# A User's Guide to Descriptors Release 1.0

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# What are Descriptors?

Descriptors are designed such that essential keyword values that describe a particular concept can be accessed from the headers of a given dataset in a consistent manner, regardless of which instrument was used to obtain the data. This is particularly useful for Gemini data, since the majority of keywords used to describe a particular concept at Gemini are not uniform between the instruments.

# **Basic Descriptor Usage**

The command typewalk -1 lists all available descriptors. As of August 3, 2011, there are 61 descriptors available (*Appendix A*).

The following commands show an example of how to use descriptors and can be entered at an interactive Python prompt (e.g., ipython, pyraf):

```
>>> from astrodata import AstroData
# Load the fits file into AstroData
>>> ad = AstroData("N20091027S0137.fits")
# Count the number of science extensions in the AstroData object
>>> ad.count_exts(extname="SCI")
# Get the airmass value using the airmass descriptor
>>> airmass = ad.airmass()
>>> print airmass
# Get the instrument name using the instrument descriptor
>>> print "My instrument is %s" % ad.instrument()
My instrument is GMOS-N
# Get the gain value for each science extension
>>> for ext in ad["SCI"]:
       print ext.gain()
. . .
2.1
2.337
>>> print ad.gain()
{('SCI', 2): 2.337000000000002, ('SCI', 1): 2.10000000000001,
('SCI', 3): 2.299999999999998}
```

In the examples above, the airmass and instrument apply to the dataset as a whole (the keywords themselves only exist in the PHU) and so only one value is returned. However, the gain applies to the pixel data extensions and so for this AstroData object, three values are returned, since there are three pixel data extensions. In this case, a Python dictionary is used to store the values, where the key of the dictionary is the ("EXTNAME", EXTVER) tuple.

# **Advanced Descriptor Usage**

#### 3.1 DescriptorValue

When a descriptor is called, what is actually returned is a DescriptorValue (DV) object:

```
>>> print type(airmass)
<type 'instance'>
```

Each descriptor has a default Python type defined:

```
>>> print ad.airmass().pytype
<type 'float'>
```

If any of the following operations are applied to the DV object, the DV object is automatically cast to the default Python type for that descriptor:

```
+ - * / // % ** << >> ^ < <= > >= ==
```

For example:

```
>>> print type(airmass*1.0)
<type 'float'>
```

When using operations with a DV object and a numpy object, care must be taken. Consider the following cases:

```
>>> ad[0].data / ad.gain()
>>> ad[0].data / ad.gain().as_pytype()
>>> ad[0].data / 12.345
```

All the above commands return the same result (assuming that ad.gain() = 12.345). However, the first command is extremely slow but the second and third commands are fast. In the first case, since both operands have overloaded operators, the operator from the operand on the left will be used. For some reason, the \_\_\_div\_\_ operator from the numpy object loops over each pixel in the numpy object and uses the DV object as an argument, which is very time consuming. Therefore, the DV object should be cast to an appropriate Python type before using it in an operation with a numpy object.

The descriptor value can be retrieved with the default Python type by using as\_pytype():

```
>>> elevation = ad.elevation().as_pytype()
>>> print elevation
48.6889222222
```

```
>>> print type(elevation)
<type 'float'>
```

The as\_pytype() member function of the DV object should only be required when the DV object can not be automatically cast to it's default Python type. For example, when it is necessary to use an actual descriptor value (e.g., string, float, etc.) rather than a DV object as a key to a Python dictionary, the DV object can be cast to it's default Python type using as\_pytype():

```
>>> my_key = ad.gain_setting.as_pytype()
>>> my_value = my_dict[my_key]
```

If an alternative Python type is required for whatever reason, the DV object can be cast to the appropriate Python type as follows:

```
>>> elevation_as_int = int(ad.elevation())
>>> print elevation_as_int
48
>>> print type(elevation_as_int)
<type 'int'>
```

In the case where a descriptor returns multiple values (one for each pixel data extension), a Python dictionary is used to store the values, where the key of the dictionary is the ("EXTNAME", EXTVER) tuple:

```
>>> print ad.gain()
{('SCI', 2): 2.3370000000000002, ('SCI', 1): 2.100000000000001,
('SCI', 3): 2.29999999999998}
```

A descriptor that is related to the pixel data extensions but has the same value for each pixel data extension will "act" as a single value that has a Python type defined by the default Python type of that descriptor:

```
>>> xbin = ad.detector_x_bin()
>>> print xbin
2
>>> print type(xbin)
<type 'instance'>
>>> print xbin.pytype
<type 'int'>
```

If the original value of the descriptor is required, it can be retrieved by using qet\_value():

```
>>> xbin = ad.detector_x_bin().get_value()
>>> print xbin
{('SCI', 2): 2, ('SCI', 1): 2, ('SCI', 3): 2}
>>> print type(xbin)
<type 'dict'>
>>> print xbin[("SCI", 1)]
2
```

#### 3.2 DescriptorUnits

The DescriptorUnits (DU) object provides a way to access and update the units for a given descriptor value. Basic implementation of this feature has been completed and development is ongoing.

