

General form of Euler Equation:

$$\frac{\partial \rho}{\partial t} + \vec{v} \cdot (\rho \vec{U}) = 0$$

$$\frac{\partial(\rho \vec{U})}{\partial t} + \vec{v} \cdot (\rho \vec{U} \otimes \vec{U}) + \vec{v} \rho = 0$$

$$\frac{\partial(\rho E)}{\partial t} + \vec{v} \cdot (\rho E \vec{U} + \rho \vec{U}) = 0$$

Then energy / mass :

$$E = e + \frac{1}{2} |\vec{U}|^2$$

With Gamma law eos: $\rho = \rho e(\gamma - 1)$

In 1D:

$$\frac{\partial \rho}{\partial t} + \frac{\partial(\rho u)}{\partial x} = 0$$

$$\frac{\partial(\rho u)}{\partial t} + \frac{\partial(\rho u^2 + p)}{\partial x} = 0$$

$$\frac{\partial(\rho E)}{\partial t} + \frac{\partial(\rho u E + up)}{\partial x} = 0$$

Write in conservative form:

$$\frac{\partial}{\partial t} u + [F(u)]_x = 0$$

$$u = \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} \quad F(u) = \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho u E + up \end{pmatrix}$$

If assume gamma law eos: $p = \rho e(\gamma - 1)$

$$F(u) = \begin{pmatrix} \rho u \\ \frac{1}{2} \frac{(\rho u)^2}{\rho} (3-\gamma) + \rho E (\gamma-1) \\ \frac{\rho^2 u E}{\rho} \gamma - \frac{1}{2} \frac{(\rho u)^3}{\rho^2} (\gamma-1) \end{pmatrix}$$

then

$$\frac{\partial \vec{F}}{\partial u} = \vec{A}(u) = \begin{pmatrix} 0 & 1 & 0 \\ -\frac{1}{2} u^2 (3-\gamma) & u(3-\gamma) & \gamma-1 \\ \frac{1}{2} (\gamma-2) u^3 - \frac{u c^2}{\gamma-1} & \frac{3-\gamma}{2} u^2 + \frac{c^2}{\gamma-1} & -u\gamma \end{pmatrix}$$

$$\text{Then } u_t + [F(u)]_x = 0$$

$$\hookrightarrow u_t + \frac{\partial F}{\partial u} \frac{\partial u}{\partial x} = 0$$

$$u_t + \vec{A}(u) u_x = 0$$

If we change to primitive variables: $q_t + \vec{A}(q) q_x = 0$

$$\vec{q} = \begin{pmatrix} p \\ -u \\ p \end{pmatrix} \quad \tilde{A}(q) = \begin{pmatrix} u & p & 0 \\ 0 & -u & \frac{1}{4}p \\ 0 & 8p & u \end{pmatrix}$$

then the eigenvalues of $\tilde{A}(q)$: $|\tilde{A} - \lambda \tilde{I}| = 0$ are:

$$\lambda = \begin{pmatrix} u-c \\ u \\ u+c \end{pmatrix}$$

We find the left and right eigenvectors:

$$\vec{L} = \begin{pmatrix} l^{(c)} \\ l^{(0)} \\ l^{(+)} \end{pmatrix} = \begin{pmatrix} 0 & \frac{-p}{2c^2} & \frac{1}{2c^2} \\ 1 & 0 & \frac{1}{c^2} \\ 0 & \frac{p}{2c^2} & \frac{1}{2c^2} \end{pmatrix}$$

$$\vec{R} = \begin{pmatrix} r^{(-)} & r^{(0)} & r^{(+)} \end{pmatrix} = \begin{pmatrix} 1 & 1 & 1 \\ \frac{-c}{p} & 0 & \frac{c}{p} \\ c^2 & 0 & c^2 \end{pmatrix}$$

