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	8.01 Exam - 2
(· a,b)	$M: N- mg \cos\theta = m(0) \Rightarrow N = mg \cos\theta$
	$x: R \times_{max} - \mathcal{U}_{S}N - mgsin\theta = 0$ $R \times_{max} = mg(\mathcal{U}_{S}\cos\theta + \sin\theta) \Rightarrow \times_{max} = mg(\mathcal{U}_{S}\cos\theta + \sin\theta)$
(d	Forces acting on the block. Spring Force (F _s) = kxmox Normal Force (N) = N = mgcoso Gravity (mg) = mg Friction (F _t) = A _s mgcoso
	After the block is gently touched, stat friction switches to kinetic and their is a net force towards the right (towards the spring)
	The black keeps accelerating until Fret. > 0 after which it starts to decellerate.

Hence, the block reaches maximum speed whom Fret = 0RX = MkN + mgsin = = mg (Mecoso + sino) d); Wg: -mg (xmax-x) sint (ii) $W_s: -ML = -\frac{1}{2}kx^2 + \frac{1}{2}kx_{max}^2$ = 1 Rxmax - 1 kx2 (iii) Wf: - Mkmgcoso (xmax-x) Work-Energy Theorem $W_{net} = \frac{1}{2} m v_2^2 - \frac{1}{2} m v_1^2$ Substitute in a) for x = 0. For the block to reach this parition, the ret work done by all the forces must be > 0 Wg + Ws + Wf > 0 - mg x max sino + 1 kx max - U k mg coso xmax > 0 1 kxmax > mg xmax (Ukcos 0 +sin0) Xmax > 2mg (Mx coso + sing)

While going up, the ycomponent of the resistive force is in the downward direction (as mg). when the object is going down, the direction is opposite. Hence the net force acting is smaller while going down and so it will take larger than 2 sec to come down. b) In normal growity, (a=0) a= 5 m/s2 $T = 2\pi \int_{\frac{Q}{2}}^{\frac{Q}{2}} = 2\pi \int_{\frac{Q}{2}}^{\frac{Q}{2}}$ When a = . 5 m/s, perceive weight = mg (m(g-a) $T' = 2\pi R = 2\pi R = \sqrt{2} \times 2\pi R$ T' = 12T The hiscous term will dominate when your K veril mg = Girvterm = Mg

2. a)

$$C_{1}\Gamma V_{L} >> C_{2}\Gamma^{2}V_{L}^{2} \Rightarrow V_{L} << C_{1}$$

$$C_{1}\Gamma V_{L} >> C_{2}\Gamma^{2}V_{L}^{2} \Rightarrow V_{L} << C_{1}$$

$$C_{2}\Gamma$$

$$m = P. 4 \pi \Gamma^{3}$$

$$4 \pi \Gamma^{3} P_{2} << C_{1} \Rightarrow \Gamma^{3} << 3c_{1}^{2}$$

$$3c_{1} \qquad C_{2} \qquad 4\pi \Gamma^{3} Q_{2}$$

$$\Gamma << \left(3c_{1}^{2}\right)^{3}$$

$$4 \pi \Gamma^{3} P_{2} << C_{1} \Rightarrow \Gamma^{3} << 3c_{1}^{2}$$

$$4 \pi \Gamma^{3} Q_{2}$$

$$A \pi^{3} P_{2} <> C_{2} \Rightarrow \Gamma^{3} << 3c_{1}^{2}$$

$$A \pi^{3} Q_{2}$$

$$A \pi$$

3 a) Since Here are no external forces, the center of mass

Genains still.

So both planets should make together; e have the same argular valarity.

b)
$$F_g = G_1 m_1 m_2$$
 $G_1 + G_2$

an $G_2 = G_2 G_1 G_2$
 $G_3 = G_4 G_2$
 $G_4 + G_2$
 $G_4 = G_4 G_4$
 $G_4 = G_4$

$$T^{2} = \frac{4\pi r_{1}^{2}}{6m_{2}r_{1}} \left(r_{1}+r_{2}\right)^{2} = \frac{4\pi r_{1}}{6m_{2}} \left(r_{1}+r_{2}\right)^{2}$$

$$\frac{r_2}{r_1} = \frac{m_1}{m_2} \Rightarrow \frac{r_2}{r_1} + 1 = \frac{m_1}{m_2} + 1$$

$$\frac{\Gamma_1 + \Gamma_2}{\Gamma_1} = \frac{m_1 + m_2}{m_2} = \frac{\Gamma_1 + \Gamma_2}{m_1 + m_2}$$

$$T^{2} = \frac{4\pi (r_{1} + r_{2})^{3}}{6(m_{1} + m_{2})} \rightarrow T = \left(\frac{4\pi (r_{1} + r_{2})^{3}}{6(m_{1} + m_{2})}\right)^{1/2}$$