## Master ENS ICFP - First Year 2020/2021

# Relativistic Quantum Mechanics and Introduction to Quantum Field Theory

#### Mid Term Homework

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#### 1 Some operator identities: 6 points

1. We have that:

$$e^A B e^{-A} = \sum_{n=0}^{\infty} \frac{A^n}{n!} B \sum_{n=0}^{+\infty} \frac{(-A)^n}{n!} = \sum_{n,m=0}^{+\infty} \frac{A^n B (-A)^m}{n!m!} = \sum_{n,m=0}^{+\infty} (-1)^m \frac{A^n B A^m}{n!m!} = \sum_{n=0}^{+\infty} \frac{A^n}{n!} \sum_{m=0}^{+\infty} (-1)^m \frac{B A^m}{m!} = \sum_{n=0}^{+\infty} \frac{A^n}{n!} \sum_{m=0}^{+\infty} (-1)^m \frac{A^n B A^m}{m!} = \sum_{n=0}^{+\infty} \frac{A^n}{n!} \sum_{m=0}^{+\infty} (-1)^m \frac{A^m B A^m}{m!} = \sum_{n=0}^{+\infty} \frac{A^n}{n!} \sum_{m=0}^{+\infty} (-1)^m \frac{A^m}{m!} = \sum_{m=0}^{+\infty} \frac{A^m}{n!} \sum_{m=0}^{+\infty} \frac{A^m}{n!} = \sum_{m=$$

2. ...

3. We have that:

$$[F, G^{\dagger}] = [\sum_{j} f_{j} a_{j}, \sum_{j} g_{j}^{\star} a_{j}^{\dagger}] = \sum_{j,k=0}^{+\infty} f_{j} g_{k}^{\star} [a_{j}, a_{k}^{\dagger}] = \sum_{j,k=0}^{+\infty} f_{j} g_{k}^{\star} \delta_{jk} = \sum_{j=0}^{+\infty} f_{j} g_{j}^{\star} \delta_{jk}$$

Furthermore we have that  $[F, G^{\dagger}] \propto \text{Id}$  and therefore we trivially have that  $[F, [F, G^{\dagger}]] = [G^{\dagger}, [F, G^{\dagger}]] = 0$ . Now applying question 2 we have that:

$$e^{G^{\dagger}}e^{F} = e^{-\frac{1}{2}\sum_{j}f_{j}g_{j}^{\star}}e^{G^{\dagger}+F} \Rightarrow e^{\frac{1}{2}\sum_{j}f_{j}g_{j}^{\star}}e^{F} = \underbrace{e^{\frac{1}{2}\sum_{j}f_{j}g_{j}^{\star}}e^{-\frac{1}{2}\sum_{j}f_{j}g_{j}^{\star}}}_{=e^{A_{e}-A}}e^{G^{\dagger}+F}$$

Now from Question 1 we have that for any A (trivially [A, Id] = 0) we get:

$$e^{A} \operatorname{Id} e^{-A} = \operatorname{Id} + \sum_{n=1}^{+\infty} \frac{1}{n!} \cdot 0 = \operatorname{Id}$$

Hence the above formula simplifies to:

$$e^{F+G^{\dagger}} = e^{\frac{1}{2}\sum_{j} f_{j} g_{j}^{\star}} e^{G^{\dagger}} e^{F}$$

4. Similarly as before let  $F = \int d^3 \mathbf{q} f(\mathbf{q}) a(\mathbf{q})$  and  $G = \int d^3 \mathbf{q} h(\mathbf{q})^{\dagger} a(\mathbf{q})$ . Then we have that:

$$[F, G^{\dagger}] = \left[ \int d^{3}\mathbf{q} f(\mathbf{q}) a(\mathbf{q}), \int d^{3}\mathbf{q} h(\mathbf{q}) a^{\dagger}(\mathbf{q}) \right] = \int d^{3}\mathbf{q} f(\mathbf{q}) h(\mathbf{q}) [a(\mathbf{q}), a^{\dagger}(\mathbf{q})] = \int d^{3}\mathbf{q} f(\mathbf{q}) [a(\mathbf{q}), a^{\dagger}(\mathbf{q})] = \int d^{3}\mathbf{q} f(\mathbf{q}$$

A similar direct application of 2 gives the desired result.

## 2 An example of an asymptotic series

We have that:

$$f(g) = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{+\infty} \mathrm{d}x e^{-x^2 - gx^4} \text{ hence } |f(g)| < \int_{-\infty}^{+\infty} \mathrm{d}x |e^{-x^2 - gx^4}| = \int_{-\infty}^{+\infty} \mathrm{d}x e^{-x^2 - x^4 \operatorname{Re}g} < \int_{-\infty}^{+\infty} \mathrm{d}x e^{-x^4 \operatorname{Re}g}$$

Hence as long as Re g > 0 this is obviously well defined from the last term and if Re g = 0 this is obviously well defined from the before last term.

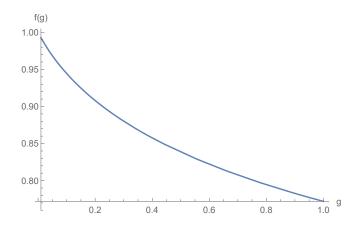


Figure 1: Plot of the numerical values of f(g) for  $g \in [0.01, 1]$ .