If we let $d\vec{w} = \vec{L} d\vec{q}$ $\leftarrow \vec{w}$ are characteristic variables

$$\vec{L}(q_t + A q_x) = 0$$

$$\vec{L} q_t + \vec{L} \vec{A} \vec{R} \vec{L} q_x = 0$$

$$\vec{L} \vec{q}_t + \vec{\lambda} \vec{L} \vec{q}_x = 0$$

$$\vec{w}_t + \vec{\lambda} \vec{w}_x = 0$$



then it can be shown that:

$$\frac{\vec{d}\vec{w}}{dt} + \vec{\lambda} \frac{\partial \vec{w}}{\partial x} = 0$$

writing things out:

$$\left. \begin{array}{l} w_t^{(+)} + \lambda^{(+)} w_x^{(+)} = 0 \\ w_t^{(0)} + \lambda^{(0)} w_x^{(0)} = 0 \\ w_t^{(-)} + \lambda^{(-)} w_x^{(-)} = 0 \end{array} \right\} \begin{array}{l} 3 \text{ decoupled nonlinear} \\ \text{advection equations.} \\ 3 \text{ waves traveling at speed} \\ u, u-c, u+c. \end{array}$$

We can then recover the primitive variables as:

$$\begin{pmatrix} \rho \\ u \\ p \end{pmatrix} = d\vec{q} = L^{-1} d\vec{w} = R d\vec{w} = R \begin{pmatrix} w^{(+)} \\ w^{(0)} \\ w^{(-)} \end{pmatrix}$$

Here we note that for $w^{(0)}$, only density proportion is non-zero, so there will only be a density jump over the middle wave.

Entropy Formulation: used to understand the condition of rarefaction since entropy is constant for rarefaction. (For shock entropy increases)

We can also write: $\tilde{q}_S = (p, u, s)^T$

We need to get rid of pressure gradient term in velocity equation:

$$\frac{\partial p(t, s)}{\partial x} = \frac{\partial p}{\partial s} \Big|_p \frac{\partial s}{\partial x} + \frac{\partial p}{\partial p} \Big|_s \frac{\partial s}{\partial x} = \frac{\partial p}{\partial s} \Big|_p \frac{\partial s}{\partial x} + \frac{p \Gamma_1}{p} \frac{\partial p}{\partial x}$$

then the velocity equation becomes:

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + \frac{1}{f} \left[\frac{\partial p}{\partial s} \Big|_p \frac{\partial s}{\partial x} + \frac{p \Gamma_1}{p} \frac{\partial p}{\partial x} \right] = 0$$

such that $q_{st} + A_S(q_S) q_{sx} = 0$

$$A_S = \begin{pmatrix} u & p & 0 \\ \frac{c^2}{f} & u & \frac{p}{f} \\ 0 & 0 & u \end{pmatrix} \quad \text{here } B = \frac{\partial p}{\partial s} \Big|_p$$

Then we still have the same eigenvalues:

$$\vec{\lambda} = \begin{pmatrix} u-c & & \\ & u & \\ & & u+c \end{pmatrix}$$

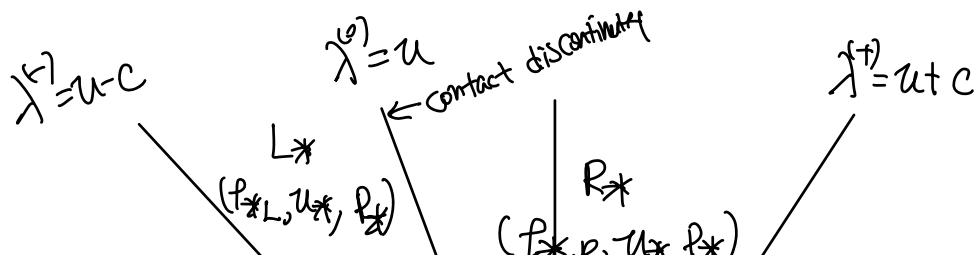
$$R = \begin{pmatrix} r^{(1)} & r^{(2)} & r^{(3)} \end{pmatrix} = \begin{pmatrix} 1 & 1 & 1 \\ -\frac{c}{P} & 0 & \frac{c}{P} \\ 0 & \frac{-c^2}{P_s} & 0 \end{pmatrix}$$

