16-720 HW6: Photometric Stereo

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Utils and Imports

Importing all necessary libraries.

```
In [1]: import numpy as np
    from matplotlib import pyplot as plt
    from skimage.color import rgb2xyz
    import warnings
    from scipy.ndimage import gaussian_filter
    from matplotlib import cm
    from skimage.io import imread
    from scipy.sparse import kron as spkron
    from scipy.sparse import eye as speye
    from scipy.sparse.linalg import lsqr as splsqr
    import os
    import shutil
```

```
Downloading the data
In [2]: if os.path.exists('data'):
         shutil.rmtree('data')
        os.mkdir('data')
        !wget 'https://docs.google.com/uc?export=download&id=13nAlHag6bJz0-h 7NmovvSRrRD76qiF0' -0 data/data.zip
        !unzip "data/data.zip"
        os.system("rm data/data.zip")
        data dir = 'data/'
       --2025-04-24 21:25:41-- https://docs.google.com/uc?export=download&id=13nA1Haq6bJz0-h 7NmovvSRrRD76qiF0
      Resolving docs.google.com (docs.google.com)... 172.253.62.138, 172.253.62.100, 172.253.62.101, ...
       Connecting to docs.google.com (docs.google.com)|172.253.62.138|:443... connected.
      HTTP request sent, awaiting response... 303 See Other
      Location: https://drive.usercontent.google.com/download?id=13nA1Haq6bJz0-h_7NmovvSRrRD76qiF0&export=downloa
      d [following]
       --2025-04-24 21:25:41-- https://drive.usercontent.google.com/download?id=13nAlHaq6bJz0-h_7NmovvSRrRD76qiF0
      &export=download
      Resolving drive.usercontent.google.com (drive.usercontent.google.com)... 142.251.16.132, 2607:f8b0:4004:c1
      7::84
      Connecting to drive.usercontent.google.com (drive.usercontent.google.com)|142.251.16.132|:443... connected.
      HTTP request sent, awaiting response... 200 OK
       Length: 6210854 (5.9M) [application/octet-stream]
       Saving to: 'data/data.zip'
                          in 0.09s
       data/data.zip
       2025-04-24 21:25:45 (62.6 MB/s) - 'data/data.zip' saved [6210854/6210854]
      Archive: data/data.zip
        inflating: data/sources.npy
         inflating: data/input 5.tif
         inflating: data/input 7.tif
         inflating: data/input_6.tif
         inflating: data/input_4.tif
         inflating: data/input_1.tif
         inflating: data/input 2.tif
         inflating: data/input 3.tif
        Utils Functions.
In [3]: def integrateFrankot(zx, zy, pad = 512):
            Question 1 (j)
```

```
Implement the Frankot-Chellappa algorithm for enforcing integrability
    and normal integration
    Parameters
    zx : numpy.ndarray
       The image of derivatives of the depth along the x image dimension
   zy : tuple
       The image of derivatives of the depth along the y image dimension
    pad : float
       The size of the full FFT used for the reconstruction
    z: numpy.ndarray
       The image, of the same size as the derivatives, of estimated depths
       at each point
    # Raise error if the shapes of the gradients don't match
   if not zx.shape == zy.shape:
        raise ValueError('Sizes of both gradients must match!')
    # Pad the array FFT with a size we specify
    h, w = 512, 512
    # Fourier transform of gradients for projection
   Zx = np.fft.fftshift(np.fft.fft2(zx, (h, w)))
   Zy = np.fft.fftshift(np.fft.fft2(zy, (h, w)))
   j = 1j
    # Frequency grid
    [wx, wy] = np.meshgrid(np.linspace(-np.pi, np.pi, w),
                           np.linspace(-np.pi, np.pi, h))
    absFreq = wx**2 + wy**2
    # Perform the actual projection
   with warnings.catch_warnings():
       warnings.simplefilter('ignore')
       z = (-j*wx*Zx-j*wy*Zy)/absFreq
    # Set (undefined) mean value of the surface depth to 0
   z[0, 0] = 0.
   z = np.fft.ifftshift(z)
   # Invert the Fourier transform for the depth
   z = np.real(np.fft.ifft2(z))
