Stereo



Many slides adapted from Steve Seitz

 Given a calibrated binocular stereo pair, fuse it to produce a depth image

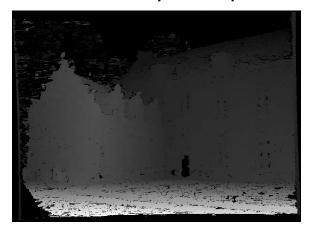
image 1



image 2



Dense depth map

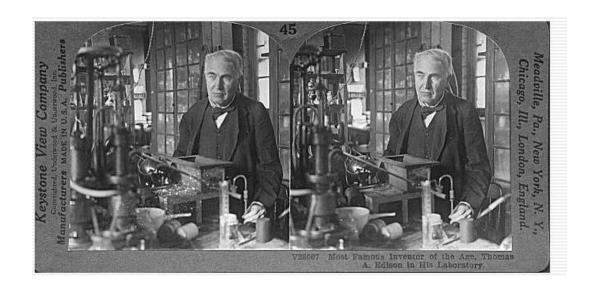


 Given a calibrated binocular stereo pair, fuse it to produce a depth image



Where does the depth information come from?

- Given a calibrated binocular stereo pair, fuse it to produce a depth image
 - Humans can do it



Stereograms: Invented by Sir Charles Wheatstone, 1838

- Given a calibrated binocular stereo pair, fuse it to produce a depth image
 - Humans can do it



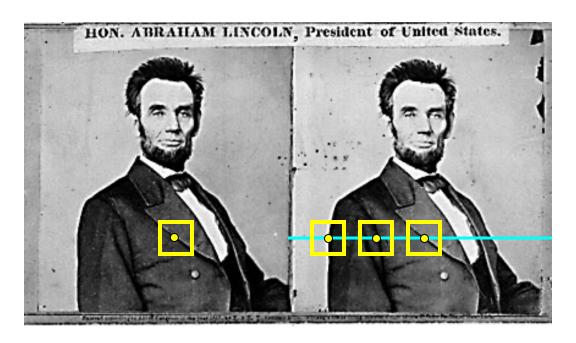
Autostereograms: www.magiceye.com

- Given a calibrated binocular stereo pair, fuse it to produce a depth image
 - Humans can do it



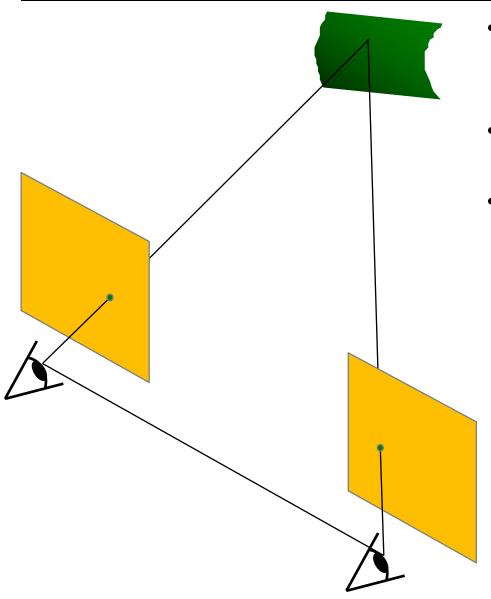
Autostereograms: www.magiceye.com

Basic stereo matching algorithm



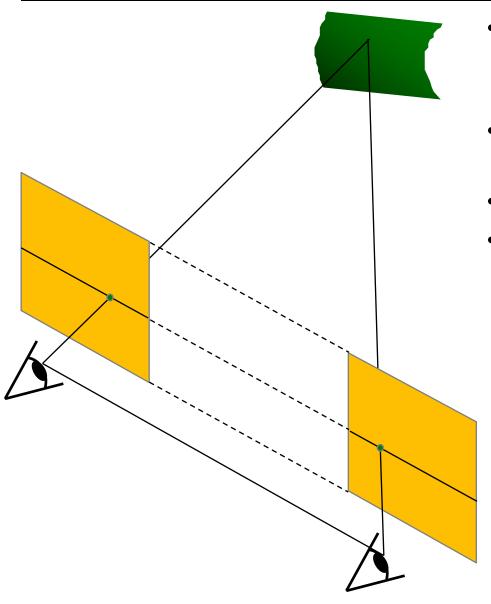
- For each pixel in the first image
 - Find corresponding epipolar line in the right image
 - Examine all pixels on the epipolar line and pick the best match
 - Triangulate the matches to get depth information
- Simplest case: epipolar lines are corresponding scanlines
 - When does this happen?