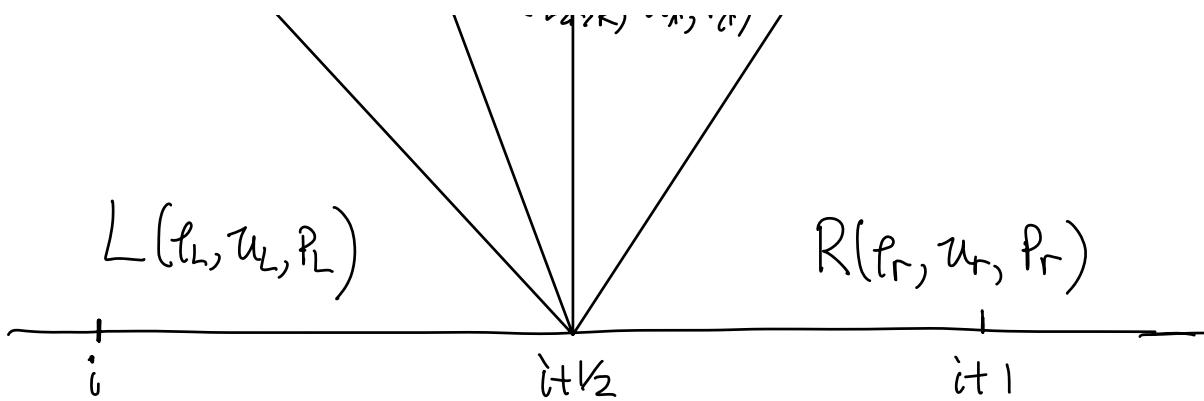
$$L = \begin{pmatrix} \tau^{(1)} \\ \tau^{(2)} \\ \tau^{(3)} \end{pmatrix} = \begin{pmatrix} 0 & -\frac{1}{c} & \frac{1}{c^2} \\ \frac{P_s}{P_s+1} & 0 & \frac{-P_s}{c^2(P_s+1)} \\ 0 & \frac{1}{c} & \frac{1}{c^2} \end{pmatrix}$$

Euler Riemann Solver:

Given left and right states in primitive variable, output an unique interface state:

$$q_{i+\frac{1}{2}} = R(q_{i+\frac{1}{2}, L}, q_{i+\frac{1}{2}, R})$$





Properties?

- 1) The middle wave, speed at u , it is neither a shock or rarefaction since velocity remains constant across this wave.
Only density changes.
- 2) The left and right waves can either be a shock or rarefaction.
- 3) L and R are the original left and right states
Since no waves have reached them, they're unchanged
- 3) Given L and R states, we have 4 unknowns,
 $(u_*, \rho_*, p_{*L}, p_{*R})$
 ~ A total of 4 unknowns

Rarefaction conditions:

\Rightarrow Since entropy remains constant across a rarefaction, $\frac{Ds}{Dt} = 0$

In the primitive variable, replace pressure with entropy. $\vec{q}_s = (f, u, s)^T$.

Then we need to express the pressure gradient, $\frac{\partial p}{\partial x}$ term in momentum Euler equations in terms of \vec{q}_s .

$$\vec{q}_s = (f, u, s)^T.$$

$$\frac{\partial p(f, s)}{\partial x} = \left. \frac{\partial p}{\partial f} \right|_s \frac{\partial f}{\partial x} + \left. \frac{\partial p}{\partial s} \right|_f \frac{\partial s}{\partial x} = \left. \frac{\partial p}{\partial s} \right|_f \frac{\partial s}{\partial x} + \frac{p \Gamma_1}{f} \frac{\partial f}{\partial x}$$

$$\hookrightarrow \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + \frac{1}{f} \left[\left. \frac{\partial p}{\partial s} \right|_f \frac{\partial s}{\partial x} + \frac{p \Gamma_1}{f} \frac{\partial f}{\partial x} \right] = 0$$

$$\text{Then } q_t + A q_x = 0 \Rightarrow q_{s,t} + A_s q_{s,x} = 0$$

where

$$A_s = \begin{pmatrix} u & f & 0 \\ c^2/f & u & \frac{1}{f} \left. \frac{\partial p}{\partial s} \right|_f \\ 0 & 0 & u \end{pmatrix}$$

$$R = \begin{pmatrix} r^{(-)} & r^{(0)} & r^{(+)} \end{pmatrix} = \begin{pmatrix} 1 & 1 & 1 \\ \frac{-c}{\rho} & 0 & \frac{\gamma}{\rho} \\ 0 & \frac{-c^2}{\rho s} & 0 \end{pmatrix}$$