## Writing and Adding New Descriptors

#### 4.1 Introduction to the Gemini Descriptor Code

The Gemini descriptor code is located in the gemini\_python package in the astrodata\_Gemini/ADCONFIG\_Gemini/descriptors directory. When writing and adding new Gemini descriptors, a user / developer will require knowledge of the following files:

- CalculatorInterface.py
- mkCalculatorInterface.py
- StandardDescriptorKeyDict.py
- StandardGenericKeyDict.py
- StandardGEMINIKeyDict.py
- Standard<INSTRUMENT>KeyDict.py
- descriptorDescriptionDict.py
- calculatorIndex.Gemini.py
- Generic\_Descriptor.py
- GEMINI\_Descriptor.py
- <INSTRUMENT> Descriptor.py

#### 4.2 Overview of the Gemini Descriptor Code

When a descriptor is called (as described in the *Basic Descriptor Usage* section and the *Advanced Descriptor Usage*) section, the CalculatorInterface.py file is accessed. This file contains a function for each available descriptor (see *Appendix C* for an example function), which first look directly in the PHU of the AstroData object for the associated descriptor keyword as defined in one of the AstroData standard key dictionary files (Standard<INSTRUMENT>KeyDict.py, StandardGemINIKeyDict.py, StandardGemericKeyDict.py or StandardDescriptorKeyDict.py). The keywords in these dictionaries are combined via inheritance to create a final dictionary containing a single set of keywords associated with a particular AstroData object. The value of the keyword in the header of the data is returned as the value of the descriptor. If the associated descriptor keyword is not found in the header of the data, the descriptor files are then searched in the order below to attempt to find an appropriate descriptor function, which will return the value of the descriptor.

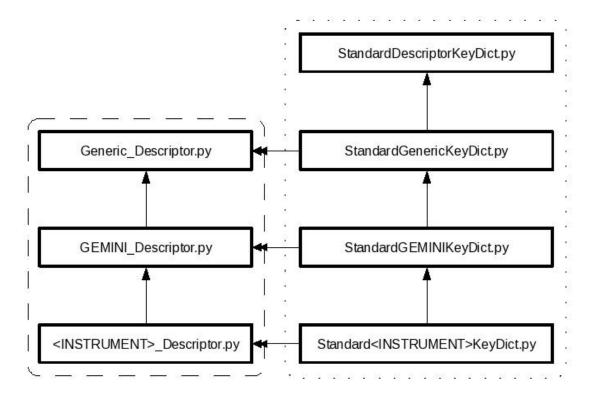


Figure 4.1: An overview showing the relationship between the descriptor files. The files located in the dashed box on the left are the descriptor files that contain descriptor functions. The files located in the dotted box on the right are the keyword files that contain dictionaries describing the one-to-one relationship between a descriptor / variable that is used in the associated descriptor files on the left (as shown by the double headed arrows) and the keyword associated with that descriptor / variable. In addition, the files within both the dashed box and the dotted box are subject to inheritance, i.e., any information contained within the files located lower in the chart will overwrite equivalent information that is contained within files located higher in the chart. The result of this inheritance is a single set of descriptor functions and keywords that will be used for a given AstroData object.

- <INSTRUMENT>\_Descriptor.py
- GEMINI\_Descriptor.py
- Generic\_Descriptor.py

A descriptor function is used if a descriptor requires access to multiple keywords, requires access to keywords in the pixel data extensions (a dictionary must be created) and / or requires some validation. If no appropriate descriptor function is found, an exception is raised (see the *Descriptor Exceptions* section). If a descriptor value is returned, either directly from the header of the data or from a descriptor function, the CalculatorInterface.py file instantiates the DV object (which contains the descriptor value) and returns this to the user.

#### 4.2.1 CalculatorInterface.py

The CalculatorInterface.py file contains the CalculatorInterface (CI) class, which contains a member function for each available descriptor (see *Appendix C* for an example function). This file is auto-generated by the mkCalculatorInterface.py file (*Appendix B*) and should never be edited directly.

#### 4.2.2 mkCalculatorInterface.py

The mkCalculatorInterface.py file contains the code required to generate the file CalculatorInterface.py. To create CalculatorInterface.py, run the following command:

shell> python mkCalculatorInterface.py > CalculatorInterface.py

#### 4.2.3 StandardDescriptorKeyDict.py

The StandardDescriptorKeyDict.py file contains a Python dictionary named globalStdkeyDict, which describes the one-to-one relationship between a descriptor and the AstroData standard keyword associated with that descriptor ( $Appendix\ D$ ).

#### 4.2.4 StandardGenericKeyDict.py

The StandardGenericKeyDict.py file contains a Python dictionary named stdkeyDictGeneric, which describes the one-to-one relationship between a descriptor / variable that is used in the generic descriptor file Generic\_Descriptor.py and the keyword associated with that descriptor / variable. The values in this dictionary overwrite (via inheritance) the AstroData standard values defined in StandardDescriptorKeyDict.py.

