Compiled: 22:46 Tuesday 27<sup>th</sup> August, 2024 (UTC) DRAFT VERSION TUESDAY 27<sup>TH</sup> AUGUST, 2024 Typeset using IAT<sub>E</sub>X **modern** style in AASTeX631

# SAUNAS: III. X-ray Scaling Relations of Diffuse Hot Gas Galactic Halos\*

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## ABSTRACT

Put the abstract here.

### 1. BACKGROUND

Here is the intro section (Aguerri et al. 1998; Bell et al. 2006b,a).

The first sentence in the intro!

1

The second sentence in the intro (Erwin & Debattista 2017; Falchi et al. 2016).

A 3rd sentence in the background.

1

Here is a comment, see if it changes color like it is supposed to.

#### 1.1. Motivation

The first sentence under motivation. (Bernardi et al. 2006)

1.2. more motivation

second sentence under motivation.

Third sentence right here right NOW!!

4th sentence

5th sentence here

1.2.1. *Goals* 

First sentence in this section underneath Goals.

2. METHODS

First sentence in methods section.

Facilities: Chandra

<sup>\*</sup> Released on May, 11th, 2023

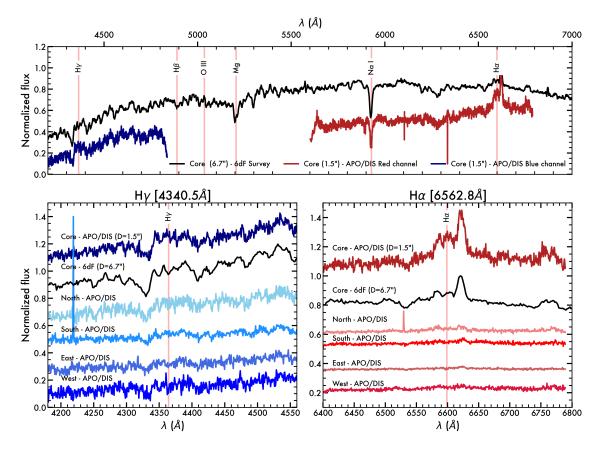


Figure 1. Optical spectral energy distribution (SED) of NGC 5084 as detected by the 6dF survey and the APO/DIS observations. Top panel: 4300-7000 Å spectrum of the central 1.5 arcsec slit (APO/DIS, blue and red channels, in color) and the 6.7 arcsec radius fiber (6dF, black). Bottom left: Detail of the H $\gamma$  spectral range (4200 – 4550 Å), showing the core as detected by 6dF and APO/DIS, as well as the North, South, East, and West subregions avoiding the core. See the labels on each spectra. Bottom right: Same as previous for the 6400-6800 Å(H $\alpha$ ) range. Vertical shadowed red lines represent the redshifted wavelengths of the typical absorption and emission lines in galaxies (H $\beta$ , OIII, Mg, Na I, H $\alpha$ ), for reference.

Software: CIAO, LIRA, VorBin

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