Since $d\vec{q}_s = \vec{R} d\vec{w}$, we see that entropy change over the (-) and (+) waves are zero.

$$L = \begin{pmatrix} l^{(+)} \\ l^{(0)} \\ l^{(-)} \end{pmatrix} = \begin{pmatrix} 0 & -\frac{f}{c} & \frac{1}{c^2} \\ \frac{\rho_s}{\rho_s + 1} & 0 & \frac{-\rho_s}{c^2(\rho_s + 1)} \\ 0 & \frac{f}{c} & \frac{1}{c^2} \end{pmatrix}$$

Now, we return to the original primitive variable \vec{q} :

Consider the (+) wave, across this wave, $w^{(+)}$ will jump, but constant across the (-) and (0) wave.

Similarly constancy of the middle and right characteristic variables across the left wave (-) gives:

$$l^{(+)} \cdot d\vec{q} = 0 = (0 \quad \frac{f}{2c} \quad \frac{1}{2c^2}) \begin{pmatrix} \frac{df}{du} \\ \frac{df}{dp} \end{pmatrix} = 0$$

$$l^{(0)} \cdot d\vec{q} = 0 = (1 \quad 0 \quad -\frac{1}{c^2}) \begin{pmatrix} \frac{df}{du} \\ \frac{df}{dp} \end{pmatrix} = 0$$

↪ Then we have:

$$du + \frac{1}{\rho C} dp = 0$$

$$dp - \frac{1}{C^2} dp = 0$$

→ Define Lagrangian sound speed: $C = \rho C$ and $\gamma = k_p$
then:

$$du = -\frac{dp}{C}$$

$$dT = -\frac{dp}{C^2} \quad \text{across the left wave.}$$

Similarly

$$du = \int_{P_{L,R}}^{P_*} \frac{dp}{C}$$

$$dT = \int_{P_{L,R}}^{P_*} -\frac{dp}{C^2} \quad \text{across the right wave.}$$

Until now, these are general. Finding the solution across the rarefaction involves integrating the system with EOS to provide $C \equiv C(T, P)$. The solutions are called Riemann Invariants.

Now consider gamma law eos: $P = K \rho^\gamma$

$$dp = \gamma K \rho^{\gamma-1} d\rho$$

then only need to integrate a single equation:

$$du = \int \frac{\gamma k \rho^{\gamma-1} d\rho}{\rho \sqrt{\frac{\gamma P}{\rho}}}$$

$$= - \int \frac{\gamma k \rho^{\gamma-1} d\rho}{\sqrt{\gamma \rho k \rho^{\gamma}}} \quad 2\gamma - 2 - (\gamma + 1) = \gamma - 3$$

$$= - \int \sqrt{\gamma k} \rho^{\gamma-3} d\rho$$

$$= - \sqrt{\gamma k} \int \rho^{\frac{\gamma}{2} - \frac{3}{2}} d\rho$$

$$= - \sqrt{\gamma k} \frac{1}{\frac{\gamma}{2} - \frac{1}{2}} \int \rho^{\frac{\gamma}{2}}$$

$$u = -\frac{2c}{\gamma-1} + \text{const}$$

then $u + \frac{2c}{\gamma-1} = \text{constant} \leftarrow \text{Riemann Invariant}$
 across left wave

Similarly $u - \frac{2c}{\gamma-1} = \text{constant} \leftarrow \text{Invariant across right wave.}$

Using the Riemann Invariant to the left waves

$$U_L + \frac{2C_L}{\gamma-1} = U_* + \frac{2C_*}{\gamma-1}$$

and

$$k = \frac{P_L}{T_L^{\gamma}} = \frac{P_*}{T_*^{\gamma}}$$

We have:

$$U_* = U_L + \frac{2C_L}{\gamma-1} \left[1 - \left(\frac{P_*}{P_L} \right)^{\frac{\gamma-1}{2\gamma}} \right] \text{ for the left rarefaction}$$

$$U_* = U_R - \frac{2C_R}{\gamma-1} \left[1 - \left(\frac{P_*}{P_R} \right)^{\frac{\gamma-1}{2\gamma}} \right] \text{ for the right rarefaction.}$$

Shock Jump Condition: (Entropy increases in shock)

When the left or right wave is a shock:

We use Rankine-Hugoniot Condition:

In general:

$$\frac{F(U_*) - F(U_s)}{U_* - U_s} = S$$

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where $s \in (L, R)$, left or right interface state.
and s is the shock speed.