#### 4.2.5 StandardGEMINIKeyDict.py

The <code>StandardGEMINIKeyDict.py</code> file contains a Python dictionary named <code>stdkeyDictGEMINI</code> and is used to overwrite (via inheritance) any of the AstroData standard keywords defined in <code>StandardDescriptorKeyDict.py</code> and any of the <code>generic</code> keywords defined in <code>StandardGenericKeyDict.py</code>. It is also used to define variables for any keywords that need to be accessed in the <code>Geminispecific</code> descriptor file <code>GEMINI\_Descriptor.py</code>.

#### 4.2.6 Standard<INSTRUMENT>KeyDict.py

The Standard<INSTRUMENT>KeyDict.py files contain a Python dictionary named stdkeyDict<INSTRUMENT> and is used to overwrite (via inheritance) any of the AstroData

standard keywords defined in StandardDescriptorKeyDict.py, any of the generic keywords defined in StandardGenericKeyDict.py and any of the Gemini specific keyword defined in StandardGEMINIKeyDict.py. It is also used to define variables for any keywords that need to be accessed in the instrument specific descriptor file <INSTRUMENT>\_Descriptor.py. These instrument specific files are located in the corresponding <INSTRUMENT> directory in the astrodata\_Gemini/ADCONFIG\_Gemini/descriptors directory.

#### 4.2.7 descriptorDescriptionDict.py

The descriptorDescriptionDict.py file contains four Python dictionaries named descriptorDescDict, detailedNameDict, asDictArgDict and stripIDArgDict, which are used by the mkCalculatorInterface.py file to automatically generate the docstrings for the descriptor functions located in the CalculatorInterface.py file. It is likely that the information in the descriptorDescriptionDict.py file will be stored in the mkCalculatorInterface.py file in the future.

#### 4.2.8 calculatorIndex.Gemini.py

The calculatorIndex.Gemini.py file contains a Python dictionary named calculatorIndex and is used to define which Python object (i.e., the descriptor class that defines the descriptor functions, in the form <module\_name>.<calculator\_class\_name>) to use as the calculator for a given <INSTRUMENT>. An example of this dictionary is shown below.

```
calculatorIndex = {
   "<INSTRUMENT>":"<INSTRUMENT>_Descriptor.<INSTRUMENT>_DescriptorCalc()",
}
```

#### 4.2.9 Generic\_Descriptor.py

The generic descriptor file <code>Generic\_Descriptor.py</code> contains descriptor functions describing those keywords that are part of the FITS standard. There are currently 53 keywords defined in the FITS standard (http://heasarc.gsfc.nasa.gov/docs/fcg/standard\_dict.html). There are 4 descriptors currently available that access these FITS standard keywords:

```
instrument [INSTRUME]object [OBJECT]telescope [TELESCOP]ut_date [DATE-OBS]
```

#### 4.2.10 GEMINI\_Descriptor.py

The Gemini specific descriptor file GEMINI\_Descriptor.py contains descriptor functions describing those keywords that are standard within Gemini. There are currently 142 Gemini standard keywords, which are relevant to the data file as a whole and exist in the PHU of the data. There are 21 descriptors currently available that access these Gemini standard keywords:

```
airmass [AIRMASS]azimuth [AZIMUTH]cass_rotator_pa [CRPA]
```

```
• data_label [DATALAB]
• dec [DEC]
• elevation [ELEVATIO]
• local_time [LT]
• observation class [OBSCLASS]
• observation id [OBSID]
• observation_type [OBSTYPE]
• observation_epoch [OBSEPOCH]
• program_id [GEMPRGID]
• qa_state [RAWPIREQ, RAWGEMQA]
• ra [RA]
• raw_bg [RAWBG]
• raw_cc [RAWCC]
raw_iq [RAWIQ]
• raw_wv [RAWWV]
• ut time [UT]
• wavefront_sensor [OIWFS_ST, PWFS2_ST, AOWFS_ST]
• x_offset [XOFFSET]
• y_offset [YOFFSET]
```

#### 4.2.11 <INSTRUMENT>\_Descriptor.py

The instrument specific descriptor files <INSTRUMENT>\_Descriptor.py contain descriptor functions specific to that <INSTRUMENT>. These instrument specific files are located in the corresponding <INSTRUMENT> directory in the astrodata\_Gemini/ADCONFIG\_Gemini/descriptors directory.

#### 4.3 How to add a new descriptor

The following instructions describe how to add a new descriptor to the system.

