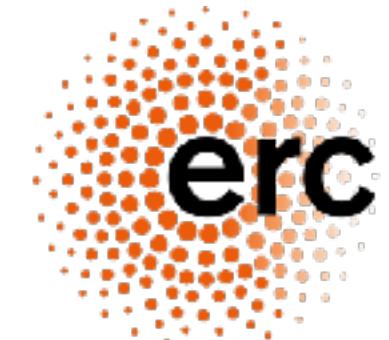


# Viable, physical and descriptive models of the early Universe

Martina Gerbino (INFN Ferrara)  
CosmoForward - 9 Feb 2026



European Research Council  
Established by the European Commission

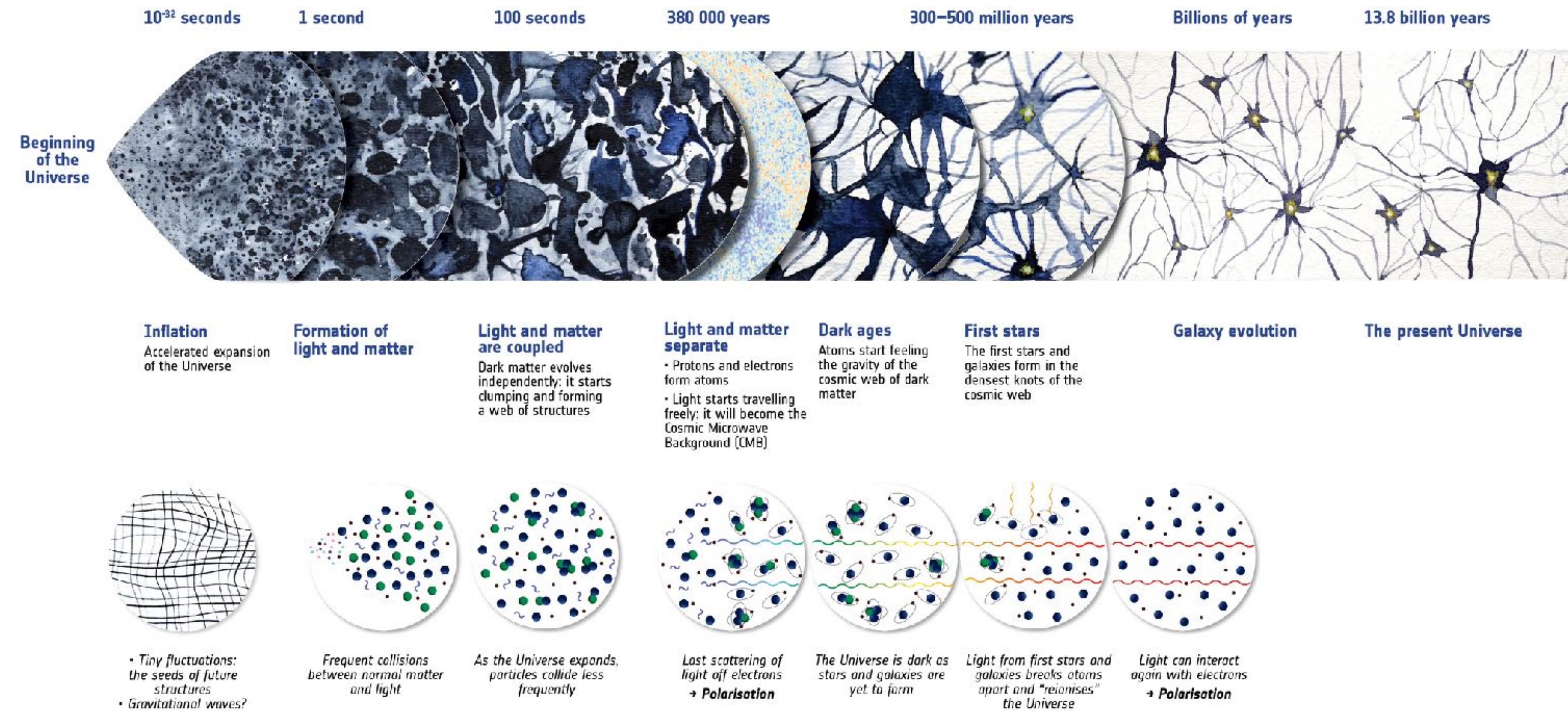


Istituto Nazionale di Fisica Nucleare



## → COSMIC HISTORY

eesa



# Where we are



**Backbones:** General Relativity; Standard Model of Particle Physics; Jeans instability; Inflation

**Observational pillars:** Light elements abundance; CMB; tracers of LSS

**Coherent, satisfying and predictive picture of the observable Universe**

# Where we are



**LCDM is phenomenological: key unknowns at the fundamental level**

**LCDM is unsatisfying: statistically significant discrepancies between experiments**

# Where we (should?) go

Beyond LCDM: new routes to better describe the Universe; to better use the Universe as a laboratory



Where and how to look for new directions?



# Where we (should?) go

Preferred directions (my taste)

**VIABLE**: obey observational constraints - not only cosmological

**PHYSICAL**: tell something about underlying physics and/or not used beyond range of applicability

**DESCRIPTIVE**: explain/predict what I observe/will observe  
(no interest in models that only live in regions beyond observational reach)



A few test cases:

- add new content
- add new interactions
- different declinations of the above

# New content: light species

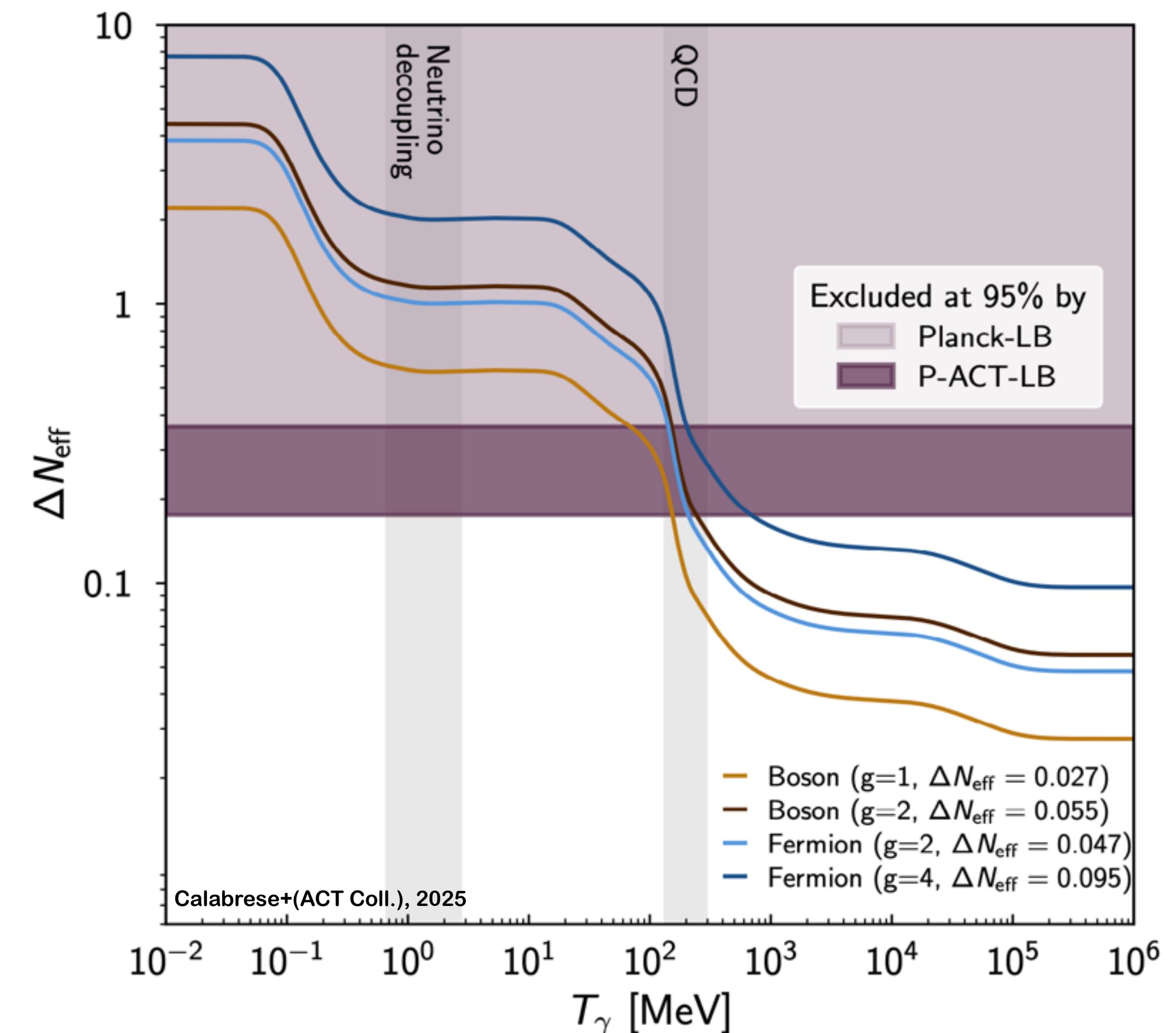
Add new content:

$N_{\text{eff}}$  as a phenomenological box (bottom-up)

Very good proxy of BSM physics at play  
at BBN, CMB and in-between times

Clearly viable, physical enough, descriptive enough

$$\Omega_\gamma \rightarrow \Omega_\gamma \left[ 1 + \frac{7}{8} \left( \frac{4}{11} \right)^{\frac{4}{3}} \Delta N_{\text{eff}} \right]$$



