Gravitational Waves

Nicholas P. Ross

February 19, 2016

Abstract

This is a simple document which will...

- 1 PreAmble
- 2 Equations
- 3 Definitions

strain = fractional change in length/distance

- 4 LIGO
- 5 References

Ross et al. (2007)

References

Ross N. P., et al., 2007, MNRAS, 381, 573

Table 1: The Lines

Name	Wavelength / Å	Transition	Rest Passband	Interpretaion
Lyman- α	1215.67	2 to 1	\sim FUV	Major QSO line
Lyman- β	1025.18	3 to 1	\sim FUV	
Lyman- γ	972.02	4 to 1	\sim FUV	
Lyman Limit	911.27	∞ to 1	\sim FUV	
$H-\alpha$	6563.	3 to 2	R,r	Recent major SF or AGN activity
H - β	4861.	4 to 2	B,V,g	
$ ext{H-}\gamma$	4341.	5 to 2	$_{\mathrm{U,B,u}}$	
$ ext{H-}\delta$	4102.	6 to 2	\sim FUV	Previous SF history
Balmer Limit	3646.	∞ to 2	\sim FUV	
HI	3646.	∞ to 2	\sim FUV	
HII	3646.	∞ to 2	\sim FUV	
HeI	3646.	∞ to 2	\sim FUV	
${ m HeII}$	3646.	∞ to 2	\sim FUV	
HeIII	3646.	∞ to 2	\sim FUV	
CIV	3646.	∞ to 2	\sim FUV	Major QSO line
OII	3646.	∞ to 2	\sim FUV	Major QSO line
OIII	3646.	∞ to 2	\sim FUV	Recent major SF line
OIII	5007.	∞ to 2	\sim FUV	Recent major SF line
Ca II H	3999.	∞ to 2	\sim FUV	Old stellar pop
Ca II K	4001.	∞ to 2	\sim FUV	Old stellar pop
NII	5007.	∞ to 2	\sim FUV	
NeV	3646.	∞ to 2	\sim FUV	Major QSO line
$\overline{\text{[OIII } \lambda 5007/\text{ H}\beta]}$				"BPT" diagram reliable tool for determining sou
[NII λ 6583/ H α]				of line emission from a galaxy visually differentia
				between Seyferts, LINERs and SF gals. However
				"low" redshifts since need $H\alpha$, (not at $z \sim 1$).
				Modified BPT with $(U - B)$ colour replacing
				[NII λ 6583/ H α] e.g. Montero-Dorta, 0801.2769
$\overline{\text{[SII } \lambda \text{ 6583/ H}\alpha]}$		∞ to 2	\sim FUV	Major QSO line
α/Fe	3646.	∞ to 2	\sim FUV	Major QSO line
NV	1???.67	2 to 1	\sim FUV	Major QSO line
SiIV	1???.67	2 to 1	\sim FUV	Major QSO line
CIV	1???.67	2 to 1	\sim FUV	Major QSO line
CIII]	1???.67	2 to 1	\sim FUV	Major QSO line
m MgII	1???67	2 to 1	\sim FUV	Major QSO line