Simplest Case: Parallel images



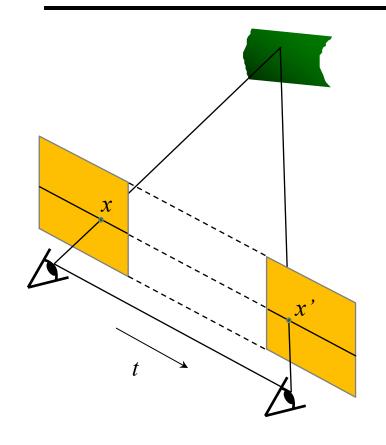
- Image planes of cameras are parallel to each other and to the baseline
- Camera centers are at same height
- Focal lengths are the same

Simplest Case: Parallel images



- Image planes of cameras are parallel to each other and to the baseline
- Camera centers are at same height
- Focal lengths are the same
- Then epipolar lines fall along the horizontal scan lines of the images

Essential matrix for parallel images



Epipolar constraint:

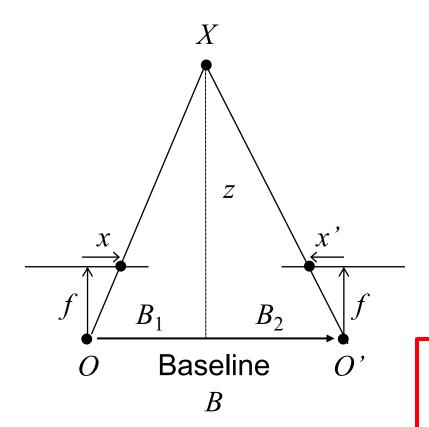
$$\mathbf{x}'^T \mathbf{E} \mathbf{x} = 0, \quad \mathbf{E} = [\mathbf{t}_{\times}] \mathbf{R}$$

$$R = I$$
 $t = (T, 0, 0)$

$$\boldsymbol{E} = [\boldsymbol{t}_{\times}] \boldsymbol{R} = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & -T \\ 0 & T & 0 \end{bmatrix}$$

The y-coordinates of corresponding points are the same!

Depth from disparity



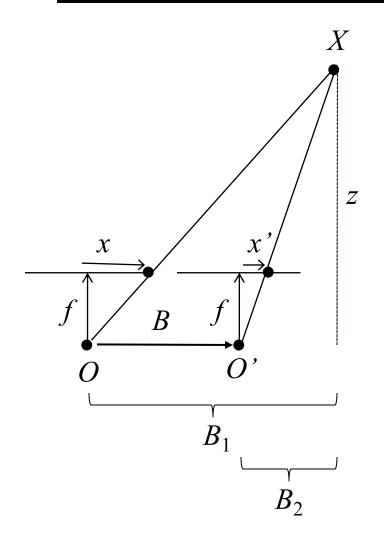
$$\frac{x}{f} = \frac{B_1}{z} \qquad \frac{-x'}{f} = \frac{B_2}{z}$$

$$\frac{x - x'}{f} = \frac{B_1 + B_2}{z}$$

$$disparity = x - x' = \frac{B \cdot f}{z}$$

Disparity is inversely proportional to depth!

Depth from disparity



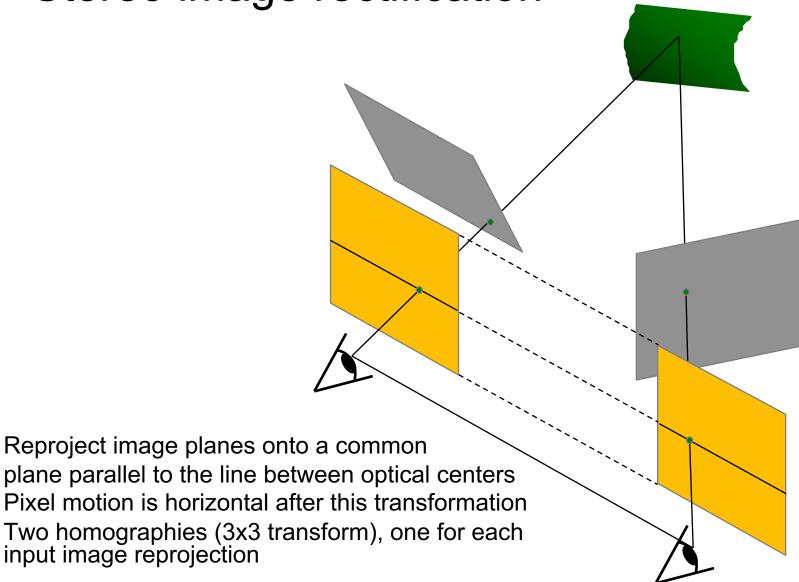
$$\frac{x}{f} = \frac{B_1}{z} \qquad \frac{x'}{f} = \frac{B_2}{z}$$

