

The Physics of Inertial Confinement Fusion at the National Ignition Facility

Ronak Desai
Candidacy Oral Exam
March 27th 2023

DEPARTMENT OF PHYSICS

Advisor

Dr. Christopher Orban



THE OHIO STATE UNIVERSITY
COLLEGE OF ARTS AND SCIENCES

Committee

**Dr. Alexandra Landsman
Dr. Douglass Schumacher
Dr. Brian Skinner**



Academic Background

- Undergraduate: Rowan University
 - B.S. Physics
 - B.A. Mathematics
- Gap Year: Brookhaven National Laboratory
 - SULI Internship Program
- 3rd Year Grad Student: OSU
 - Plasma Physics
 - Particle In Cell Simulations
 - Machine Learning



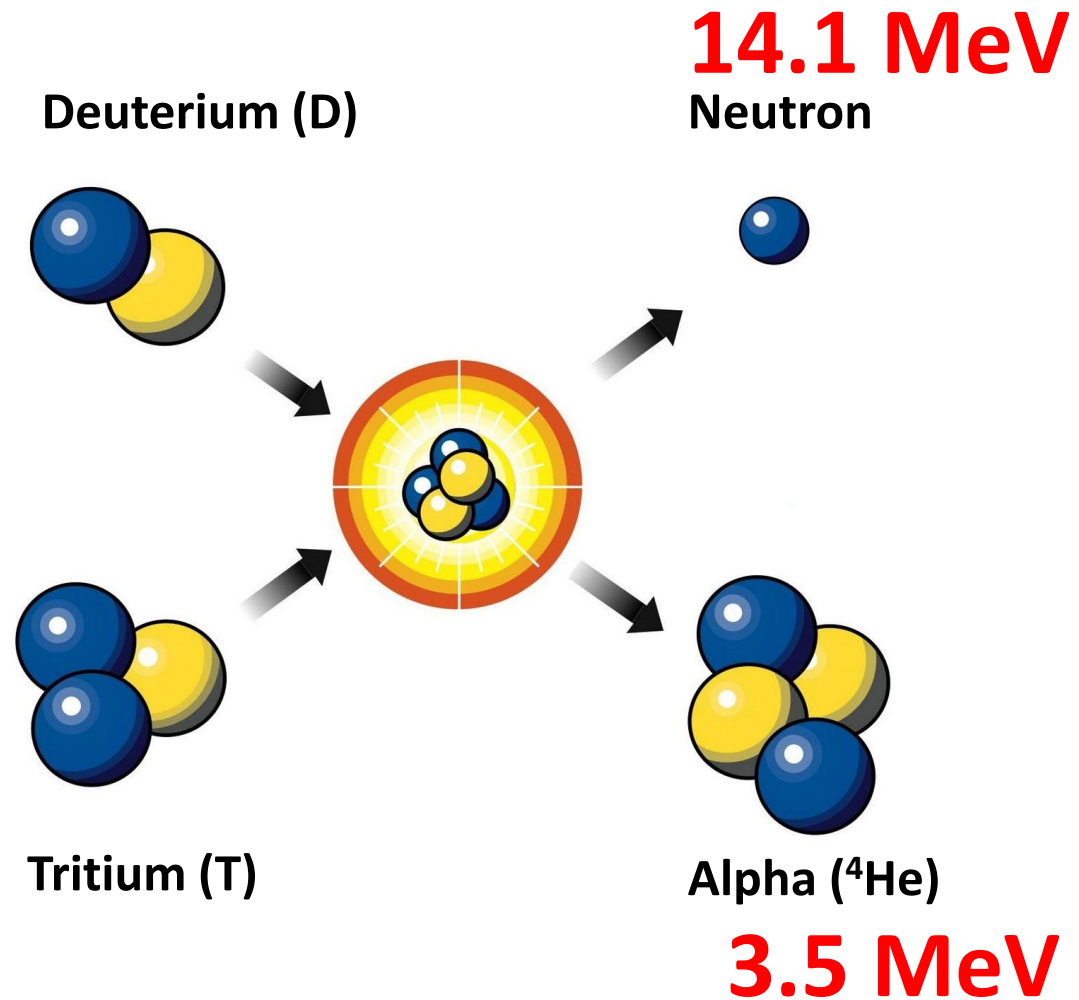
Brookhaven[®]
National Laboratory





What is Nuclear Fusion?

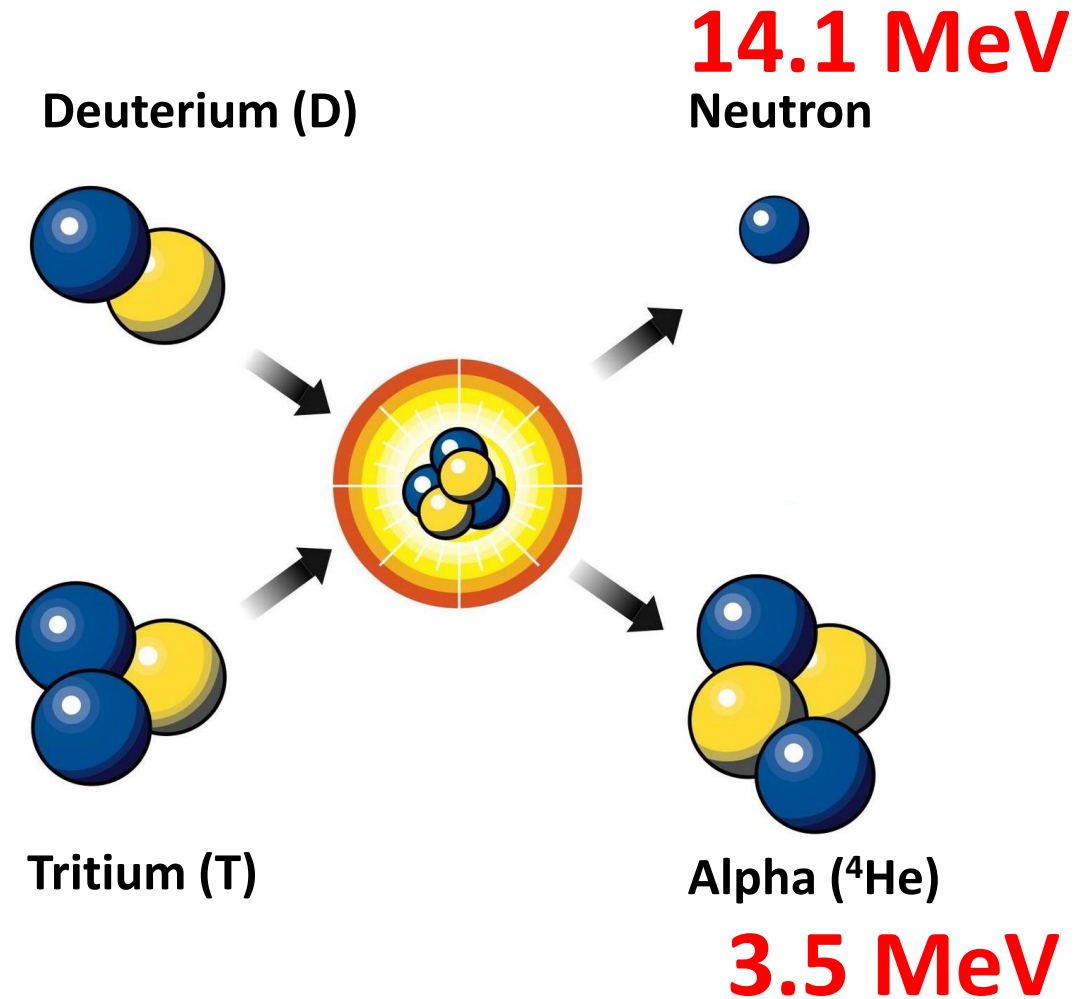
- Two lighter nuclei combine to form heavier nucleus
- $Q = (\Delta m)c^2$
 - $Q_{DT} = 17.6 \text{ MeV}$





What is Nuclear Fusion?

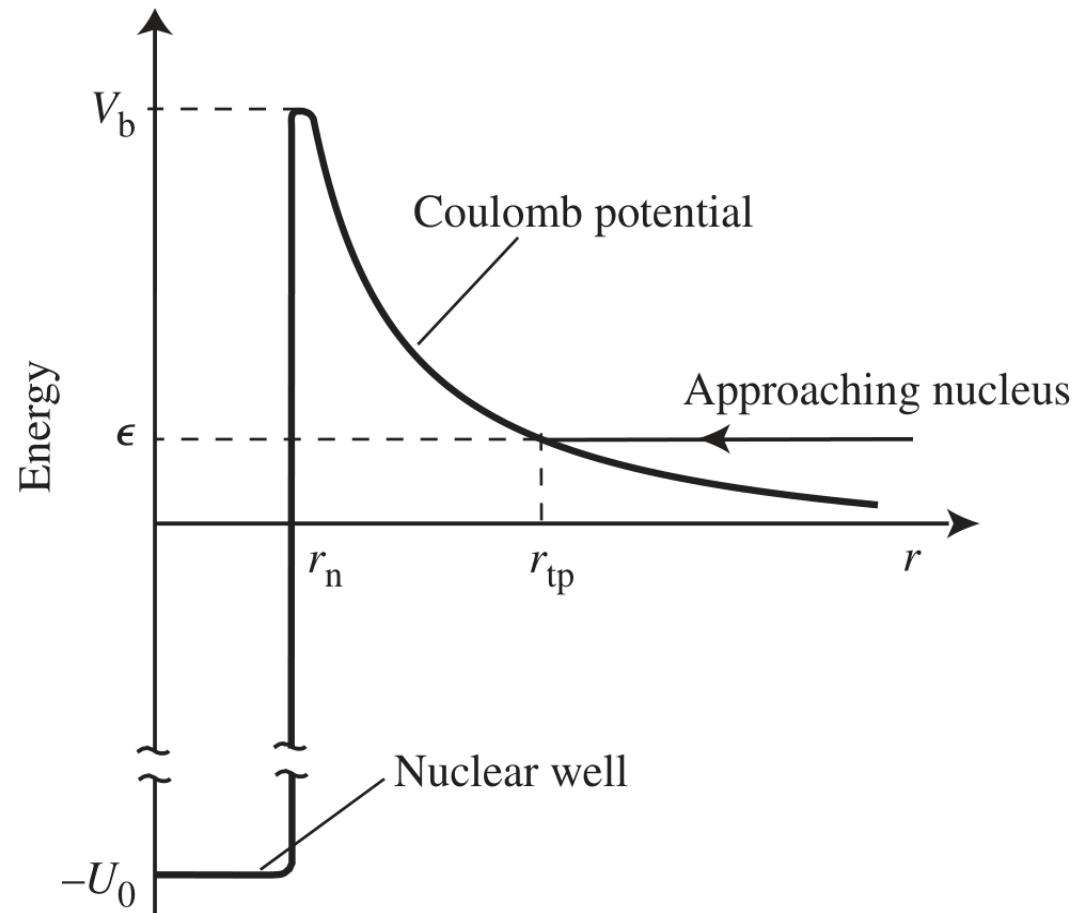
- Two lighter nuclei combine to form heavier nucleus
- $Q = (\Delta m)c^2$
 - $Q_{DT} = 17.6 \text{ MeV}$
- Compare DT and Coal:
 - DT: 300 GJ/g
 - Coal: 30 kJ/g
 - Factor of 10 Million!





How to Fuse Nuclei

- Need to overcome repulsive coulomb barrier
- $V_b \approx 1 \text{ MeV}$
 - 500x hotter than solar core

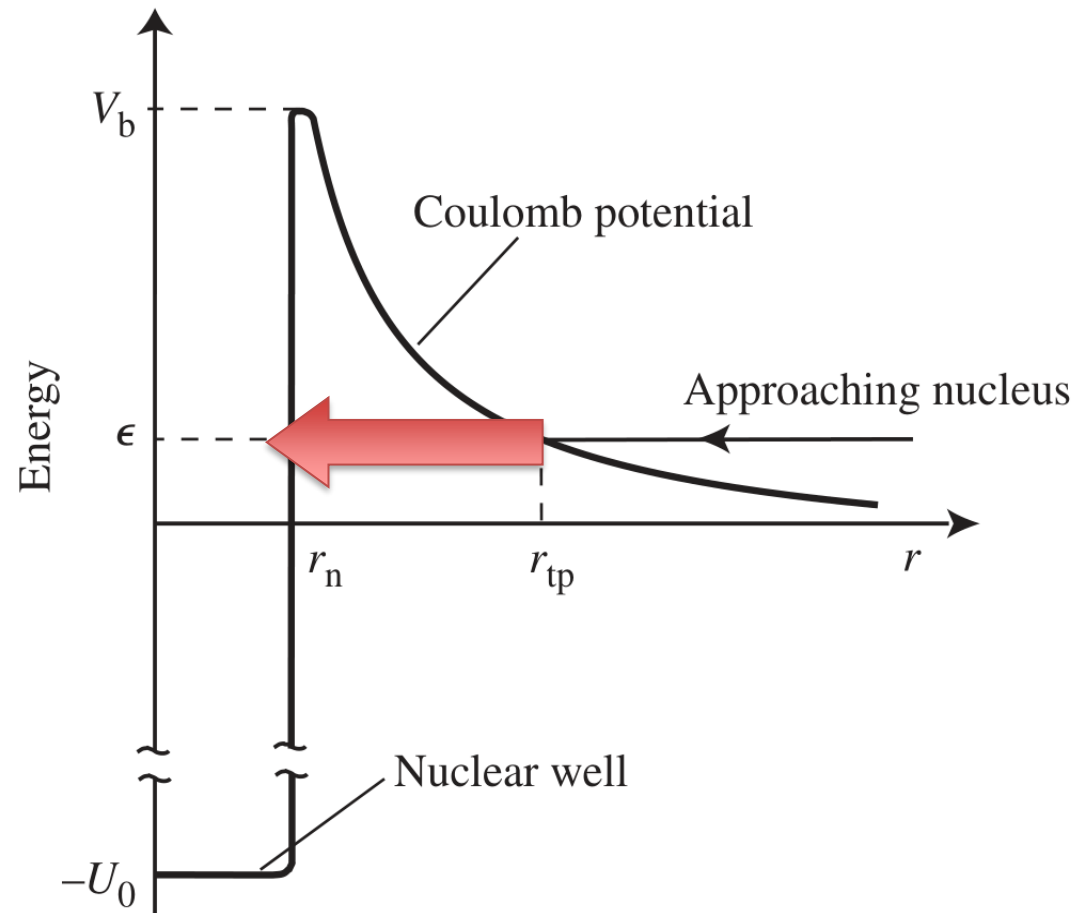


S. Atzeni, *The Physics of Inertial Fusion: Beam Plasma Interaction, Hydrodynamics, Hot Dense Matter* (2004)



How to Fuse Nuclei

- Need to overcome repulsive coulomb barrier
- $V_b \approx 1 \text{ MeV}$
 - 500x hotter than solar core
- Can Tunnel through barrier Quantum Mechanically