We define velocity in the shock frame:

$$L, R \text{ interface} \rightarrow \hat{u}_s = u_s - s$$

velocity

$$\text{star state} \rightarrow \hat{u}_* = u_* - s$$

velocity

Using velocity in shock frame:

$$\frac{F(\hat{u}_*) - F(\hat{u}_s)}{\hat{u}_* - \hat{u}_s} = 0$$

For 1D Euler: where $s \in [L, R]$

$$① \quad f_* \hat{u}_* = f_s \hat{u}_s$$

$$② \quad f_* \hat{u}_*^2 + p_* = f_s \hat{u}_s^2 + p_s$$

$$③ \quad f_* \hat{u}_* e_* + \frac{1}{2} f_* \hat{u}_*^3 + \hat{u}_* p_* = f_s \hat{u}_s e_s + \frac{1}{2} f_s \hat{u}_s^3 + \hat{u}_s p_s$$

For a general EOS:

$$\text{from Eq. (1): } \hat{u}_* = \frac{f_s}{f_*} \hat{u}_s$$

Plug above into Eq. (2):

$$f_s^2 \left(\frac{1}{f_s} - \frac{1}{f_*} \right) \hat{u}_s^2 = P_* - P_s$$

$$\begin{cases} \text{let } [P] = P_* - P_s \\ \text{then} \end{cases}$$

$$\rightarrow -[T] = \frac{1}{f_s} - \frac{1}{f_*} = \frac{[P]}{f_s^2 \hat{u}_s^2}$$

$$T = \frac{1}{f}$$

Now introduce mass flux, w_s .

For the shock separating L , L_* , mass will be moving through the shock to the right. Then in the frame of the shock:

$$w_L = f_L \hat{u}_L = f_* \hat{u}_*$$

Likewise, for the shock separating R/R_* , mass will move to the left passing the shock:

$$w_R = -f_R \hat{u}_R = -f_* \hat{u}_*$$

Then our first jump condition becomes:

$$-\left[\gamma\right] = \frac{[p]}{W_e^2} \quad (1) \text{ density relation}$$

Now going back to the momentum equation

$$t_* \hat{u}_*^2 + p_* = t_s \hat{u}_s^2 + p_s$$

For left shock:

$$W_L \hat{u}_* + p_* = W_L \hat{u}_L + p_L$$

$$\hookrightarrow \hat{u}_* - \hat{u}_L = \frac{-p_* + p_L}{W_L}$$

$$\hookrightarrow [u] = \frac{-[p]}{W_L} \quad (2) \text{ velocity relation of left shock}$$

For right shock:

$$-W_R \hat{u}_* + p_* = -W_R \hat{u}_R + p_R$$

$$[u] = \frac{[p]}{W_R} \quad (3) \text{ velocity relation of right shock.}$$

Need condition for energy:

Use energy equation with mass flux: W_s

$$W_s e_* + \frac{1}{2} W_s \hat{u}_*^2 + \frac{W_s}{f_*} P_* = W_s e_s + \frac{1}{2} W_s \hat{u}_s^2 + \frac{W_s}{f_s} P_s$$

Cancelling W_s :

$$[e] + \frac{P_*}{f_*} - \frac{P_s}{f_s} + \frac{1}{2} (\hat{u}_*^2 - \hat{u}_s^2) = 0$$

$$\text{Substitute } \hat{u}_*^2 = \frac{W_s^2}{f_*^2} \text{ and } \hat{u}_s^2 = \frac{W_s^2}{f_s^2}$$

$$[e] + \frac{P_*}{f_*} - \frac{P_s}{f_s} + \frac{1}{2} W_s^2 \left(\frac{1}{f_*^2} - \frac{1}{f_s^2} \right) = 0$$

$$\text{Using the first jump condition: } W_s^2 = \frac{-[P]}{[\tau]}$$

$$[e] + \frac{P_*}{f_*} - \frac{P_s}{f_s} + \frac{1}{2} (P_s - P_*) \left(\frac{1}{f_*} - \frac{1}{f_s} \right)^{-1} \left(\frac{1}{f_*^2} - \frac{1}{f_s^2} \right) = 0$$