   z = z[:zx.shape[0], :zx.shape[1]]
    return z
def enforceIntegrability(N, s, sig = 3):
   Question 2 (e)
   Find a transform Q that makes the normals integrable and transform them
   by it
   Parameters
   N : numpy.ndarray
        The 3 x P matrix of (possibly) non-integrable normals
    s : tuple
       Image shape
   Returns
    Nt : numpy.ndarray
       The 3 x P matrix of transformed, integrable normals
```

```
N1 = N[0, :].reshape(s)
   N2 = N[1, :].reshape(s)
N3 = N[2, :].reshape(s)
    N1y, N1x = np.gradient(gaussian filter(N1, sig), edge order = 2)
    N2y, N2x = np.gradient(gaussian_filter(N2, sig), edge_order = 2)
    N3y, N3x = np.gradient(gaussian_filter(N3, sig), edge_order = 2)
    A1 = N1*N2x-N2*N1x
    A2 = N1*N3x-N3*N1x
    A3 = N2*N3x-N3*N2x
    A4 = N2*N1y-N1*N2y
    A5 = N3*N1y-N1*N3y
   A6 = N3*N2y-N2*N3y
    A = np.hstack((A1.reshape(-1, 1),
                   A2.reshape(-1, 1),
                   A3.reshape(-1, 1),
                   A4.reshape(-1, 1),
                   A5.reshape(-1, 1),
                   A6.reshape(-1, 1)))
    AtA = A.T.dot(A)
   W, V = np.linalg.eig(AtA)
   h = V[:, np.argmin(np.abs(W))]
    delta = np.asarray([[-h[2], h[5], 1],
                         [ h[1], -h[4], 0],
[-h[0], h[3], 0]])
    Nt = np.linalg.inv(delta).dot(N)
    return Nt
def plotSurface(surface, suffix=''):
    Plot the depth map as a surface
    Parameters
    surface : numpy.ndarray
        The depth map to be plotted
    suffix: str
        suffix for save file
    Returns
        None
   x, y = np.meshgrid(np.arange(surface.shape[1]),
                       np.arange(surface.shape[0]))
    fig = plt.figure()
    #ax = fig.gca(projection='3d')
    ax = fig.add_subplot(111, projection='3d')
    surf = ax.plot_surface(x, y, -surface, cmap = cm.coolwarm,
                           linewidth = 0, antialiased = False)
    ax.view_init(elev = 60., azim = 75.)
    plt.savefig(f'figs/faceCalibrated{suffix}.png')
    plt.show()
def plotSurfaceGrid(surface, suffix='', ax = None):
    Plot the depth map as a surface
    Parameters
    surface : numpy.ndarray
        The depth map to be plotted
    suffix: str
```

```
suffix for save file
    Returns
        None
    render = ax is None
    x, y = np.meshgrid(np.arange(surface.shape[1]),
                       np.arange(surface.shape[0]))
    if render:
       fig = plt.figure()
       #ax = fig.gca(projection='3d')
       ax = fig.add_subplot(111, projection='3d')
    surf = ax.plot_surface(x, y, -surface, cmap = cm.coolwarm,
                           linewidth = 0, antialiased = False)
    ax.view_init(elev = 60., azim = 75.)
    ax.set_title(suffix)
    if render:
        plt.savefig(f'figs/faceCalibrated{suffix}.png')
        plt.show()
def loadData(path = "../data/"):
    Question 1 (c)
    Load data from the path given. The images are stored as input_n.tif
    for n = \{1...7\}. The source lighting directions are stored in
    sources.mat.
    Paramters
    path: str
        Path of the data directory
   Returns
   I : numpy.ndarray
        The 7 x P matrix of vectorized images
    L : numpy.ndarray
        The 3 \times 7 matrix of lighting directions
    s: tuple
        Image shape
   I = None
   L = None
   s = None
   L = np.load(path + 'sources.npy').T
    im = imread(path + 'input_1.tif')
   P = im[:, :, 0].size
s = im[:, :, 0].shape
    I = np.zeros((7, P))
    for i in range(1, 8):
        im = imread(path + 'input_' + str(i) + '.tif')
        im = rgb2xyz(im)[:, :, 1]
        I[i-1, :] = im.reshape(-1,)
    return I, L, s
def displayAlbedosNormals(albedos, normals, s):
    Question 1 (e)
    From the estimated pseudonormals, display the albedo and normal maps
    Please make sure to use the `coolwarm` colormap for the albedo image
```