1. This integral admits an exact solution given by:

$$f(g) = \frac{e^{\frac{1}{8g}} K_{\frac{1}{4}}(\frac{1}{8g})}{2\sqrt{\pi g}} \delta_{\text{Re }g>0} + \delta_{\text{Re }g=0} \text{ where } K_n(z) \text{ is the modified Bessel function of the second kind.}$$

The plot of the numerical values for  $g \in [0.01, 1]$  is given in Figure 1. Then f(g) decreases monotonically when g > 0 increases since:

$$\frac{\mathrm{d}}{\mathrm{d}g}f(g) = \frac{1}{\sqrt{\pi}} \int_{-\infty}^{+\infty} \mathrm{d}x (-x^4) e^{-x^2 - gx^4} = \frac{-1}{\sqrt{\pi}} \underbrace{\int_{-\infty}^{+\infty} x^4 e^{-x^2 - gx^4}}_{>0 \text{ when } g \in \mathbb{R}^+} < 0$$

2. We have that:

$$e^{-gx^4} = \sum_{n=0}^{+\infty} \frac{(-gx^4)^n}{n!}$$

And plugging this in the expression of f and inverting the sum and the integral gives:

$$\frac{1}{\sqrt{\pi}} \sum_{n=0}^{+\infty} \frac{(-g)^n}{n!} \int_{-\infty}^{+\infty} dx \ x^{4n} e^{-x^2}$$

We notice the integral ressembles strongly the gamma function hence we change variables by taking  $u=x^2$  ( $\mathrm{d}u=2\sqrt{u}\mathrm{d}x$ ) and we get:

$$2^{-1}\int_{-\infty}^{+\infty}\mathrm{d} u u^{2n-\frac{1}{2}}e^{-u}=2^{-1}\Gamma(2n+\frac{1}{2})=2^{-4n}\sqrt{\pi}\frac{\Gamma(4n)}{\Gamma(2n)} \ \text{ from the Legendre duplication formula}.$$

Hence plugging it back up top we obtain:

$$\tilde{f}(g) = \sum_{n=0}^{+\infty} \left( \frac{(-1)^n (4n)!}{n! 2^{4n} (2n)!} \right) g^n$$

Notice that the terms  $f_n$  are monotonically increasing in norm and diverge hence the sum does not converge absolutely and R=0 and it also does not converge conditionally. The order of magnitude of the first few terms is ...

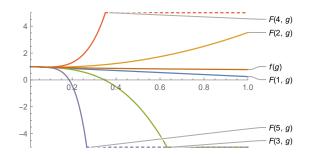
3.

#### 3 A relation between Dirac spinors

1. We have that:

...

2. ...



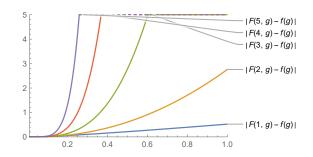


Figure 2: Series approximations of f(g) and their errors for  $g \in [0.01, 1]$ .

#### 4 Some traces of products of $\gamma$ -matrices

We have that:

$$\operatorname{tr} \gamma_{\mu} \gamma_{\nu} = \operatorname{tr} \{ \gamma_{\mu}, \gamma_{\nu} \} - \gamma_{\nu} \gamma_{\mu} = 2 \operatorname{tr} \eta_{\mu\nu} I_4 - \operatorname{tr} \gamma_{\nu} \gamma_{\mu} = 2 \operatorname{tr} \eta_{\mu\nu} I_4 - \operatorname{tr} \gamma_{\mu} \gamma_{\nu}$$

Hence adding on both side we obtain the desired equality:

$$\operatorname{tr} \gamma_{\mu} \gamma_{\nu} = \eta_{\mu\nu} \operatorname{tr} I_4 = 4\eta_{\mu\nu}$$

Similarly we have that:

$$\operatorname{tr} \gamma_{\mu} \gamma_{\nu} \gamma_{\rho} \gamma_{\sigma} = \dots$$

# 5 Energy levels of a relativistic charged spin-0 particle in a harmonic electrostatic potential

1. From these relation we have that:

$$\begin{split} X^2 \left| n \right\rangle &= \frac{1}{2m\Omega} \left( a^2 + (a^\dagger)^2 + \left\{ a, a^\dagger \right\} \right) \left| n \right\rangle \\ &= \frac{1}{2m\Omega} (\sqrt{n}\sqrt{n-1} \left| n-2 \right\rangle + \sqrt{n+1}\sqrt{n+2} \left| n+2 \right\rangle + (n+1) \left| n \right\rangle + n \left| n \right\rangle ) \end{split}$$

Hence we get that:

$$\langle n|\,X^4\,|n\rangle = (\langle n|\,X^2)(X^2\,|n\rangle) = \frac{1}{(2m\Omega)^2}(n^2-n+n^2+3n+2+n+1+n) = \frac{2n^2+4n+3}{(2m\Omega)^2}$$

2. ...

#### 6 The axial current

#### 7 Supersymmetry