



COLLECTIVE EFFECTS

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Space Charge

So far, we have not considered the effect that particles in a bunch might have on each other, or on particles in another bunch.

Consider the effect off space charge on the transverse distribution of the beam.

If we look at the field at a radius r, we have

$$\oint \vec{E} \cdot d\vec{A} = E(2\pi rL) = \frac{Q_{encl}}{\epsilon_0} = \frac{Ne}{\sigma^2} \int_0^r re^{-r^2/2\sigma^2} dr$$

$$= Ne\left(1 - e^{-r^2/2\sigma^2}\right)$$

$$\vec{E} = \frac{Ne}{2\pi\epsilon_0 rL} \left(1 - e^{-r^2/2\sigma^2}\right) \hat{r}$$





Similarly, Ampere's Law gives

$$\oint \vec{B} \cdot d\vec{l} = 2\pi r B = \mu_0 I_{enclosed} = \mu_0 \frac{Nev}{\sigma^2 L} \int_0^r r e^{-r^2/2\sigma^2} dr$$

$$\longrightarrow \vec{B} = \mu_0 \frac{Nev}{2\pi r L} \left(1 - e^{-r^2/2\sigma^2}\right) \hat{\theta}$$

$$\vec{F} = e(\vec{E} + \vec{v} \times \vec{B}) = -\hat{r}$$

$$= \frac{Ne^2}{2\pi L} \left(1 - e^{-r^2/2\sigma^2}\right) \left(\frac{1}{\epsilon_0} \hat{r} + v^2 \mu_0 \left(\hat{s} \times \hat{\theta}\right)\right)$$

$$= \frac{1}{\epsilon_0} (\epsilon_0 \mu_0) = \frac{1}{\epsilon_0} \frac{1}{c^2}$$

$$= \hat{r} \frac{Ne^2}{2\pi r L \epsilon_0} \left(1 - e^{-r^2/2\sigma^2}\right) \left(1 - \beta^2\right)$$

$$= \hat{r} \frac{ne^2}{2\pi r \epsilon_0 \gamma^2} \left(1 - e^{-r^2/2\sigma^2}\right); \quad n \equiv \frac{N}{L} = \frac{dN}{ds} \quad \text{Linear charge density}$$





We can break this into components in x and y

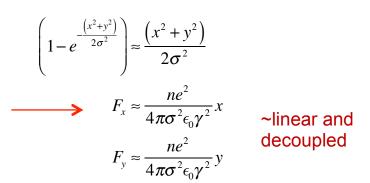
$$F_{x} = |F| \frac{x}{r}$$

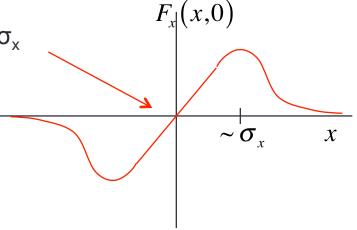
$$= \frac{ne^{2}}{2\pi\epsilon_{0}\gamma^{2}} \frac{x}{r^{2}} \left(1 - e^{-r^{2}/2\sigma^{2}}\right)$$

$$= \frac{ne^{2}}{2\pi r\epsilon_{0}\gamma^{2}} \frac{x}{\left(x^{2} + y^{2}\right)} \left(1 - e^{-\frac{\left(x^{2} + y^{2}\right)}{2\sigma^{2}}}\right)$$

$$F_{y} = \frac{ne^{2}}{2\pi r\epsilon_{0}\gamma^{2}} \frac{y}{\left(x^{2} + y^{2}\right)} \left(1 - e^{-\frac{\left(x^{2} + y^{2}\right)}{2\sigma^{2}}}\right)$$

Non-linear and coupled \rightarrow ouch! but for x<< σ_x









$$F_{x} = \frac{dp_{x}}{dt}$$

$$\Rightarrow \Delta x' = \frac{\Delta p_{x}}{p} = \frac{1}{p} \int F_{x} dt = \frac{1}{p} \int F_{x} \frac{dt}{ds} ds$$

$$= \frac{1}{vp} \int F_{x} ds$$

$$= \frac{r_{0}}{\beta^{2} \gamma^{3} \sigma^{2}} nx;$$

$$r_{0} = \frac{e^{2}}{4\pi \epsilon_{0} m_{0} c^{2}}$$
"classical radius"
$$= 1.53 \times 10^{-18} \text{ m for protons}$$

This looks like a distributed defocusing quad of strength

so the total tuneshift is $\Delta v_x = \frac{1}{4\pi} \oint k \beta_x(s) ds$

particles" lecture)

the total tuneshift is
$$\Delta v_{x} = \frac{1}{4\pi} \oint k\beta_{x}(s) ds$$

$$= -\frac{r_{0}}{4\pi\beta^{2}\gamma^{3}} \oint n\frac{\beta_{x}(s)}{\sigma_{x}^{2}} ds = -\frac{r_{0}}{4\pi\beta^{2}\gamma^{3}\epsilon} \oint n ds$$

$$= -\frac{r_{0}}{4\pi\beta^{2}\gamma^{3}} \frac{NB}{\epsilon_{x}}; \quad B = \frac{n_{peak}}{\langle n \rangle}$$
"Bunching factor"
$$= -\frac{NBr_{0}}{4\pi\beta\gamma^{2}L} \frac{L}{(\beta\gamma\epsilon_{x})}$$

$$= -\frac{NBr_{0}}{4\pi\beta\gamma^{2}L} \frac{L}{(\beta\gamma\epsilon_{x})}$$
Maximum tuneshift for particles near core of beam