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```
and the `rainbow` colormap for the normals.
Parameters
albedos : numpy.ndarray
    The vector of albedos
normals : numpy.ndarray
    The 3 x P matrix of normals
s : tuple
    Image shape
Returns
albedoIm : numpy.ndarray
    Albedo image of shape s
normalIm : numpy.ndarray
    Normals reshaped as an s \times 3 image
albedoIm = None
normalIm = None
albedoIm = albedos.reshape(s)
normalIm = (normals.T.reshape((s[0], s[1], 3))+1)/2
plt.figure()
plt.imshow(albedoIm, cmap = 'gray')
plt.figure()
plt.imshow(normalIm, cmap = 'rainbow')
plt.show()
return albedoIm, normalIm
```

Q1: Calibrated photometric stereo (75 points)

Q 1 (a): Understanding n-dot-I lighting (5 points)

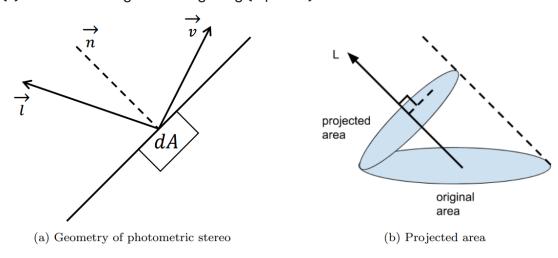


Fig 2.a depicts the interaction of vectors respect to the surface of an object. Vector \vec{l} represents direction of the lighting source of the scene; \vec{n} is the normal of the surface and \vec{v} is the camera view vector. The area iluminated depends on the angle between the light source vector and the surface normal.

The dot product between the vectors arises as it precisely describes the cosine of the angle of the two vectors, scaled by their magnitude, allowing to compactly use them to model the radiance. In particular if \vec{l} and \vec{n} are unit vectors then $\vec{n} \cdot \vec{l} = \cos \theta_{\vec{l} \vec{n}}$.

The cosine of the angle, and thus the dot product $\vec{n} \cdot \vec{l}$, geometrically represents the scaling factor between the original surface dA and the surface projected in the view of the light vector \hat{dA} .

$$\hat{dA} = dA\vec{n} \cdot \vec{l}$$

Notice that the radiance does not dependend on where we view the object from, rather it only depends of the relationship between the object's surface and the light source. As regardless of the viewpoint, the scattered light will be the same.

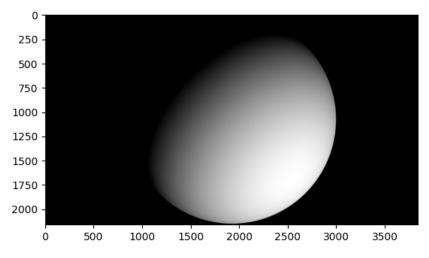
Q 1 (b): Rendering the n-dot-I lighting (10 points)

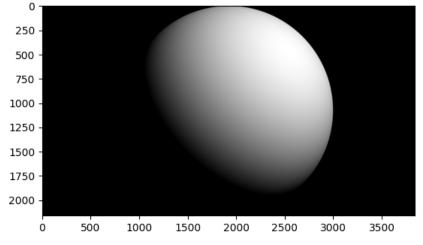
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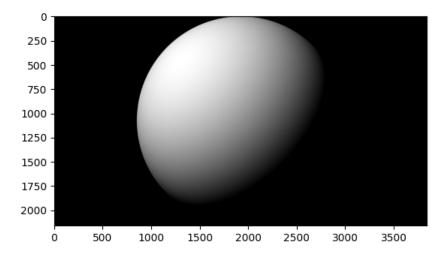
```
In [ ]: def renderNDotLSphere(center, rad, light, pxSize, res):
            Question 1 (b)
            Render a hemispherical bowl with a given center and radius. Assume that
            the hollow end of the bowl faces in the positive z direction, and the
            camera looks towards the hollow end in the negative z direction. The
            camera's sensor axes are aligned with the x- and y-axes.
            Parameters
            center : numpy.ndarray
                The center of the hemispherical bowl in an array of size (3,)
            rad : float
                The radius of the bowl
            light : numpy.ndarray
                The direction of incoming light
            pxSize : float
                Pixel size
            res : numpy.ndarray
                The resolution of the camera frame
            Returns
            image : numpy.ndarray
                The rendered image of the hemispherical bowl
            [X, Y] = np.meshgrid(np.arange(res[0]), np.arange(res[1]))
            X = (X - res[0]/2) * pxSize*1.e-4
            Y = (Y - res[1]/2) * pxSize*1.e-4
            Z = np.sqrt(rad**2+0j-X**2-Y**2)
            X[np.real(Z) == 0] = 0
            Y[np.real(Z) == 0] = 0
            Z = np.real(Z)
            image = None
            # Get light vector
            light vec = light/np.linalg.norm(light)
            # Get surface normals
            cx , cy , cz = center
            X = (X - cx)/rad