1. First, check to see whether a descriptor already exists that has the same concept as the new descriptor to be added (*Appendix A*). If a new descriptor is required, edit the mkCalculatorInterface.py file and add the new descriptor to the DD constructor in the descriptors list in alphabetical order. Ensure that the default Python type for the descriptor is defined:

2. Regenerate the CalculatorInterface.py file:

```
shell> python mkCalculatorInterface.py > CalculatorInterface.py
```

- 3. Edit the globalStdkeyDict dictionary in the StandardDescriptorKeyDict.py file to include the AstroData standard keyword associated with the new descriptor (*Appendix D*).
- 4. If the new descriptor simply requires access to the AstroData standard keyword in the header of the data and returns the value, the descriptor can now be tested; go to step 7.
  - However, if the new descriptor requires access to a keyword that is different from the AstroData standard keyword (perhaps specific to a particular <INSTRUMENT>), go to step 5. If the new descriptor requires access to multiple keywords and / or requires some validation, a descriptor function must be created; go to step 6.
- 5. If the new descriptor requires access to a keyword that is different from the AstroData standard keyword, edit either the StandardGenericKeyDict.py file, the StandardGEMINIKeyDict.py file or the Standard<INSTRUMENT>KeyDict.py file (as appropriate) to include the keyword associated with the new descriptor. The descriptor can now be tested; go to step 7.
- 6. If the new descriptor requires access to multiple keywords and / or requires some validation, a descriptor function should be created. Depending on the type of information the new descriptor will provide, edit one of the following files to include the new descriptor function:

```
• Generic_Descriptor.py
```

- GEMINI\_Descriptor.py
- <INSTRUMENT>\_Descriptor.py

The <INSTRUMENT>\_Descriptor.py descriptor file is located in the <INSTRUMENT> directory. If the new descriptor is for a new <INSTRUMENT>, create an <INSTRUMENT> directory and edit the calculatorIndex.Gemini.py appropriately. An example descriptor function (detector\_x\_bin) from GMOS\_Descriptor.py can be found in *Appendix E*. If the descriptor should return more than one value, i.e., one value for each pixel data extension, a dictionary should be returned by the descriptor function, where the key is the ("EXTNAME", EXTVER) tuple. If access to a particular keyword is required, first check the appropriate keyword files (StandardDescriptorKeyDict.py, StandardGenericKeyDict.py, StandardGEMINIKeyDict.py and Standard<INSTRUMENT>KeyDict.py to see if it has already been defined. If required, the Standard<INSTRUMENT>KeyDict.py file should be edited to contain any new keywords required for this new descriptor function.

- 7. Update the Python dictionaries in descriptorDescriptionDict.py so that a docstring can be automatically generated for the new descriptor.
- 8. Test the descriptor:

```
>>> from astrodata import AstroData
>>> ad = AstroData("N20091027S0137.fits")
>>> print ad.<my_descriptor_name>()
```

#### 4.4 Descriptor Coding Guidelines

When creating descriptor functions, the guidelines below should be followed:

- 1. Return value
  - The descriptors will return the correct value, regardless of the data processing status of the AstroData object.
  - The descriptors will not write keywords to the headers of the AstroData object or cache any information, since it is no effort to use the descriptors to obtain the correct value as and when it is required.

• The descriptor values can be written to the history, for information only.

#### 2. Return value Python type

- The descriptors will always return a DV object to the user.
- The DV object is instantiated by the CI for descriptors that obtain their values directly from the headers of the AstroData object. For descriptors that obtain their values from the descriptor functions (i.e., those functions located in the *descriptor files*), the descriptor functions should be coded to return a DV object. The DV object contains information related to the descriptor, including the descriptor value, the default Python type for that descriptor and the units of the descriptor.

#### 3. Keyword access

• The phu\_get\_key\_value and get\_key\_value AstroData member functions should be used in the descriptor functions to access keywords in the header of an AstroData object.

#### 4. Logging

· Descriptors will not log any messages.

#### 5. Raising exceptions

- If a descriptor value can not be determined for whatever reason, the descriptor function should raise an exception.
- The descriptor functions should never be coded to return None. Instead, a descriptor function should throw an exception with a message explaining why a value could not be returned (e.g., if the concept does not directly apply to the data). An exception thrown from a descriptor function will be caught by the CI.

#### 6. Exception rule

- Descriptors should throw exceptions on fatal errors.
- Exceptions thrown on fatal errors (e.g., if a descriptor function is not found in a loaded calculator) should never be caught by the CI. The high level code, such as script or primitive, should catch any relevant exceptions.

#### 7. Descriptor names

- Descriptor names will be:
  - all lower case
  - terms separated with ""
  - not instrument specific
  - not mode specific, mostly
- A descriptor should describe a particular concept and apply for all instrument modes.

#### 8. Standard arguments

- Descriptors accept arguments, some with general purposes are standardized.
- It is especially important for descriptor arguments to follow the Standard Parameter Names as they are front-facing to the user and should therefore be consistent.

For example, for raw GMOS data, the gain descriptor uses the raw keywords in the header and a look up table to determine the gain value. During the first data processing step for Gemini data (which includes standardizing the headers of the data), the value of the raw GAIN keyword is overwritten (since it was incorrect in the raw data) and the value of the raw GAIN keyword is written to the HISTORY keyword (for information only). If a descriptor is then called after the first processing step, the gain descriptor reads the value directly from the GAIN keyword. This way, keyword values are always correct, regardless of the processing state of the data (and any external system that wishes to work on that data will also access the correct values).