# Background modified: light species

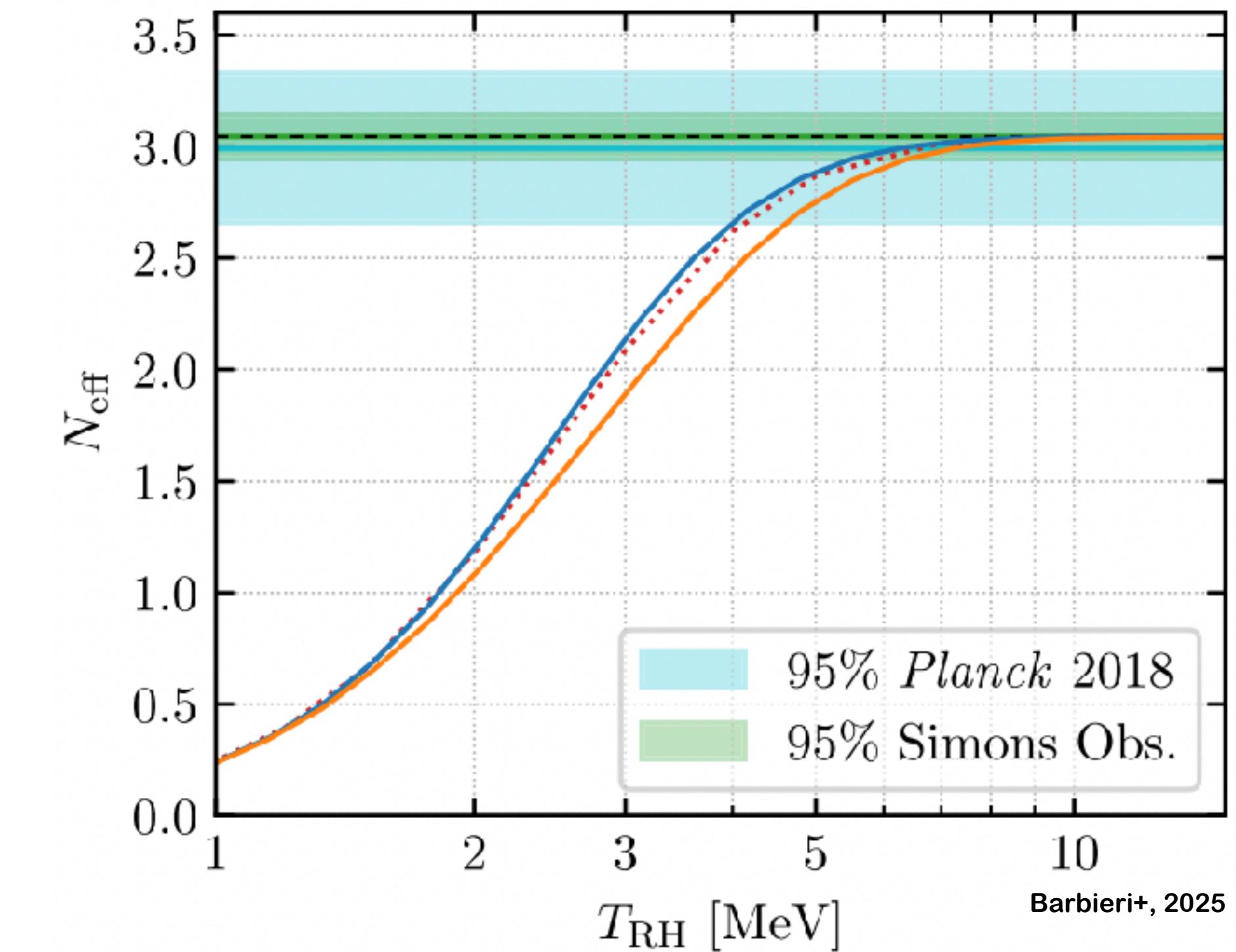
Modify the background:

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# New content: light species

Neff as a more physical and descriptive model: top-down  
 Abandon catch-all “cozyness”

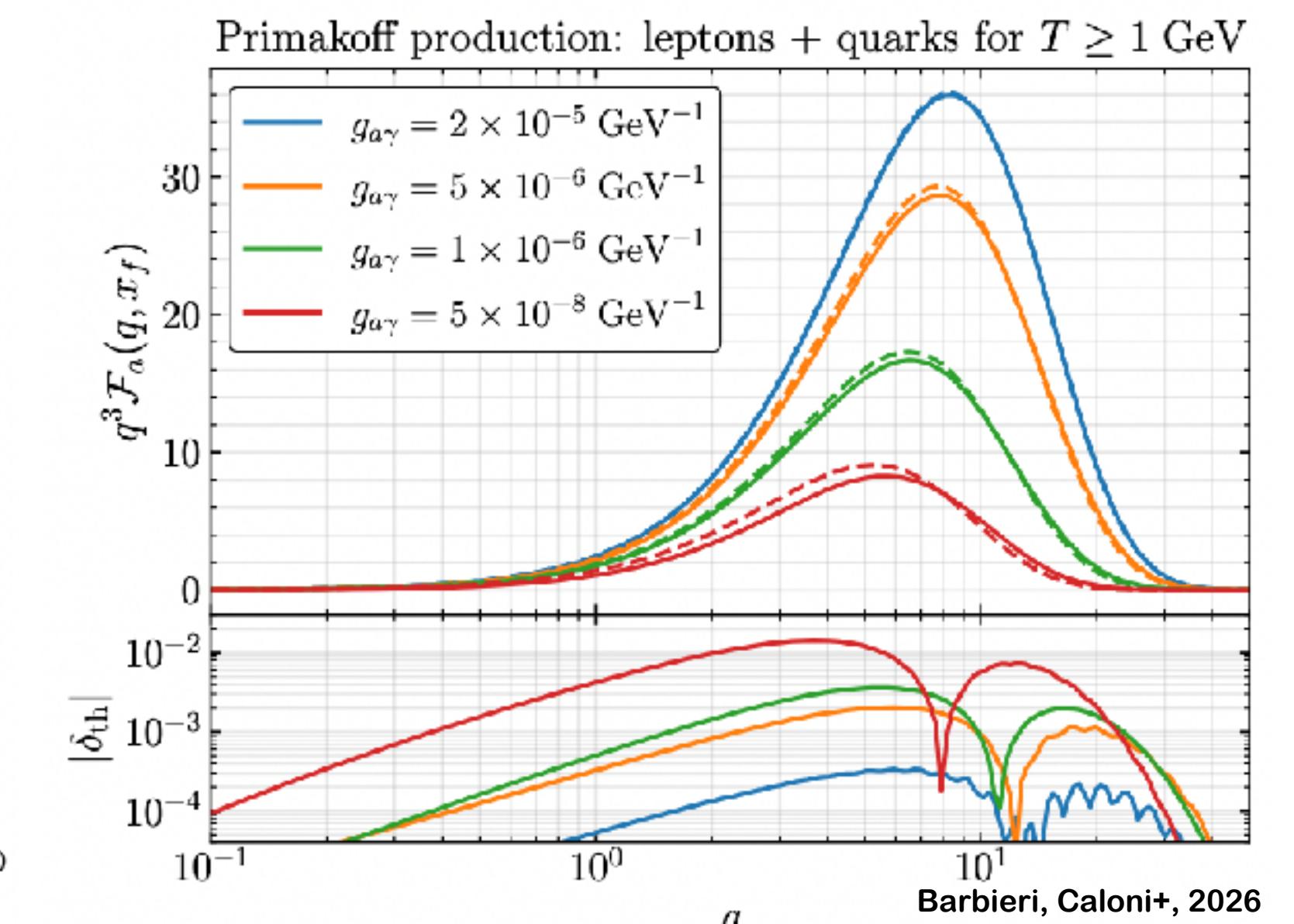
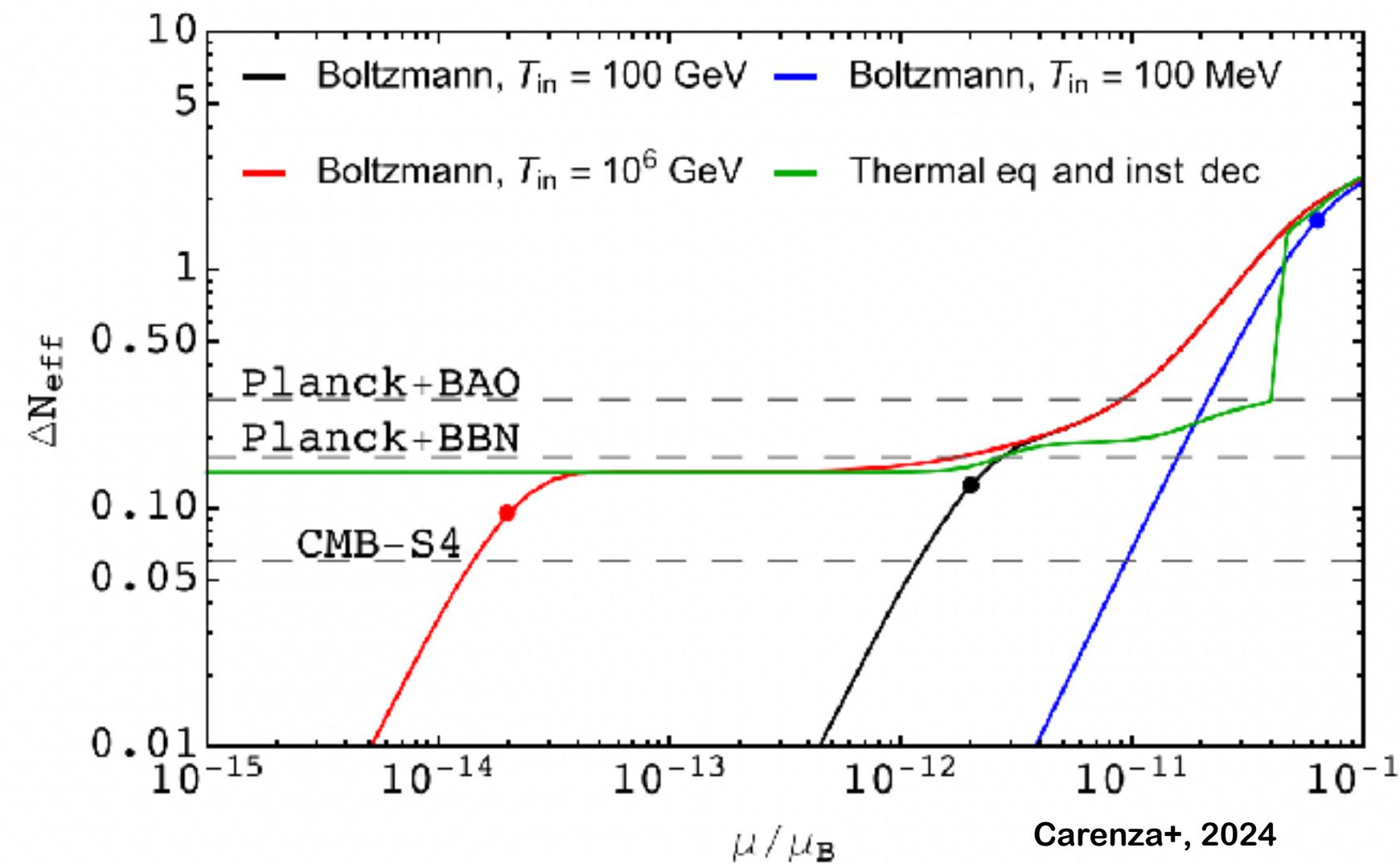
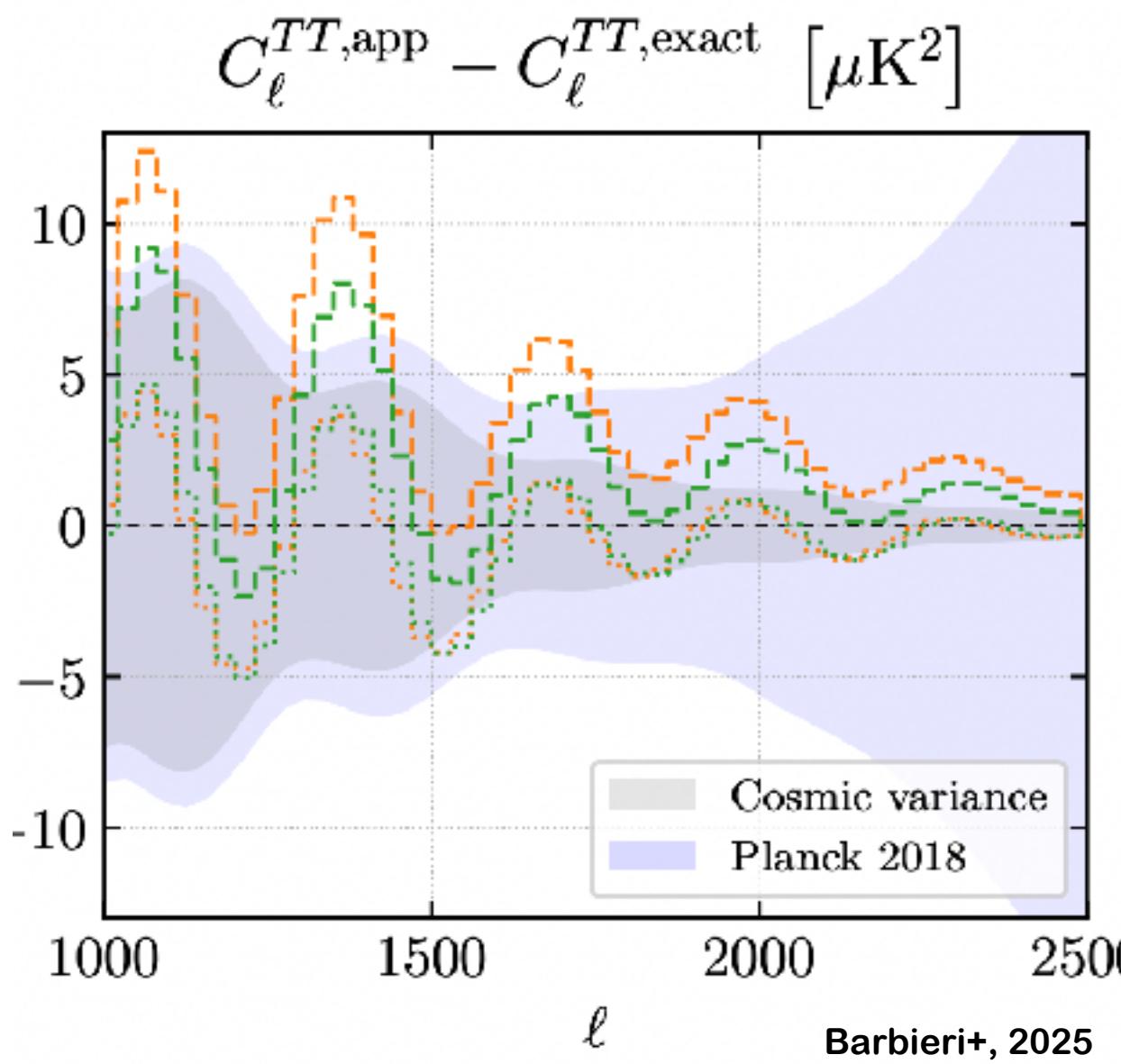
Question: real need for top-down or ok to map to individual models a posteriori?