Y = (Y - cy)/rad
            Z = (Z - cz)/rad
            normals = np.array([X, Y, Z])
            normals /= np.linalg.norm(normals)
            # Get dot product
            n_dot_l = np.dot(normals.T, light_vec)
            n_dot_l[n_dot_l < 0] = 0
            # Build the image
            image = n_dot_l.astype(np.float32)
            image = image.T
            return image
        # Part 1(b)
        radius = 0.75 # cm
        center = np.asarray([0, 0, 0]) # cm
```

```
pxSize = 7 # um
res = (3840, 2160)
light = np.asarray([1, 1, 1])/np.sqrt(3)
image = renderNDotLSphere(center, radius, light, pxSize, res)
plt.figure()
plt.imshow(image, cmap = 'gray')
plt.imsave('figs/1b-a.png', image, cmap = 'gray')
light = np.asarray([1, -1, 1])/np.sqrt(3)
image = renderNDotLSphere(center, radius, light, pxSize, res)
plt.figure()
plt.imshow(image, cmap = 'gray')
plt.imsave('figs/lb-b.png', image, cmap = 'gray')
light = np.asarray([-1, -1, 1])/np.sqrt(3)
image = renderNDotLSphere(center, radius, light, pxSize, res)
plt.figure()
plt.imshow(image, cmap = 'gray')
plt.imsave('figs/lb-c.png', image, cmap = 'gray')
I, L, s = loadData(data_dir)
```







Q1(c): Initials (10 points)

```
In [5]: #NOTE: I = L.T B
U, S ,Vh = np.linalg.svd(I, full_matrices=False)
print("Singular values: ", S)

Singular values: [79.36348099 13.16260675 9.22148403 2.414729 1.61659626 1.26289066 0.89368302]
```

1. Why should I be rank 3?

Theoritically given that the lighting direction have dimensions $L \in \mathbb{R}^{3 \times N}$ and the psuedo normal matrix has one column per point making it shape $B \in \mathbb{R}^{3 \times P}$, then there are a maximum of three linearly independent components forming I meaning we should expect a rank of 3.

2. Do the singular values agree with this?

A: We can see that the all 7 singular values are different than 0, meaning that it is not purely speaking rank 3. However we can see that the first 3 values are significantly larger than the rest; while the remaining singular values are closer to 0. This occurs due to inherent noise in real world photographs; as well as departures from the Lambertain assumptions made during modelling.

Q 1 (d) Estimating pseudonormals (20 points)

```
B, _, _, _ = np.linalg.lstsq(A, y, rcond=None)
return B

# Part 1(e)
B = estimatePseudonormalsCalibrated(I, L)
print("Pseudonormals shape: ", B.shape)
```

Pseudonormals shape: (3, 159039)

Α

Given that $I=L^TB$ we have a linear system Ax=y that we can solve directly, where $A=L^T$ and y=I. Note that in this simple implementation one could run into memory issues because of the size of the matrixes. In the case of my system, I was able to load everything without a problem.

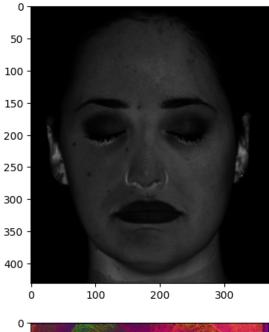
Q 1 (e) Albedos and normals (10 points)

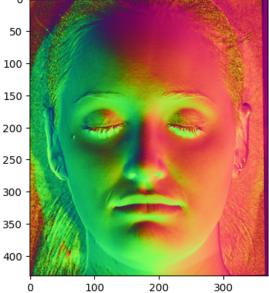
A:

Surprisingly, higher albedo values are observed in areas where we typically would see shadows caused by the nose, ears and neck. Given that we are estimating the albedo from the pseudonormals by taking their scale as the the albedo value (B=an : a=||B|| and $b=\frac{B}{||B||}$); this means regions with high magnitude of the pseudonorm, have high albedo.

When solving for the psuedo, there are multiple factors that could cause the norm to be large. Given that we are using a simple Lambertian model, this might produce incorrect estimation of shadows while omitting other considerations that impact realworld lighting, leading to an overestimation of the pseudonormal to fit the perceived intensities in shadows and noisy regions.