Example: Fermilab Booster@Injection

$$K = 400 \text{ MeV}$$

$$N=5 \times 10^{12}$$

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$$\epsilon_N = 2 \pi$$
-mm-mr

B=1 (unbunched beam)

$$\Delta_{v} = -\frac{Nr_0}{4\pi\beta\gamma^2\epsilon_N} = -.247$$

This is pretty large, but because this is a rapid cycling machine, it is less sensitive to resonances

Because this affects individual particles, it's referred to as an "incoherent tune shift", which results in a tune spread. There is also a "coherent tune shift", caused by images charges in the walls of the beam pipe and/or magnets, which affects the entire bunch more or less equally.

This is an important effect, but beyond the scope of this lecture.

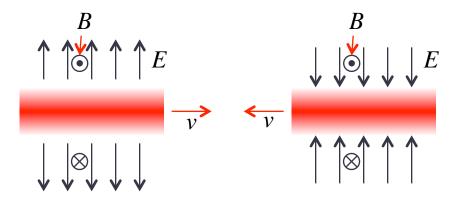




Beam-beam Interaction

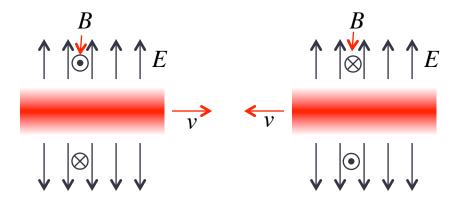
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If two oppositely charged bunches pass through each other...



Both E and B fields are *attractive* to the particles in the other bunch

If two bunches with the same sign pass through each other...



Both E and B fields are *repulsive* to the particles in the other bunch

In either case, the forces add

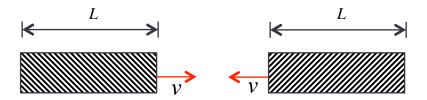




$$\vec{F} = -\hat{r} \frac{e^2}{2\pi\epsilon_0 r} \frac{N}{L} \left(1 - e^{-r^2/2\sigma^2} \right) \left(1 + \beta^2 \right)^2$$

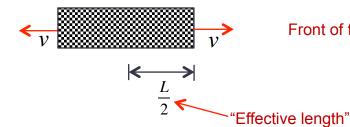
$$\approx -\hat{r} \frac{e^2}{\pi\epsilon_0 r} \frac{N}{L} \left(1 - e^{-r^2/2\sigma^2} \right)$$

Effective Length





Front of first bunch encounters front of second bunch



Front of first bunch exits second bunch.



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$$\Delta x' = \frac{F_x}{vp} \Delta s = \frac{F_x}{vp} \left(\frac{L}{2}\right)$$

$$= -\frac{N_b e^2}{2\pi \epsilon_0 r \gamma \beta^2 m c^2} \frac{x}{r} \left(1 - e^{-r^2/2\sigma^2}\right)$$

$$\approx -\frac{2N_b r_0}{\gamma} \frac{x}{r^2} \left(1 - e^{-r^2/2\sigma^2}\right)$$

$$= -\frac{2N_b r_0}{\gamma} \frac{x}{\left(x^2 + y^2\right)} \left(1 - e^{-\frac{\left(x^2 + y^2\right)}{2\sigma^2}\right)$$

$$\approx -\frac{N_b r_0}{\gamma \sigma^2} x = -\frac{1}{f_{eff}} x \quad \text{Small x and y}$$

$$\Rightarrow \Delta y' \approx -\frac{N_b r_0}{\gamma \sigma^2} y = -\frac{1}{f_{eff}} y$$

$$\Delta v = \frac{1}{4\pi} \frac{\beta^*}{f_{eff}}$$

$$= \frac{N_b r_0}{4\pi} \frac{\beta^*}{\gamma \sigma^2} = \frac{N_b r_0}{4\pi \gamma \epsilon}$$

$$= \frac{N_b r_0}{4\pi \epsilon_N} \qquad \text{normalized emittance (protons)}$$

$$\equiv \xi$$
"Tuneshift Parameter"

Maximum tuneshift for particles near center of bunch





Luminosity and Tuneshift

The total tuneshift will ultimately limit the performance of any collider, by driving the beam onto an unstable resonance. Values of on the order ~.02 are typically the limit. However, we have seen the somewhat surprising result that the tuneshift

$$\xi = \frac{N_b r_0}{2\pi\epsilon\gamma}$$

does not depend on β^* , but only on

$$\frac{N_b}{\epsilon} \equiv$$
 "brightness"

For a collider, we have

$$\mathcal{L} = \frac{f n_b N_b^2}{4\pi\sigma^2} = \frac{f n_b N_b^2}{4\pi \left(\frac{\beta^* \epsilon_N}{\gamma}\right)} = \frac{f n_b N_b \gamma}{r_0 \beta^*} \left(\frac{r_0}{4\pi} \frac{N_b}{\epsilon_N}\right)$$

$$= f \frac{n_b N_b \gamma}{r_0 \beta^*} \xi$$

We assume we will run the collider at the "tuneshift limit", in which case we can increase luminosity by

- Making β* as small as possible
- Increasing N_b and ε proportionally.