Your favourite HEP model →

$$\mathcal{L}_a \supset \frac{1}{2}(\partial^\mu a)(\partial_\mu a) - \frac{1}{2}m_a^2 a^2 + \sum_\ell c_\ell \frac{\partial_\mu a}{2f_a} \bar{\ell} \gamma^\mu \gamma^5 \ell + \frac{1}{4}g_{a\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

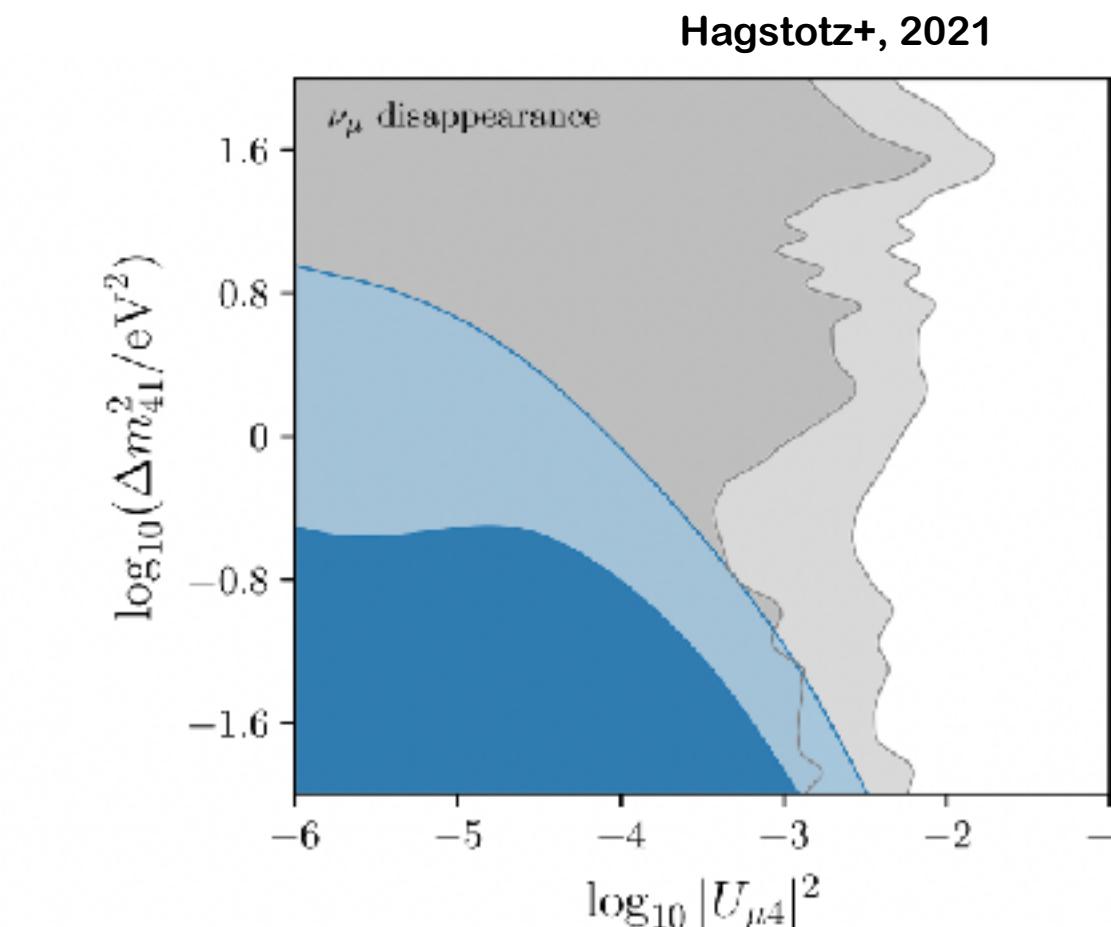
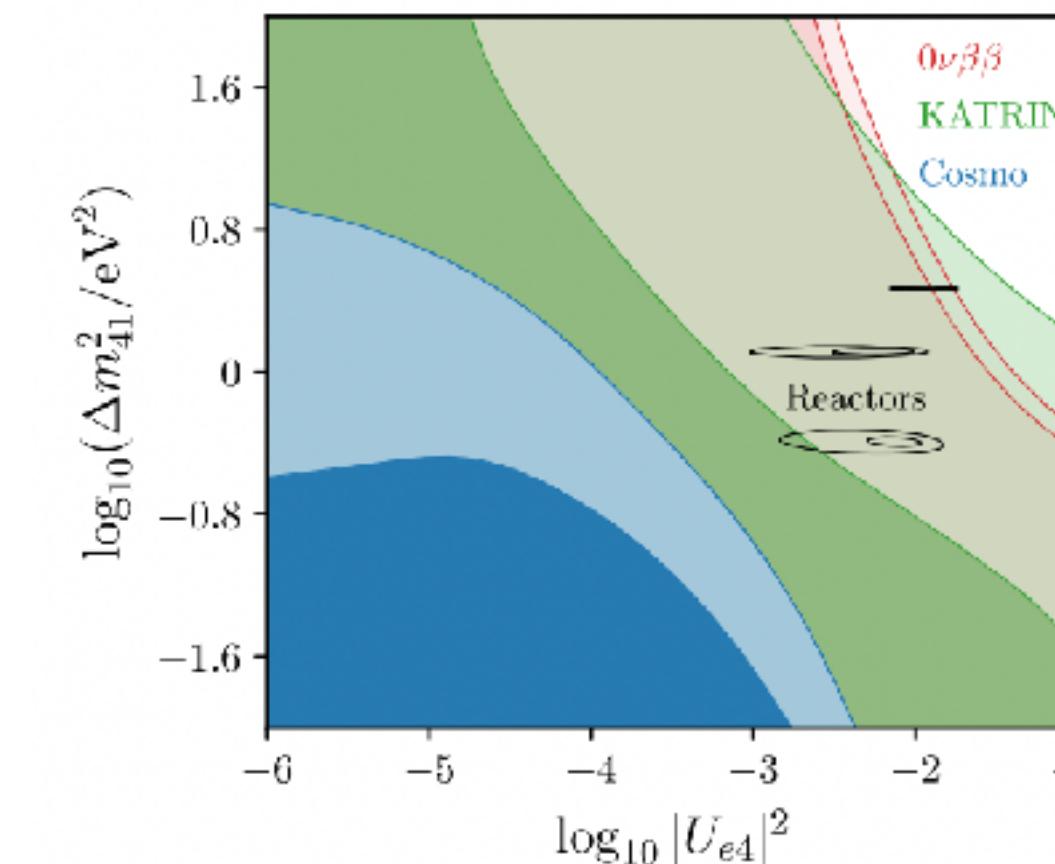
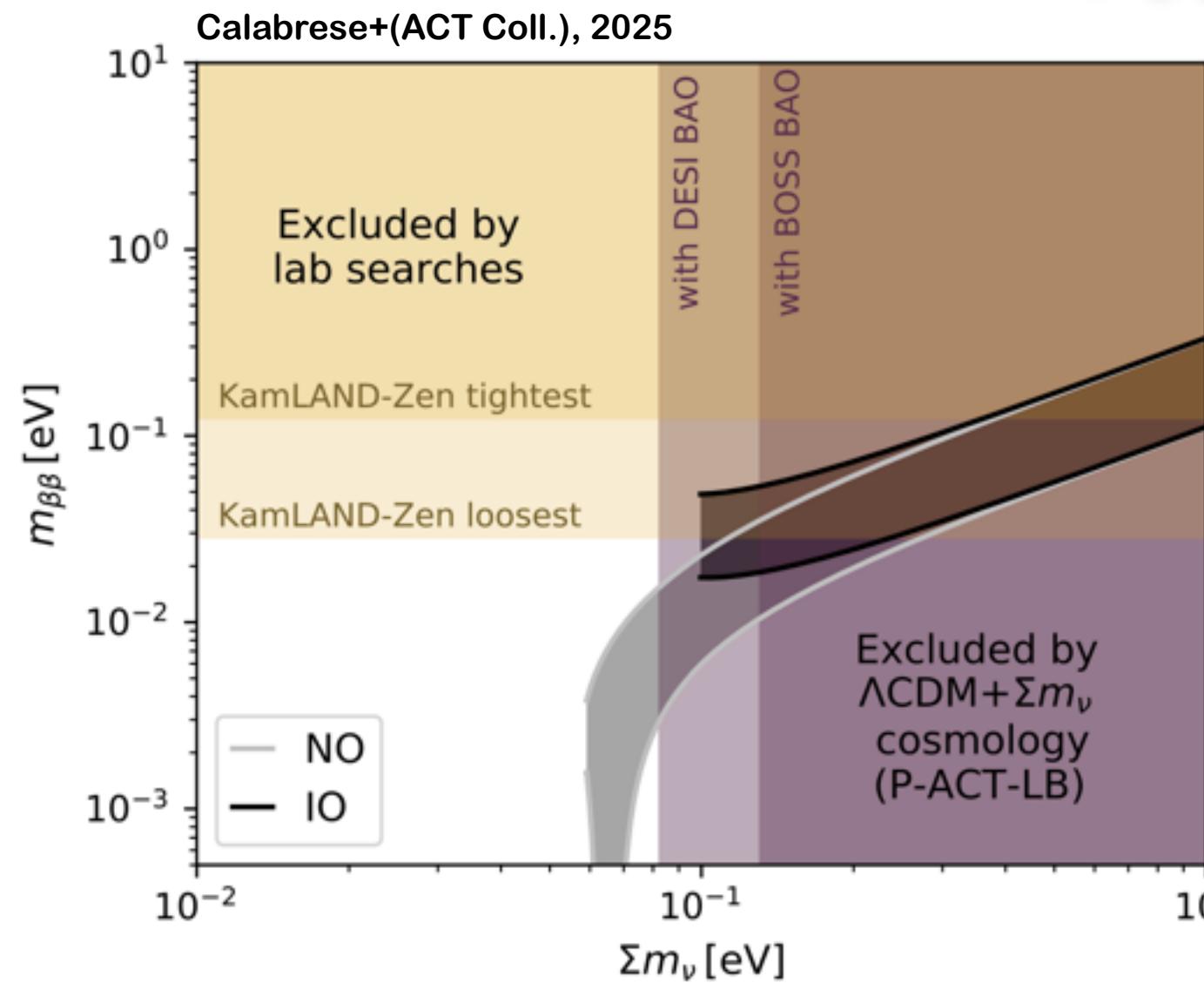
↓↑