With algebra:

$$\hookrightarrow [e] = -\frac{P_* + P_s}{2} [\tau]$$

Summary of jump conditions across shock:

$$\textcircled{1} \quad [\tau] = -\frac{[P]}{W_s^2}$$

$$\textcircled{2} \quad [u] = \mp \frac{[P]}{W_s} \quad '-' \text{ for left, '+' for right.}$$

$$\textcircled{3} \quad [e] = -\frac{P_s + P_*}{2} [\tau]$$

Procedure:

① root find to find f_* corresponding to P_* :

- guess a value for f_*
- use eos, express $e_* = e(P_*, f_*)$
- use Newton or other methods

② Compute: $\frac{1}{w_s^2} = -\frac{[\gamma]}{[P]}$

③ Find the star velocity:

$$u_* = u_s \mp \frac{[P]}{w_s} \quad \begin{array}{l} '-' \text{ for left shock} \\ '+' \text{ for right shock} \end{array}$$

Shock speed then can be calculated by:

$$S = u_s \mp \frac{w_s}{P_s} \quad '-' \text{ for left, '+' for right.}$$

Consider a simple case with gamma law eos:

First, find f_* :

Introduce: $e = \frac{P}{\rho} \frac{1}{r-1}$

Plug into $[e] = -\left(\frac{P_* + P_s}{2}\right) [\gamma]$

then : $t_{*,S} = t_S \left[\frac{\frac{P_*}{P_S}(\gamma+1) + (\gamma-1)}{(\gamma+1) + \frac{P_*}{P_S}(\gamma-1)} \right]$

Second: Compute $\frac{1}{w_s^2} = -\frac{[T]}{[P]}$

$$w_s^2 = \frac{1}{2} P_S t_S \left[\left(\frac{P_*}{P_S} \right) (\gamma+1) + (\gamma-1) \right]$$

Third: Find star velocity: u_* :

$$u_* = u_s \pm c_s \left[\frac{2}{\gamma(\gamma-1)} \right]^{1/2} \frac{1 - \frac{P_*}{P_S}}{\left(\frac{P_*}{P_S} \frac{\gamma+1}{\gamma-1} + 1 \right)^{1/2}}$$

(+) for left shock
(-) for right shock

Last: Find shock speed:

$$S = u_s \mp c_s \left[\left(\frac{P_*}{P_S} \right) \frac{\gamma+1}{2\gamma} + \frac{\gamma-1}{2\gamma} \right]^{1/2}$$

(-) for left shock
(+) for right shock

Now that we have both star velocity, u_* , from rarefaction and shock condition, we determine which to we based on pressure.

$$u_{*,L}(P) = \begin{cases} u_{*,L}^{\text{shock}} & P > P_L \\ u_{*,L}^{\text{rare}} & P \leq P_L \end{cases}$$

$$u_{*,R}(P) = \begin{cases} u_{*,R}^{\text{shock}} & P > P_R \\ u_{*,R}^{\text{rare}} & P \leq P_R \end{cases}$$

Therefore, we need to solve for pressure, P_* using condition:

$$\rightarrow u_{*,L}(P_*) - u_{*,R}(P_*) = 0$$

This means that velocity and pressure do not jump across the contact wave.

For gamma law gas:

$$u_s \pm \frac{2c}{\gamma-1} \left[1 - \left(\frac{P_*}{P_s} \right)^{\frac{\gamma-1}{\gamma}} \right] \quad \text{for } P_* \leq P_s$$

$$u_{*,S}(P_*) = \begin{cases} u_s \pm c \left[\frac{2}{\gamma(\gamma-1)} \right]^{\frac{1}{2}} \frac{1 - \frac{P_*}{P_s}}{\left(\frac{P_*}{P_s} \frac{\gamma+1}{\gamma-1} + 1 \right)^{\frac{1}{2}}} & \text{for } P_* > P_s \end{cases}$$

Solve for P_* such that

$$u_{*,L}(P_*) - u_{*,R}(P_*) = 0$$