Statistical thermodynamics

- Link to macroscopic thermodynamics from the statistical analysis of molecular properties.
 - Valid for systems in thermal equilibrium
- "System" refers to the macroscopic term
- Microscopic level is populated by "particules" (molecules actually in our case)
- State refers to the "thermodynamic state" or to the "microscopic state", the number of molecules and their energy levels
 - Many microstates (the "ensemble") are compatible with a single thermodynamic state.
 - Ensembles are in constant fluctuation, but the derived macroscopic properties are constant.
 - The value of a macroscopic property is the average of the corresponding microscopic property in the ensemble. Ex. $E = \langle \varepsilon \rangle = \sum p_i \, \varepsilon_i$ i: all ensemble microstates. p_i probability of a particle being in the microstate i

Statistical thermodynamics

Two key ideas

- Boltzmann distribution predicts populations of states in systems at thermal equilibrium.
- The derivation of Boltzmann distribution leads naturally to the introduction of partition function.

Questions:

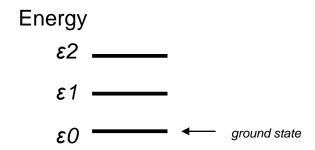
- How to interpret partition function and calculate it.
- How to extract thermodynamic information from the partition function

Statistical thermodynamics

- System composed of N molecules.
 - Total internal energy is constant at E
 - It is not possible to be definite about how that energy is shared between the molecules.
 - Collisions result in the ceaseless redistribution of energy, not only between the molecules but also among their different modes of motion.
- Population of a state
 - the average number of molecules that occupy it. On average there are n_i molecules in a state of energy ϵ_i .
 - The populations of the states depend only on ε_i that is a function of T
- All states having the same Energy level (and T) are equally probable.

Configurations and weights

Instantaneous configurations



At any instant molecules could occupy one energy level. In this example: $\varepsilon 0$, $\varepsilon 1$, or $\varepsilon 2$

The specification of the set of populations n0, n1, n2, ... in the form $\{n0, n1, n2, ...\}$ is a statement of the **instantaneous configuration** of the system.

Weight of a configuration (W) = N possible microstates for a configuration

$$W = \frac{N!}{n_0! n_1! n_2! n_3! \dots}$$

For Large N can be written as : $\ln W = N \ln N - \sum_{i} n_{i} \ln n_{i}$

Example: 5 molecules, 2 energy levels

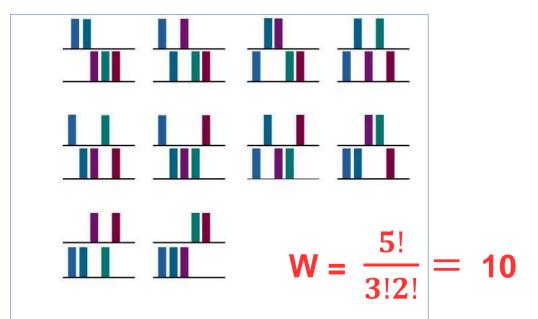


Fig. 16.1 Whereas a configuration {5,0,0,...} can be achieved in only one way, a configuration {3,2,0,...} can be achieved in the ten different ways shown here, where the tinted blocks represent different molecules.

Configurations and weights: The optimal distribution

We could find the most probable distribution imposing a system with:

- Constant total number of molecules: $N = \sum_i n_i$
- Constant total energy: $E = \sum_{i} p_{i} \varepsilon_{i}$
- A compatible configuration having the maximum weight $\delta W = 0$

$$p_i = \frac{n_i}{N} = \frac{e^{-\beta \varepsilon_i}}{\sum_i e^{-\beta \varepsilon_j}} \qquad \beta = \frac{1}{k_B T}$$

where T is the thermodynamic temperature and k is Boltzmann's constant.

Identical distribution can be obtained from both Kinetic Gas Theory and Statistical Mechanics!!

The molecular partition function

- Boltzmann distribution is a **probability distribution**. It gives the probability of a certain state in function of that state's Energy and the Temperature of the system .
- The normalization term is called Molecular (canonical)
 Partition Function
- Molecular partition function defines the statistical properties of the system from the microscopic states
- The molecular partition function gives an indication of the number of states that are thermally accessible to a molecule at the temperature of the system.
 - At T = 0, only the ground level is accessible and q = g
 - At very high temperatures, virtually all states are accessible, and q is correspondingly large

$$p_i = \frac{n_i}{N} = \frac{e^{-\beta \varepsilon_i}}{\sum_j e^{-\beta \varepsilon_j}}$$
 $\beta = \frac{1}{k_B T}$

$$Z = \sum_{j} e^{-\beta \varepsilon_{j}} = \sum_{l} g_{j} e^{-\beta \varepsilon_{l}}$$

Counting either for "j" microstates or for "l" energy levels g_j degenerated

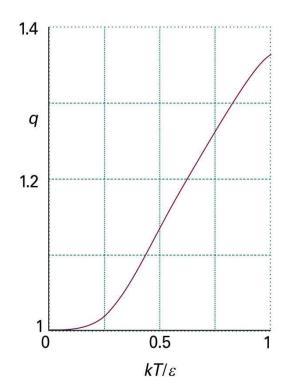
Example: two-level system

$$\varepsilon = \varepsilon$$

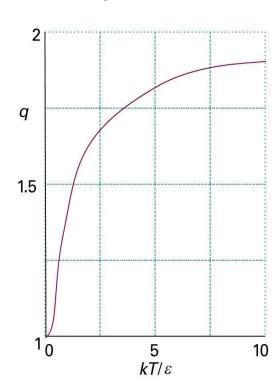
$$\varepsilon = 0$$

$$Z = \sum_{i=0}^{\varepsilon} e^{-\beta \varepsilon_i} = 1 + e^{-\beta \varepsilon}$$





big scale



The partition function for a two-level system as a function of temperature.