$$\frac{x - x'}{f} = \frac{B_1 - B_2}{z}$$

$$disparity = x - x' = \frac{B \cdot f}{z}$$

Stereo image rectification

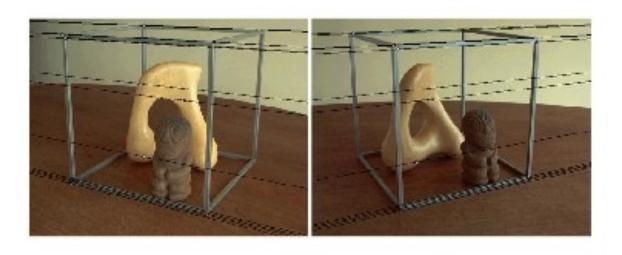
Stereo image rectification



C. Loop and Z. Zhang. <u>Computing Rectifying Homographies for Stereo Vision</u>. IEEE Conf. Computer Vision and Pattern Recognition, 1999.

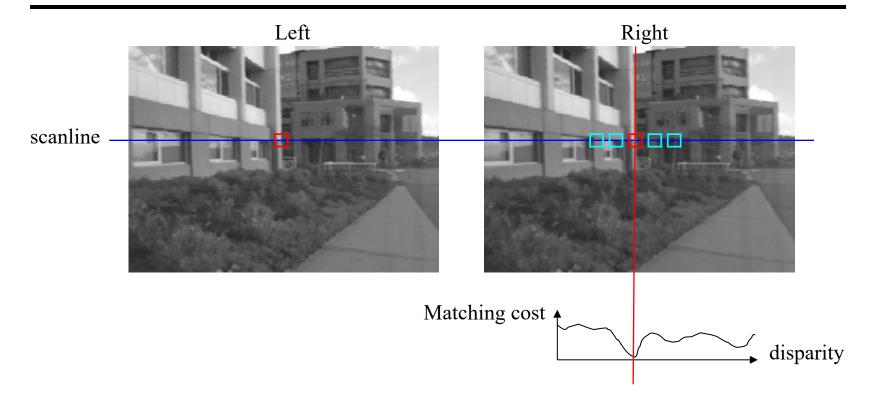
input image reprojection

Rectification example



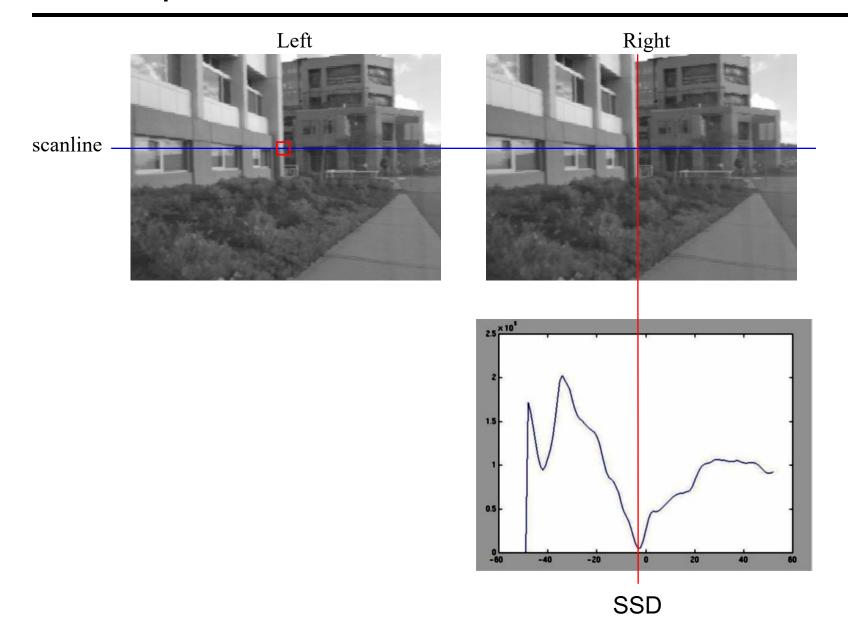


Correspondence search

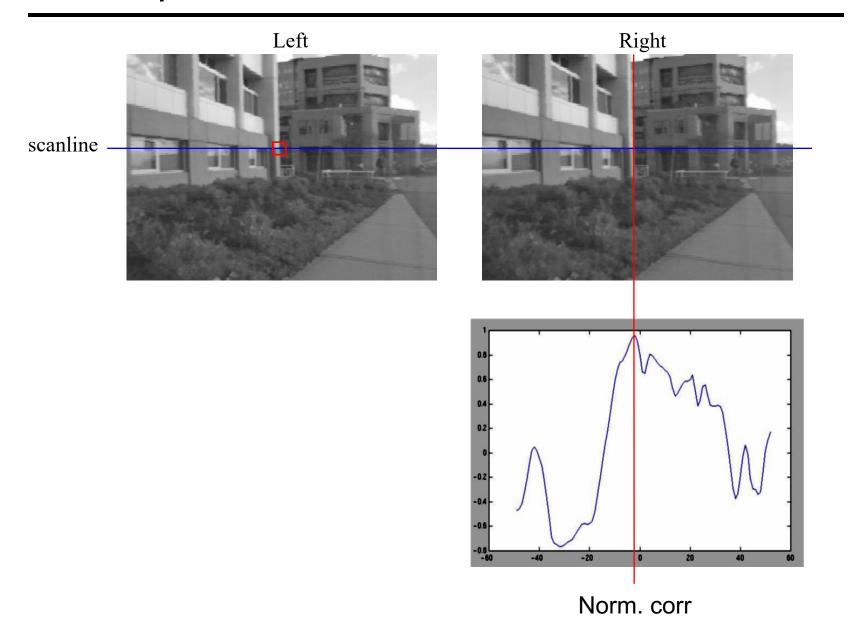


- Slide a window along the right scanline and compare contents of that window with the reference window in the left image
- Matching cost: SSD or normalized correlation

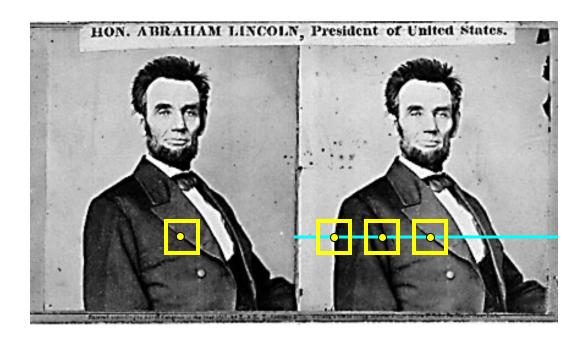
Correspondence search



Correspondence search

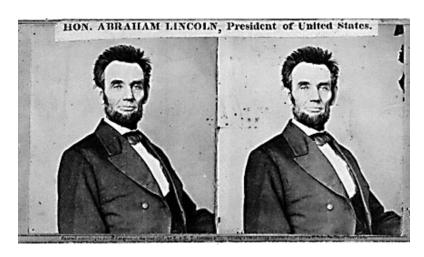


Basic stereo matching algorithm



- If necessary, rectify the two stereo images to transform epipolar lines into scanlines