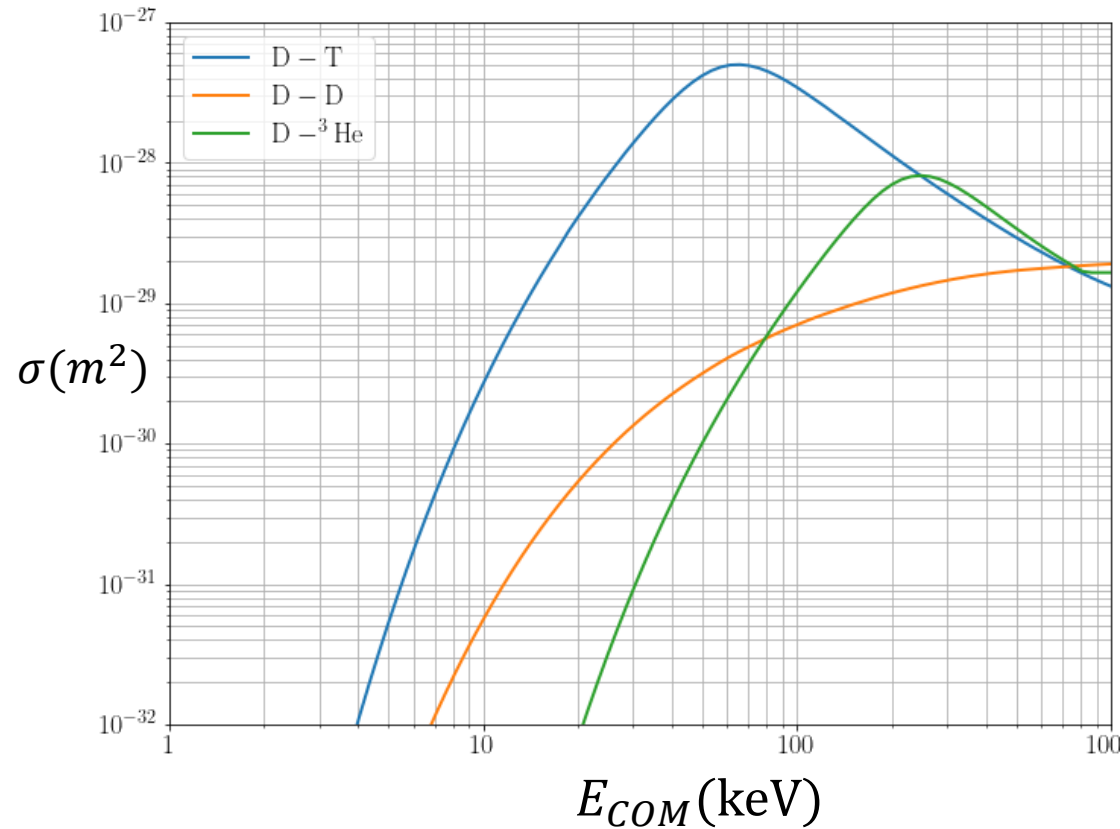


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Why DT Fusion?

- High Fusion Cross Section (FCS): σ
 - Dependent on Geometric Cross Section and Tunneling Probability

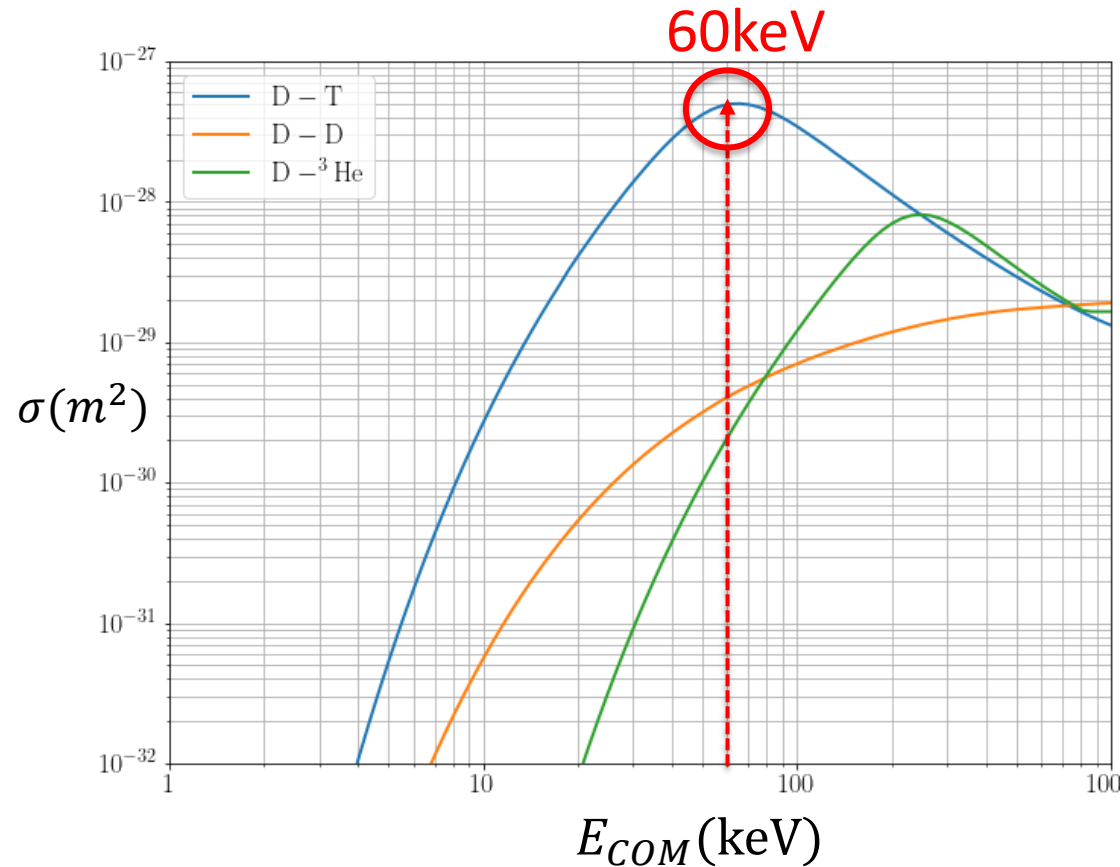


<https://scipython.com/blog/plotting-nuclear-fusion-cross-sections/>



Why DT Fusion?

- High Fusion Cross Section (FCS): σ
 - Dependent on Geometric Cross Section and Tunneling Probability
- Low Temperature
 - $E_{COM} \sim 60 \text{ keV}$
 - $T \sim 10 \text{ keV}$
 - $1 \text{ keV} \sim 11 \text{ Million K}$



<https://scipython.com/blog/plotting-nuclear-fusion-cross-sections/>



Lawson Criterion

- Conditions for sustained fusion?
 - n : Number density high enough for frequent collisions
 - τ : Long confinement time for fuel to fully burn



Lawson Criterion

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$$n\tau > 10^{15} \text{ s/cm}^3$$

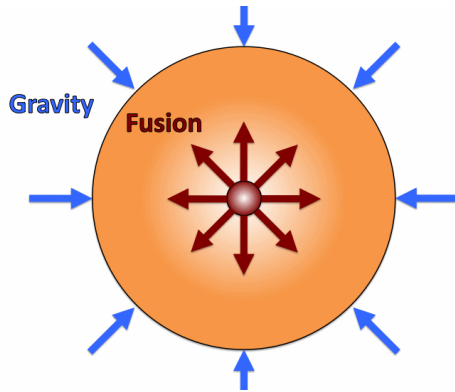
- for DT fusion

$$n\tau > \frac{12k_B T}{\langle \sigma v \rangle Q}$$

- T = Temperature
- Q = Energy Released
- $\langle \sigma v \rangle$ = Fusion Cross Section integrated over Maxwell-Boltzmann Distribution



Types of Confinement

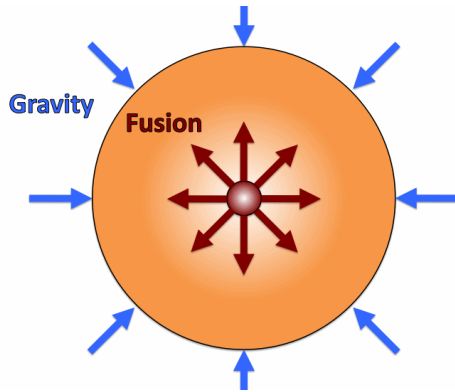


<http://large.stanford.edu/courses/2011/ph241/olson1/>

Gravitational

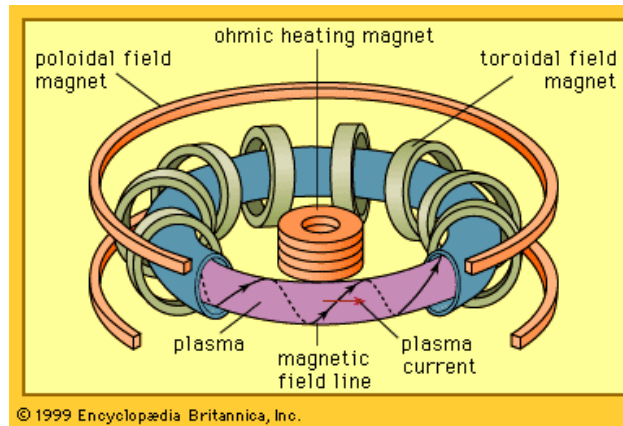


Types of Confinement



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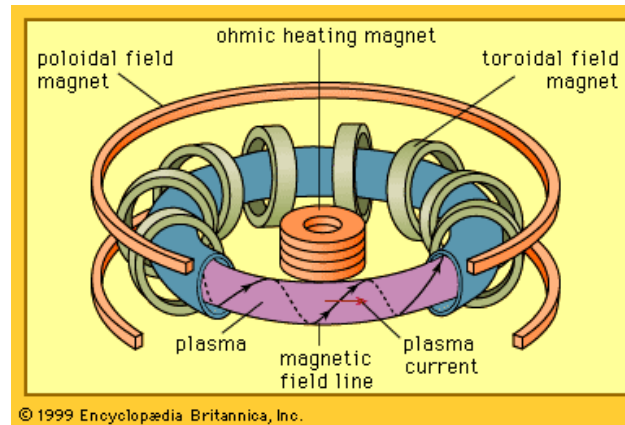
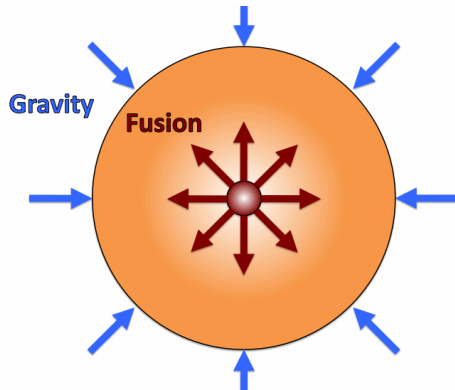


<https://www.britannica.com/technology/fusion-reactor/Principles-of-magnetic-confinement>

Magnetic



Types of Confinement



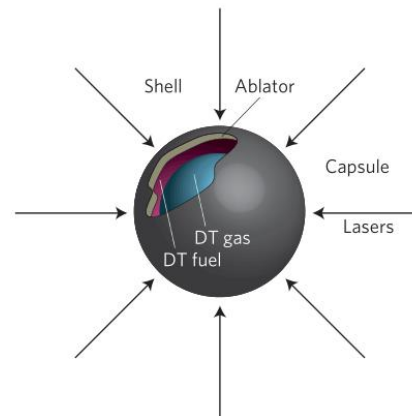
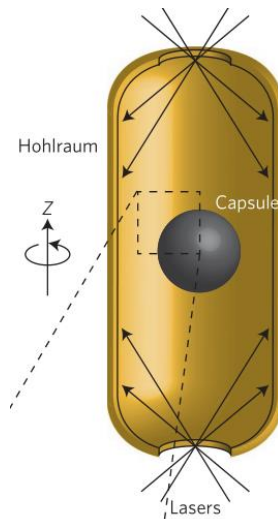
Magnetic

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Gravitational

Inertial (ICF)

Indirect Drive



Nature Phys **12**, 435–448 (2016)

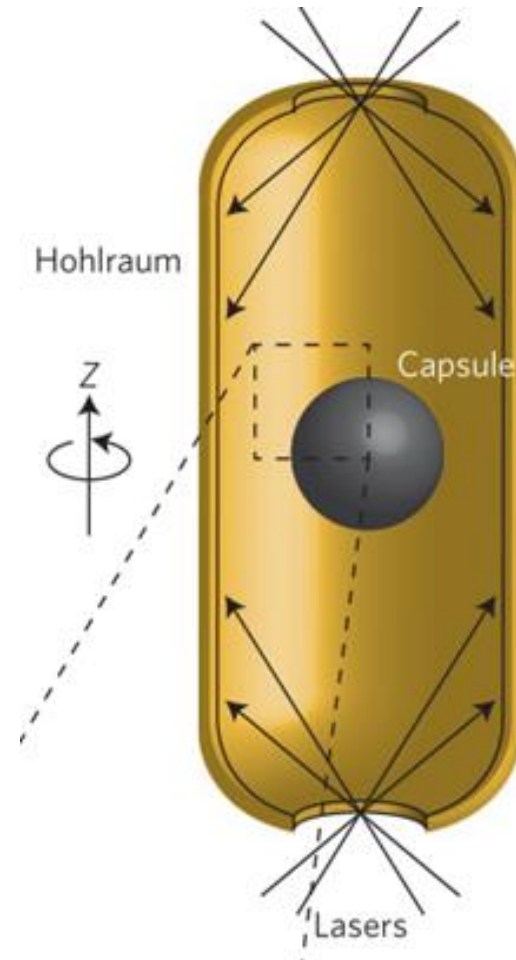
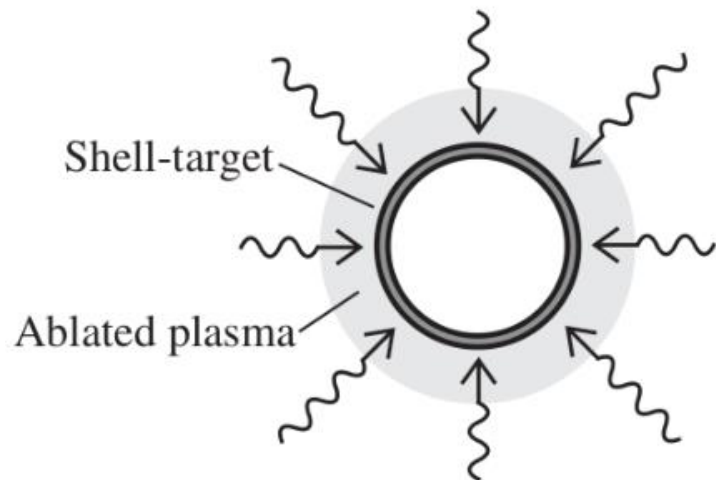
Direct Drive



ICF Stages

1. Laser Heating

Nature Phys **12**, 435–448 (2016)