$$\Delta N_{\text{eff}}$$



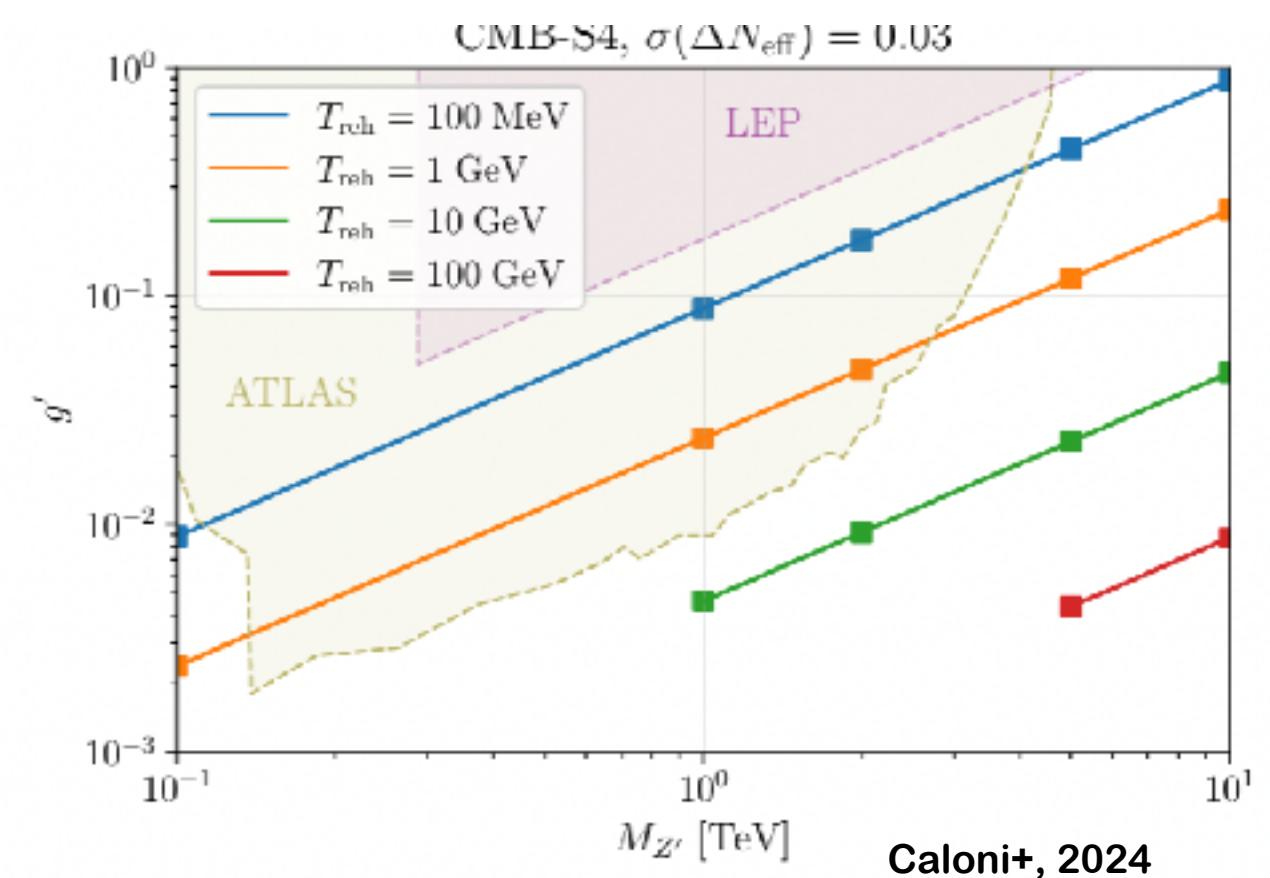
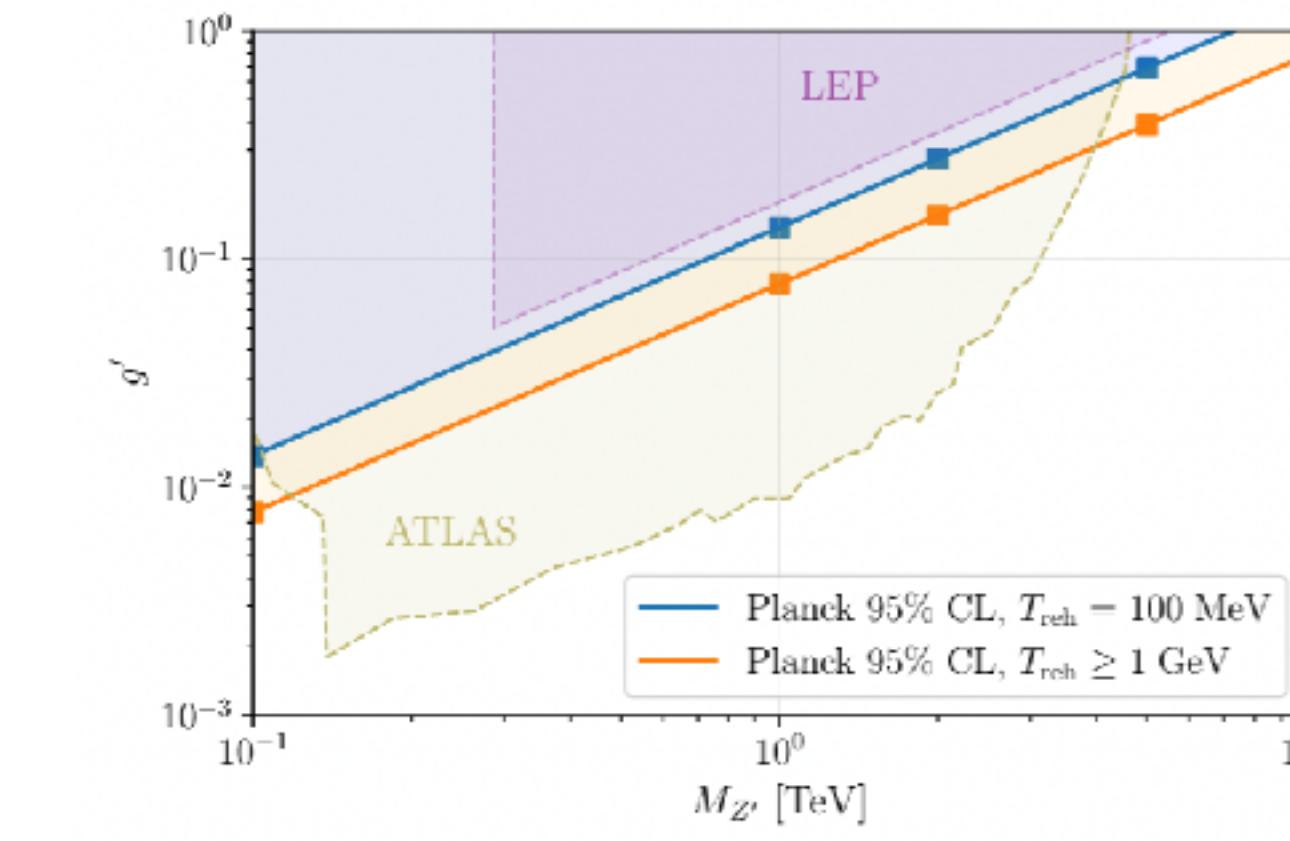
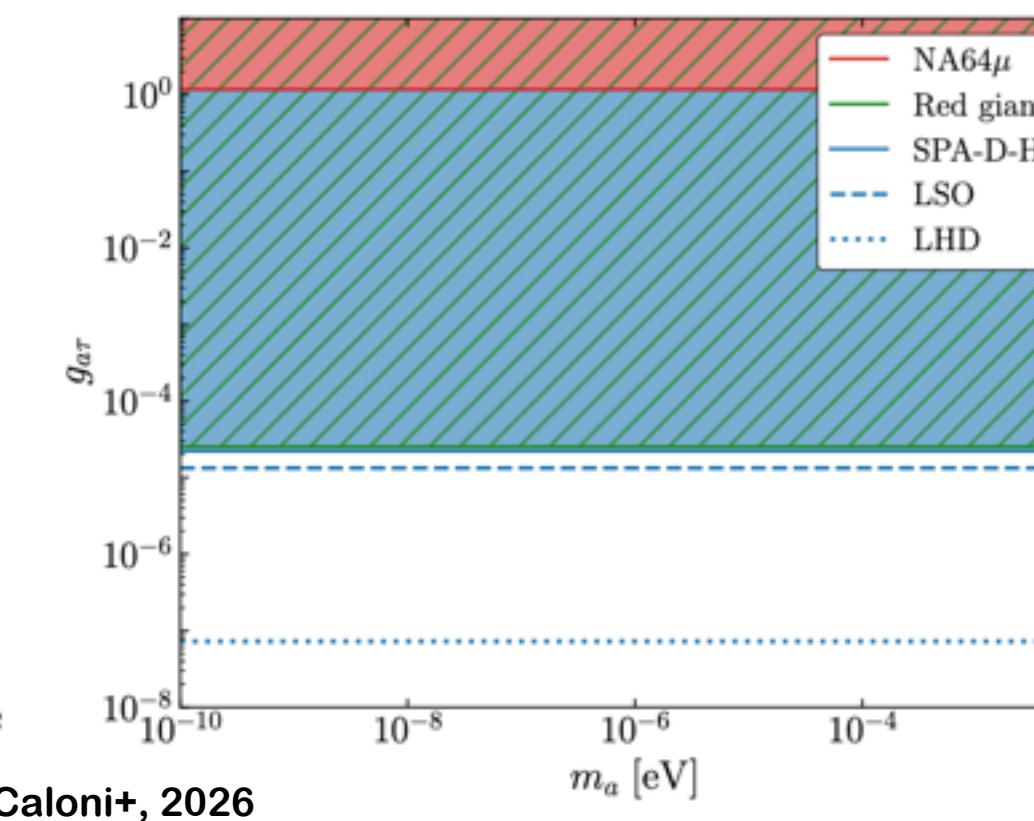
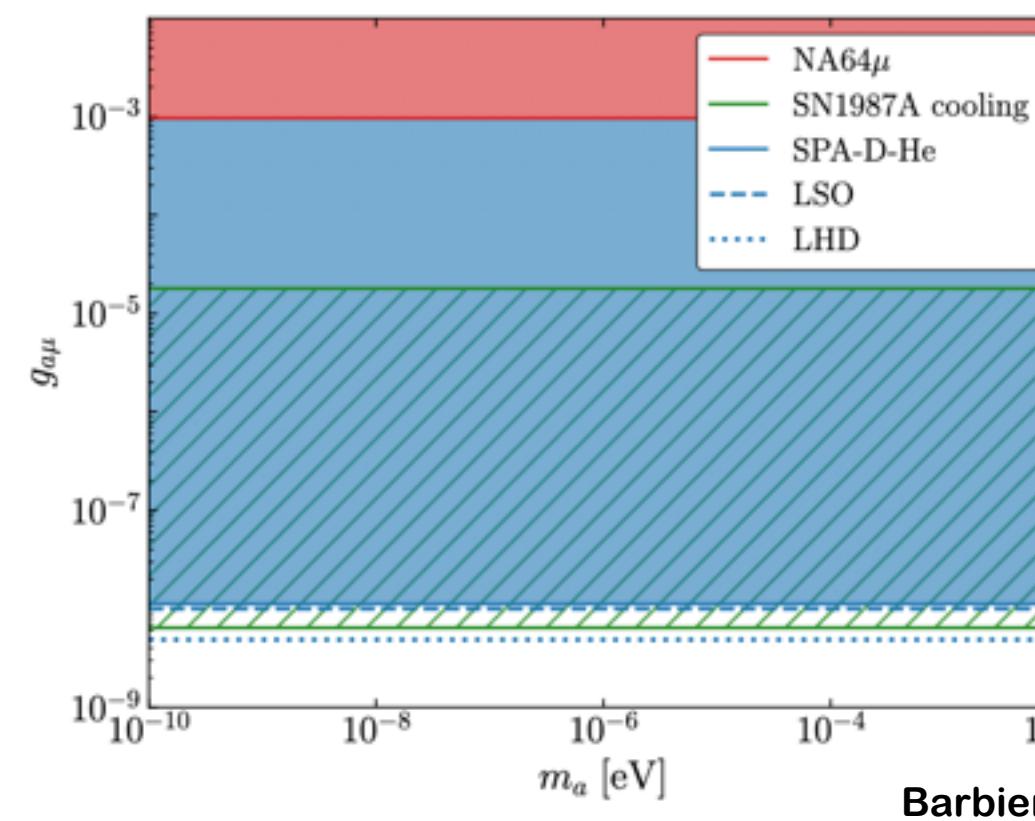
# New content: light species