```
In [7]: def estimateAlbedosNormals(B):
            Question 1 (e)
            From the estimated pseudonormals, estimate the albedos and normals
            Parameters
            B : numpy.ndarray
                The 3 x P matrix of estimated pseudonormals
            Returns
            albedos : numpy.ndarray
                The vector of albedos
            normals : numpy.ndarray
                The 3 x P matrix of normals
            albedos = None
            normals = None
            # Albedos are the scale of the pseudonormals
            albedos = np.linalg.norm(B, axis=0)
            # Normals are the normalized pseudonormals
            normals = B/albedos
            return albedos, normals
        # Part 1(e)
        albedos, normals = estimateAlbedosNormals(B)
        albedoIm, normalIm = displayAlbedosNormals(albedos, normals, s)
        plt.imsave('figs/1f-a.png', albedoIm, cmap = 'gray')
        plt.imsave('figs/1f-b.png', normalIm, cmap = 'rainbow')
```





Q 1 (f): Normals and depth (5 points)

A:

The face is defined as a surface $S=\left[egin{array}{c} x \\ y \\ f(x,y) \end{array}
ight]$. We wish to establish a relationship between the surface normal and the

partial derivatives of the f(x). We know that the normals will always be perpendicular to the tangent planes of the surface; the partial derivatives describe vectors on these tangent plane. We can exploit this relationship to express the normals in terms of the partials.

1. The tangent vectors in the x and y are obtained by taking the corresponding partials,

$$S_x = egin{bmatrix} 1 \ 0 \ f_x \end{bmatrix}$$

$$S_y = \left[egin{array}{c} 0 \ 1 \ f_y \end{array}
ight]$$

2. To obtain a vector in the direction of the normal, that is perpendicular to both S_x and S_y we take their cross product.

$$S_x imes S_y = egin{bmatrix} (0)(f_y) - f_x(1) \ f_x(0) - (1)(f_y) \ (1)(1) - (0)(0) \end{bmatrix}$$

$$S_x imes S_y = egin{bmatrix} -f_x \ -f_y \ 1 \end{bmatrix}$$

3. We know that n is proportional to $S_x \times S_y$, to ger rid of the notion of scale we can normalize $S_x \times S_y$ such that $n = \text{norm}(S_x \times S_y)$

$$n = rac{(-f_x - f_y + 1)}{\sqrt{f_x^2 + f_y^2 + 1}}$$

$$n_1=rac{(-f_x)}{\sqrt{f_x^2+f_y^2+1}}$$

$$n_2=rac{(-f_y)}{\sqrt{f_x^2+f_y^2+1}}$$

$$n_3 = rac{(1)}{\sqrt{f_x^2 + f_y^2 + 1}}$$

4. We can use n_3 to write the partials in terms of the components of the normal

$$rac{n_1}{n_3} = rac{(-f_x)}{\sqrt{f_x^2 + f_y^2 + 1}} rac{\sqrt{f_x^2 + f_y^2 + 1}}{1} = -f_x$$

$$rac{n_2}{n_3} = rac{(-f_y)}{\sqrt{f_x^2 + f_y^2 + 1}} rac{\sqrt{f_x^2 + f_y^2 + 1}}{1} = -f_y$$

5. Therfore,

$$f_x = -rac{n_1}{n_3}$$

$$f_y = -rac{n_2}{n_3}$$

Q1(g): Understanding integrability of gradients (5 points)

Given
$$g = \left[egin{array}{cccc} 1 & 2 & 3 & 4 \ 5 & 6 & 7 & 8 \ 9 & 10 & 11 & 12 \ 13 & 14 & 15 & 16 \ \end{array}
ight].$$

0. Calculate the $g_x(x_i,y_j)=g(x_i+1,y_j)-g(x_i,y_j)$ and $g_y(x_i,y_j)=g(x_i,y_j+1)-g(x_i,y_j)$

$$g_x = egin{bmatrix} 2-1 & 3-2 & 4-3 \ 6-5 & 7-6 & 8-7 \ 10-9 & 11-10 & 12-11 \ 14-13 & 15-14 & 16-15 \end{bmatrix}$$

1. Use g_x to construct the first row of g, then use g_y to construct the rest of g

For the first row, and using $g(0,0)=g_x(0,0)$ we obtain

For the rest,

$$\hat{g} = \begin{bmatrix} 1 & 1+1=2 & 2+1=3 & 3+1=4 \\ 1+4=5 & 2+4=6 & 3+4=7 & 4+4=8 \\ 5+4=9 & 6+4=10 & 7+4=11 & 8+4-12 \\ 9+4=13 & 10+4=14 & 11+4=15 & 12+4=16 \end{bmatrix}$$

We see that .

2. Use g_y to construct the first column of g, then use g_x to construct the rest of g.

$$\hat{g} = egin{bmatrix} 1 & 0 & 0 & 0 \ 1+4=5 & 0 & 0 & 0 \ 5+4=9 & 0 & 0 & 0 \ 9+4=13 & 0 & 0 & 0 \end{bmatrix}$$

Now applying g_x iteratively,

$$\hat{g} = \begin{bmatrix} 1 & 2 & 3 & 4 \\ 1+4=5 & 6 & 7 & 8 \\ 5+4=9 & 10 & 11 & 12 \\ 9+4=13 & 14 & 15 & 16 \end{bmatrix}$$

In this case, $\hat{g} = g$. Confirming that g_x and g_y are integrable.

• What could make gradients non-integrable?

In general, gradients are integrable if they are derived from a function that follows Clairut's Theorem, that is that **the second** order mixed partial derivatives are equal

$$\frac{\partial^2 f}{\partial x} = \frac{\partial^2 f}{\partial y}$$

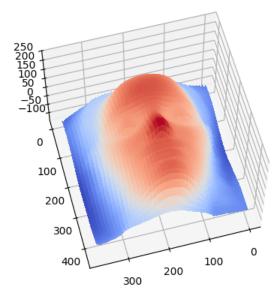
This allows us to disregard the order in which we integrate, which is we start with the row or column yielded the same result when reconstructing g. We see that indeed for the provided example g if we again differentiate g_x and g_y these are equal.

To make it non-integrable, we could asume some noise in the derivation process that results in noisy gradientes, such that their seconf order mixed derivates are no longer similar.

Q 1 (h): Shape estimation (10 points)

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```
In [8]: def estimateShape(normals, s):
            Question 1 (h)
            Integrate the estimated normals to get an estimate of the depth map
            of the surface.
            Parameters
            normals : numpy.ndarray
                The 3 x P matrix of normals
            s : tuple
                Image shape
            Returns
            surface: numpy.ndarray
                The image, of size s, of estimated depths at each point
            surface = None
            n1, n2, n3 = normals[0, :], normals[1, :], normals[2, :]
            # Estimate depth partials
            depth_dx = -n1/n3
            depth dy = -n2/n3
            depth_dx = depth_dx.reshape(s)
            depth_dy = depth_dy.reshape(s)
            # Use Frankot-Chellappa to get surface
            surface = integrateFrankot(depth_dx, depth_dy)
            return surface
        # Part 1(h)
        surface = estimateShape(normals, s)
        plotSurface(surface)
```