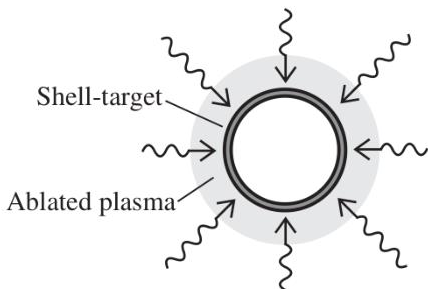
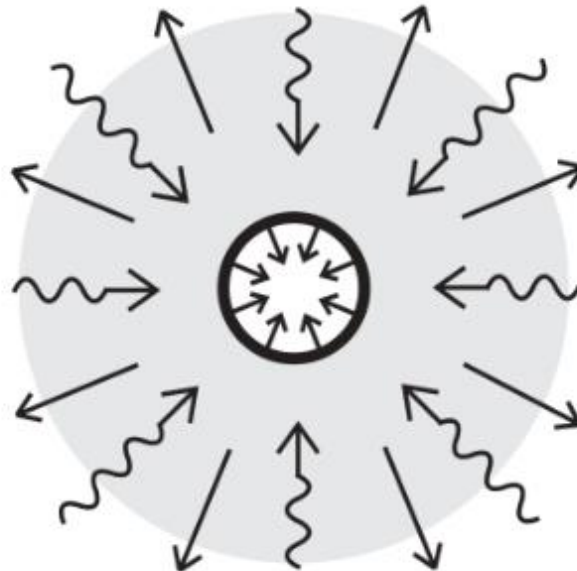
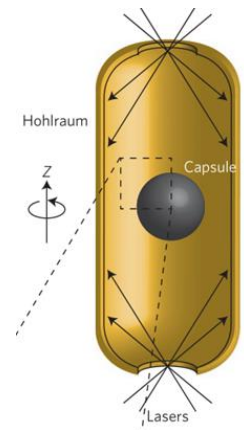


ICF Stages

1. Laser Heating

2. Capsule Expansion (Ablation)

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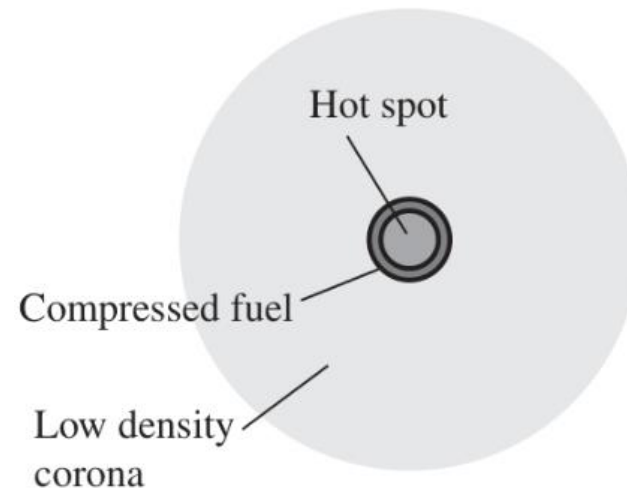
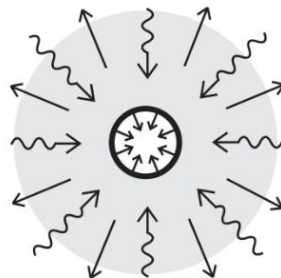
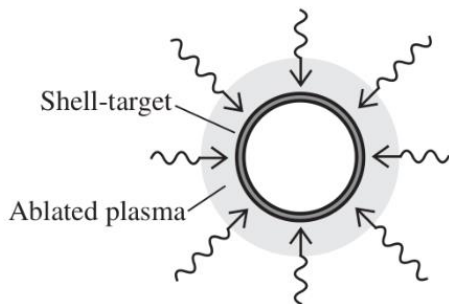
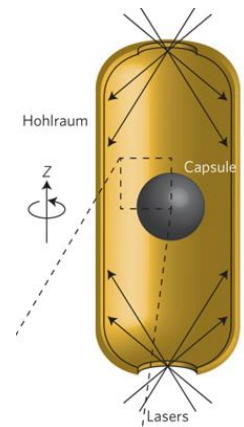
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ICF Stages

1. Laser Heating
2. Capsule Expansion (Ablation)
3. Compression and Ignition

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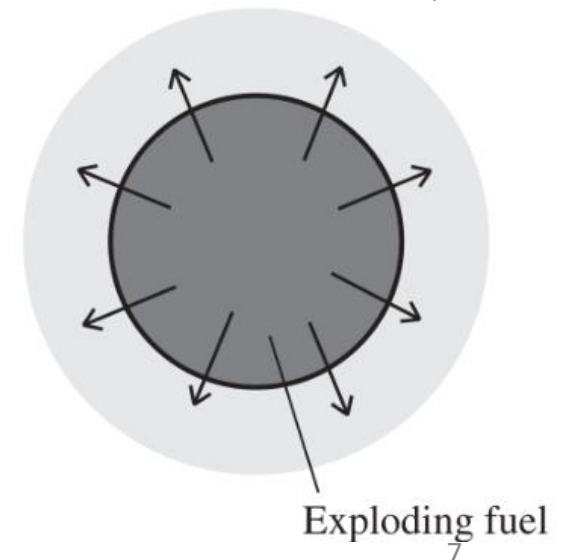
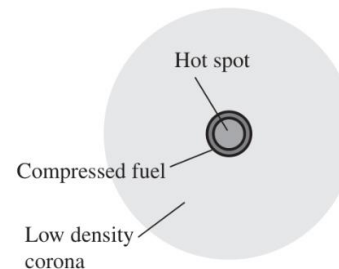
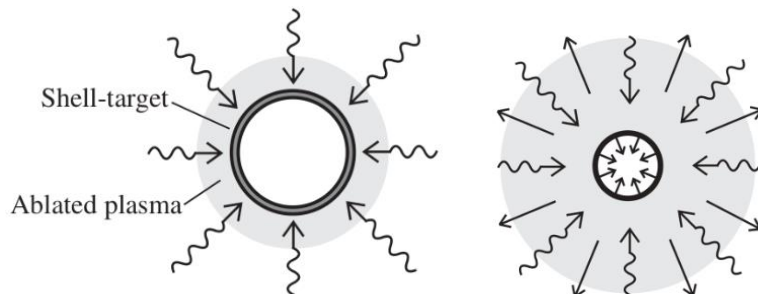
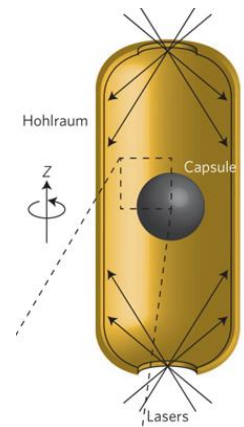
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ICF Stages

1. Laser Heating
2. Capsule Expansion (Ablation)
3. Compression and Ignition
4. Fuel Burn

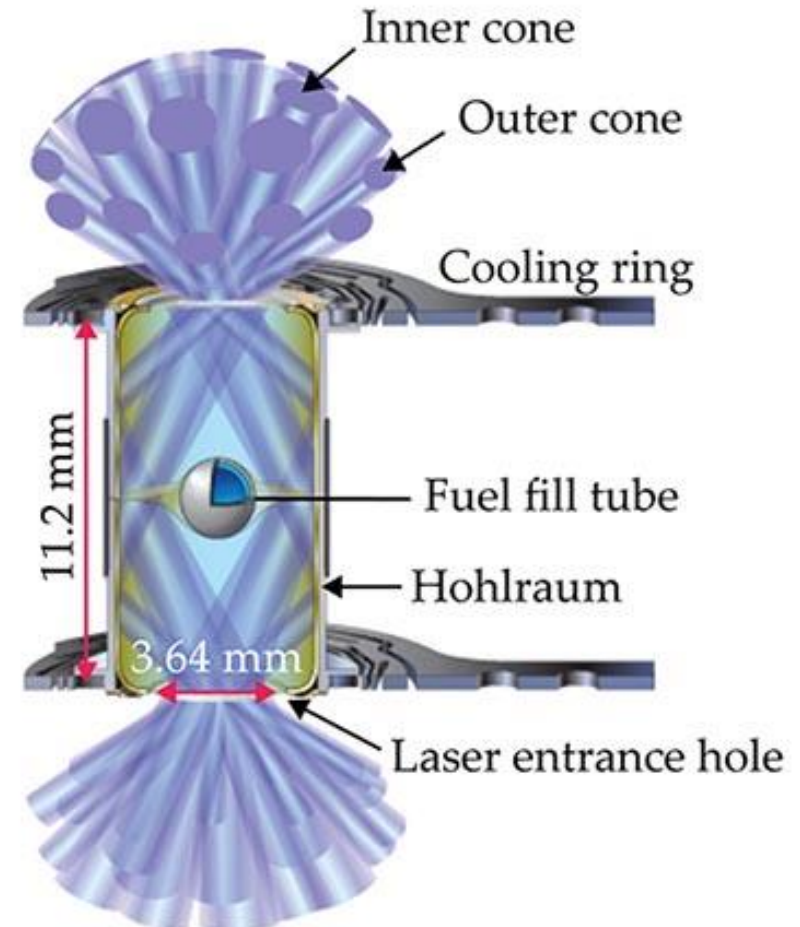
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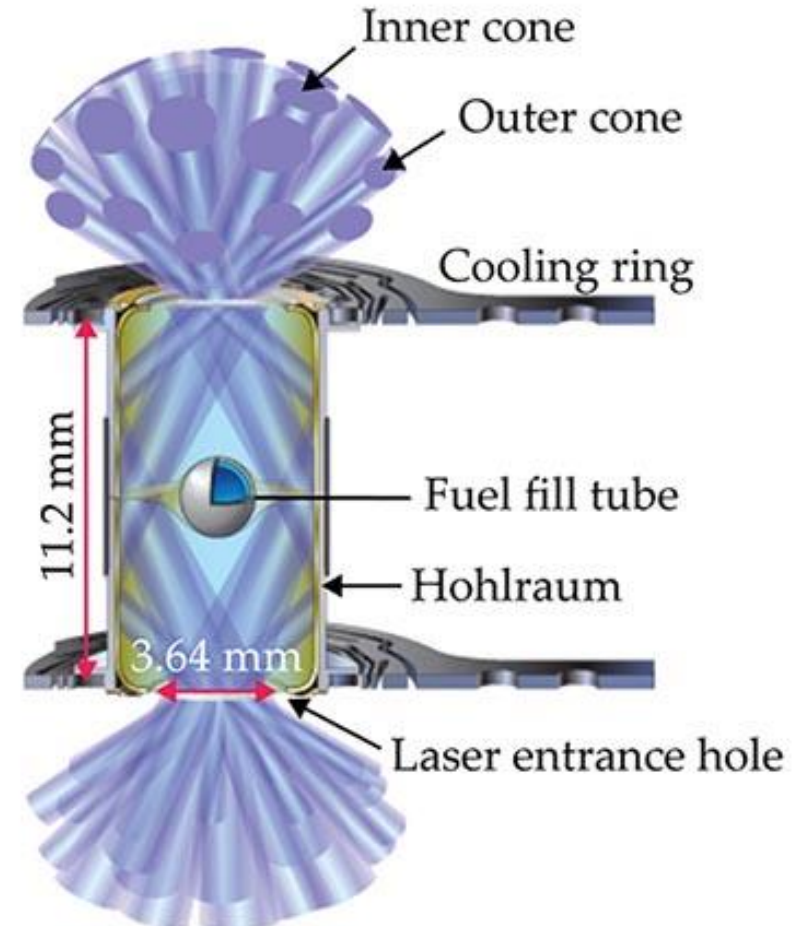
Hohlraum Temperature

- Laser enters through Laser Entrance Hole (LEH)
- Heats inner surface of cylinder



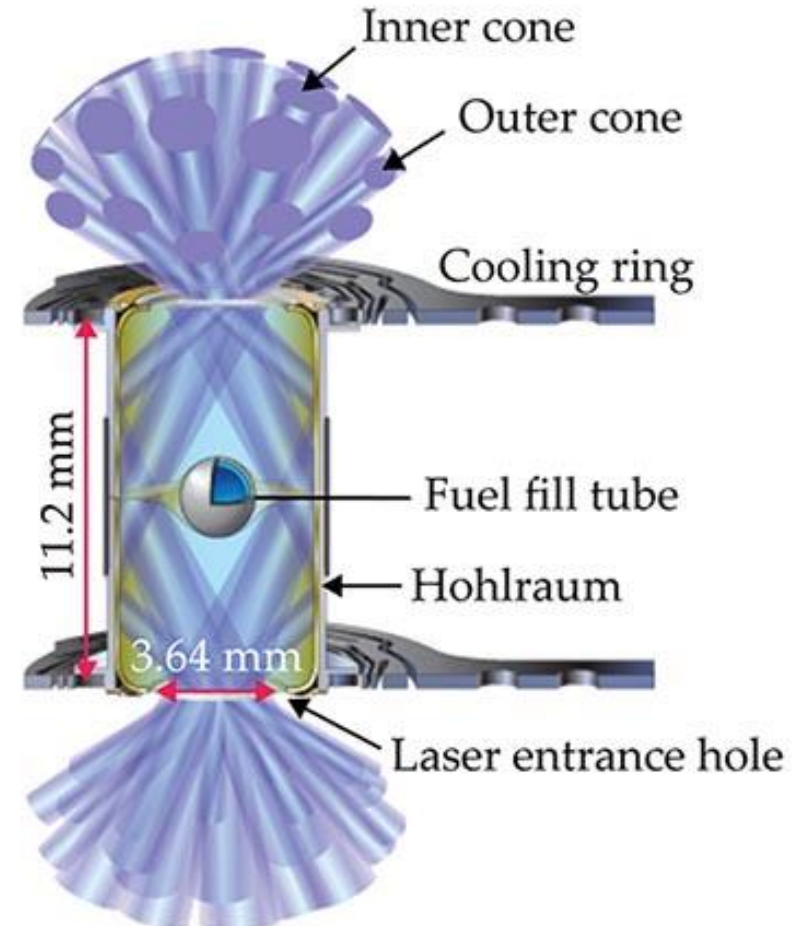
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- Laser enters through Laser Entrance Hole (LEH)
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- $I \sim \sigma_{SB} T^4$
 - $I = \frac{P \sim 500 \text{ TW}}{A_H \sim 1 \text{ cm}^2} \sim 10^{15} \text{ W/cm}^2$
 - $T_r \sim 250 \text{ eV}$