See Massi's talk tomorrow



## Descriptive:

The synergy with non-cosmo constraints is key in supporting/ruling out models  
(sterile neutrinos, ALPs, ...)



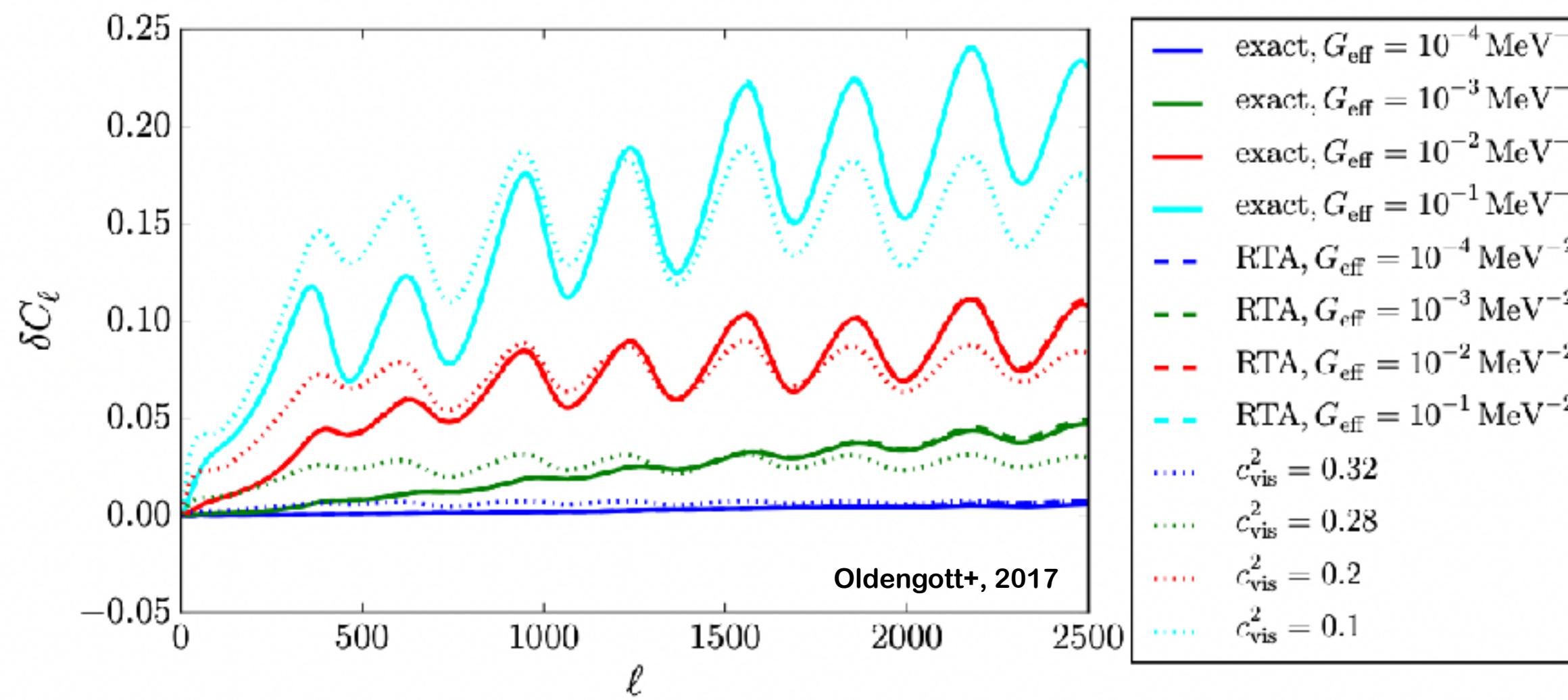
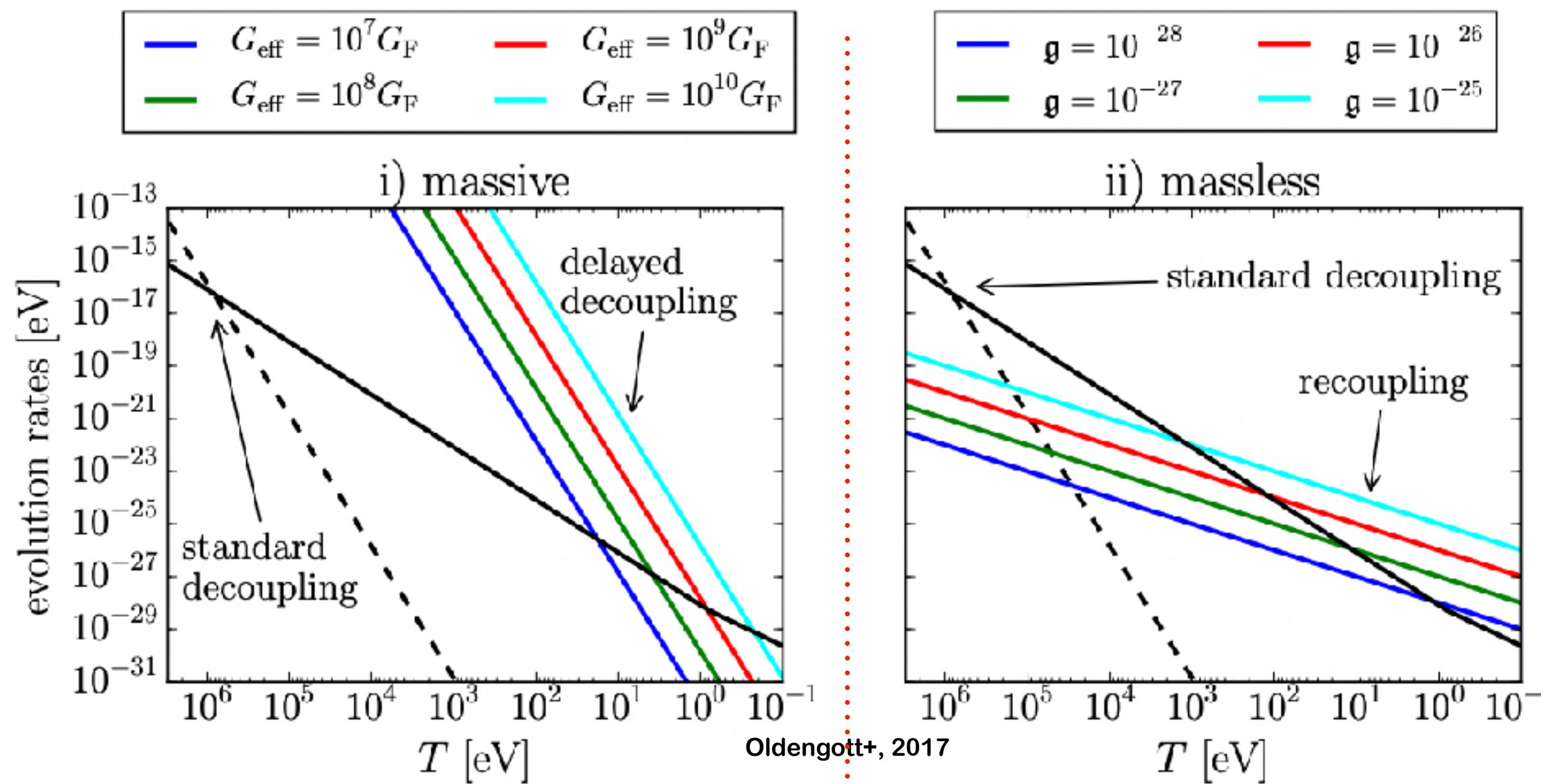
# New interactions: neutrino NSI

$$\Gamma_{\text{NSI}} = G_{\text{eff}}^2 T^5$$

$$\log_{10}(G_{\text{eff}} \text{ MeV}^2) \lesssim -2.1 \quad \text{MI}\nu$$

$$\log_{10}(G_{\text{eff}} \text{ MeV}^2) = -1.33^{+0.21}_{-0.14} \quad \text{SI}\nu$$