Q2: Uncalibrated photometric stereo (50 points)

Q 2 (a): Uncalibrated normal estimation (10 points)

Given that in the uncalibrated scenario we do not have acess to the ligh sources, we now have to factorize I to estimate the light sources \hat{L} and the psuedo normals \hat{B} .

We can leverage single value decomposition to factor ${\it I}$ as

$$I = U\Sigma V^T$$

To get the best rank 3 aproximation of I, we simply pick the top 3 singular values from Σ , giving us that

$$\Sigma_3 = egin{bmatrix} \sigma_1 & 0 & 0 \ 0 & \sigma_2 & 0 \ 0 & 0 & \sigma_3 \end{bmatrix}.$$

To enable us to express I as two factors, we can write $\Sigma_3 = \Sigma^{1/2} \Sigma^{1/2}$. Thus,

$$I = U \Sigma^{1/2} \Sigma^{1/2} V^T$$

By grouping terms we have,

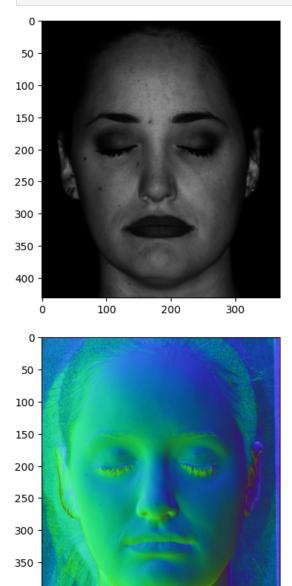
$$I = (U\Sigma^{1/2})(\Sigma^{1/2}V^T)$$

The light sources are aproximate by $\hat{L}^T=(U\Sigma^{1/2})$, while the psuedonormals are approximated by $\hat{B}=(\Sigma^{1/2}V^T)$.