Hohlraum Temperature

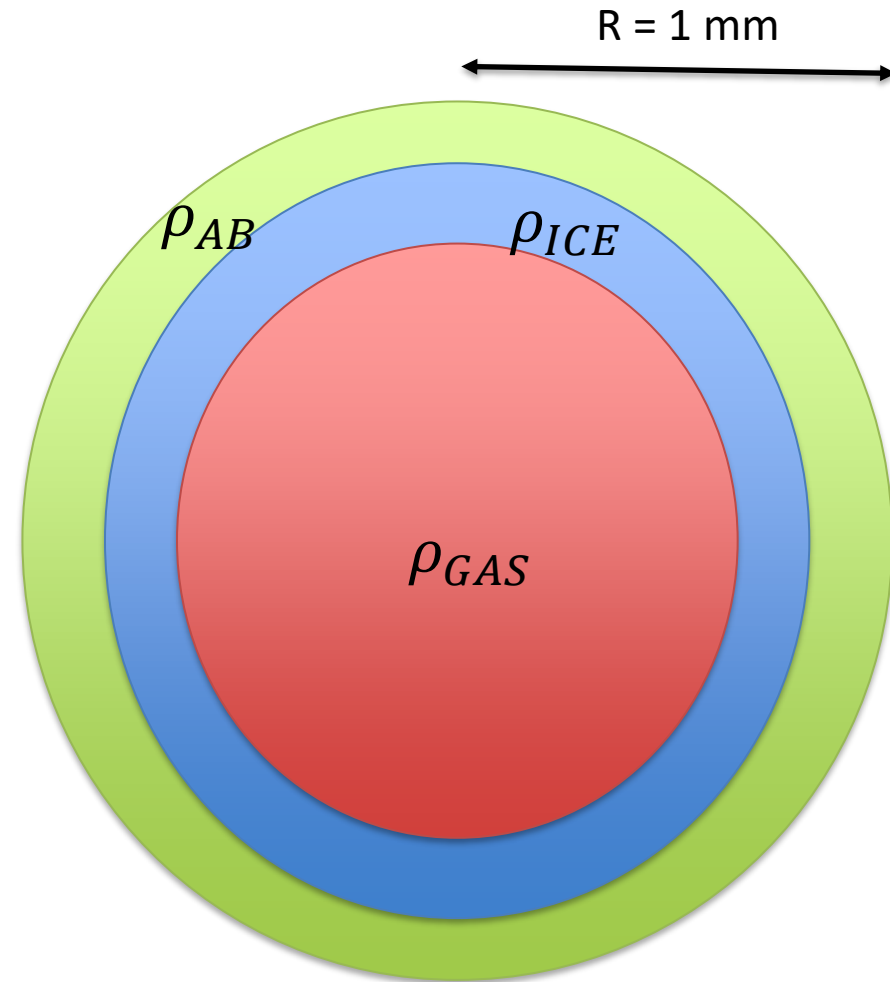
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 - $T_r \sim 250 \text{ eV}$
- Want High Temperature
 - Small Hohlraum
 - Small LEH





Capsule

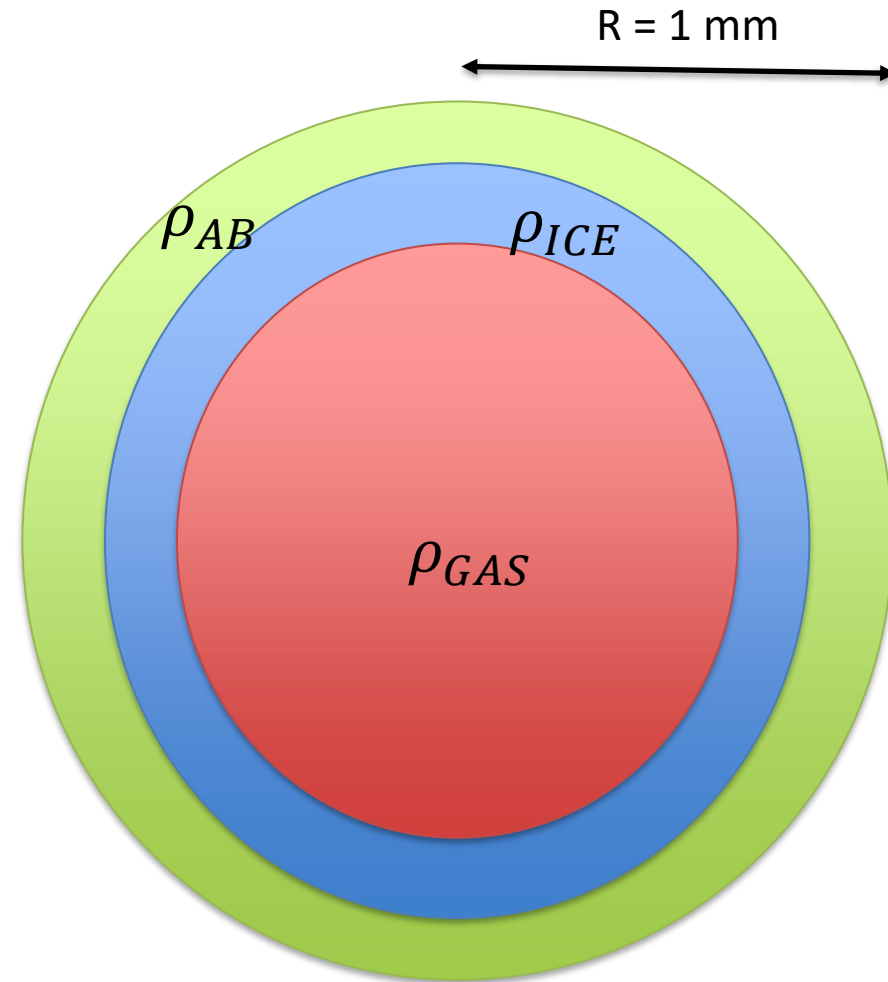
- Outer Shell: Ablator
 - Vaporized by Laser





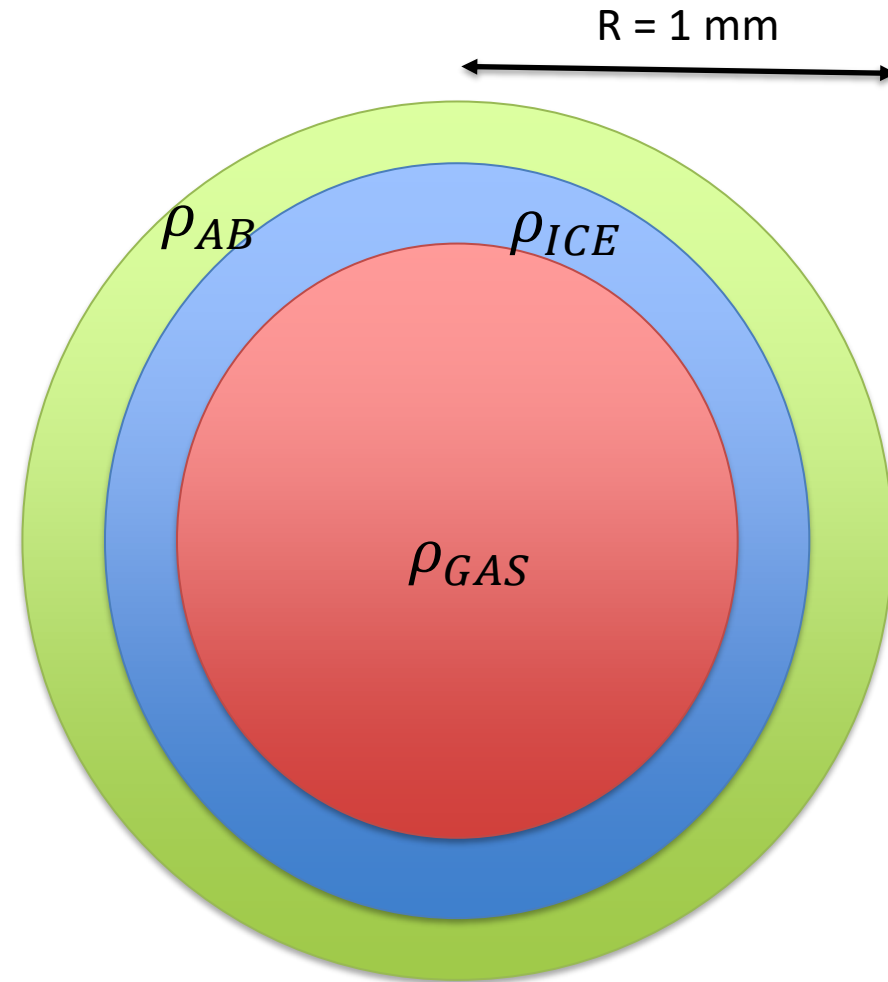
Capsule

- Outer Shell: Ablator
 - Vaporized by Laser
- Inner Shell: DT Ice
 - $T \sim 18\text{K}$
 - Most of Fuel Mass



Capsule

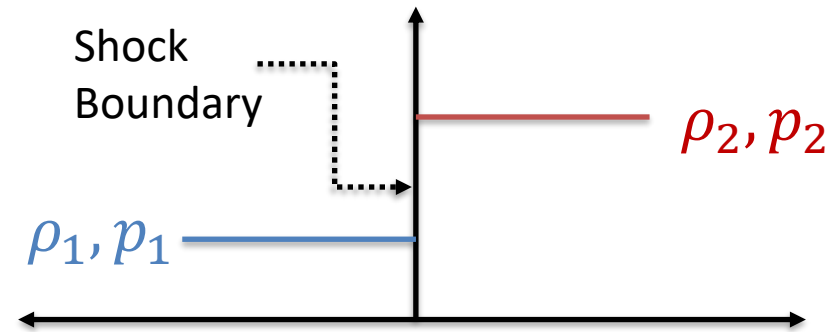
- Outer Shell: Ablator
 - Vaporized by Laser
- Inner Shell: DT Ice
 - $T \sim 18\text{K}$
 - Most of Fuel Mass
- Inner Core: DT gas
 - Low Density
 - Reaches Highest Temperature





1D Hydrodynamics

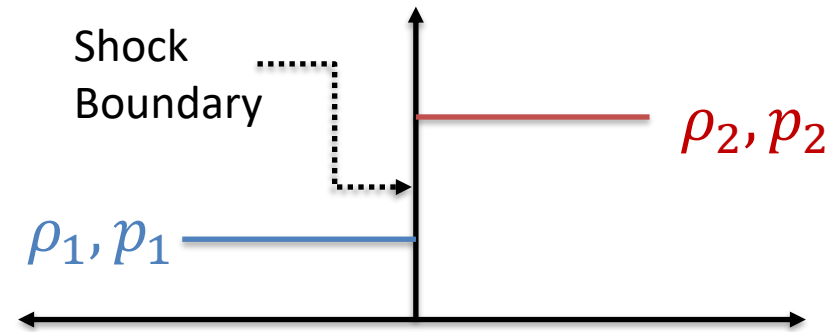
- Shock Waves Drive Compression
 - Sharp pressure changes from short pulse laser





1D Hydrodynamics

- Shock Waves Drive Compression
 - Sharp pressure changes from short pulse laser
- Euler Equations of Hydrodynamics
 - mass, momentum, energy



$$\text{Specific Volume } V \equiv \frac{1}{\rho}$$

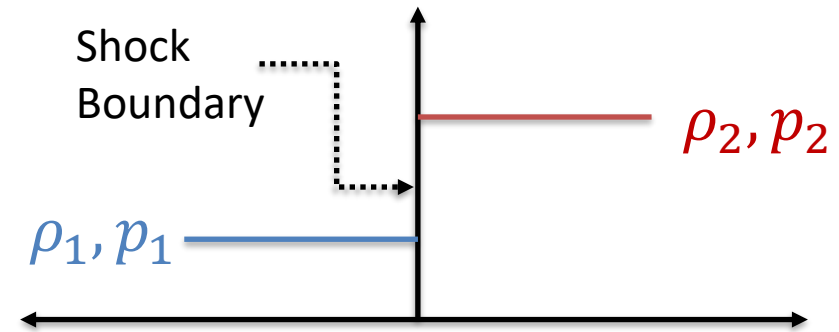
$$\text{Shock Compression} \quad \frac{p_2}{p_1} = \frac{4 - V_2/V_1}{4V_2/V_1 - 1}$$

$$\text{Isentropic Compression} \quad \frac{p_2}{p_1} = \left(\frac{V_1}{V_2} \right)^{5/3}$$



1D Hydrodynamics

- Shock Waves Drive Compression
 - Sharp pressure changes from short pulse laser
- Euler Equations of Hydrodynamics
 - mass, momentum, energy
- Ex) Want $\frac{V_2}{V_1} = \frac{1}{4}$:
 - $\frac{p_2}{p_1} \rightarrow \infty$ for 1 shock!
 - $\frac{p_2}{p_1} \approx 10$ isentropically



$$\text{Specific Volume } V \equiv \frac{1}{\rho}$$

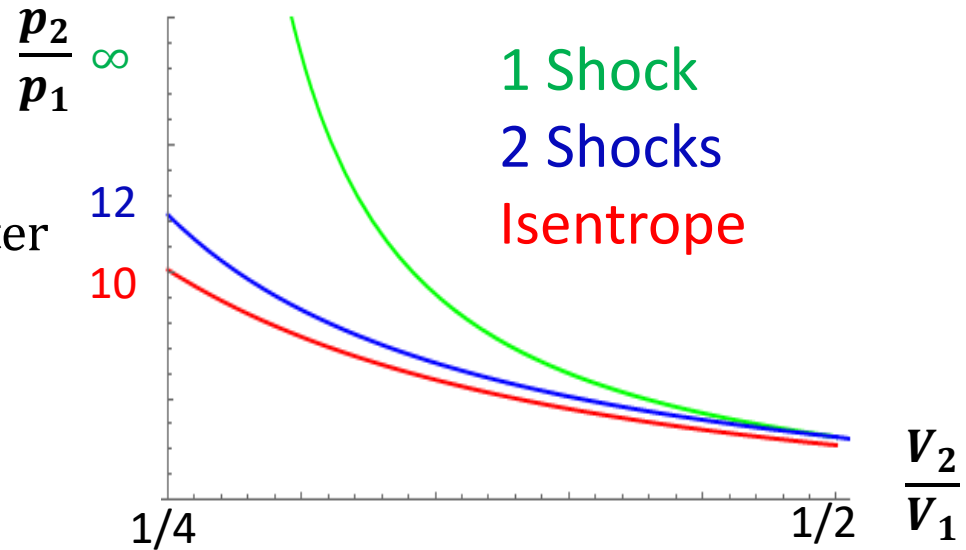
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Shocks