**ACT+Planck+DESI**  
**LCDM+NSI**  
**Calabrese+, 2025**



**Detailed comparison still missing  
due to more complicated physics scenario  
and evolution equations**

$$\Gamma_{\text{NSI}} \propto g_{\text{eff}}^4 T$$

$$g_{\text{eff}}^4 < 1.3 \times 10^{-27}$$

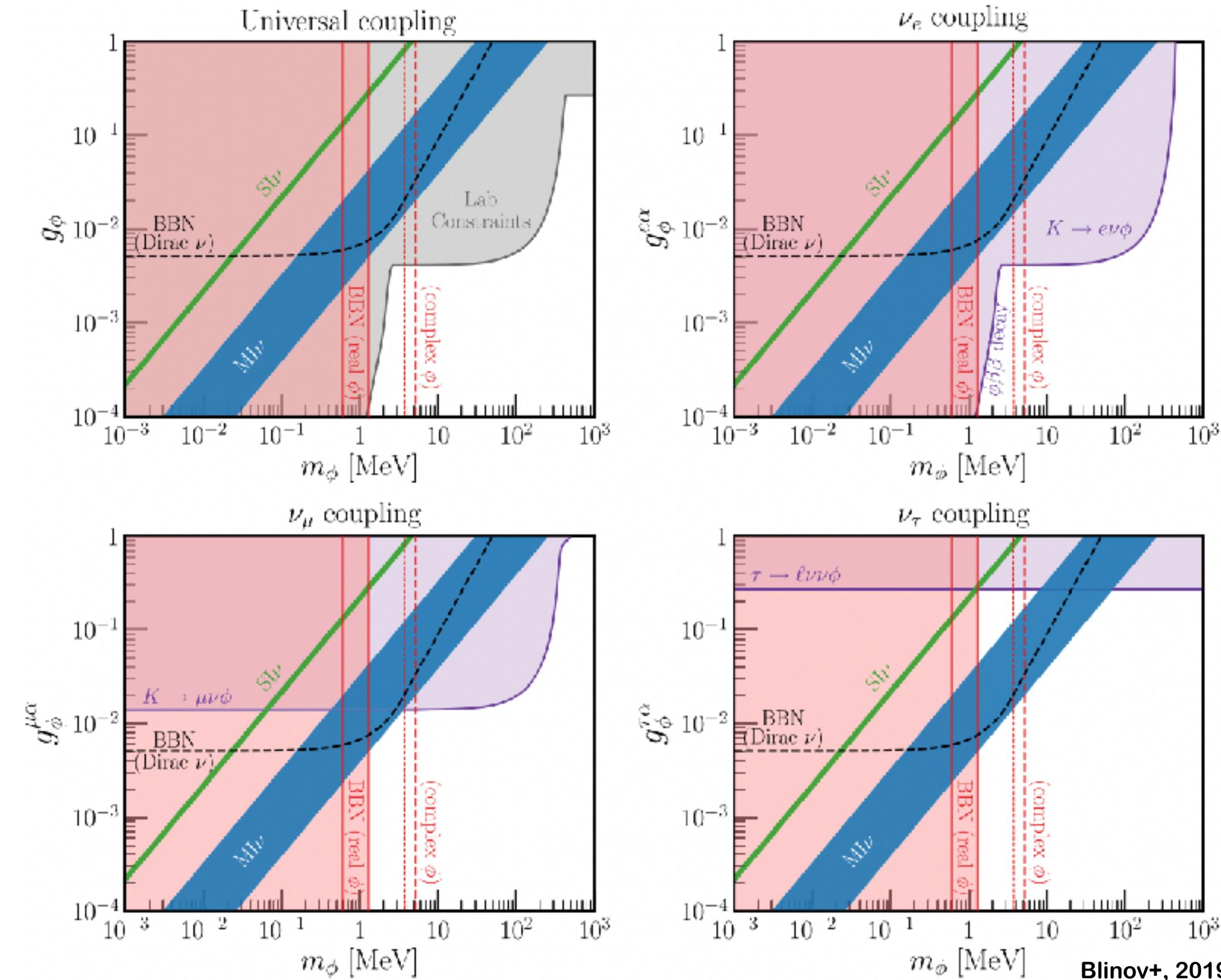
**ACT+Planck+DESI**  
**LCDM+NSI**  
**Calabrese+, 2025**

# New interactions: neutrino NSI

See Massi's talk tomorrow

## Physical and descriptive

Consider assumptions in the model: coupling, neutrino “spectrum”, etc



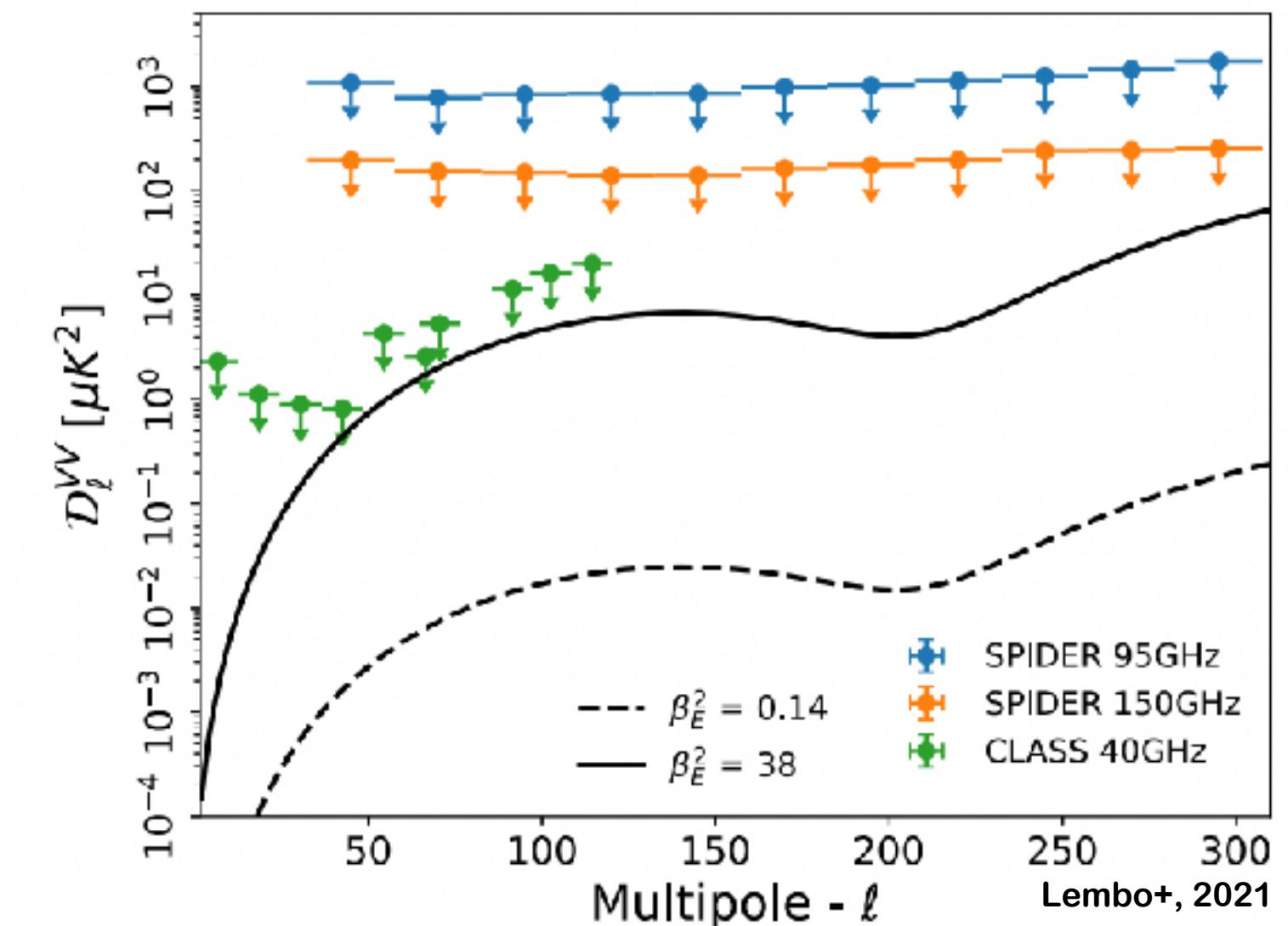
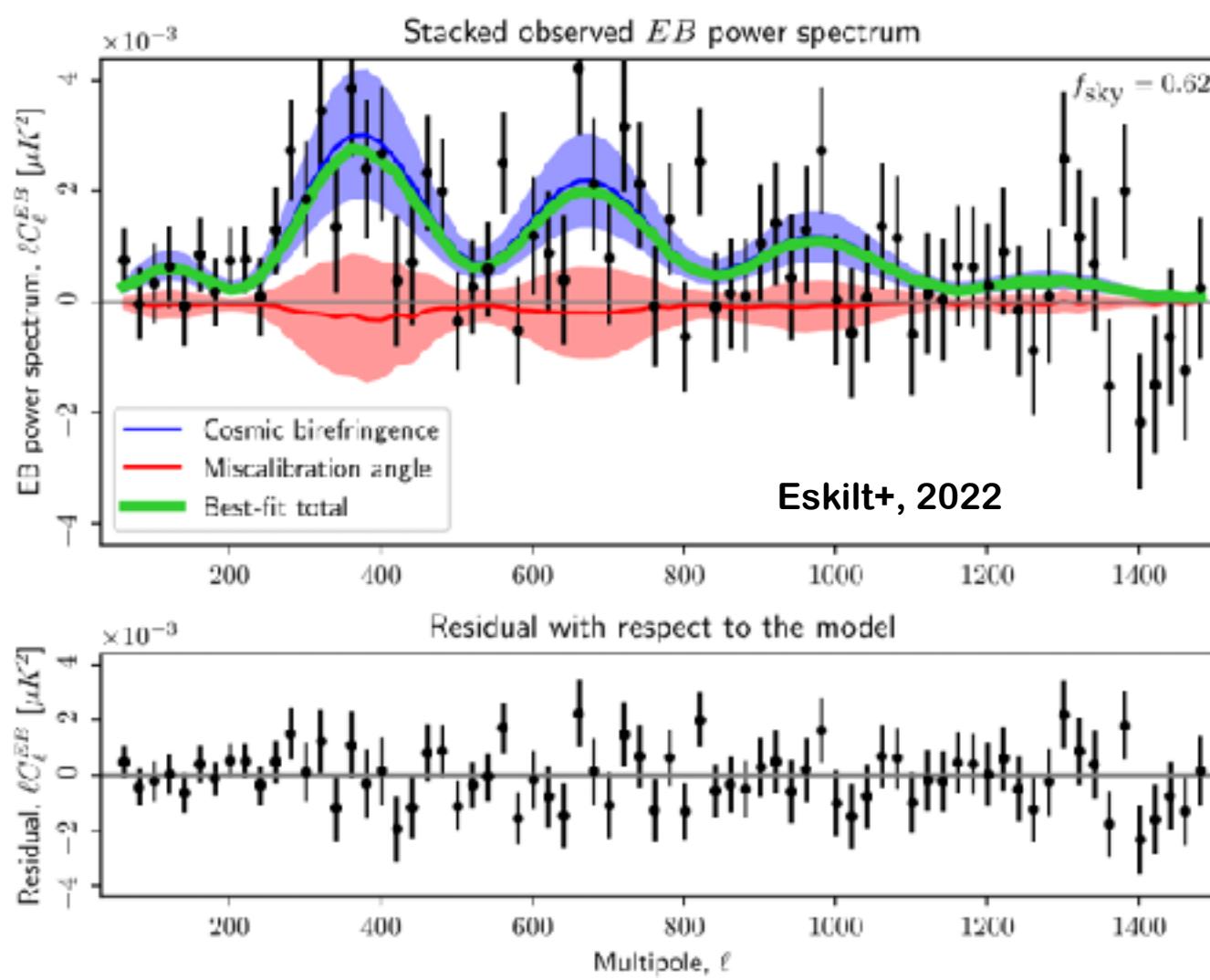
Blinov+, 2019