#### 4.5 Descriptor Exceptions

Normally, if a descriptor is unable to return a value, None is returned instead. However, exceptions that describe exactly why a value could not be returned (where applicable) are stored so that a user can access that information, if they wish to do so.

When writing descriptor functions, exceptions should be raised in the code with an appropriate, explicit error message, so that it is clear to the user exactly what went wrong. The exception is caught by the CI and if throwExceptions = False (line 62 in astrodata/Calculator.py), the exception information is stored in exception\_info and None is returned. Otherwise the exception is thrown. During development, throwExceptions = True so that exceptions are thrown. When the code is released, throwExceptions = False and the exception information will be available in exception\_info. Available astrodata exceptions can be found in astrodata/Errors.py. Additional required exceptions can be added to this file, if necessary.

## Appendix A

## A Complete List of Available Descriptors

Descriptors are designed such that essential keyword values that describe a particular concept can be accessed from the headers of a given dataset in a consistent manner, regardless of which instrument was used to obtain the data. This is particularly useful for Gemini data, since the majority of keywords used to describe a particular concept at Gemini are not uniform between the instruments.

airmass

• the mean airmass of the observation

amp\_read\_area

• the composite string containing the name of the detector amplifier (ampname) and the readout area of the CCD (detsec) used for the observation.

azimuth

• the azimuth (in degrees between 0 and 360) of the observation

camera

· the camera used for the observation

cass\_rotator\_pa

• the cassegrain rotator position angle (in degrees between -360 and 360) of the observation

central wavelength

• the central wavelength (in meters as default) of the observation

coadds

• the number of coadds used for the observation

data\_label

• the data label of the observation

data section

• the section of the data of the observation

dec

• the declination (in decimal degrees) of the observation

decker

• the decker position used for the observation

detector section

• the detector section of the observation

detector\_x\_bin

• the binning of the x-axis of the detector used for the observation

detector\_y\_bin

• the binning of the y-axis of the detector used for the observation

disperser

• the disperser used for the observation

dispersion

• the dispersion (in meters per pixel as default) of the observation

dispersion\_axis

• the dispersion axis (x = 1; y = 2; z = 3) of the observation

elevation

• the elevation (in degrees) of the observation

exposure\_time

• the total exposure time (in seconds) of the observation

filter\_name

• the unique, sorted filter name idenifier string used for the observation

focal\_plane\_mask

• the focal plane mask used for the observation

gain

• the gain (in electrons per ADU) of the observation

gain\_setting

• the gain setting of the observation

grating

• the grating used for the observation

group\_id

• the group\_id

instrument

• the instrument used for the observation

local\_time

• the local time (in HH:MM:SS.S) at the start of the observation

mdf\_row\_id

• the corresponding reference row in the MDF

nod\_count

• the number of nod and shuffle cycles in the nod and shuffle observation

nod\_pixels

• the number of pixel rows the charge is shuffled by in the nod and shuffle observation

non\_linear\_level

• the non linear level in the raw images (in ADU) of the observation

object

• the name of the target object observed

observation\_class

• the class (either 'science', 'progCal', 'partnerCal', 'acq', 'acqCal' or 'dayCal') of the observation

observation\_epoch

• the epoch (in years) at the start of the observation

observation\_id

• the ID (e.g., GN-2011A-Q-123-45) of the observation

observation\_type

• the type (either 'OBJECT', 'DARK', 'FLAT', 'ARC', 'BIAS' or 'MASK') of the observation

overscan\_section

• the overscan\_section

pixel\_scale

• the pixel scale (in arcsec per pixel) of the observation

prism

• the prism used for the observation

program\_id

• the Gemini program ID (e.g., GN-2011A-Q-123) of the observation

pupil\_mask

• the pupil mask used for the observation

qa\_state

• the quality assessment state (either 'Undefined', 'Pass', 'Usable', 'Fail' or 'CHECK') of the observation

ra

• the Right Ascension (in decimal degrees) of the observation

raw\_bg

• the raw background (either '20-percentile', '50-percentile', '80-percentile' or 'Any') of the observation

raw\_cc

• the raw cloud cover (either '50-percentile', '70-percentile', '80-percentile', '90-percentile' or 'Any') of the observation

raw\_iq

• the raw image quality (either '20-percentile', '70-percentile', '85-percentile' or 'Any') of the observation

raw\_wv

• the raw water vapour (either '20-percentile', '50-percentile', '80-percentile' or 'Any') of the observation

read mode

• the read mode (either 'Very Faint Object(s)', 'Faint Object(s)', 'Medium Object', 'Bright Object(s)', 'Very Bright Object(s)', 'Low Background', 'Medium Background', 'High Background' or 'Invalid') of the observation

read\_noise

• the estimated readout noise (in electrons) of the observation

read\_speed\_setting

• the read speed setting (either 'fast' or 'slow') of the observation

saturation\_level

• the saturation level in the raw images (in ADU) of the observation

slit

• the slit used for the observation

telescope

• the telescope used for the observation

ut\_date

• the UT date at the start of the observation

ut\_datetime

• the UT date and time at the start of the observation

ut\_time

• the UT time at the start of the observation

wavefront\_sensor

• the wavefront sensor (either 'AOWFS', 'OIWFS', 'PWFS1', 'PWFS2', some combination in alphebetic order separated with an ampersand or None) used for the observation