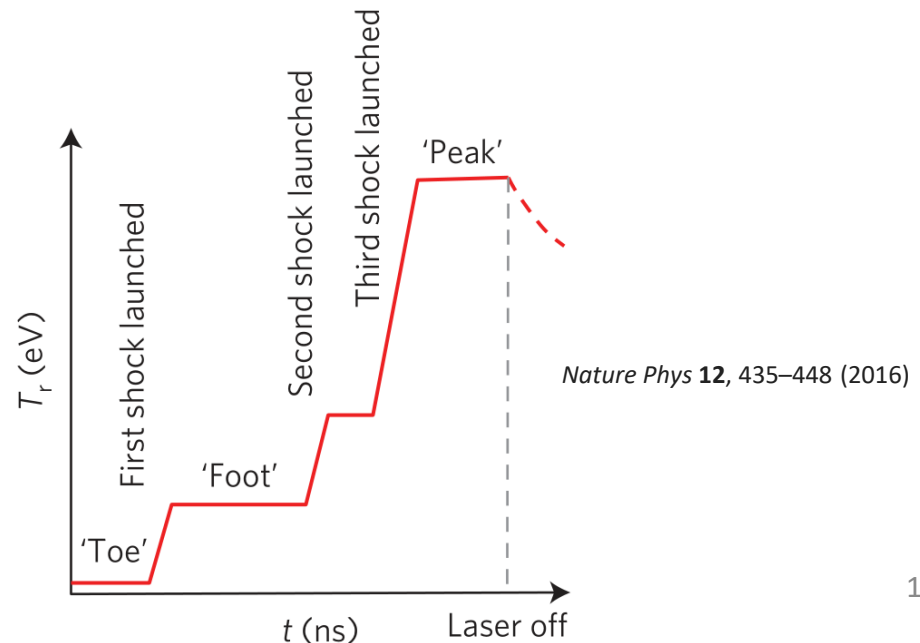
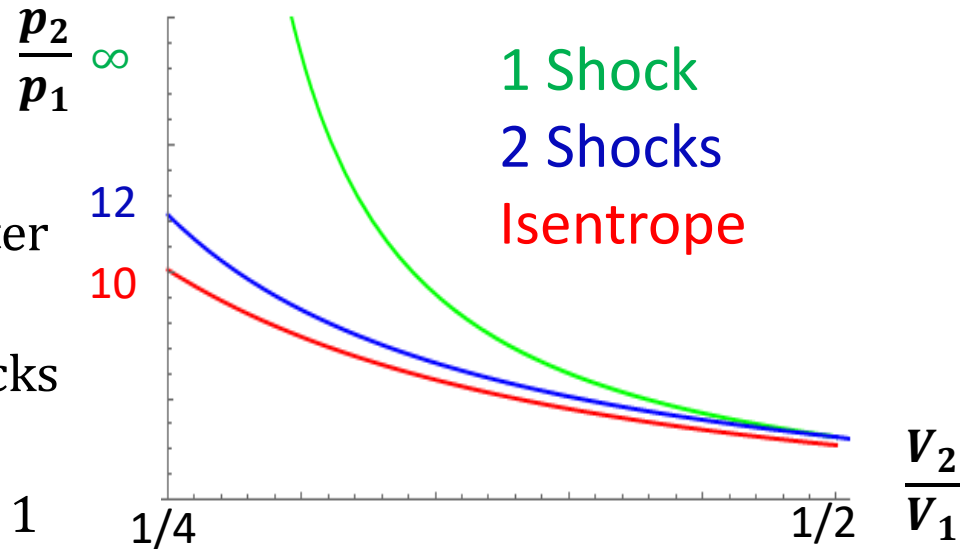
- Several Weak Shocks better than One Strong Shock





Shocks

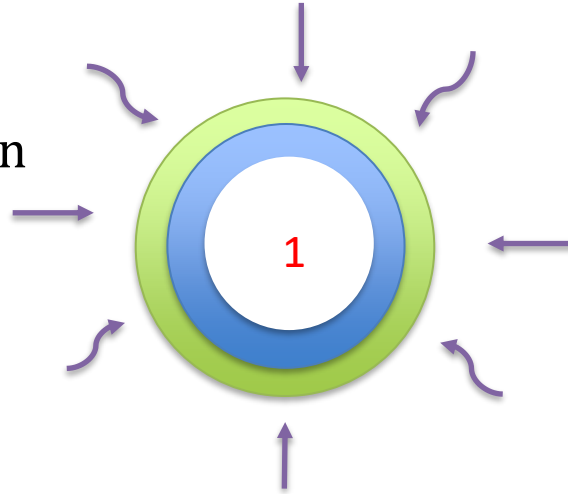
- Several Weak Shocks better than One Strong Shock
- Laser pulse tuned so shocks converge at center
- Isentrope Parameter $\alpha > 1$
 - Strong Shocks Increase
 - Want to minimize





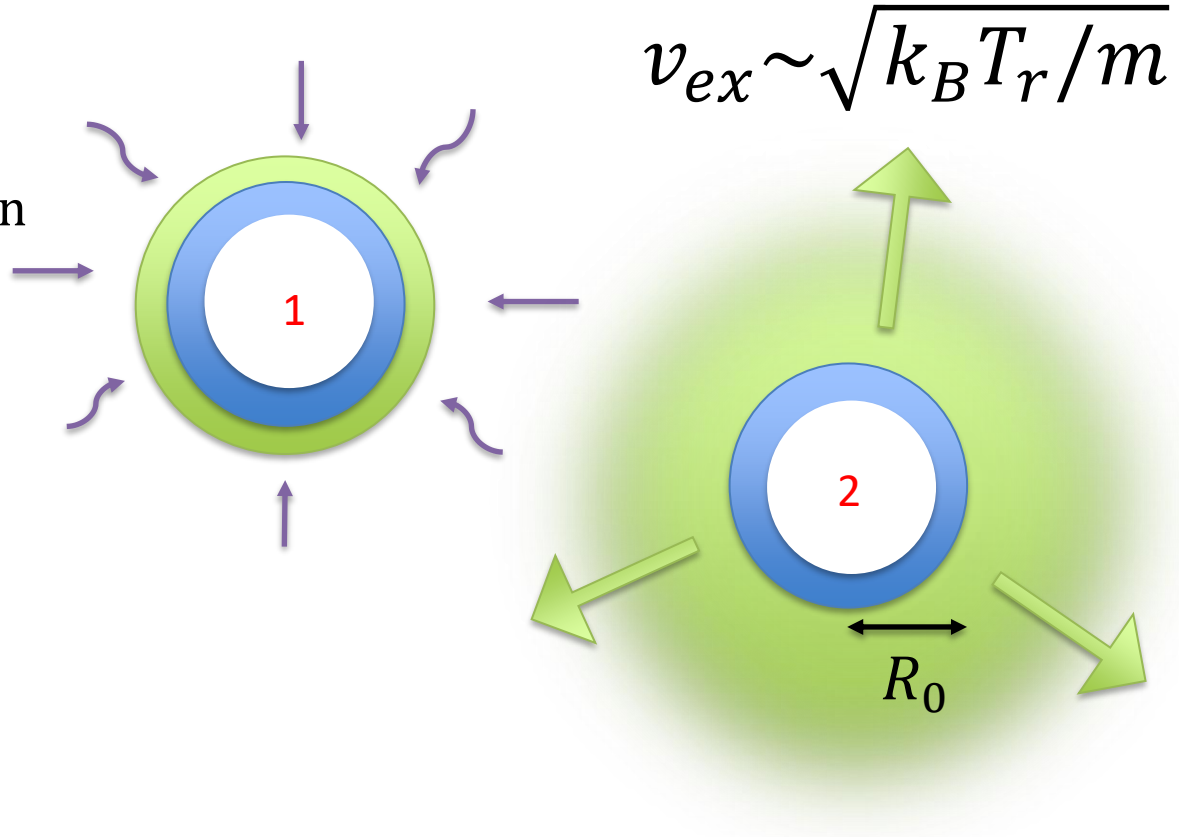
Shell Ablation

- Momentum Conservation
- (1) X-rays heat Ablator
 - Shell gets vaporized



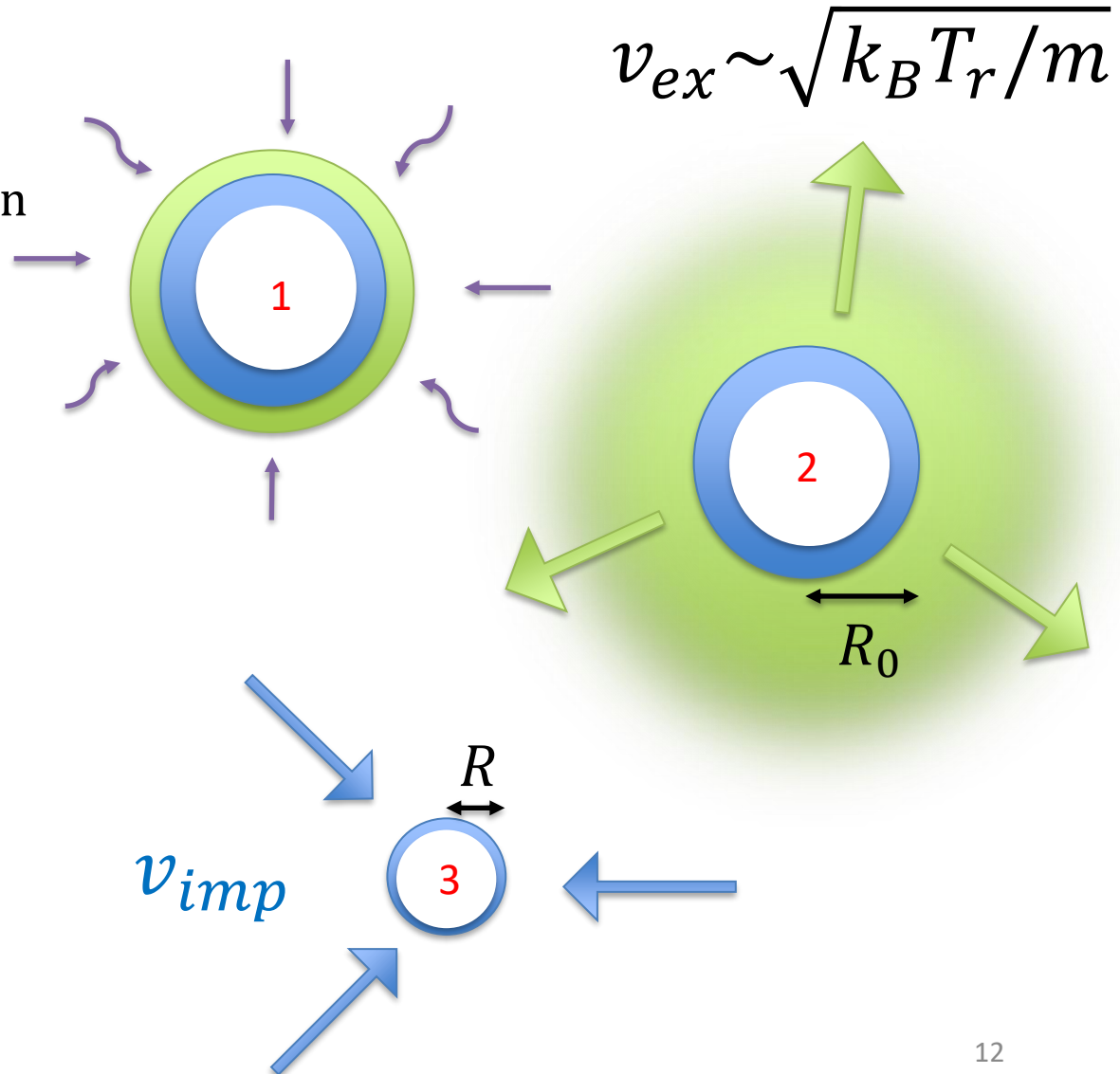
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 - outward
 - speed: v_{ex}



Shell Ablation

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- (1) X-rays heat Ablator
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- (2) Material ejected
 - outward
 - speed: v_{ex}
- (3) Implosion of shell
 - inward
 - speed: v_{imp}



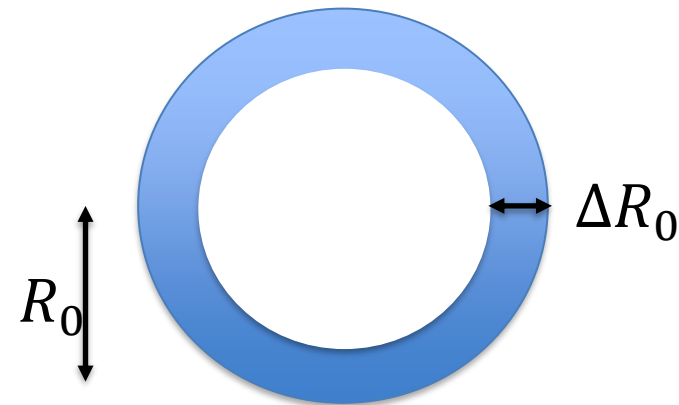


Spherical Rocket

- 1D Rocket Model
 - $M \frac{dv_{imp}}{dt} = v_{ex} \frac{dM}{dt} \rightarrow v_{imp} = v_{ex} \ln \left(\frac{M_0}{M} \right)$ Standard Rocket Equation
 - Radius of shell changes as fuel implodes inwards

Spherical Rocket

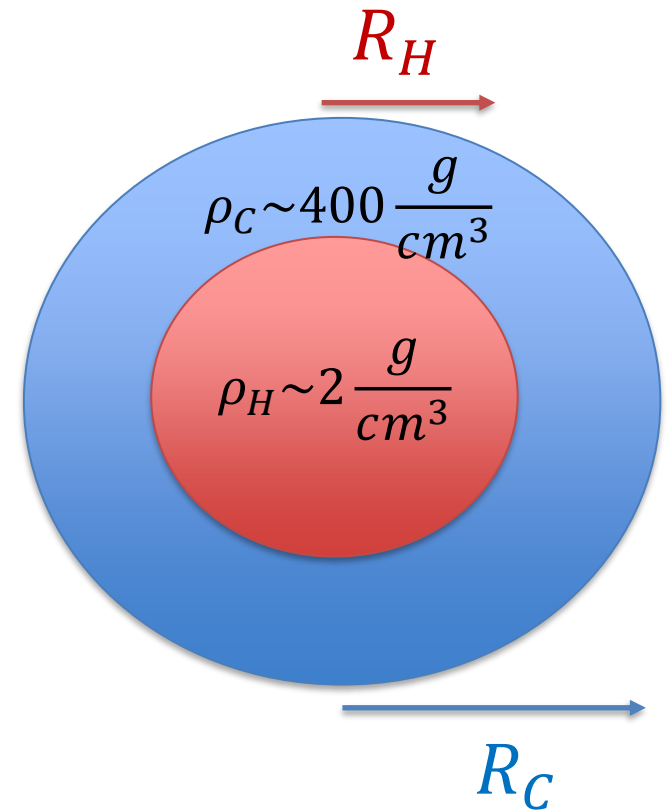
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 - $M \frac{dv_{imp}}{dt} = v_{ex} \frac{dM}{dt} \rightarrow v_{imp} = v_{ex} \ln \left(\frac{M_0}{M} \right)$ **Standard Rocket Equation**
 - Radius of shell changes as fuel implodes inwards
- Implosion Velocity Scaling
 - $v_{imp} \sim v_a A$
 - Aspect Ratio: $A \approx \frac{R_0}{\Delta R_0}$
 - Thin shell drives faster implosions
 - Ablation Velocity: v_a
 - speed at which shell recedes
 - related to hohlraum temperature





Ignition and Burn

- Kinetic Energy of Imploding Shell goes to Internal Energy of DT Fuel.
 - $KE = \frac{1}{2} M v_{imp}^2$



Ignition and Burn

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- $KE = \frac{1}{2} M v_{imp}^2$