Viable  
Comparison with lab is key

# New interactions: BSM EM

$$\frac{dS}{ds} = \rho \wedge S, \quad S = (Q, U, V)$$

BSM EM interactions can induce several cosmological consequences:  
birefringence, circular polarisation, Faraday rotation, ...



## Viable

Models can be tested and constrained. Some appear to be BSM hints...  
Very good complementarity with astro/lab

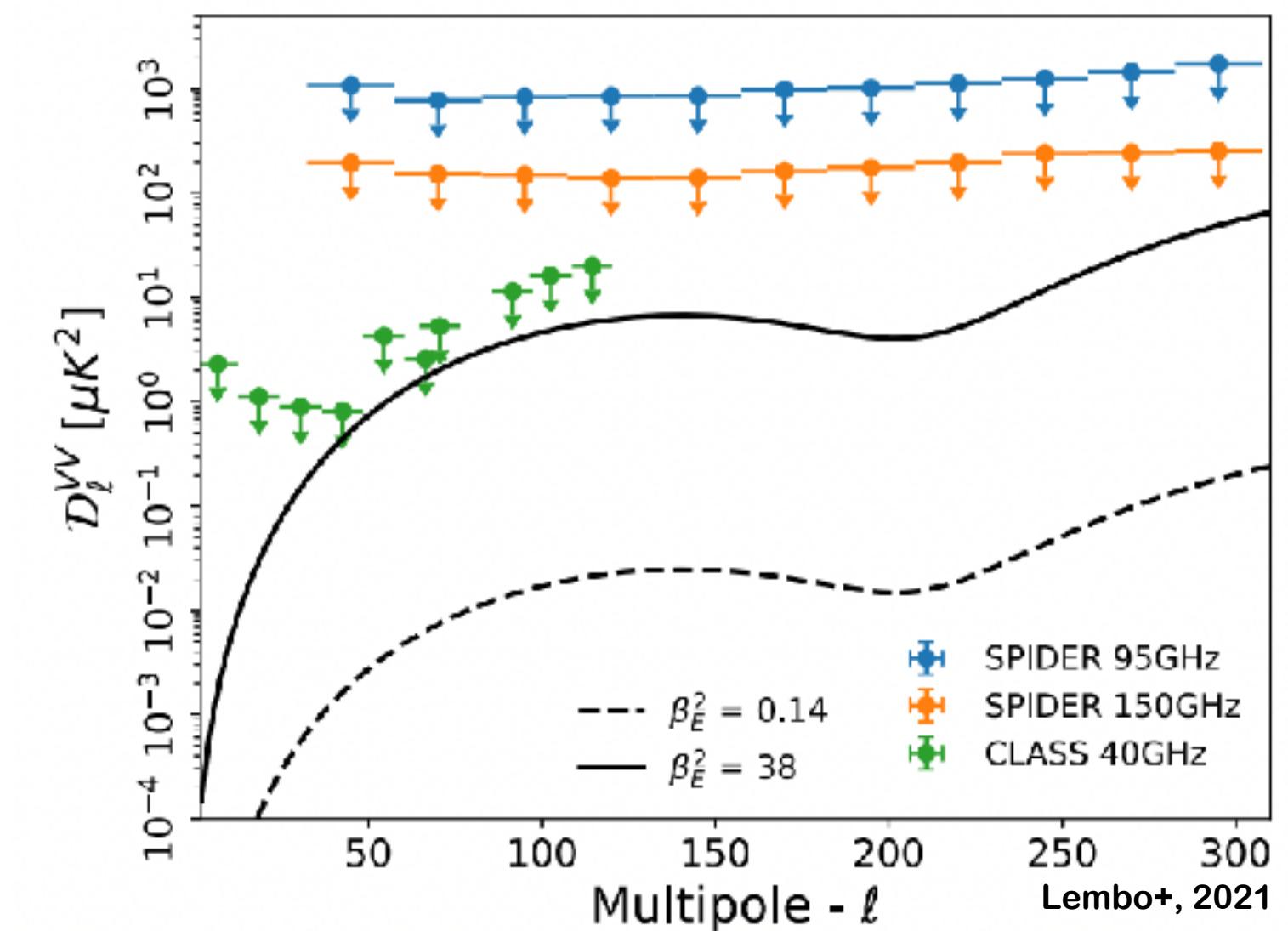
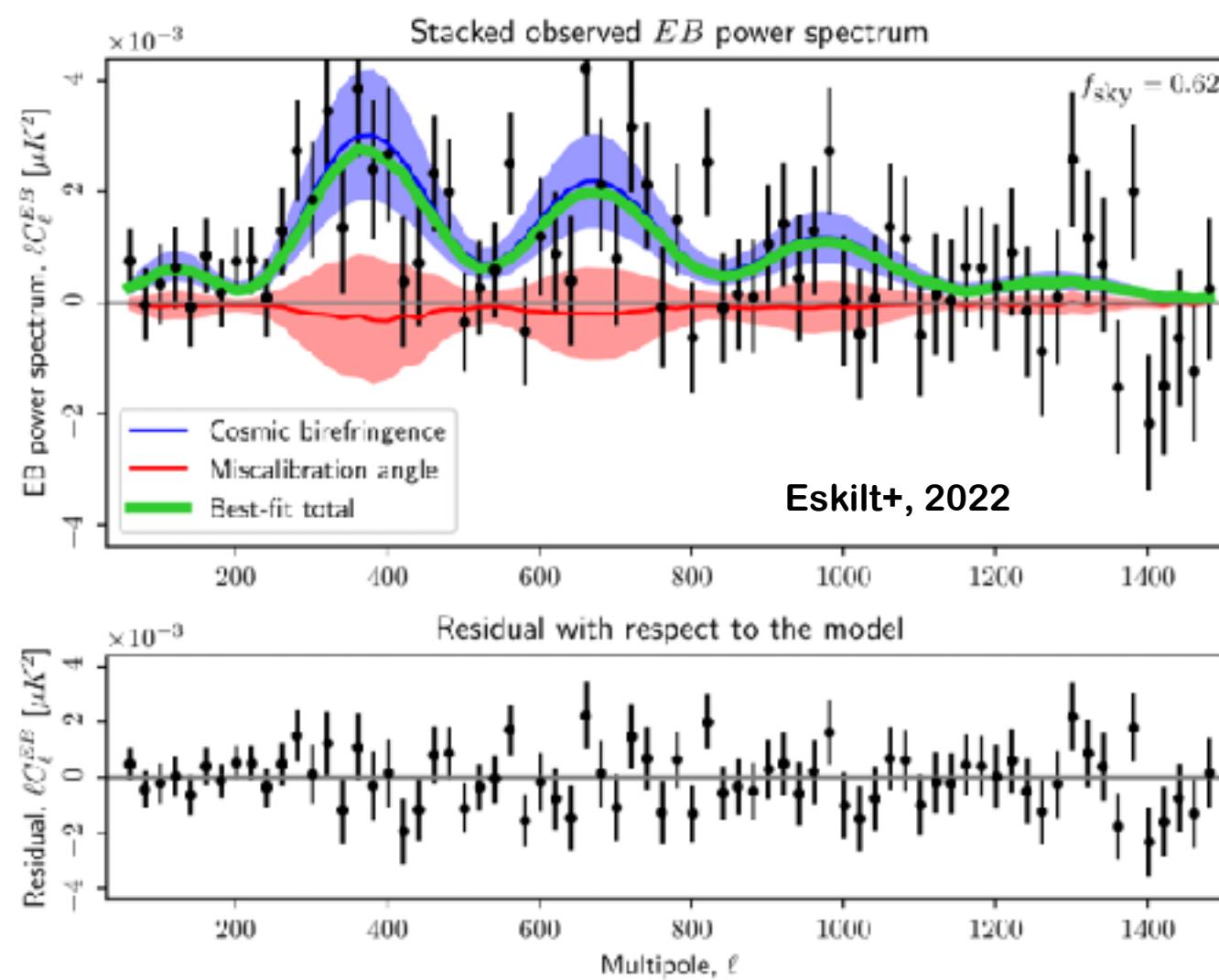
## Physical and descriptive

Phenomenological approach can/must be linked to high-energy models (Chern-Simons, QG, ...)  
New observational windows (cross-corr, V-mode searches)