wavelength\_reference\_pixel

• the reference pixel of the central wavelength of the observation

well\_depth\_setting

• the well depth setting (either 'Shallow', 'Deep' or 'Invalid') of the observation

x offset

• the x offset of the observation

y\_offset

• the y offset of the observation

## Appendix B

## mkCalculatorInterface.py

```
from datetime import datetime
from descriptorDescriptionDict import asDictArgDict
from descriptorDescriptionDict import descriptorDescDict
from descriptorDescriptionDict import detailedNameDict
from descriptorDescriptionDict import stripIDArgDict
class DescriptorDescriptor:
   name = None
    description = None
   pytype = None
   unit = None
    thunkfuncbuff = """
    def %(name)s(self, format=None, **args):
        \"\"\"
        %(description)s
        \"\"\"
        try:
            self._lazyloadCalculator()
            keydict = self.descriptor_calculator._specifickey_dict
            #print hasattr(self.descriptor_calculator, "%(name)s")
            if not hasattr(self.descriptor_calculator, "%(name)s"):
                key = "key_" + "% (name) s"
                #print "mkCI10:", key, repr(keydict)
                #print "mkCI12:", key in keydict
                if key in keydict.keys():
                    retval = self.phu_get_key_value(keydict[key])
                    if retval is None:
                        if hasattr(self, "exception_info"):
                            raise self.exception_info
                else:
                    msg = "Unable to find an appropriate descriptor function "
                    msg += "or a default keyword for %(name)s"
                    raise KeyError(msg)
                retval = self.descriptor_calculator.%(name)s(self, **args)
            %(pytypeimport)s
            ret = DescriptorValue( retval,
                                    format = format,
                                   name = "% (name) s",
```

```
ad = self,
                                pytype = %(pytype)s)
        return ret
    except:
        if not hasattr(self, "exception_info"):
            setattr(self, "exception_info", sys.exc_info()[1])
        if (self.descriptor_calculator is None
            or self.descriptor_calculator.throwExceptions == True):
            raise
        else:
            #print "NONE BY EXCEPTION"
            self.exception_info = sys.exc_info()[1]
            return None
0.00
def __init__(self, name=None, pytype=None):
    self.name = name
    if pytype:
        self.pytype = pytype
        rtype = pytype.__name__
    try:
        desc = descriptorDescDict[name]
    except:
        if rtype == 'str':
            rtype = 'string'
        if rtype == 'int':
            rtype = 'integer'
        try:
            dname = detailedNameDict[name]
        except:
            dname = name
        try:
            asDictArg = asDictArgDict[name]
        except:
            asDictArg = 'no'
        try:
            stripIDArg = stripIDArgDict[name]
        except:
            stripIDArg = 'no'
        if stripIDArg == 'ves':
            desc = 'Return the %(name)s value\n' % {'name':name} + \
                            :param dataset: the data set\n' + \
                             :type dataset: AstroData\n' + \
                             :param stripID: set to True to remove the ^\prime + \setminus
                   'component ID from the \n
                    'returned % (name) s value \n' % {'name':name} + \
                             :type stripID: Python boolean\n' + \
                             :param pretty: set to True to return a ^\prime + ^{\setminus}
                    'human meaningful \n' + \
                                             % (name) s ' % {'name':name} + \
                   'value\n' + \
                            :type pretty: Python boolean\n' + \
                            :rtype: %(rtype)s ' % {'rtype':rtype} + \
                   'as default (i.e., format=None)\n' + \
                            :return: the %(dname)s' \
                   % {'dname':dname}
        elif asDictArg == 'yes':
            desc = 'Return the %(name)s value\n' % {'name':name} + \
```

```
:param dataset: the data set\n' + \
                                :type dataset: AstroData\n' + \
                                :param format: the return format\n' + \
                                               set to as_dict to return a ' + \
                       'dictionary, where the number ' + \
                                                of dictionary elements ' + \
                       'equals the number of pixel data ' + \
                                                 extensions in the image. ' + \
                       'The key of the dictionary is ' + \
                                                 an (EXTNAME, EXTVER) ' + \
                       'tuple, if available. Otherwise, ' + \
                                                 the key is the integer ' + \
                       'index of the extension.\n' + \n'
                                :type format: string\n' + \
                                :rtype: %(rtype)s ' % {'rtype':rtype} + \
                       'as default (i.e., format=None)\n' + \n'
                                :rtype: dictionary containing one or more ' + \
                       '%(rtype)s(s) '% {'rtype':rtype} + \
                       '(format=as_dict)\n' + \
                                :return: the %(dname)s' \
                       % {'dname':dname}
            else:
                desc = 'Return the % (name) s value \n' % {'name':name} + \
                               :param dataset: the data set\n' + \
                                :type dataset: AstroData\n' + \
                                :param format: the return format\n' + \
                                :type format: string\n' + \
                                :rtype: %(rtype)s ' % {'rtype':rtype} + \
                       'as default (i.e., format=None)\n' + \
                               :return: the %(dname)s' \
                       % {'dname':dname}
        self.description = desc
    def funcbody(self):
        if self.pytype:
           pytypestr = self.pytype.__name__
        else:
           pytypestr = "None"
        if pytypestr == "datetime":
           pti = "from datetime import datetime"
        else:
           pti = ""
        #print "mkC150:", pti
        ret = self.thunkfuncbuff % {'name':self.name,
                                    'pytypeimport': pti,
                                    'pytype': pytypestr,
                                    'description':self.description}
        return ret
DD = DescriptorDescriptor
                    DD("airmass", pytype=float),
descriptors = [
                    DD("amp_read_area", pytype=str),
                    DD("azimuth", pytype=float),
                    DD("camera", pytype=str),
                    DD("cass_rotator_pa", pytype=float),
```

```
DD("central_wavelength", pytype=float),
DD("coadds", pytype=int),
DD("data_label", pytype=str),
DD("data_section", pytype=list),
DD("dec", pytype=float),
DD ("decker", pytype=str),
DD("detector_section", pytype=list),
DD("detector_x_bin", pytype=int),
DD("detector_y_bin", pytype=int),
DD("disperser", pytype=str),
DD("dispersion", pytype=float),
DD("dispersion_axis", pytype=int),
DD("elevation", pytype=float),
DD("exposure_time", pytype=float),
DD("filter_name", pytype=str),
DD("focal_plane_mask", pytype=str),
DD("gain", pytype=float),
DD("grating", pytype=str),
DD("group_id", pytype=str),
DD("gain_setting", pytype=str),
DD("instrument", pytype=str),
DD("local_time", pytype=str),
DD("mdf_row_id", pytype=int),
DD("nod_count", pytype=int),
DD("nod_pixels", pytype=int),
DD("non_linear_level", pytype=int),
DD("object", pytype=str),
DD("observation_class", pytype=str),
DD("observation_epoch", pytype=str),
DD("observation_id", pytype=str),
DD("observation_type", pytype=str),
DD("overscan_section", pytype=list),
DD("pixel_scale", pytype=float),
DD("prism", pytype=str),
DD("program_id", pytype=str),
DD("pupil_mask", pytype=str),
DD("qa_state", pytype=str),
DD("ra", pytype=float),
DD("raw_bg", pytype=str),
DD("raw_cc", pytype=str),
DD("raw_iq", pytype=str),
DD("raw_wv", pytype=str),
DD("read_mode", pytype=str),
DD("read_noise", pytype=float),
DD("read_speed_setting", pytype=str),
DD("saturation_level", pytype=int),
DD("slit", pytype=str),
DD("telescope", pytype=str),
DD("ut_date", pytype=datetime),
DD("ut_datetime", pytype=datetime),
DD("ut_time", pytype=datetime),
DD("wavefront_sensor", pytype=str),
DD("wavelength_reference_pixel", pytype=float),
DD("well_depth_setting", pytype=str),
DD("x_offset", pytype=float),
DD("y_offset", pytype=float),
```