- For each pixel x in the first image
 - Find corresponding epipolar scanline in the right image
 - Examine all pixels on the scanline and pick the best match x'
 - Compute disparity x-x' and set $depth(x) = B^*f/(x-x')$

Failures of correspondence search



Textureless surfaces



Occlusions, repetition







Non-Lambertian surfaces, specularities

Effect of window size







$$W = 3$$

W = 20

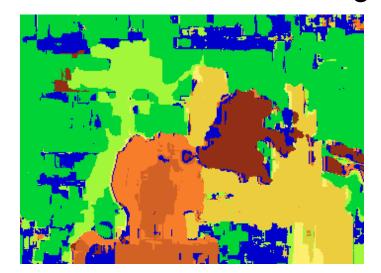
- Smaller window
 - + More detail
 - More noise
- Larger window
 - + Smoother disparity maps
 - Less detail

Results with window search

Data



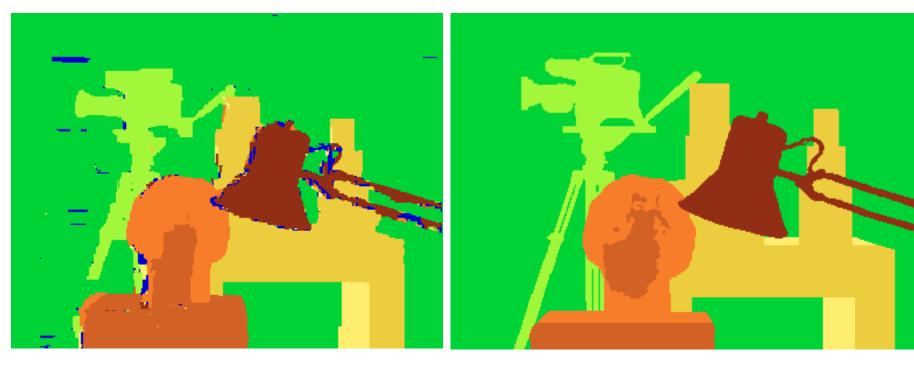
Window-based matching



Ground truth



Better methods exist...



Graph cuts

Ground truth

Y. Boykov, O. Veksler, and R. Zabih, <u>Fast Approximate Energy</u> <u>Minimization via Graph Cuts</u>, PAMI 2001

For the latest and greatest: http://www.middlebury.edu/stereo/