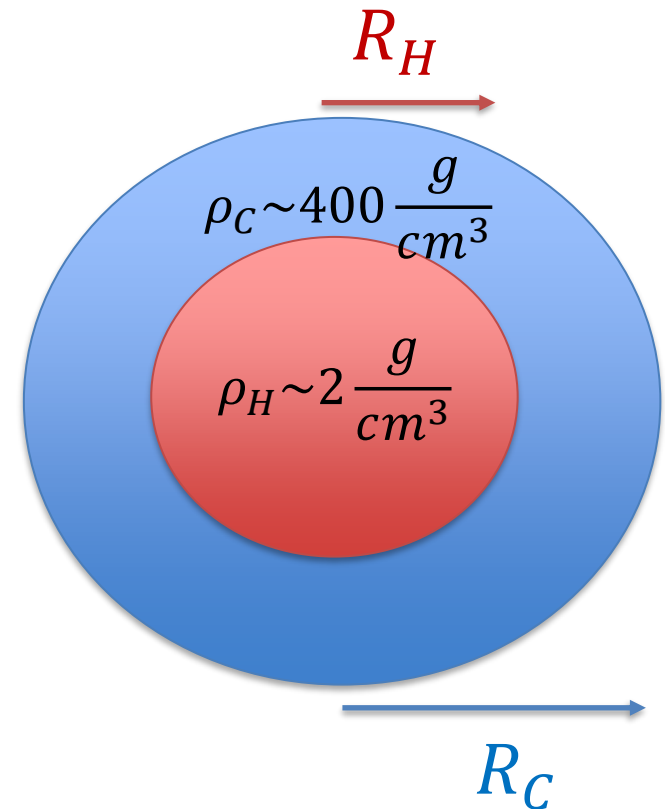
- Ignition Condition: $n\tau > \frac{10^{15} s}{cm^3}$

- $n\tau = \frac{\rho}{m} \frac{R}{v_{th}} \sim \rho R : \text{in } \frac{g}{cm^2}$

- How to Quantify how much is burned?: Φ

- Burn Efficiency: $\Phi \equiv \frac{\rho R}{\rho R + H_B}$

- H_B is the burn parameter





Gain and Yield

- Given ρR and T_H , we know fraction of fuel that gets burned Φ
 - $\Phi \equiv \frac{\rho R}{\rho R + H_B}$



Gain and Yield

- Given ρR and T_H , we know fraction of fuel that gets burned Φ
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- Alpha Particles cause heating:
 - $Q = 3.5$ MeV per DT pair or 67 MJ/mg
- Multiply by total fuel mass M_f

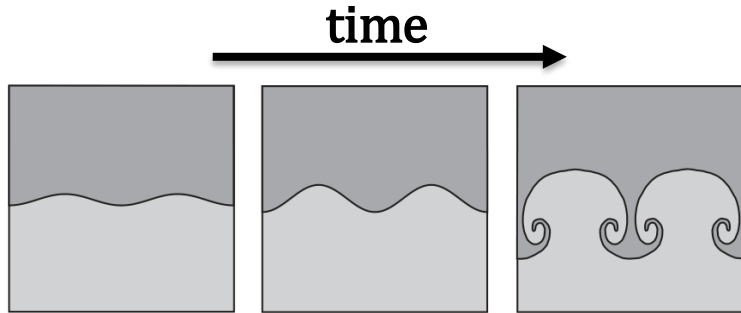


Gain and Yield

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 - $\Phi \equiv \frac{\rho R}{\rho R + H_B}$
- Alpha Particles cause heating:
 - $Q = 3.5$ MeV per DT pair or 67 MJ/mg
- Multiply by total fuel mass M_f
- Fusion Energy Yield in MJ is “Y”
 - $Y = M_f Q \Phi$
- Gain: $G \equiv \frac{Y}{E_L}$
 - E_L is the laser energy on target

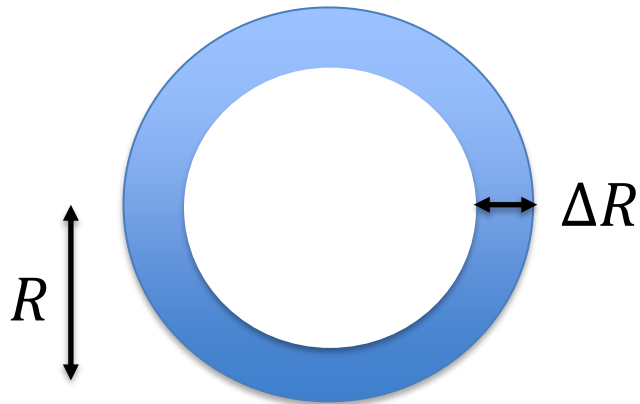


Instabilities and Symmetry



S. Atzeni, *The Physics of Inertial Fusion: Beam Plasma Interaction, Hydrodynamics, Hot Dense Matter* (2004)

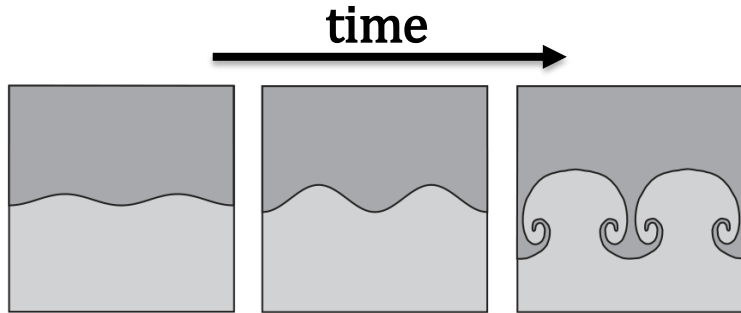
Rayleigh-Taylor Instability



Aspect Ratio: $A \equiv R/\Delta R$

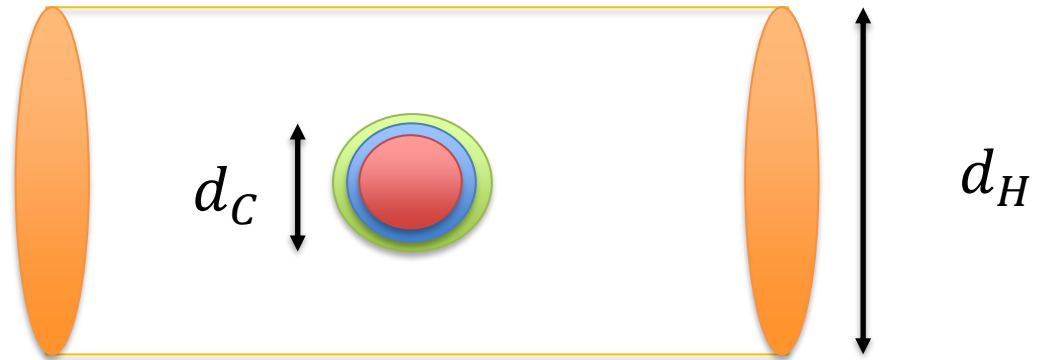


Instabilities and Symmetry

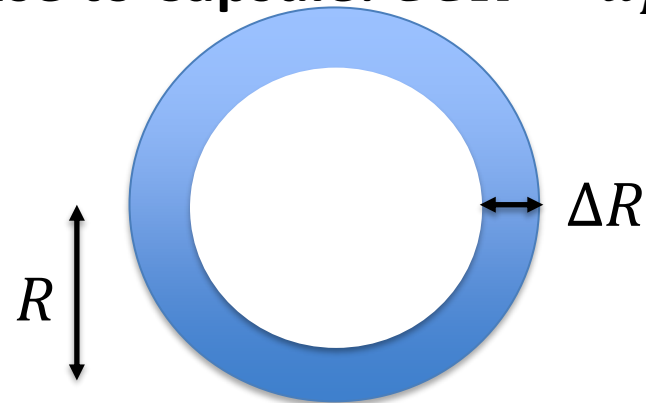


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Rayleigh-Taylor Instability



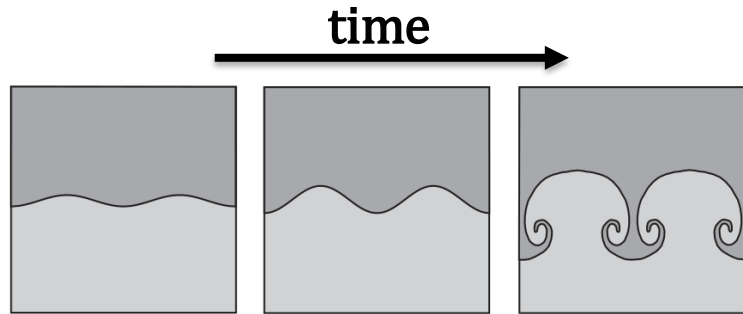
Case to Capsule: $CCR = d_H/d_c$



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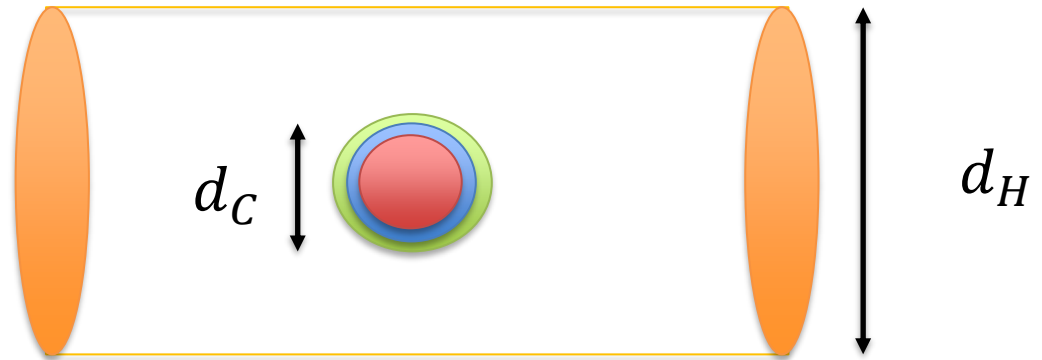


Instabilities and Symmetry

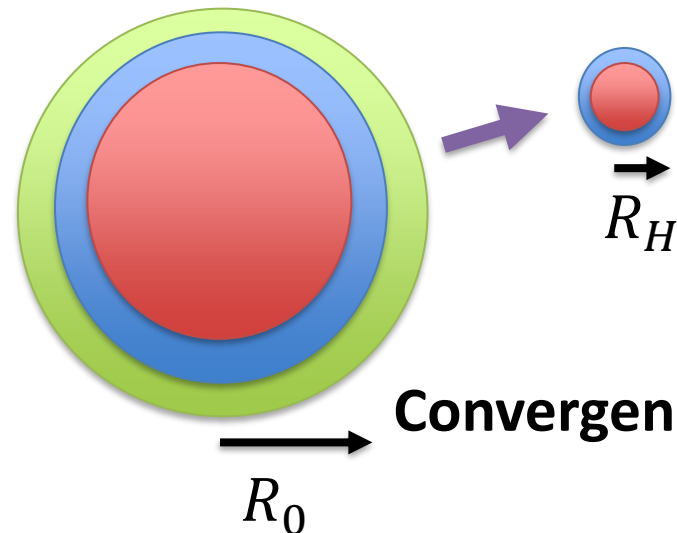


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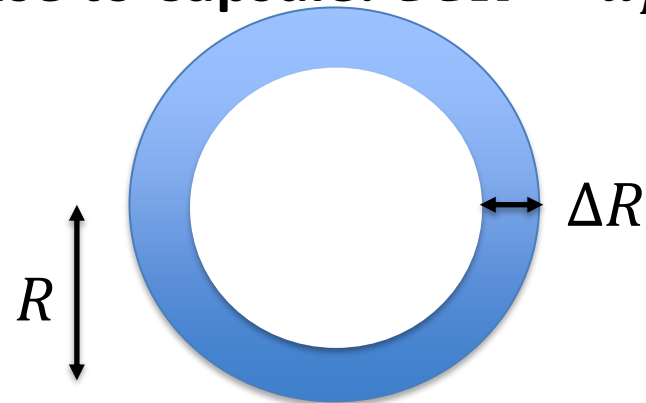
Rayleigh-Taylor Instability



Case to Capsule: $CCR = d_H/d_c$



Convergence: $C \equiv R_0/R_H$

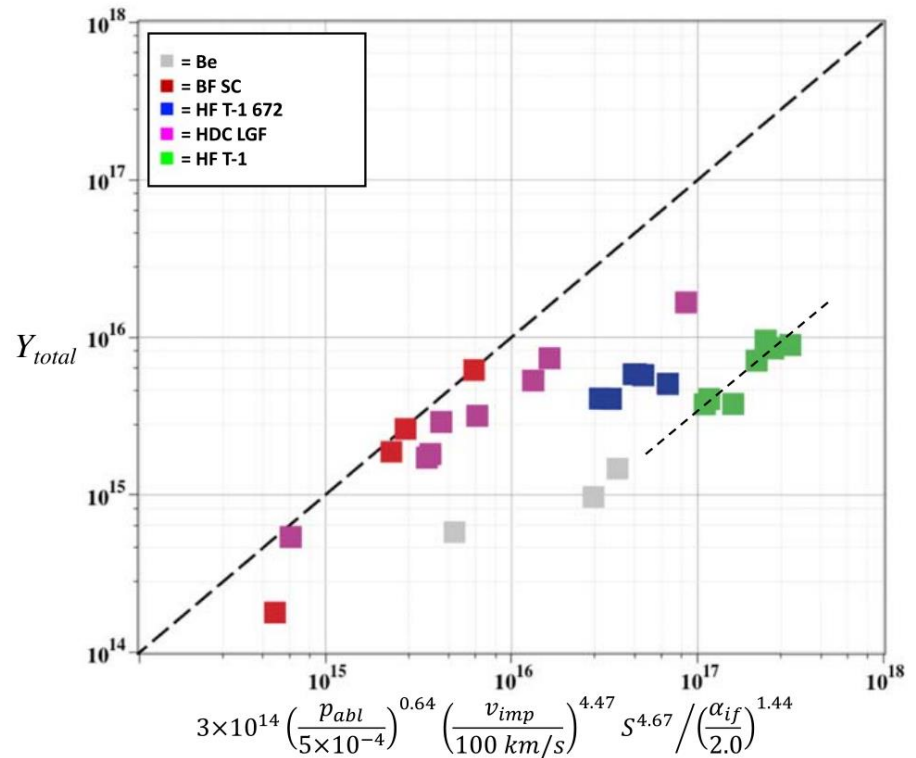


Aspect Ratio: $A \equiv R/\Delta R$



Neutron Yield Scaling

$$Y \sim P_{abl}^{0.64} \frac{v_{imp}^{4.47}}{\alpha^{1.44}} S^{4.67}$$



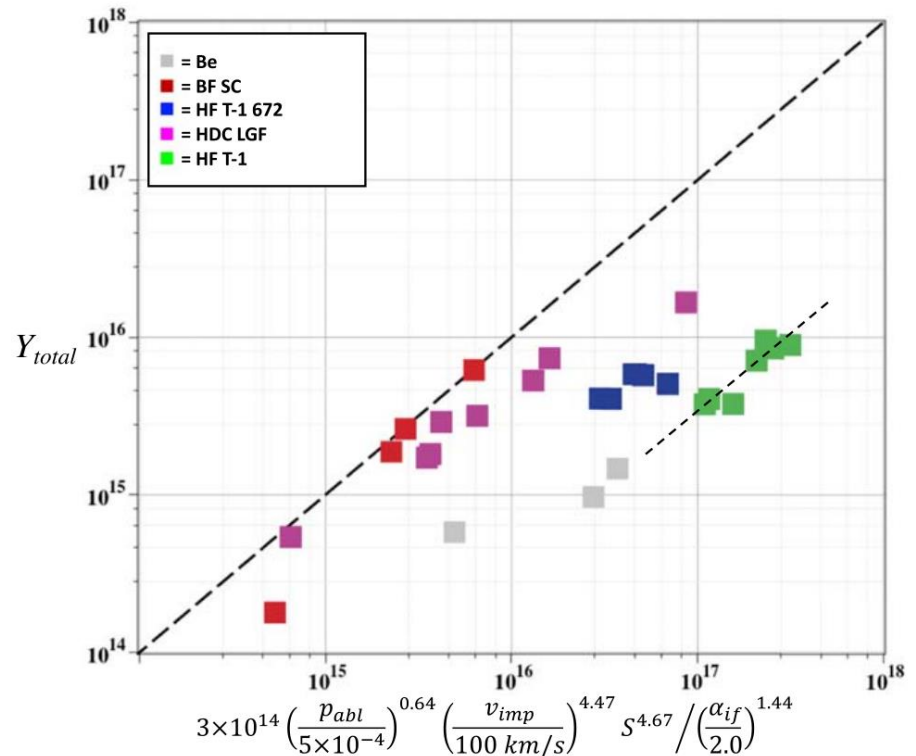
Plasma Phys. Control. Fusion 61 (2019)