1

```
wholeout = """import sys
from astrodata import Descriptors
from astrodata. Descriptors import Descriptor Value
from astrodata import Errors
class CalculatorInterface:
   descriptor_calculator = None
%(descriptors)s
# UTILITY FUNCTIONS, above are descriptor thunks
    def _lazyloadCalculator(self, **args):
        '''Function to put at top of all descriptor members
        to ensure the descriptor is loaded. This way we avoid
        loading it if it is not needed.'''
        if self.descriptor_calculator is None:
            self.descriptor_calculator = Descriptors.get_calculator(self, **args)
out = ""
for dd in descriptors:
   out += dd.funcbody()
finalout = wholeout % {"descriptors": out}
print finalout
```

## **Appendix C**

# An Example Function from CalculatorInterface.py

The example function below is auto-generated by the mkCalculatorInterface.py file. The CalculatorInterface.py file should never be edited directly.

```
def airmass(self, format=None, **args):
    Return the airmass value
    :param dataset: the data set
    :type dataset: AstroData
    :param format: the return format
    :type format: string
    :rtype: float as default (i.e., format=None)
    :return: the mean airmass of the observation
    try:
        self._lazyloadCalculator()
        keydict = self.descriptor_calculator._specifickey_dict
        if not hasattr(self.descriptor_calculator, "airmass"):
            key = "key_"+"airmass"
            if key in keydict.keys():
                retval = self.phu_get_key_value(keydict[key])
                if retval is None:
                    if hasattr(self, "exception_info"):
                        raise self.exception_info
            else:
                msg = "Unable to find an appropriate descriptor function "
                msq += "or a default keyword for airmass"
                raise KeyError(msg)
        else:
            retval = self.descriptor_calculator.airmass(self, **args)
        ret = DescriptorValue( retval,
                               format = format,
                               name = "airmass",
                               ad = self,
                               pytype = float )
        return ret
    except:
```

```
if (self.descriptor_calculator is None
    or self.descriptor_calculator.throwExceptions == True):
    raise
else:
    #print "NONE BY EXCEPTION"
    self.exception_info = sys.exc_info()[1]
    return None
```