Example: two-level system

$$\varepsilon = \varepsilon$$

$$\varepsilon = 0$$

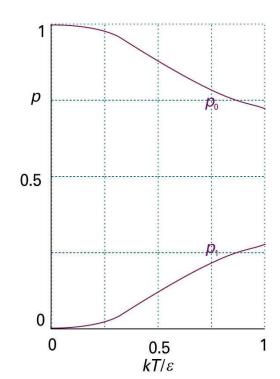
$$p_i = \frac{n_i}{N} = \frac{e^{-\beta \varepsilon_i}}{q} = \frac{e^{-\beta \varepsilon_i}}{1 + e^{-\beta \varepsilon}}$$

$$Z = \sum_{i=0}^{1} e^{-\beta \varepsilon_i} = 1 + e^{-\beta \varepsilon}$$

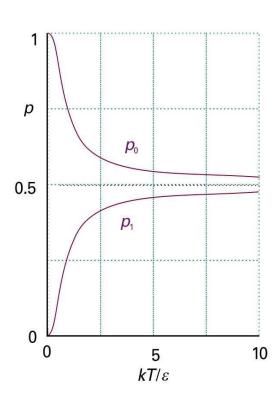
$$p_0 = \frac{n_i}{N} = \frac{1}{1 + e^{-\beta \varepsilon}}$$

$$p_1 = \frac{n_i}{N} = \frac{e^{-\beta \varepsilon}}{1 + e^{-\beta \varepsilon}}$$





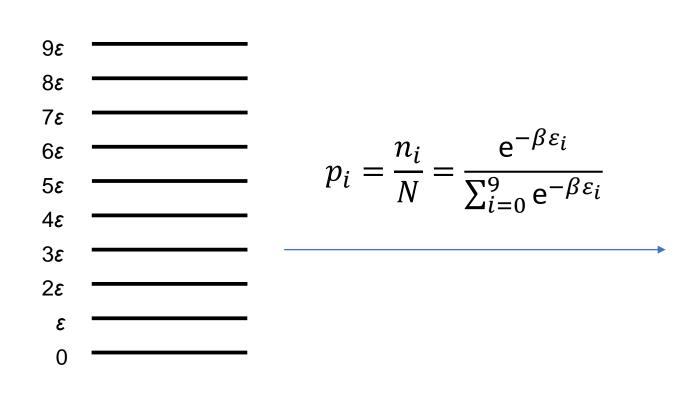
big scale



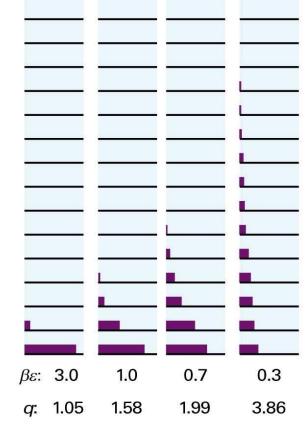
The fraction of populations of the two states of a twolevel system as a function of temperature.

Example: uniform distribution of energy levels

A system with equally spaced non-degenerate energy levels.



The equally spaced array of energy levels used in the calculation of the partition function.



Low

temperature

High

temperature

Temperature (in βε)
Partition fuction

Derived Thermodynamic functions

$$E = \frac{-N}{Z} \left(\frac{\partial Z}{\partial \beta} \right) = \frac{N}{Z} \sum_{i} \varepsilon_{i} e^{-\beta \varepsilon_{i}}$$

$$P = k_B NT \left(\frac{\partial \ln Z}{\partial V} \right)$$

$$S = E/_T + k_B \operatorname{Nln} Z$$

$$G = -k_B NT \left[\ln Z - \left(\frac{\partial \ln Z}{\partial V} \right) \right]$$

• "Statistical" entropy

$$S = -k_B N \sum_i p_i \ln p_i$$
$$S = k_B \ln W$$

$$p_{i} = \frac{1}{Z}e^{-\beta\varepsilon_{i}}$$

$$\ln p_{i} = -\beta\varepsilon_{i} - \ln Z$$

$$-\sum_{i} p_{i} \ln p_{i} = \sum_{i} (p_{i}\beta\varepsilon_{i} + p_{i} \ln Z) =$$

$$\beta \sum_{i} p_{i}\varepsilon_{i} + \ln Z = \beta E/N + \ln Z$$

$$\operatorname{Nln} Z = -\beta E - N \sum_{i} p_{i} \ln p_{i}$$

$$S = \frac{E}{T} + k_{B} \operatorname{Nln} Z = -k_{B} N \sum_{i} p_{i} \ln p_{i}$$

$$S = -k_B N \sum_{i} \frac{n_i}{N} \ln \frac{n_i}{N} = -k_B N \left(\frac{1}{N} \sum_{i} n_i \ln n_i - \ln N \right)$$

$$\ln W = N \ln N - \sum_{i} n_{i} \ln n_{i}$$

Summary

- Gas theory and statistical thermodynamics can be used to understand the microscopic level
- At the microscopic level, molecules form a complex ensemble of many microstates of different populations and energies
- Macroscopic properties are just averages over such microstates
- Boltzmann distribution gives the probability of particles (molecules) to populate specific microstates
- System thermodynamics functions can be obtained from the molecular partition function.