# New interactions: BSM EM

$$\frac{dS}{ds} = \rho \wedge S, \quad S = (Q, U, V)$$

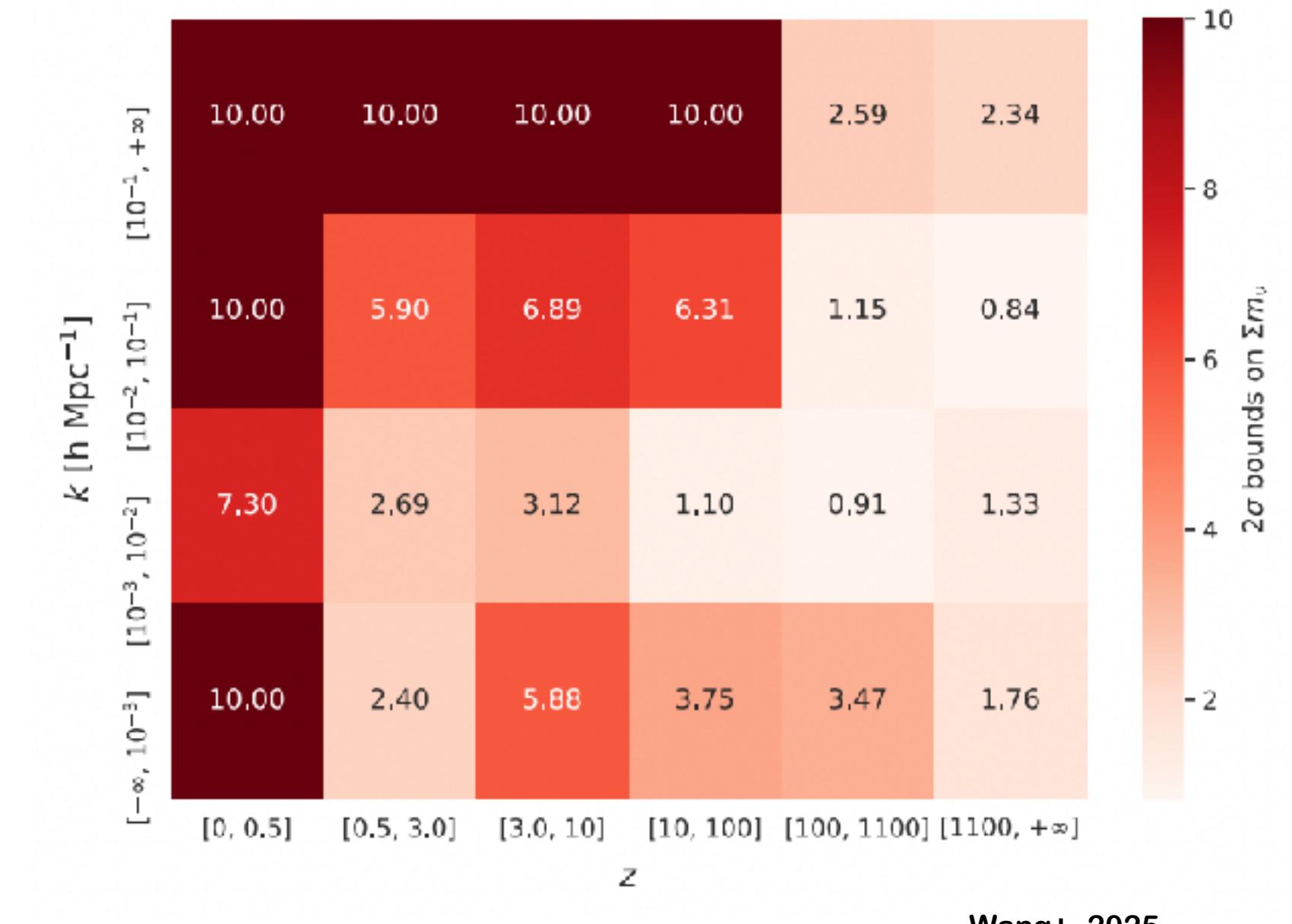
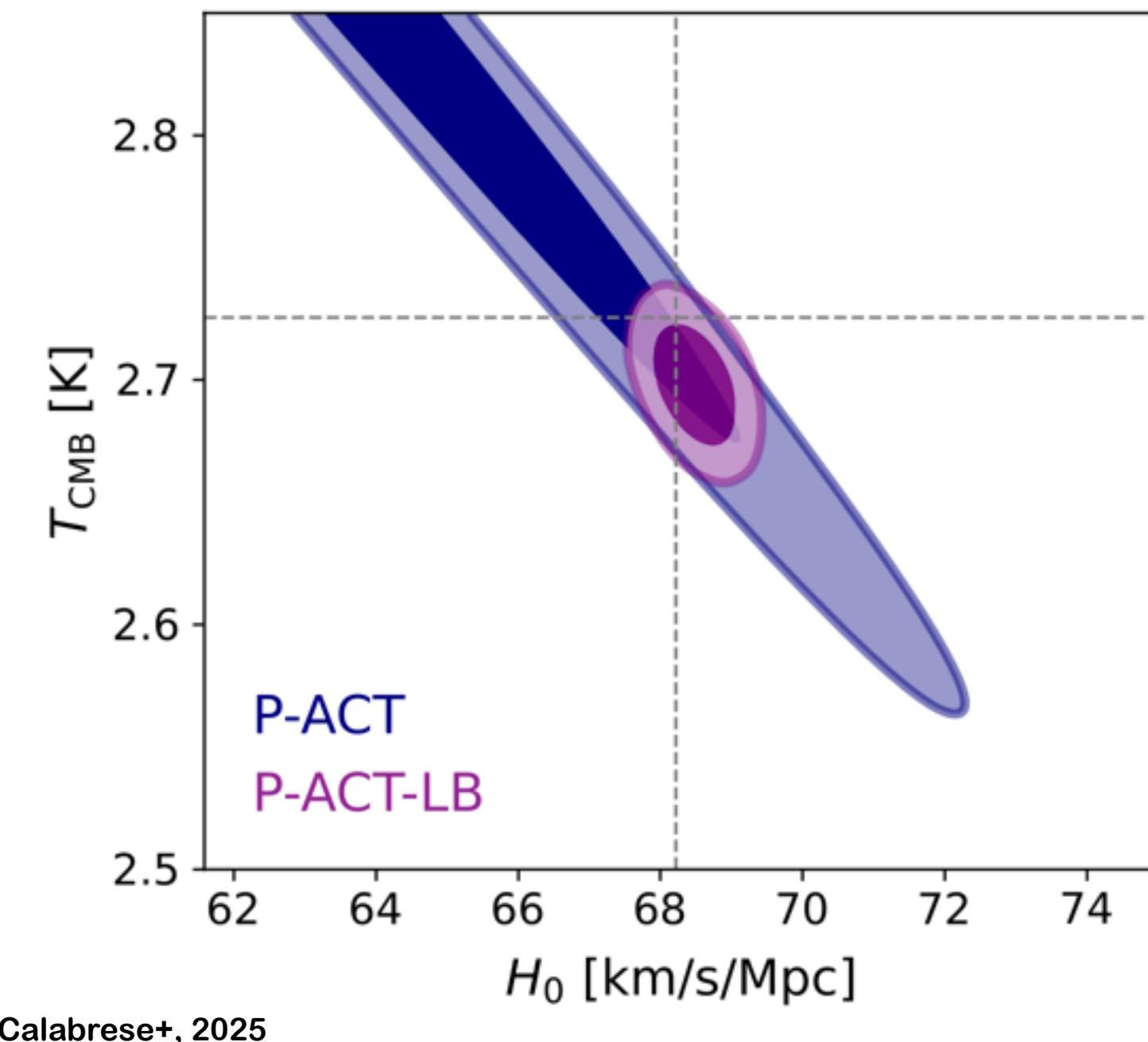
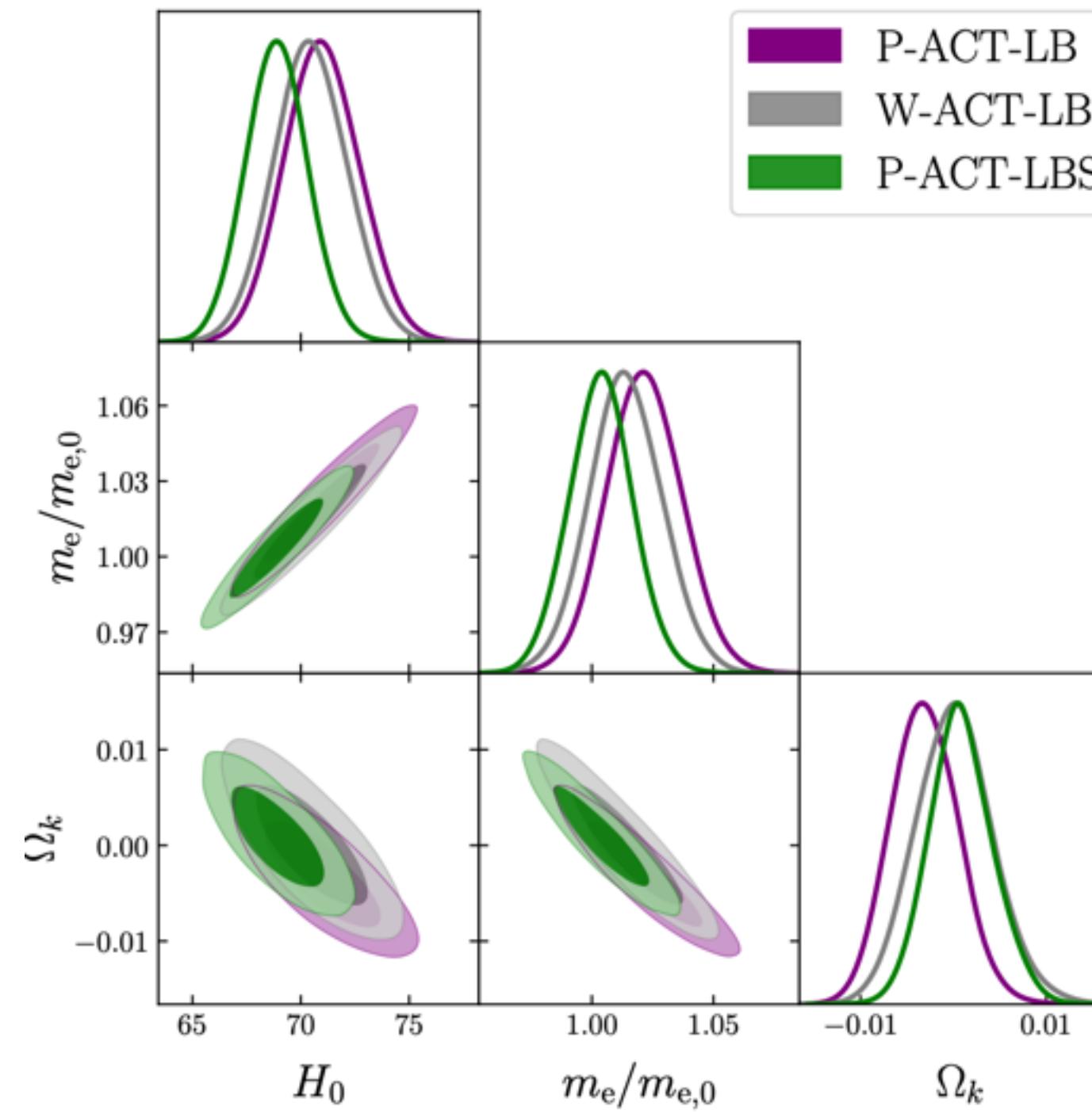
BSM EM interactions can induce several cosmological consequences:  
birefringence, circular polarisation, Faraday rotation, ...



**Key question is:**  
**What if we the signal is confirmed? What is it telling to us?**  
**Link to HEP and/or external evidence cannot be avoided**

# (Kind of) New interactions

Phenomenological variation of fundamental constants/masses/TCMB



## Physical and descriptive

**Only if tied to specific high-energy model** (i.e., cannot exist only as a solution to H0 tension)

Non-phenomenological implementation sometimes very different from pheno approach

Useful as a diagnostic tool

# Conclusions

Sure thing: we need to explore beyond-LCDM territory

Exploration can/must be phenomenological and data-driven (i.e, predictive)

Conclusions must be physical: link to fundamental physics must exist (viable at large)

Synergy with complementary probes is key to confirm/rule out findings