## **Appendix D**

## StandardDescriptorKeyDict.py

```
globalStdkeyDict = {
    "key_airmass":"AIRMASS",
    "key_amp_read_area": "AMPROA",
    "key_azimuth": "AZIMUTH",
    "key_camera":"CAMERA",
    "key_cass_rotator_pa":"CRPA",
    "key_central_wavelength": "CWAVE",
    "key_coadds": "COADDS",
    "key_data_label": "DATALAB",
    "key_data_section": "DATASEC",
    "key_dec":"DEC",
    "key_decker":"DECKER",
    "key_detector_section": "DETSEC",
    "key_detector_x_bin":"XCCDBIN",
    "key_detector_y_bin":"YCCDBIN",
    "key_disperser": "DISPERSR",
    "key_dispersion": "WDELTA",
    "key_dispersion_axis":"DISPAXIS",
    "key_elevation":"ELEVATIO",
    "key_exposure_time": "EXPTIME",
    "key_filter_name":"FILTNAME",
    "key_focal_plane_mask": "FPMASK",
    "key_gain": "GAIN",
    "key_gain_setting": "GAINSET",
    "key_grating": "GRATING",
    "key_group_id": "GROUPID",
    "key_instrument": "INSTRUME",
    "key_local_time":"LT",
    "key_mdf_row_id": "MDFROW",
    "key_nod_count": "NODCOUNT",
    "key_nod_pixels": "NODPIX",
    "key_non_linear_level": "NONLINEA",
    "key_object":"OBJECT",
    "key_observation_class": "OBSCLASS",
    "key_observation_epoch": "OBSEPOCH",
    "key_observation_id": "OBSID",
    "key_observation_type":"OBSTYPE",
    "key_overscan_section":"OVERSSEC",
    "key_pixel_scale":"PIXSCALE",
    "key_prism": "PRISM",
    "key_program_id": "GEMPRGID",
```

```
"key_pupil_mask":"PUPILMSK",
"key_qa_state":"QASTATE",
"key_ra":"RA",
"key_raw_bg":"RAWBG",
"key_raw_cc": "RAWCC",
"key_raw_iq": "RAWIQ",
"key_raw_wv":"RAWWV",
"key_read_mode":"READMODE",
"key_read_noise":"RDNOISE",
"key_read_speed_setting":"RDSPDSET",
"key_saturation_level": "SATLEVEL",
"key_slit":"SLIT",
"key_telescope":"TELESCOP",
"key_ut_date":"DATE-OBS",
"key_ut_datetime": "DATETIME",
"key_ut_time":"UT",
"key_wavefront_sensor":"WFS",
"key_wavelength_reference_pixel":"WREFPIX",
"key_well_depth_setting":"WELDEPTH",
"key_x_offset":"XOFFSET",
"key_y_offset":"YOFFSET",
```

## **Appendix E**

# An Example Descriptor Function from GMOS\_Descriptor.py

```
from astrodata import Errors
from StandardGMOSKeyDict import stdkeyDictGMOS
from GEMINI_Descriptor import GEMINI_DescriptorCalc
class GMOS_DescriptorCalc(GEMINI_DescriptorCalc):
    # Updating the global key dictionary with the local key dictionary
    # associated with this descriptor class
   _update_stdkey_dict = stdkeyDictGMOS
    def __init__(self):
       GEMINI_DescriptorCalc.__init__(self)
    def detector_x_bin(self, dataset, **args):
        # Since this descriptor function accesses keywords in the headers of
        # the pixel data extensions, always return a dictionary where the key
        # of the dictionary is an (EXTNAME, EXTVER) tuple.
        ret_detector_x_bin = {}
        # Loop over the science extensions in the dataset
        for ext in dataset["SCI"]:
            # Get the ccdsum value from the header of each pixel data
            # extension. The ccdsum keyword is defined in the local key
            # dictionary (stdkeyDictGMOS) but is read from the updated global
            # key dictionary (self.get_descriptor_key())
            ccdsum = ext.get_key_value(self.get_descriptor_key("key_ccdsum"))
            if ccdsum is None:
                # The get_key_value() function returns None if a value cannot
                # be found and stores the exception info. Re-raise the
                # exception. It will be dealt with by the CalculatorInterface
                if hasattr(ext, "exception_info"):
                    raise ext.exception_info
            detector_x_bin, detector_y_bin = ccdsum.split()
            # Return a dictionary with the binning of the x-axis integer as
```

```
# the value
ret_detector_x_bin.update({
          (ext.extname(), ext.extver()):int(detector_x_bin)})

if ret_detector_x_bin == {}:
    # If the dictionary is still empty, the AstroData object was not
    # autmatically assigned a "SCI" extension and so the above for
    # loop was not entered
    raise Errors.CorruptDataError()
return ret_detector_x_bin
```