



# Nomenclature

<b>AC</b>	Alternating Circuit
<b>APA</b>	Anode Plane Assemblies
<b>ArgoNeuT</b>	Argon Neutrino Teststand
<b>BDT</b>	Boosted Decision Tree
<b>BNB</b>	Booster Neutrino Beam
<b>BSM</b>	Beyond Standard Model
<b>C.L.</b>	Confidence Level
<b>CC</b>	Charged Current
<b>CE</b>	Cold Electronics
<b>CERN</b>	European Organisation for Nuclear Research
<b>CPA</b>	Cathode Plane Assembly
<b>CPT</b>	Charge Parity Time symmetries
<b>CRT</b>	Cosmic Ray Tagger
<b>CRUMBS</b>	Cosmic Rejection Using a Multi-system Boosted decision tree Score
<b>DAQ</b>	Data AcQuisition
<b>DCA</b>	Distance of Closest Approach
<b>DUNE</b>	Deep Underground Neutrino Experiment

<b>ETRIG</b>	Event TRIGger
<b>FD</b>	Fine Delay
<b>FEB</b>	Front End Board
<b>FMC</b>	Field programmable gate arrays Mezzanine Card
<b>FTRIG</b>	Flash TRIGger
<b>HNL</b>	Heavy Neutral Lepton
<b>ICARUS</b>	Imaging Cosmic And Rare Underground Signals
<b>KDAR</b>	Kaon Decay At Rest
<b>KEK</b>	High Energy Accelerator Research Organization (Kō Enerugi Kasokuki Kenkyū Kikō)
<b>LArTPC</b>	Liquid Argon Time Projection Chamber
<b>LSND</b>	Liquid Scintillator Neutrino Detector
<b>MC</b>	Monte Carlo
<b>MCS</b>	Multiple Coulomb Scattering
<b>MicroBooNE</b>	Micro Booster Neutrino Experiment
<b>MiniBooNe</b>	Mini Booster Neutrino Experiment
<b>MIP</b>	Minimum Ionising Particle
<b>MTC/A</b>	Master Trigger Card Analog
<b>NC</b>	Neutral Current
<b>NTB</b>	Nevis Trigger Board
<b>NTP</b>	Network Time Protocol
<b>NuMI</b>	Neutrinos at the Main Injector beam
<b>NuTeV</b>	Neutrino at the Tevatron
<b>PDF</b>	Probability Density Function
<b>PDS</b>	Photon Detection System

## Nomenclature

---

<b>PE</b>	PhotoElectron
<b>PID</b>	Particle IDentification
<b>PIENU</b>	<b>P</b> ion $\rightarrow$ <b>E</b> lectron + <b>N</b> eutrino
<b>PMNS</b>	Pontecorvo-Maki-Nakagawa-Sakata
<b>PMT</b>	PhotonMultiplier Tube
<b>POT</b>	Protons On Target
<b>PPS</b>	Pulse Per Second
<b>PTB</b>	Penn Trigger Board
<b>QE</b>	Quantum Efficiency
<b>RWM</b>	Resistor Wall Monitor
<b>s.d.</b>	standard deviation
<b>SBN</b>	Short-Baseline Neutrino
<b>SBND</b>	Short-Baseline Near Detector
<b>SCE</b>	Space Charge Effect
<b>SER</b>	Single Electron Response
<b>SIN</b>	Swiss Institute for Nuclear Research
<b>SiPM</b>	Silicon PhotoMultiplier
<b>SM</b>	Standard Model
<b>SPEC</b>	Simple PCIe FMC Carrier
<b>SVEC</b>	Simple VME FMC Carrier
<b>T2K</b>	Tokai to Kamioka
<b>TAI</b>	Coordinated Universal Time
<b>TDC</b>	Time to Digital Converter
<b>TPB</b>	TertraPhenyl Butadiene

<b>TPC</b>	Time Projection Chamber
<b>UTC</b>	International Atomic Time
<b>VUV</b>	Vacuum Ultraviolet
<b>WIB</b>	Warm Interface Board
<b>WR</b>	White Rabbit

# Chapter 1

## Introduction

Neutrino oscillation implies that neutrinos have mass, which cannot be currently explained by the Standard Model (SM) due to an absence of right-handed partners to the left-handed neutrinos. Chapter 2 begins with the motivation of a right-handed heavy neutrino state, also referred to as *Heavy Neutral Leptons* (HNLs), that allows for the construction of neutrino mass. An overview of the theoretical models of HNLs is additionally given, covering their productions and decays. The focus is on kinematically-allowed channels that can be produced from the Booster Neutrino Beam (BNB) and subsequently decay inside the Short-Baseline Near Detector (SBND).

In following, a description of Liquid Argon Time Projection Chamber (LArTPC) is provided in Chapter 3, which is the main detector technology of SBND. The operating principles of a LArTPC is presented, identifying key physical processes of the two main detection signals, ionisation electrons and scintillation photons, that underpin the performance of a LArTPC.

Chapter 4 then provides an overview of the SBND and the BNB. The chapter begins with the physics program of SBND, followed by the detector design, describing each subsystem that comprise the detector. The BNB is discussed next, detailing the beam design and presenting the secondary meson fluxes and neutrino fluxes arriving at SBND.

The simulation framework at SBND is outlined in Chapter 5, to produce Monte Carlo (MC) samples ideally representing data. A description on different generators to simulate SM neutrinos, cosmic muons and HNLs is first provided. The HNL generator is presented in details to illustrate the physics behind the lateness of HNLs compared to SM neutrinos produced from the BNB, that the thesis work presented in later chapters rely upon. Finally, the simulation of the particle propagation and the detection response is summarised.

Following that, the reconstruction framework is provided in Chapter 6, covering the reconstruction for each detection subsystem: (1) TPC, (2) Photon Detection System (PDS) and (3) Cosmic Ray Tagger (CRT). Specifically in the TPC reconstruction workflow, an update to an algorithm separating track-like and shower-like reconstructed object is detailed. An overview of some high-level analysis tools, combining complementary signals from all subsystems, is discussed next.

Chapter 7 outlines the timing performance of the Data Acquisition (DAQ) at SBND. The chapter begins with a description of the White Rabbit timing system setup to maintain timing synchronisation across different DAQ subsystems. The timing precision of the readout electronics of the CRT and PDS are then accessed, which are the two detection subsystems with signals  $\mathcal{O}(1\text{ ns})$ .

Some studies within the scope of charge calibration are discussed in Chapter 8. The first study is on the measurement of electron lifetime, performed on MC samples of anode-to-cathode crossing cosmic muon tracks that fully traverse the detector volume. The second study is to assess the impacts of delta ray fluctuations on recombination, via simulation studies varying delta ray thresholds.

A selection procedure has been developed to select HNLs and reject backgrounds from SM neutrinos and cosmic muons, which will be presented in Chapter 9. Firstly, a signal and background definition is provided, followed by a description of MC samples used to perform the selection. Dedicated cuts for cosmic background and SM neutrinos rejection and HNL showers selection are detailed subsequently. The result of the selection is analysed next, identifying the most impactful cuts of the selection. Finally, a *truth*-based study is presented, proposing a hypothetical scenario if the resolution of the timing reconstruction is improved.

Finally, the capabilities of SBND to search for HNLs is assessed in Chapter 10. The systematic uncertainties in the analysis are discussed, covering their sources and treatments. The procedure used to set an upper limit at the 90% confidence level is detailed next. The result of the limit setting is then presented, covering three result projections that can be achieved at SBND. A discussion of the results is finally given, including suggestions for future developments.