Neutron Yield Scaling

$$Y \sim P_{abl}^{0.64} \frac{v_{imp}^{4.47}}{\alpha^{1.44}} S^{4.67}$$

- P_{abl} : Ablation Pressure
 - Laser Energy
- v_{imp} : Implosion Velocity
 - Aspect Ratio



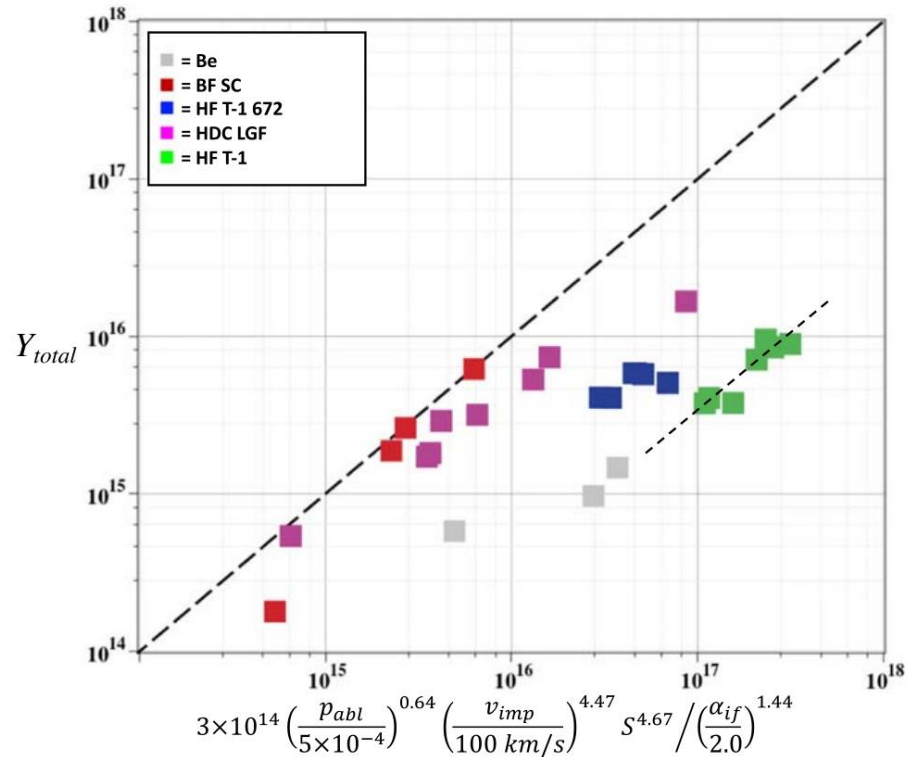
Plasma Phys. Control. Fusion 61 (2019)



Neutron Yield Scaling

$$Y \sim P_{abl}^{0.64} \frac{v_{imp}^{4.47}}{\alpha^{1.44}} S^{4.67}$$

- P_{abl} : Ablation Pressure
 - Laser Energy
- v_{imp} : Implosion Velocity
 - Aspect Ratio
- S : Scale
 - Mass and Radius
- α : Isentrope Parameter
 - Laser Pulse Profile

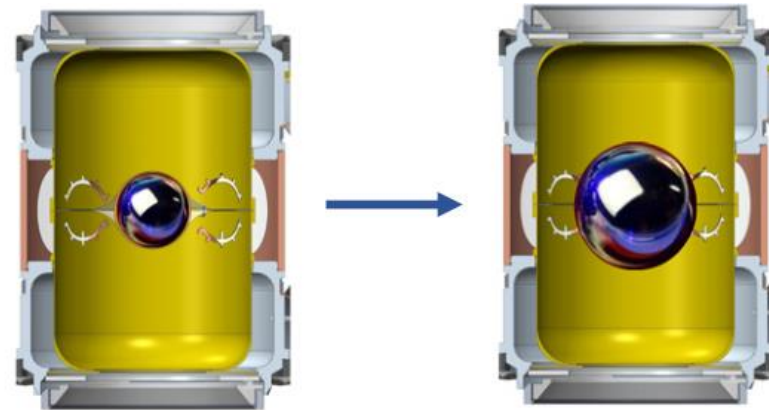


Plasma Phys. Control. Fusion 61 (2019)

Hybrid-E Campaign

- High Yield Big Radius Implosion Design
 - Increased Scale of NIF capsules $\sim 15\%$
 - Kept Hohlraum Size Same
 - Differences in P_{abl} , v_{imp} , α negligible

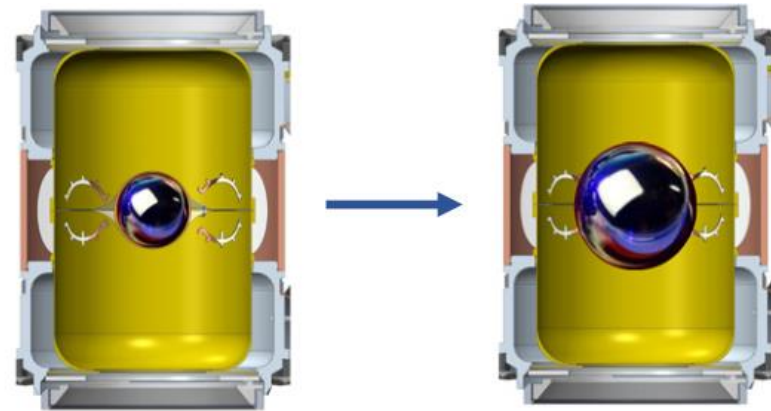
Phys. Plasmas 26, 052704 (2019)



Hybrid-E Campaign

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- N170827 (HDC Campaign)
 - $R \approx 910 \mu m$
 - $Y = 0.053 \text{ MJ}$
- N210207 (HYBRID-E Campaign)
 - $R \approx 1050 \mu m$
 - $Y = 0.174 \text{ MJ}$

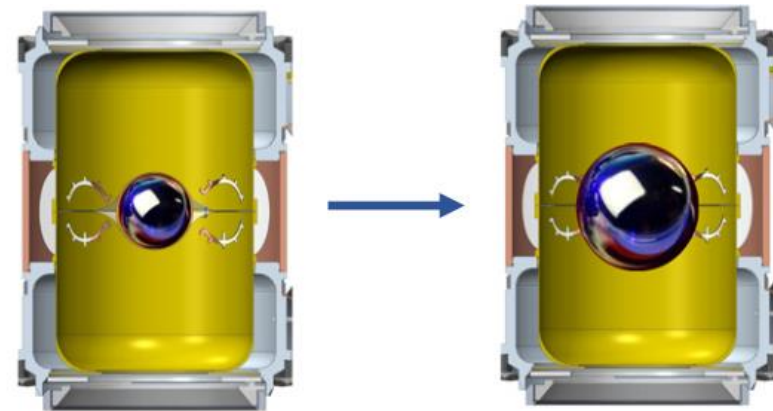
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 - $R \approx 1050 \mu m$
 - $Y = 0.174 \text{ MJ}$
- Over 3x increase in yield, scaling predicts 2x increase
 - Scaling works best within same campaign

Phys. Plasmas 26, 052704 (2019)



$$\frac{Y_{21}}{Y_{17}} \sim \left(\frac{R_{21}}{R_{17}} \right)^{4.67} \\ = (1.15)^{4.67} \approx 2$$



Conclusion

- DT most viable candidate for controlled fusion



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- Physical Considerations
 - Laser Energy, Hohlraum/Capsule Size
- Engineering Considerations
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 - Capsule Smoothness, Laser Efficiency
- N210207->N210808 Shot
 - 8x gain increase: same capsule size
 - mainly due to engineering advances
- Future of NIF
 - N210808->N221204 had $G = 0.72 \rightarrow 1.5$
 - N221204->N23???? has $G=1.5 \rightarrow ???$
 - 8% thicker ablator, 8% increase in laser energy
 - Symmetry Improvements



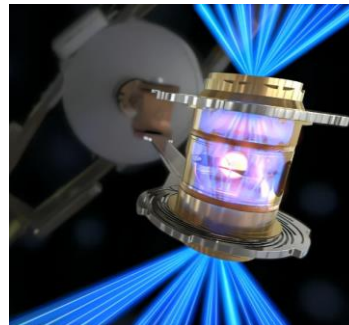
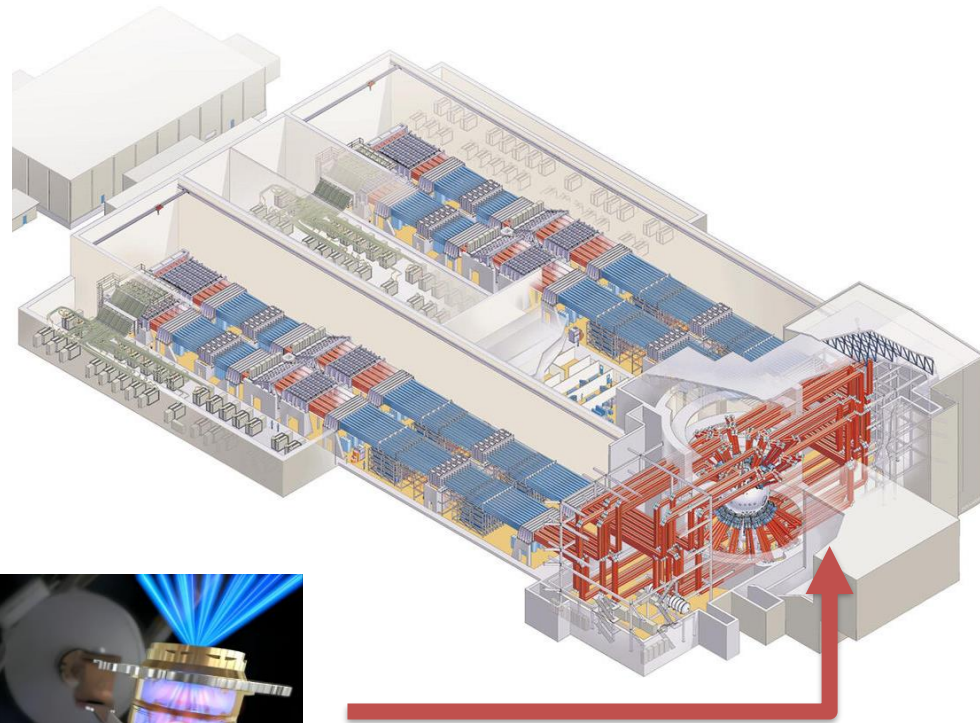
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11. [IAEA Webinar Explores NIF's Ignition and Energy Gain Breakthroughs \(Inl.gov\)](#)



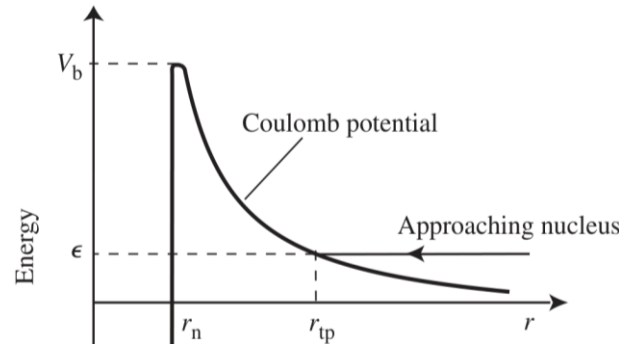
National Ignition Facility

- Size of a Sports Stadium
- 192 laser beams
 - angles: 23, 30, 40, 50
 - ~2 MJ peak energy
 - ~500 TW Peak Power
 - Nd-glass laser
 - $351\mu\text{m}$ (frequency tripled)
 - 1 shot every 8 hours





Tunneling Coulomb Barrier



- $\Psi'' = \frac{2m(V-E)}{\hbar^2} \Psi$
- $\Psi = e^{-\phi(x)}$
- $-\phi''(x) + \phi'(x)^2 = \frac{2m(V-E)}{\hbar^2}$
- $\phi(x) \approx \int_{x_0}^x \sqrt{\frac{2m(V-E)}{\hbar^2}} dx'$
- $V(r) = \frac{e^2}{4\pi\epsilon_0 r}, E = \frac{e^2}{4\pi\epsilon_0 r_{tp}}$
- $\phi(r_{tp}) \sim \sqrt{r_{tp}} \sim 1/\sqrt{E}$
- $\Psi(r_{tp}) \sim e^{1/\sqrt{E}}$

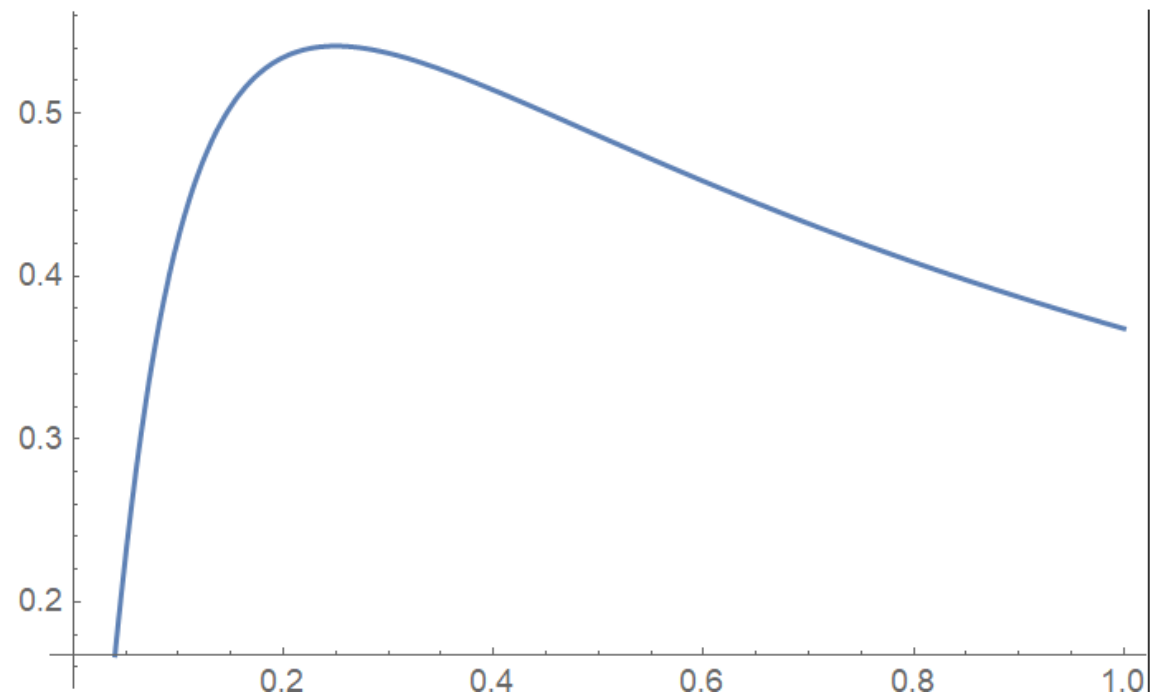
- (1) Schrodinger Equation
- (2) Assume form of Ψ
- (3) $\phi''(x) = 0$ (slowly varies)
- (4) WKB (w/ eq. (2))
- (5) For 1D Z=1 Barrier
- (6) Apply eq. (4) to eq. (5)
- (7) Apply eq. (2) to eq. (6)