Q 2 (b): Calculation and visualization (10 points)

```
In [9]: def estimatePseudonormalsUncalibrated(I):
            Question 2 (b)
            Estimate pseudonormals without the help of light source directions.
            Parameters
            I : numpy.ndarray
                The 7 x P matrix of loaded images
            Returns
            B : numpy.ndarray
                The 3 x P matrix of pesudonormals
            L : numpy.ndarray
                 The 3 \times 7 array of lighting directions
            B = None
            L = None
            U, S, Vh = np.linalg.svd(I, full_matrices=False)
            # Pick the first 3 singular values
                    = U[:, :3]
            S rank3 = np.diag(S[:3])
                    = Vh[:3, :]
            # Split singular value matrix in two matrix
            s_sqrt = np.sqrt(S_rank3)
            L hat = U @ s sqrt \# U \setminus Sigma^{1/2}
            B_hat = s_sqrt @ Vh # \Sigma^{1/2}V^T
            aprox_I = L_hat @ B_hat
            error = np.linalg.norm(I - aprox_I)
            assert error < 5, "Error is too high"</pre>
            L = L_hat.T
            B = B hat
            return B, L
```

```
# Part 2 (b)
I, L, s = loadData(data_dir)
B, LEst = estimatePseudonormalsUncalibrated(I)
albedos, normals = estimateAlbedosNormals(B)
albedoIm, normalIm = displayAlbedosNormals(albedos, normals, s)
plt.imsave('figs/2b-a.png', albedoIm, cmap = 'gray')
plt.imsave('figs/2b-b.png', normalIm, cmap = 'rainbow')
```



Q 2 (c): Comparing to ground truth lighting

200

300

400

0

100

A: As seen in the print below, the true lighting directions are vastly different than the estimated ones. This is because the choosen factorization scheme only garantees that the aproximated I image will be the same, however it offers no garantees on the choosen factors.

To obtain factors that more closely resemble the true lightning, we can play around with the factorization. For instance, we can apply symetric transformations to each of the factors, such that we permute and scale the the matrixes but still mantain the same output. Note that for this to work, the choosen transformation must be invertible and a 3×3 matrix.

$$I = (\hat{L}T)^T (T^{-1}\hat{B})$$

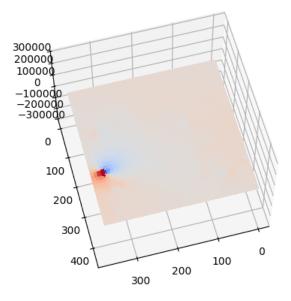
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```
In [10]: np.set_printoptions(precision=4, suppress=True)
        print("True light directions (L):\n", L)
        print("\nEstimated light directions (LEst):\n", LEst)
       True light directions (L):
        [[-0.1418  0.1215  -0.069
                               0.067 -0.1627 0.
                                                     0.14781
        [-0.1804 -0.2026 -0.0345 -0.0402 0.122
                                            0.1194
                                                    0.1209]
        \hbox{\tt [-0.9267-0.9717-0.838-0.9772-0.979-0.9648-0.9713]]}
       Estimated light directions (LEst):
        [[-2.9927 -3.87 -2.408 -3.745 -3.5914 -3.3867 -3.3525]
        [ 0.9478 -2.3171  0.4991 -0.626
                                     2.3257 0.4661 -0.7927]
```

Q 2 (d): Reconstructing the shape, attempt 1 (5 points)

It does not look like a face. As explained in c), given that the aproximation has no garantees the quantities will ressemble the real conditions, there is ambiguity in the process. In this case, the estimated normal did not yield integrable partials.

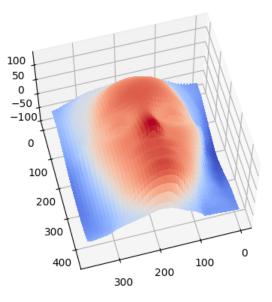
```
In [11]: # Part 2 (d)
surface = estimateShape(normals, s)
plotSurface(surface)
```



Q 2 (e): Reconstructing the shape, attempt 2 (5 points)

A: The reconstructed face more closely resembles the calibrated reconstruction. However, in general there is less distinction with the background, and in the edges the shape of face is less defined. For example, looking at the forhead, in the calibrated reconstruction the "roundness" is much more apparent; while in the uncalibrated case, it appears more flat.

