in 0, so at (t) = = = = = A free- folking object occelerates with 3, so R'(t) = 3 tou.

From (3): 34(E) = 0

All the previous equations still hold

or as long as cie with the exception of (7),

that requires a clarification about the

coordinates used:

$$\frac{1}{2}(t) = \begin{bmatrix} 0 \\ -9 \end{bmatrix} + \begin{bmatrix} \cos(\omega t) \\ \cos(\omega t) \end{bmatrix}$$

$$\frac{\cos(\omega t)}{\cos(\omega t)}$$

The coordinate systems of both frames of reference are related

by:  $\begin{bmatrix} \times' \\ y' \end{bmatrix} = \begin{bmatrix} \cos(\omega t) & \sin(\omega t) \\ -\sin(\omega t) & \cos(\omega t) \end{bmatrix} \begin{bmatrix} \times \\ y \end{bmatrix} = \Omega(t) \begin{bmatrix} \times \\ y \end{bmatrix}$ 

$$\int_{0}^{\infty} = \Omega(t) \begin{bmatrix} \times \\ y \end{bmatrix}_{0}^{(8)}$$

So we con express (7.1) in 0' coordinates as:

$$\frac{\partial}{\partial t}(t) = \Omega(t) \begin{bmatrix} 0 \\ -9 \end{bmatrix}_0 + \frac{v^2}{R_0} \begin{bmatrix} \cos(ut) \\ \sin(ut) \end{bmatrix}_0 =$$

$$= 9 \begin{bmatrix} -\sin(ut) \\ -\cos(ut) \end{bmatrix}_0 + \frac{v^2}{R_0} \begin{bmatrix} 1 \\ 0 \end{bmatrix}_0$$
(7.2)

Summarizing:

If our accelerometer moves according to R(t) and rotates according to 12(t), from (3) and (8) we can conclude that it will read:

$$\begin{bmatrix} \frac{\partial x}{\partial y} \\ \frac{\partial y}{\partial z} \end{bmatrix}_{0} = \Omega(t) \begin{bmatrix} \frac{\partial x}{\partial y} \\ \frac{\partial y}{\partial z} \\ \frac{\partial z}{\partial z} \end{bmatrix}_{0} \begin{bmatrix} \frac{\partial x}{\partial z} \\ \frac{\partial z}{\partial z} \\ \frac{\partial z}{\partial z} \end{bmatrix}_{0}$$
(2)

as registered by

Accolerations Accoleration Relative present in O acceleration the accelerometer growity) and to motive stations

Trick: wait a second. Aven't rotations a pain in the \*\*s? Is there any way around it? (Luckily yes hall (du) 200 (du) 12-

We can exploit the fact that rotation matrices are orthogonal:

$$\Omega^{t}(t) = \Omega^{-1}(t) \Longrightarrow \Omega^{t}(t) \Omega(t) = 1$$

and apply the norm (to (2): 07 + (0) (1) = (3) (5)

$$||\vec{a}'||^2 = ||\vec{a}|^2 - ||\vec{R}||^2 - ||\vec{a}||^2 + ||\vec{R}||^2 - 2\vec{a}\cdot\vec{R}$$
 (11)

The price we pay is that we destroy the information about orientation, but we keep the magnitudes.