Fusion Cross Section

- $\sigma_{geo} \sim \lambda^2 \sim p^{-2} \sim \frac{1}{\epsilon}$
- $P_{tun} \sim e^{-\sqrt{\frac{1}{\epsilon}}}$

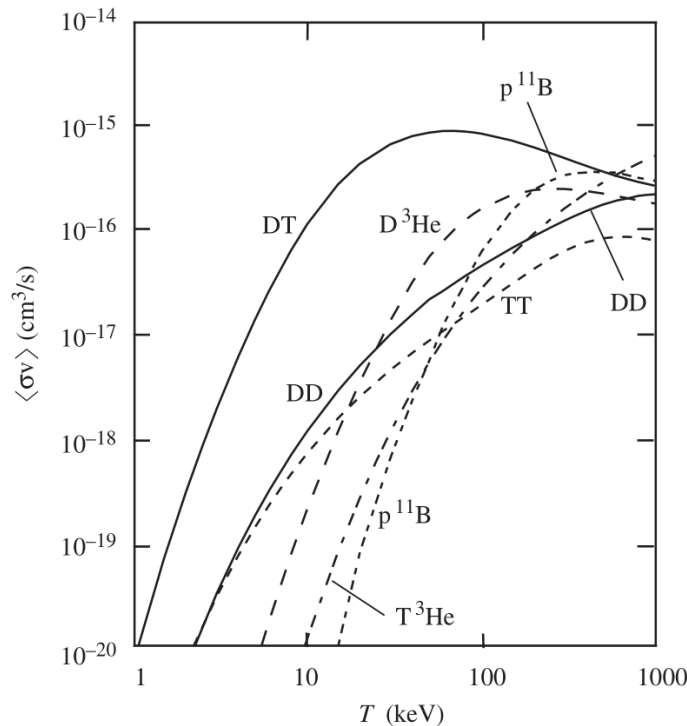
Plot $\left[\frac{\text{Exp}[-1/\sqrt{\epsilon}]}{\epsilon}, \{\epsilon, 0, 1\} \right]$





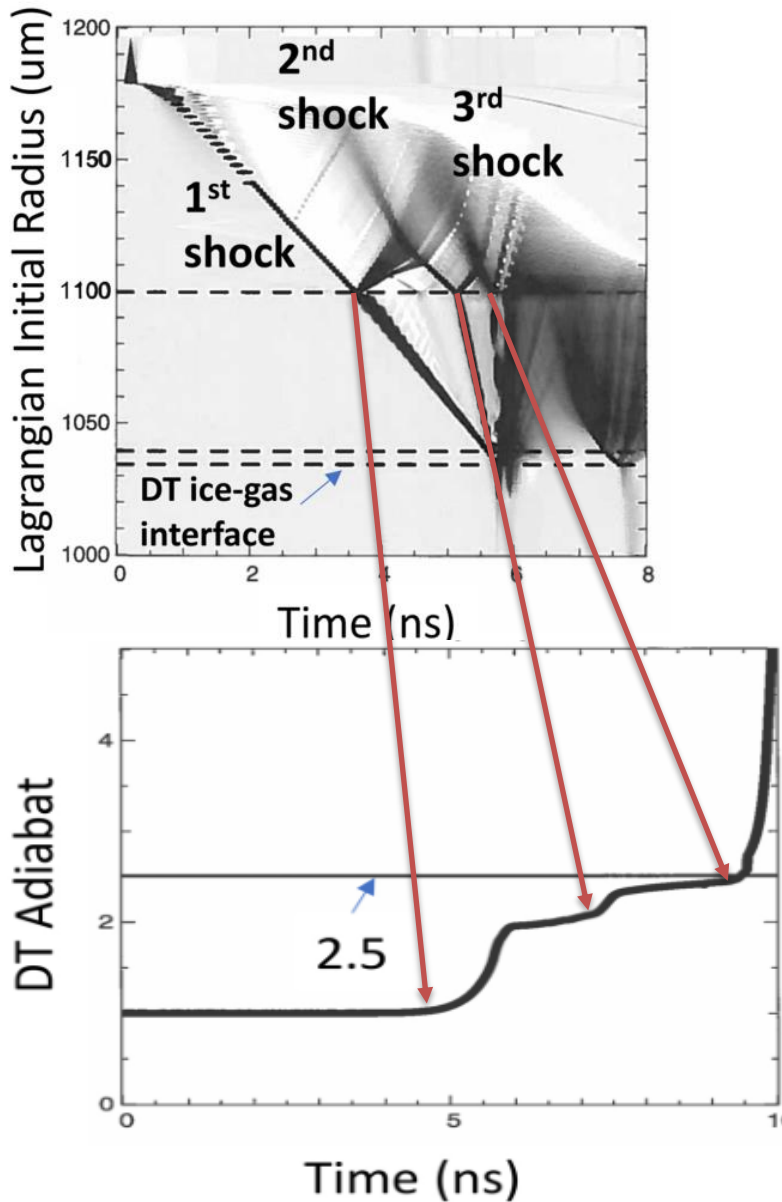
Reactivity

$$\langle \sigma v \rangle_{DT} = \begin{cases} 4.2 \times 10^{-20} (T_{keV})^4 \text{ cm}^3 \text{ s}^{-1} & \text{if } 3 < T_{keV} < 6 \\ 1.1 \times 10^{-18} (T_{keV})^2 \text{ cm}^3 \text{ s}^{-1} & \text{if } 8 < T_{keV} < 25 \end{cases}$$



S. Atzeni, *The Physics of Inertial Fusion: Beam Plasma Interaction, Hydrodynamics, Hot Dense Matter* (2004)

$$\langle \sigma v \rangle = \frac{4\pi}{(2\pi m_r)^{1/2}} \frac{1}{(k_B T)^{3/2}} \int_0^\infty \sigma(\epsilon) \epsilon \exp(-\epsilon/k_B T) d\epsilon.$$



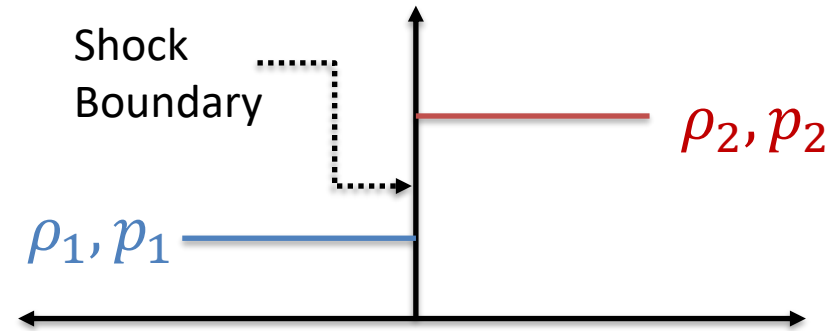
Shocks Increasing DT Fuel Adiabat (Isentrope Parameter)

Phys. Plasmas 28, 072706 (2021)



1D Hydrodynamics

- Laser pulse is short and intense: Can't compress adiabatically!
 - Shock Waves Drive Compression
- Euler Equations of Hydrodynamics across shock interface
- Ex) Want 4x compression:
 - $\frac{p_2}{p_1} \approx 10$ isentropically
 - $\frac{p_2}{p_1} \rightarrow \infty$ for 1 shock!



Mass:

$$\rho_1 v_1 = \rho_2 v_2$$

Momentum: $p_1 + \rho_1 v_1^2 = p_2 + \rho_2 v_2^2$

Energy: $e_1 + p_1/\rho_1 + v_1^2/2 = e_2 + p_2/\rho_2 + v_2^2/2$

$$V \equiv 1/\rho \qquad e = \frac{3}{2} pV$$

$$\frac{p_2}{p_1} = \frac{4 - V_2/V_1}{4V_2/V_1 - 1}$$

$$\frac{p_2}{p_1} = \left(\frac{V_1}{V_2} \right)^{5/3}$$



Spherical Rocket

- 1D Rocket Model

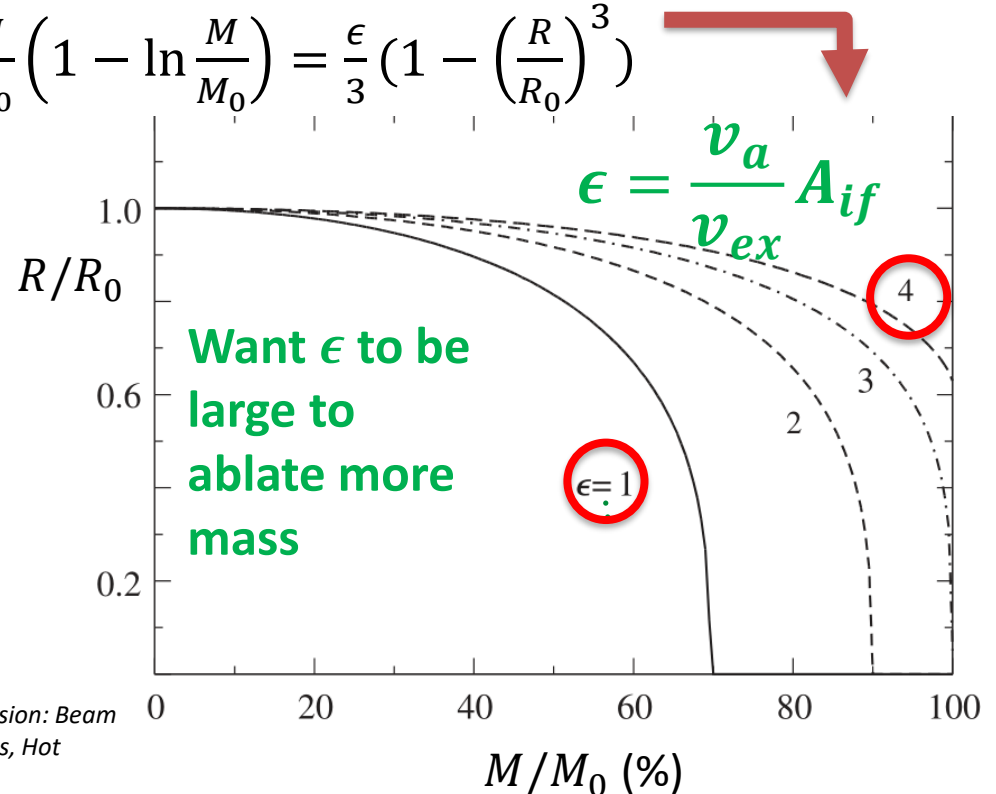
- $$M \frac{dv_{imp}}{dt} = v_{ex} \frac{dM}{dt} \rightarrow v_{imp} = v_{ex} \ln \left(\frac{M_0}{M} \right) \quad \text{Standard Rocket Equation}$$

- $$\frac{dM}{dt} = -4\pi R^2 \dot{m}_a \rightarrow 1 - \frac{M}{M_0} \left(1 - \ln \frac{M}{M_0} \right) = \frac{\epsilon}{3} \left(1 - \left(\frac{R}{R_0} \right)^3 \right)$$

- $$v_{imp} \sim v_a A_{if}$$

- Aspect Ratio: $A_{if} \approx \frac{R_0}{\Delta R_0}$

- Ablation Velocity: v_a
 - speed at which shell recedes





Example: Yield/Gain Calculation

EX) $\rho R = 2 \frac{g}{cm^2}, H_B = 6 \frac{g}{cm^2}, M_f = 0.2 \text{ mg}, E_L = 1.9 \text{ MJ}$

- $\Phi = \frac{2}{2+6} = 0.25$
- $Y = 0.2 \times 67 \times 0.25 = 3.35 \text{ MJ}$
- $G = \frac{3.35}{1.9} = 1.76$
 - More energy in than out!
- But: 400 MJ Laser \rightarrow 1.9MJ on target
- Need larger size/mass target for practical fusion power plant



Simplified Scaling Estimate

1. Energy Balance (Assume $p_H = p_S$)

$$\frac{3}{2}pV_H + \frac{3}{2}pV_S = \frac{1}{2}M_C v_{imp}^2$$

$$p = \rho \frac{\Delta R}{R_H} v_{imp}^2$$

2. Partially Fermi Degenerate Shell

$$p \propto \rho_C^{5/3}$$

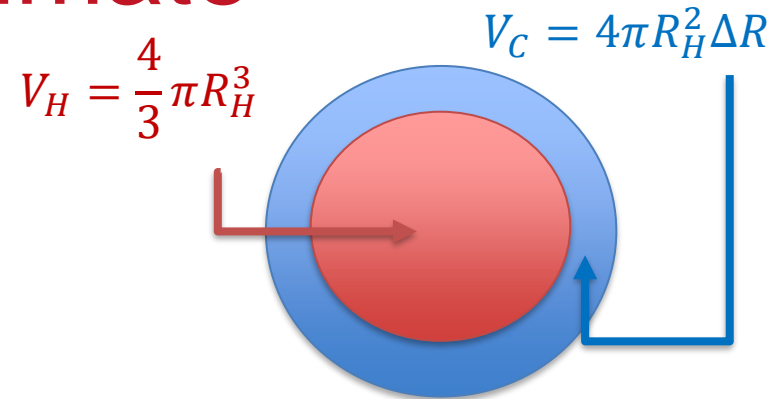
$$\rho_C \propto \left(\frac{\Delta R}{R_H}\right)^{3/2} \frac{1}{\alpha^{3/2}} v_{imp}^3$$

3. Areal Density

$$\rho_C R_H \sim \frac{v_{imp}^3}{\alpha^{3/2}} S$$

4. Yield

$$Y \sim \Phi M_C \sim \rho_C R_H M_C \sim \frac{v_{imp}^3}{\alpha^{1.5}} S^4$$



$$\frac{p}{p_D} \equiv \alpha, \quad p_D \equiv \frac{(3\pi^2)^{2/3} \hbar^2}{5m_e} (\rho)^{5/3}$$

$$M_C \sim R_0^3 \sim S^3$$

Compare to

$$Y \sim P_{abl}^{0.64} \frac{v_{imp}^{4.47}}{\alpha^{1.44}} S^{4.67}$$