```
In [12]: # Part 2 (e)
    # Your code here
    int_pseudonormals = enforceIntegrability(B, s)
    albedos, normals = estimateAlbedosNormals(int_pseudonormals)
    surface_integrable = estimateShape(normals, s)
    plotSurface(surface_integrable, suffix='integrable')
```



Q 2 (f): Why low relief? (5 points)

Α

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As described in [1], the bas-relief ambiguity refers to the difficulty to distinguish the true depth (or relief) of a surface. In uncalibrated photometric stereo, this can cause reconstructed surfaces to appear more or less flat than the actually are. To remedy this, one can apply transformations that warp the surface such that

$$\hat{f}(x,y) = \lambda f(x,y) + \mu x + \nu y$$

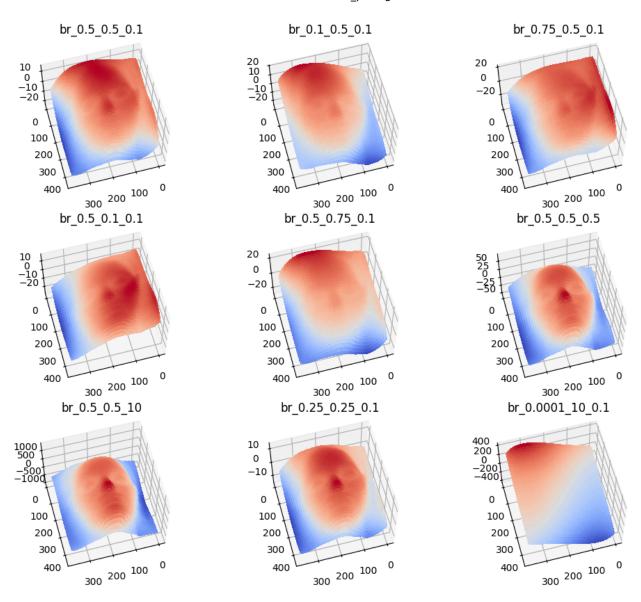
- ullet λ affects the height of the surfaces, changing the overall scale of the relief.
- $\,\mu$ scales the inclination of the surface, biasing towards the x-axis.
- + ν scales the inclination, biasing towards the y-axis.

[1] P. N. Belhumeur, D. J. Kriegman, and A. L. Yuille, "The bas-relief ambiguity," in Proceedings of IEEE Computer Society Conference on Computer Vision and Pattern Recognition, San Juan, Puerto Rico: IEEE Comput. Soc, 1997, pp. 1060–1066. doi: 10.1109/CVPR.1997.609461.

```
In [13]: def plotBasRelief(B, mu, nu, lam, ax):
             Question 2 (f)
             Make a 3D plot of of a bas-relief transformation with the given parameters.
             Parameters
             B : numpy.ndarray
                 The 3 x P matrix of pseudonormals
             mu : float
                 bas-relief parameter
             nu : float
                 bas-relief parameter
             lambda : float
                 bas-relief parameter
             Returns
                 None
             P = np.asarray([[1, 0, -mu/lam],
                                                  [0, 1, -nu/lam],
                                                 [0, 0, 1/lam]])
```

```
Bp = P.dot(B)
   surface = estimateShape(Bp, s)
   plotSurfaceGrid(surface, suffix=f'br_{mu}_{nu}_{lam}', ax=ax)
# keep all outputs visible
from IPython.display import Javascript
display(Javascript('''google.colab.output.setIframeHeight(0, true, {maxHeight: 5000})'''))
cols = 3
fig, axes = plt.subplots(3, cols, figsize=(12, 10),
                         subplot_kw={'projection': '3d'})
# Base case
mu, nu, lam = 0.5, 0.5, 0.1
plotBasRelief(int_pseudonormals, mu, nu, lam, axes[0, 0])
# Vary mu (x-axis)
plotBasRelief(int pseudonormals, 0.1, nu, lam, axes[0, 1])
plotBasRelief(int_pseudonormals, 0.75, nu, lam, axes[0, 2])
# Vary nu (y-axis)
plotBasRelief(int_pseudonormals, mu, 0.1, lam, axes[1, 0])
plotBasRelief(int_pseudonormals, mu, 0.75, lam, axes[1, 1])
# Vary lam (z-axis)
plotBasRelief(int pseudonormals, mu, nu, 0.5, axes[1, 2])
plotBasRelief(int_pseudonormals, mu, nu, 10, axes[2, 0])
# Vary mu and nu (inclination)
plotBasRelief(int_pseudonormals, 0.25, 0.25, lam, axes[2, 1])
plotBasRelief(int_pseudonormals, 0.0001, 10, lam, axes[2, 2])
```

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Q 2 (g): Flattest surface possible (5 points)

A: for the flattest surface posible we want no biasing in the inclinations so $\mu=0$ and $\nu=0$. We want to reduce the scale of the image as much as possible, without setting it to 0, as this would result in a non-exisiting surface. For example, we can achieve a very falt surface with $\mu=0, \nu=0, \lambda=1e^{-100}$

Q 2 (h): More measurements

A: Acquiring more images will not resolve the ambiguity. Though adding more images might result in a better shape estimation, the ambuity arises from the fact that in uncalibrated photometric stereo, the surface can only be determined up to a certain rotation and skew.

This is because the underlying mechanism to estimate the lightning sources and the pseudonormals remains the same. While adding more images might produce less noisy singular valeus, the factors themselves are not guaranteed to be faithfull to the real conditions under which the photographs where captured.